## CODATA recommended values of the fundamental physical constants: 2006

Cite as: J. Phys. Chem. Ref. Data 37, 1187 (2008); https://doi.org/10.1063/1.2844785 Submitted: 28 December 2007 . Accepted: 08 January 2008 . Published Online: 15 July 2008

Peter J. Mohr, Barry N. Taylor, and David B. Newell

## ARTICLES YOU MAY BE INTERESTED IN

CODATA Recommended Values of the Fundamental Physical Constants: 2010 Journal of Physical and Chemical Reference Data 41, 043109 (2012); https:// doi.org/10.1063/1.4724320

CODATA Recommended Values of the Fundamental Physical Constants: 2014 Journal of Physical and Chemical Reference Data 45, 043102 (2016); https:// doi.org/10.1063/1.4954402

CODATA Recommended Values of the Fundamental Physical Constants: 1998
Journal of Physical and Chemical Reference Data 28, 1713 (1999); https:// doi.org/10.1063/1.556049


# CODATA recommended values of the fundamental physical constants: 2006* 

Peter J. Mohr, ${ }^{\dagger}$ Barry N. Taylor, ${ }^{\ddagger}$ and David B. Newell ${ }^{\text {§ }}$<br>National Institute of Standards and Technology, Gaithersburg, Maryland 20899-8420, USA

(Received 28 December 2007; accepted 8 January 2008; published online 15 July 2008)


#### Abstract

This paper gives the 2006 self-consistent set of values of the basic constants and conversion factors of physics and chemistry recommended by the Committee on Data for Science and Technology (CODATA) for international use. Further, it describes in detail the adjustment of the values of the constants, including the selection of the final set of input data based on the results of least-squares analyses. The 2006 adjustment takes into account the data considered in the 2002 adjustment as well as the data that became available between 31 December 2002, the closing date of that adjustment, and 31 December 2006, the closing date of the new adjustment. The new data have led to a significant reduction in the uncertainties of many recommended values. The 2006 set replaces the previously recommended 2002 CODATA set and may also be found on the World Wide Web at http://physics.nist.gov/constants. © 2008 by the U.S. Secretary of Commerce on behalf of the United States. All rights reserved. DOI: 10.1063/1.2844785


Key words: Avogadro constant; Boltzmann constant; elementary charge; fine-structure constant; fundamental constants; Josephson effect; Newtonian gravitational constant; particle magnetic moments; particle masses; Planck constant; quantum electrodynamics; quantum Hall effect.

## CONTENTS

I. Introduction
A. Background 1189
B. Time variation of the constants
C. Outline of paper

[^0]II. Special Quantities and Units ..... 1190
III. Relative Atomic Masses ..... 1191
A. Relative atomic masses of atoms ..... 1191
B. Relative atomic masses of ions and nuclei ..... 1192
C. Cyclotron resonance measurement of the electron relative atomic mass $A_{\mathrm{r}}(\mathrm{e})$ ..... 1193
IV. Atomic Transition Frequencies ..... 1193
A. Hydrogen and deuterium transition frequencies, the Rydberg constant $R_{\infty}$, and the proton and deuteron charge radii $R_{\mathrm{p}}, R_{\mathrm{d}}$ ..... 1193

1. Theory relevant to the Rydberg constant ..... 1193
a. Dirac eigenvalue ..... 1193
b. Relativistic recoil ..... 1194
c. Nuclear polarization ..... 1194
d. Self energy ..... 1194
e. Vacuum polarization ..... 1195
f. Two-photon corrections ..... 1196
g. Three-photon corrections ..... 1198
h. Finite nuclear size ..... 1198
i. Nuclear-size correction to self energy and vacuum polarization ..... 1199
j. Radiative-recoil corrections ..... 1199
k. Nucleus self energy ..... 1199
l. Total energy and uncertainty ..... 1199
m . Transition frequencies between levels with $n=2$ ..... 1200
2. Experiments on hydrogen and deuterium ..... 1200
3. Nuclear radii ..... 1201
B. Antiprotonic helium transition frequencies and $A_{\mathrm{r}}(\mathrm{e})$ ..... 12011. Theory relevant to antiprotonic helium
4. Experiments on antiprotonic helium1202
5. Values of $A_{\mathrm{r}}(\mathrm{e})$ inferred from antiprotonic helium ..... 1204
C. Hyperfine structure and fine structure ..... 1204
6. Hyperfine structure ..... 1204
7. Fine structure ..... 1205
V. Magnetic Moment Anomalies and $g$-Factors ..... 1205
A. Electron magnetic moment anomaly $a_{\mathrm{e}}$ and the
fine-structure constant $\alpha$

1206

1. Theory of $a_{\mathrm{e}}$
2. Measurements of $a_{\mathrm{e}}$
a. University of Washington
b. Harvard University
3. Values of $\alpha$ inferred from $a_{\text {e }}$
B. Muon magnetic moment anomaly $a_{\mu}$
4. Theory of $a_{\mu}$
5. Measurement of $a_{\mu}$ : Brookhaven
a. Theoretical value of $a_{\mu}$ and inferred value of $\alpha$

1211
C. Bound electron $g$-factor in ${ }^{12} \mathrm{C}^{5+}$ and in ${ }^{16} \mathrm{O}^{7+}$ and $A_{\mathrm{r}}(\mathrm{e})$

1. Theory of the bound electron $g$-factor
2. Measurements of $g_{e}\left({ }^{12} \mathrm{C}^{5+}\right)$ and $g_{\mathrm{e}}\left({ }^{16} \mathrm{O}^{7+}\right)$

$$
1213
$$

a. Experiment on $g_{e}\left({ }^{12} \mathrm{C}^{5+}\right)$ and inferred value of $A_{\mathrm{r}}(\mathrm{e})$
b. Experiment on $g_{e}\left({ }^{16} \mathrm{O}^{7+}\right)$ and inferred value of $A_{\mathrm{r}}(\mathrm{e})$
c. Relations between $g_{\mathrm{e}}\left({ }^{12} \mathrm{C}^{5+}\right)$ and $g_{\mathrm{e}}\left({ }^{16} \mathrm{O}^{7+}\right)$

1215
1216
VI. Magnetic Moment Ratios and the Muon-Electron Mass
Ratio
A. Magnetic moment ratios

1. Theoretical ratios of atomic bound-particle to free-particle $g$-factors
2. Ratio measurements
$\begin{array}{ll}\text { a. Electron to proton magnetic moment ratio } \\ \mu_{\mathrm{e}} / \mu_{\mathrm{p}} & 1217\end{array}$
b. Deuteron to electron magnetic moment ratio $\mu_{\mathrm{d}} / \mu_{\mathrm{e}}$
c. Proton to deuteron and triton to proton magnetic moment ratios $\mu_{\mathrm{p}} / \mu_{\mathrm{d}}$ and $\mu_{\mathrm{t}} / \mu_{\mathrm{p}}$

1218
$\begin{array}{ll}\text { d. Electron to shielded proton magnetic } & \\ \text { moment ratio } \mu_{\mathrm{e}} / \mu_{\mathrm{p}}^{\prime} & 1219\end{array}$
$\begin{array}{ll}\text { e. Shielded helion to shielded proton } & \\ \text { magnetic moment ratio } \mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime} & 1220\end{array}$
$\begin{array}{ll}\text { f. Neutron to shielded proton magnetic } & \\ \text { moment ratio } \mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime} & 1220\end{array}$
B. Muonium transition frequencies, the muon-proton magnetic moment ratio $\mu_{\mu} / \mu_{\mathrm{p}}$, and muon-electron mass ratio $m_{\mu} / m_{\mathrm{e}}$

1. Theory of the muonium ground-state hyperfine splitting
2. Measurements of muonium transition frequencies and values of $\mu_{\mu} / \mu_{\mathrm{p}}$ and $m_{\mu} / m_{\mathrm{e}}$ 1222
a. LAMPF 19821222
b. LAMPF 1999 1222
c. Combined LAMPF results 1223
VII. Electrical Measurements
A. Shielded gyromagnetic ratios $\gamma^{\prime}$, the fine-structure constant $\alpha$, and the Planck constant $h$
3. Low-field measurements
a. NIST: Low field
b. NIM: Low field
c. KRISS/VNIIM: Low field
4. High-field measurements
a. NIM: High field
b. NPL: High field

1224
B. von Klitzing constant $R_{\mathrm{K}}$ and $\alpha \quad 1225$

1. NIST: Calculable capacitor

1225
2. NMI: Calculable capacitor 1226
3. NPL: Calculable capacitor 1226
4. NIM: Calculable capacitor 1226
5. LNE: Calculable capacitor 1226
C. Josephson constant $K_{\mathrm{J}}$ and $h \quad 1226$

1. NMI: Hg electrometer 1226
2. PTB: Capacitor voltage balance 1227
D. Product $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ and $h \quad 1227$
3. NPL: Watt balance 1227
4. NIST: Watt balance 1228
a. 1998 measurement 1228
b. 2007 measurement 1228
5. Other values 1229
6. Inferred value of $K_{\mathrm{J}} \quad 1229$
E. Faraday constant $F$ and $h \quad 1229$
7. NIST: Ag coulometer 1230
VIII. Measurements Involving Silicon Crystals 1230
A. $\{220\}$ lattice spacing of silicon $d_{220} \quad 1230$
$\begin{array}{ll}\text { 1. X-ray optical interferometer measurements } \\ \text { of } d_{220}(X) & 1231\end{array}$
a. PTB measurement of $d_{220}(\mathrm{~W} 4.2 \mathrm{a}) \quad 1232$
b. NMIJ measurement of $d_{220}$ (NR3) 1232
c. INRIM measurement of $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ and
$d_{220}(\mathrm{MO} *)$
8. $d_{220}$ difference measurements 1233
a. NIST difference measurements 1233
b. PTB difference measurements 1234
B. Molar volume of silicon $V_{\mathrm{m}}(\mathrm{Si})$ and the Avogadro
constant $N_{\mathrm{A}}$
C. Gamma-ray determination of the neutron relative
atomic mass $A_{\mathrm{r}}(\mathrm{n})$
$\begin{array}{ll}\text { D. Quotient of Planck constant and particle mass } \\ h / m(X) \text { and } \alpha & 1236\end{array}$
9. Quotient $h / m_{\mathrm{n}} 1236$
10. Quotient $h / m\left({ }^{133} \mathrm{Cs}\right) \quad 1236$
11. Quotient $h / m\left({ }^{87} \mathrm{Rb}\right) \quad 1237$
IX. Thermal Physical Quantities 1238
A. Molar gas constant $R \quad 1238$
12. NIST: Speed of sound in argon 1239
13. NPL: Speed of sound in argon 1239
14. Other values 1239
B. Boltzmann constant $k \quad 1239$
C. Stefan-Boltzmann constant $\sigma \quad 1240$
X. Newtonian Constant of Gravitation $G \quad 1240$
A. Updated values 1240
15. Huazhong University of Science and
Technology 1240
16. University of Zurich 1241
B. Determination of 2006 recommended value of $G \quad 1242$
C. Prospective values 1243
XI. X-ray and Electroweak Quantities 1243
A. X-ray units 1243
B. Particle Data Group input 1243
XII. Analysis of Data 1244
A. Comparison of data 1244
B. Multivariate analysis of data 1253
17. Summary of adjustments 1253
18. Test of the Josephson and quantum Hall effect relations
XIII. The 2006 CODATA Recommended Values
A. Calculational details
B. Tables of values
XIV. Summary and Conclusion

1260
1260
1261
1270
A. Comparison of 2006 and 2002 CODATA recommended values
B. Some implications of the 2006 CODATA recommended values and adjustment for physics and metrology
C. Outlook and suggestions for future work

Acknowledgments
Nomenclature
References
the periodic adjustment of the values of the constants and for descriptions of how units, quantity symbols, numerical values, numerical calculations, and uncertainties are treated, in addition to how the data are characterized, selected, and evaluated. Since the calculations are carried out with more significant figures than are displayed in the text to avoid rounding errors, data with more digits are available on the FCDC web site for possible independent analysis.

However, because of their importance, we recall in detail the following two points also discussed in these references. First, although it is generally agreed that the correctness and overall consistency of the basic theories and experimental methods of physics can be tested by comparing values of particular fundamental constants obtained from widely differing experiments, throughout this adjustment, as a working principle, we assume the validity of the underlying physical theory. This includes special relativity, quantum mechanics, quantum electrodynamics (QED), the standard model of particle physics, including combined charge conjugation, parity inversion, and time reversal ( $C P T$ ) invariance, and the theory of the Josephson and quantum Hall effects, especially the exactness of the relationships between the Josephson and von Klitzing constants $K_{\mathrm{J}}$ and $R_{\mathrm{K}}$ and the elementary charge $e$ and Planck constant $h$. In fact, tests of the relations $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ using the input data of the 2006 adjustment are discussed in Sec. XII.B.2.

The second point has to do with the 31 December 2006 closing date for data to be considered for inclusion in the 2006 adjustment. A datum was considered to have met this date, even though not yet reported in an archival journal, as long as a description of the work was available that allowed the Task Group to assign a valid standard uncertainty $u\left(x_{i}\right)$ to the datum. Thus, any input datum labeled with an 07 identifier because it was published in 2007 was, in fact, available by the cutoff date. Also, some references to results that became available after the deadline are included, even though the results were not used in the adjustment.

## B. Time variation of the constants

This subject, which was briefly touched upon in CODATA-02, continues to be an active field of experimental and theoretical research, because of its importance to our understanding of physics at the most fundamental level. Indeed, a large number of papers relevant to the topic have appeared in the last four years; see the FCDC bibliographic database on the fundamental constants using the keyword time variation at http://physics.nist.gov/constantsbib. For example, see Fortier et al. (2007) and Lea (2007). However, there has been no laboratory observation of the time dependence of any constant that might be relevant to the recommended values.

## C. Outline of paper

Section II touches on special quantities and units, that is, those that have exact values by definition.

Sections III-XI review all available experimental and theoretical data that might be relevant to the 2006 adjustment of the values of the constants. As discussed in Appendix E of CODATA-98, in a least-squares analysis of the fundamental constants, the numerical data, both experimental and theoretical, also called observational data or input data, are expressed as functions of a set of independent variables called adjusted constants. The functions that relate the input data to the adjusted constants are called observational equations, and the leastsquares procedure provides best estimated values, in the least-squares sense, of the adjusted constants. The focus of the review-of-data sections is thus the identification and discussion of the input data and observational equations of interest for the 2006 adjustment. Although not all observational equations that we use are explicitly given in the text, all are summarized in Tables XXXVIII, XL, and XLII of Sec. XII.B.

As part of our discussion of a particular datum, we often deduce from it an inferred value of a constant, such as the fine-structure constant $\alpha$ or Planck constant $h$. It should be understood, however, that these inferred values are for comparison purposes only; the datum from which it is obtained, not the inferred value, is the input datum in the adjustment.

Although just four years separate the 31 December closing dates of the 2002 and 2006 adjustments, there are a number of important new results to consider. Experimental advances include the 2003 Atomic Mass Evaluation from the Atomic Mass Data Center (AMDC), which provides new values for the relative atomic masses $A_{\mathrm{r}}(X)$ of a number of relevant atoms; a new value of the electron magnetic moment anomaly $a_{\mathrm{e}}$ from measurements on a single electron in a cylindrical Penning trap, which provides a value of the fine-structure constant $\alpha$; better measurements of the relative atomic masses of ${ }^{2} \mathrm{H},{ }^{3} \mathrm{H}$, and ${ }^{4} \mathrm{He}$; new measurements of transition frequencies in antiprotonic helium ( $\overline{\mathrm{p}}^{A} \mathrm{He}^{+}$atom $)$ that provide a competitive value of the relative atomic mass of the electron $A_{\mathrm{r}}(\mathrm{e})$; improved measurements of the nuclear magnetic resonance (NMR) frequencies of the proton and deuteron in the HD molecule and of the proton and triton in the HT molecule; a highly accurate value of the Planck constant obtained from an improved measurement of the product $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ using a moving-coil watt balance; new results using combined x-ray and optical interferometers for the $\{220\}$ lattice spacing of single crystals of naturally occurring silicon; and an accurate value of the quotient $h / m\left({ }^{87} \mathrm{Rb}\right)$ obtained by measuring the recoil velocity of rubidium- 87 atoms upon absorption or emission of photons-a result that provides an accurate value of $\alpha$ that is virtually independent of the electron magnetic moment anomaly.
Theoretical advances include improvements in certain aspects of the theory of the energy levels of hydrogen
and deuterium; improvements in the theory of antiprotonic helium transition frequencies that, together with the new transition frequency measurements, have led to the aforementioned competitive value of $A_{\mathrm{r}}(\mathrm{e})$; a new theoretical expression for $a_{\mathrm{e}}$ that, together with the new experimental value of $a_{\mathrm{e}}$, has led to the aforementioned value of $\alpha$; improvements in the theory for the $g$-factor of the bound electron in hydrogenic ions with nuclear spin quantum number $i=0$ relevant to the determination of $A_{\mathrm{r}}(\mathrm{e})$; and improved theory of the ground-state hyperfine splitting of muonium $\Delta \nu_{\mathrm{Mu}}$ (the $\mu^{+} \mathrm{e}^{-}$atom).

Section XII describes the analysis of the data, with the exception of the Newtonian constant of gravitation, which is analyzed in Sec. X. The consistency of the data and potential impact on the determination of the 2006 recommended values were appraised by comparing measured values of the same quantity, comparing measured values of different quantities through inferred values of a third quantity such as $\alpha$ or $h$, and finally by using the method of least squares. Based on these investigations, the final set of input data used in the 2006 adjustment was selected.

Section XIII provides, in several tables, the 2006 CODATA recommended values of the basic constants and conversion factors of physics and chemistry, including the covariance matrix of a selected group of constants.

Section XIV concludes the paper with a comparison of the 2006 and 2002 recommended values of the constants, a survey of implications for physics and metrology of the 2006 values and adjustment, and suggestions for future work that can advance our knowledge of the values of the constants.

## II. SPECIAL QUANTITIES AND UNITS

Table I lists those quantities whose numerical values are exactly defined. In the International System of Units (SI) (BIPM, 2006), used throughout this paper, the definition of the meter fixes the speed of light in vacuum $c$, the definition of the ampere fixes the magnetic constant (also called the permeability of vacuum) $\mu_{0}$, and the definition of the mole fixes the molar mass of the carbon 12 atom $M\left({ }^{12} \mathrm{C}\right)$ to have the exact values given in the table. Since the electric constant (also called the permittivity of vacuum) is related to $\mu_{0}$ by $\epsilon_{0}=1 / \mu_{0} c^{2}$, it too is known exactly.

The relative atomic mass $A_{\mathrm{r}}(X)$ of an entity $X$ is defined by $A_{\mathrm{r}}(X)=m(X) / m_{\mathrm{u}}$, where $m(X)$ is the mass of $X$ and $m_{\mathrm{u}}$ is the atomic mass constant defined by

$$
\begin{equation*}
m_{\mathrm{u}}=\frac{1}{12} m\left({ }^{12} \mathrm{C}\right)=1 \mathrm{u} \approx 1.66 \times 10^{-27} \mathrm{~kg} \tag{1}
\end{equation*}
$$

where $m\left({ }^{12} \mathrm{C}\right)$ is the mass of the carbon 12 atom and u is the unified atomic mass unit (also called the dalton, Da ). Clearly, $A_{\mathrm{r}}(X)$ is a dimensionless quantity and $A_{\mathrm{r}}\left({ }^{12} \mathrm{C}\right)$ $=12$ exactly. The molar mass $M(X)$ of entity $X$, which is the mass of one mole of $X$ with SI unit $\mathrm{kg} / \mathrm{mol}$, is given by

TABLE I. Some exact quantities relevant to the 2006 adjustment.

| Quantity | Symbol | Value |
| :--- | :--- | :--- |
| speed of light in vacuum | $c, c_{0}$ | $299792458 \mathrm{~m} \mathrm{~s}^{-1}$ |
| magnetic constant | $\mu_{0}$ | $4 \pi \times 10^{-7} \mathrm{~N} \mathrm{~A}^{-2}=12.566370614 \ldots \times 10^{-7} \mathrm{~N} \mathrm{~A}^{-2}$ |
| electric constant | $\epsilon_{0}$ | $\left(\mu_{0} \mathrm{c}^{2}\right)^{-1}=8.854187817 \ldots \times 10^{-12} \mathrm{~F} \mathrm{~m}^{-1}$ |
| relative atomic mass of ${ }^{12} \mathrm{C}$ | $A_{\mathrm{r}}\left({ }^{12} \mathrm{C}\right)$ | 12 |
| molar mass constant | $M_{\mathrm{u}}$ | $10^{-3} \mathrm{~kg} \mathrm{~mol}^{-1}$ |
| molar mass of ${ }^{12} \mathrm{C} A_{\mathrm{r}}\left({ }^{12} \mathrm{C}\right) M_{\mathrm{u}}$ | $M\left({ }^{12} \mathrm{C}\right)$ | $12 \times 10^{-3} \mathrm{~kg} \mathrm{~mol}^{-1}$ |
| conventional value of Josephson constant | $K_{\mathrm{J}-90}$ | $483597.9 \mathrm{GHz} \mathrm{V}^{-1}$ |
| conventional value of von Klitzing constant | $R_{\mathrm{K}-90}$ | $25812.807 \Omega$ |

$$
\begin{equation*}
M(X)=N_{\mathrm{A}} m(X)=A_{\mathrm{r}}(X) M_{\mathrm{u}} \tag{2}
\end{equation*}
$$

where $N_{\mathrm{A}} \approx 6.02 \times 10^{23} / \mathrm{mol}$ is the Avogadro constant and $M_{\mathrm{u}}=10^{-3} \mathrm{~kg} / \mathrm{mol}$ is the molar mass constant. The numerical value of $N_{\mathrm{A}}$ is the number of entities in one mole, and since the definition of the mole states that one mole contains the same number of entities as there are in 0.012 kg of carbon $12, M\left({ }^{12} \mathrm{C}\right)=0.012 \mathrm{~kg} / \mathrm{mol}$ exactly.

The Josephson and quantum Hall effects have played and continue to play important roles in adjustments of the values of the constants, because the Josephson and von Klitzing constants $K_{\mathrm{J}}$ and $R_{\mathrm{K}}$, which underlie these two effects, are related to $e$ and $h$ by

$$
\begin{equation*}
K_{\mathrm{J}}=\frac{2 e}{h}, \quad R_{\mathrm{K}}=\frac{h}{e^{2}}=\frac{\mu_{0} c}{2 \alpha} . \tag{3}
\end{equation*}
$$

Although we assume these relations are exact, and no evidence-either theoretical or experimental-has been put forward that challenges this assumption, the consequences of relaxing it are explored in Sec. XII.B.2. Some references to recent work related to the Josephson and quantum Hall effects may be found in the FCDC bibliographic database (see Sec. I.B).
The next-to-last two entries in Table I are the conventional values of the Josephson and von Klitzing constants adopted by the International Committee for Weights and Measures (CIPM) and introduced on January 1, 1990 to establish worldwide uniformity in the measurement of electrical quantities. In this paper, all electrical quantities are expressed in SI units. However, those measured in terms of the Josephson and quantum Hall effects with the assumption that $K_{\mathrm{J}}$ and $R_{\mathrm{K}}$ have these conventional values are labeled with a subscript 90.

For high-accuracy experiments involving the force of gravity, such as the watt balance, an accurate measurement of the local acceleration of free fall at the site of the experiment is required. Fortunately, portable and easy-to-use commercial absolute gravimeters are available that can provide a local value of $g$ with a relative standard uncertainty of a few parts in $10^{9}$. That these instruments can achieve such a small uncertainty if properly used is demonstrated by a periodic international comparison of absolute gravimeters (ICAG) carried out at the International Bureau of Weights and Measures
(BIPM), Sèvres, France; the seventh and most recent, denoted ICAG-2005, was completed in September 2005 (Vitushkin, 2007); the next is scheduled for 2009. In the future, atom interferometry or Bloch oscillations using ultracold atoms could provide a competitive or possibly more accurate method for determining a local value of $g$ (Peters et al., 2001; McGuirk et al., 2002; Cladé et al., 2005).

## III. RELATIVE ATOMIC MASSES

Included in the set of adjusted constants are the relative atomic masses $A_{\mathrm{r}}(X)$ of a number of particles, atoms, and ions. Tables II-VI and the following sections summarize the relevant data.

## A. Relative atomic masses of atoms

Most values of the relative atomic masses of neutral atoms used in this adjustment are taken from the 2003 atomic mass evaluation (AME2003) of the Atomic Mass Data Center, Centre de Spectrométrie Nucléaire et de Spectrométrie de Masse (CSNSM), Orsay, France (Audi et al., 2003; Wapstra et al., 2003; AMDC, 2006). The results of AME2003 supersede those of both the 1993 atomic mass evaluation and the 1995 update. Table II lists the values from AME2003 of interest here, while Table III gives the covariance for hydrogen and deuterium (AMDC, 2003). Other non-negligible covariances of these values are discussed in the appropriate sections.

Table IV gives six values of $A_{\mathrm{r}}(X)$ relevant to the 2006 adjustment reported since the completion and publication of AME2003 in late 2003 that we use in place of the corresponding values in Table II.

The ${ }^{3} \mathrm{H}$ and ${ }^{3} \mathrm{He}$ values are those reported by the SMILETRAP group at the Manne Siegbahn Laboratory (MSL), Stockholm, Sweden (Nagy et al., 2006) using a Penning trap and a time-of-flight technique to detect cyclotron resonances. This new ${ }^{3} \mathrm{He}$ result is in good agreement with a more accurate, but still preliminary, result from the University of Washington group in Seattle, WA, USA (Van Dyck, 2006). The AME2003 values for ${ }^{3} \mathrm{H}$ and ${ }^{3} \mathrm{He}$ were influenced by an earlier result for

TABLE II. Values of the relative atomic masses of the neutron and various atoms as given in the 2003 atomic mass evaluation together with the defined value for ${ }^{12} \mathrm{C}$.

| Atom | Relative atomic mass $A_{\mathrm{r}}(X)$ | Relative standard uncertainty $u_{\mathrm{r}}$ |
| :---: | :---: | :---: |
| n | $1.00866491574(56)$ | $5.6 \times 10^{-10}$ |
| ${ }^{1} \mathrm{H}$ | $1.00782503207(10)$ | $1.0 \times 10^{-10}$ |
| ${ }^{2} \mathrm{H}$ | 2.014101777 85(36) | $1.8 \times 10^{-10}$ |
| ${ }^{3} \mathrm{H}$ | 3.016049 2777(25) | $8.2 \times 10^{-10}$ |
| ${ }^{3} \mathrm{He}$ | $3.0160293191(26)$ | $8.6 \times 10^{-10}$ |
| ${ }^{4} \mathrm{He}$ | $4.002603254153(63)$ | $1.6 \times 10^{-11}$ |
| ${ }^{12} \mathrm{C}$ | 12 | (Exact) |
| ${ }^{16} \mathrm{O}$ | 15.994914619 56(16) | $1.0 \times 10^{-11}$ |
| ${ }^{28} \mathrm{Si}$ | $27.9769265325(19)$ | $6.9 \times 10^{-11}$ |
| ${ }^{29} \mathrm{Si}$ | $28.976494700(22)$ | $7.6 \times 10^{-10}$ |
| ${ }^{30} \mathrm{Si}$ | $29.973770171(32)$ | $1.1 \times 10^{-9}$ |
| ${ }^{36} \mathrm{Ar}$ | $35.967545105(28)$ | $7.8 \times 10^{-10}$ |
| ${ }^{38} \mathrm{Ar}$ | $37.96273239(36)$ | $9.5 \times 10^{-9}$ |
| ${ }^{40} \mathrm{Ar}$ | $39.9623831225(29)$ | $7.2 \times 10^{-11}$ |
| ${ }^{87} \mathrm{Rb}$ | $86.909180526(12)$ | $1.4 \times 10^{-10}$ |
| ${ }^{107} \mathrm{Ag}$ | $106.9050968(46)$ | $4.3 \times 10^{-8}$ |
| ${ }^{109} \mathrm{Ag}$ | 108.904 7523(31) | $2.9 \times 10^{-8}$ |
| ${ }^{133} \mathrm{Cs}$ | $132.905451932(24)$ | $1.8 \times 10^{-10}$ |

${ }^{3} \mathrm{He}$ from the University of Washington group, which is in disagreement with their new result.
The values for ${ }^{4} \mathrm{He}$ and ${ }^{16} \mathrm{O}$ are those reported by the University of Washington group (Van Dyck et al., 2006) using their improved mass spectrometer; they are based on a thorough reanalysis of data that yielded preliminary results for these atoms that were used in AME2003. They include an experimentally determined image-charge correction with a relative standard uncertainty $u_{\mathrm{r}}=7.9 \times 10^{-12}$ in the case of ${ }^{4} \mathrm{He}$ and $u_{\mathrm{r}}=4.0$ $\times 10^{-12}$ in the case of ${ }^{16} \mathrm{O}$. The value of $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right)$ is also from this group and is a near-final result based on the analysis of ten runs carried out over a 4 year period (Van Dyck, 2006). Because the result is not yet final, the total uncertainty is conservatively assigned; $u_{\mathrm{r}}=9.9$ $\times 10^{-12}$ for the image-charge correction. This value of $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right)$ is consistent with the preliminary value reported

TABLE III. The variances, covariance, and correlation coefficient of the AME2003 values of the relative atomic masses of hydrogen and deuterium. The number in bold above the main diagonal is $10^{18}$ times the numerical value of the covariance, the numbers in bold on the main diagonal are $10^{18}$ times the numerical values of the variances, and the number in italics below the main diagonal is the correlation coefficient.

|  | $A_{\mathrm{r}}\left({ }^{1} \mathrm{H}\right)$ | $\left.A_{\mathrm{r}}{ }^{2} \mathrm{H}\right)$ |
| :--- | :---: | :---: |
| $\left.A_{\mathrm{r}}{ }^{1} \mathrm{H}\right)$ | $\mathbf{0 . 0 1 0 7}$ | $\mathbf{0 . 0 0 2 7}$ |
| $\left.A_{\mathrm{r}}{ }^{2} \mathrm{H}\right)$ | 0.0735 | $\mathbf{0 . 1 2 7 2}$ |

TABLE IV. Values of the relative atomic masses of various atoms that have become available since the 2003 atomic mass evaluation.

| Atom | Relative atomic <br> mass $A_{\mathrm{r}}(X)$ | Relative standard <br> uncertainty $u_{\mathrm{r}}$ |
| :--- | :---: | :---: |
| ${ }^{2} \mathrm{H}$ | $2.014101778040(80)$ | $4.0 \times 10^{-11}$ |
| ${ }^{3} \mathrm{H}$ | $3.0160492787(25)$ | $8.3 \times 10^{-10}$ |
| ${ }^{3} \mathrm{He}$ | $3.0160293217(26)$ | $8.6 \times 10^{-10}$ |
| ${ }^{4} \mathrm{He}$ | $4.002603254131(62)$ | $1.5 \times 10^{-11}$ |
| ${ }^{16} \mathrm{O}$ | $15.99491461957(18)$ | $1.1 \times 10^{-11}$ |
| ${ }^{29} \mathrm{Si}$ | $28.9764946625(20)$ | $6.9 \times 10^{-11}$ |

by Van Dyck et al. (2006) based on the analysis of only three runs.
The covariance and correlation coefficient of $\left.A_{\mathrm{r}}{ }^{3} \mathrm{H}\right)$ and $A_{\mathrm{r}}\left({ }^{3} \mathrm{He}\right)$ given in Table V are due to the common component of uncertainty $u_{\mathrm{r}}=1.4 \times 10^{-10}$ of the relative atomic mass of the $\mathrm{H}_{2}{ }^{+}$reference ion used in the SMILETRAP measurements; the covariances and correlation coefficients of the University of Washington values of $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right), A_{\mathrm{r}}\left({ }^{4} \mathrm{He}\right)$, and $\left.A_{\mathrm{r}}{ }^{16} \mathrm{O}\right)$ given in Table VI are due to the uncertainties of the image-charge corrections, which are based on the same experimentally determined relation.
The ${ }^{29} \mathrm{Si}$ value is that implied by the ratio $A_{\mathrm{r}}\left({ }^{29} \mathrm{Si}^{+}\right) / A_{\mathrm{r}}\left({ }^{28} \mathrm{Si} \mathrm{H}^{+}\right)=0.9997151241812(65)$ obtained at the Massachusetts Institute of Technology (MIT), Cambridge, MA, USA, using a recently developed technique of determining mass ratios by directly comparing the cyclotron frequencies of two different ions simultaneously confined in a Penning trap (Rainville et al., 2005). (The relative atomic mass work of the MIT group has now been transferred to Florida State University, Tallahassee, FL, USA.) This approach eliminates many components of uncertainty arising from systematic effects. The value for $\left.A_{\mathrm{r}}{ }^{(29} \mathrm{Si}\right)$ is given in the supplementary information to Rainville et al. (2005) and has a significantly smaller uncertainty than the corresponding AME2003 value.

## B. Relative atomic masses of ions and nuclei

The relative atomic mass $A_{\mathrm{r}}(X)$ of a neutral atom $X$ is given in terms of the relative atomic mass of an ion of the atom formed by the removal of $n$ electrons by

TABLE V. The variances, covariance, and correlation coefficient of the values of the SMILETRAP relative atomic masses of tritium and helium 3. The number in bold above the main diagonal is $10^{18}$ times the numerical value of the covariance, the numbers in bold on the main diagonal are $10^{18}$ times the numerical values of the variances, and the number in italics below the main diagonal is the correlation coefficient.

|  | $A_{\mathrm{r}}\left({ }^{3} \mathrm{H}\right)$ | $A_{\mathrm{r}}\left({ }^{3} \mathrm{He}\right)$ |
| :--- | :---: | :---: |
| $\left.A_{\mathrm{r}}{ }^{3} \mathrm{H}\right)$ | $\mathbf{6 . 2 5 0 0}$ | $\mathbf{0 . 1 7 8 3}$ |
| $A_{\mathrm{r}}\left({ }^{3} \mathrm{He}\right)$ | 0.0274 | $\mathbf{6 . 7 6 0 0}$ |

TABLE VI. The variances, covariances, and correlation coefficients of the University of Washington values of the relative atomic masses of deuterium, helium 4, and oxygen 16. The numbers in bold above the main diagonal are $10^{20}$ times the numerical values of the covariances, the numbers in bold on the main diagonal are $10^{20}$ times the numerical values of the variances, and the numbers in italics below the main diagonal are the correlation coefficients.

|  | $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right)$ | $A_{\mathrm{r}}\left({ }^{4} \mathrm{He}\right)$ | $A_{\mathrm{r}}\left({ }^{16} \mathrm{O}\right)$ |
| :--- | :---: | :---: | :---: |
| $\left.A_{\mathrm{r}}{ }^{2} \mathrm{H}\right)$ | $\mathbf{0 . 6 4 0 0}$ | $\mathbf{0 . 0 6 3 1}$ | $\mathbf{0 . 1 2 7 6}$ |
| $\left.A_{\mathrm{r}}{ }^{4} \mathrm{He}\right)$ | 0.1271 | $\mathbf{0 . 3 8 4 4}$ | $\mathbf{0 . 2 0 2 3}$ |
| $\left.A_{\mathrm{r}}{ }^{16} \mathrm{O}\right)$ | 0.0886 | 0.1813 | $\mathbf{3 . 2 4 0 0}$ |

$$
\begin{equation*}
A_{\mathrm{r}}(X)=A_{\mathrm{r}}\left(X^{n+}\right)+n A_{\mathrm{r}}(\mathrm{e})-\frac{E_{\mathrm{b}}(X)-E_{\mathrm{b}}\left(X^{n+}\right)}{m_{\mathrm{u}} c^{2}} \tag{4}
\end{equation*}
$$

Here $E_{\mathrm{b}}(X) / m_{\mathrm{u}} c^{2}$ is the relative-atomic-mass equivalent of the total binding energy of the $Z$ electrons of the atom, where $Z$ is the atomic number (proton number), and $E_{\mathrm{b}}\left(X^{n+}\right) / m_{\mathrm{u}} c^{2}$ is the relative-atomic-mass equivalent of the binding energy of the $Z-n$ electrons of the $X^{n+}$ ion. For a fully stripped atom, that is, for $n=Z, X^{Z+}$ is $N$, where $N$ represents the nucleus of the atom, and $E_{\mathrm{b}}\left(X^{Z+}\right) / m_{\mathrm{u}} c^{2}=0$, which yields the first few equations of Table XL in Sec. XII.B.

The binding energies $E_{\mathrm{b}}$ used in this work are the same as those used in the 2002 adjustment; see Table IV of CODATA-02. For tritium, which is not included there, we use the value $1.097185439 \times 10^{7} \mathrm{~m}^{-1}$ (Kotochigova, 2006). The uncertainties of the binding energies are negligible for our application.

## C. Cyclotron resonance measurement of the electron relative atomic mass $\boldsymbol{A}_{\mathrm{r}}(\mathrm{e})$

A value of $A_{\mathrm{r}}(\mathrm{e})$ is available from a Penning-trap measurement carried out by the University of Washington group (Farnham et al., 1995); it is used as an input datum in the 2006 adjustment, as it was in the 2002 adjustment:

$$
\begin{equation*}
A_{\mathrm{r}}(\mathrm{e})=0.0005485799111(12) \quad\left[2.1 \times 10^{-9}\right] \tag{5}
\end{equation*}
$$

## IV. ATOMIC TRANSITION FREQUENCIES

Atomic transition frequencies in hydrogen, deuterium, and antiprotonic helium yield information on the Rydberg constant, the proton and deuteron charge radii, and the relative atomic mass of the electron. The hyperfine splitting in hydrogen and fine-structure splitting in helium do not yield a competitive value of any constant at the current level of accuracy of the relevant experiment and/or theory. All of these topics are discussed in this section.
A. Hydrogen and deuterium transition frequencies, the Rydberg constant $\boldsymbol{R}_{\infty}$, and the proton and deuteron charge radii $R_{\mathrm{p}}, R_{\mathrm{d}}$

The Rydberg constant is related to other constants by the definition

$$
\begin{equation*}
R_{\infty}=\alpha^{2} \frac{m_{\mathrm{e}} c}{2 h} . \tag{6}
\end{equation*}
$$

It can be accurately determined by comparing measured resonant frequencies of transitions in hydrogen $(\mathrm{H})$ and deuterium (D) to the theoretical expressions for the energy level differences in which it is a multiplicative factor.

## 1. Theory relevant to the Rydberg constant

The theory of the energy levels of hydrogen and deuterium atoms relevant to the determination of the Rydberg constant $R_{\infty}$, based on measurements of transition frequencies, is summarized in this section. Complete information necessary to determine the theoretical values of the relevant energy levels is provided, with an emphasis on results that have become available since the previous adjustment described in CODATA-02. For brevity, references to earlier work, which can be found in Eides et al. (2001b) for example, are not included here.

An important consideration is that the theoretical values of the energy levels of different states are highly correlated. For example, for S states, the uncalculated terms are primarily of the form of an unknown common constant divided by $n^{3}$. This fact is taken into account by calculating covariances between energy levels in addition to the uncertainties of the individual levels as discussed in detail in Sec. IV.A.1.1. In order to take these correlations into account, we distinguish between components of uncertainty that are proportional to $1 / n^{3}$, denoted by $u_{0}$, and components of uncertainty that are essentially random functions of $n$, denoted by $u_{n}$.

The energy levels of hydrogenlike atoms are determined mainly by the Dirac eigenvalue, QED effects such as self energy and vacuum polarization, and nuclear size and motion effects, all of which are summarized in the following sections.

## a. Dirac eigenvalue

The binding energy of an electron in a static Coulomb field (the external electric field of a point nucleus of charge $Z e$ with infinite mass) is determined predominantly by the Dirac eigenvalue

$$
\begin{equation*}
E_{\mathrm{D}}=f(n, j) m_{\mathrm{e}} c^{2} \tag{7}
\end{equation*}
$$

where

$$
\begin{equation*}
f(n, j)=\left[1+\frac{(Z \alpha)^{2}}{(n-\delta)^{2}}\right]^{-1 / 2} \tag{8}
\end{equation*}
$$

$n$ and $j$ are the principal quantum number and total angular momentum of the state, respectively, and

$$
\begin{equation*}
\delta=j+\frac{1}{2}-\left[\left(j+\frac{1}{2}\right)^{2}-(Z \alpha)^{2}\right]^{1 / 2} \tag{9}
\end{equation*}
$$

Although we are interested only in the case in which the nuclear charge is $e$, we retain the atomic number $Z$ in order to indicate the nature of various terms.

Corrections to the Dirac eigenvalue that approximately take into account the finite mass of the nucleus $m_{\mathrm{N}}$ are included in the more general expression for atomic energy levels, which replaces Eq. (7) (Barker and Glover, 1955; Sapirstein and Yennie, 1990),

$$
\begin{align*}
E_{M}= & M c^{2}+[f(n, j)-1] m_{\mathrm{r}} c^{2}-[f(n, j)-1]^{2} \frac{m_{\mathrm{r}}^{2} c^{2}}{2 M} \\
& +\frac{1-\delta_{l 0}}{\kappa(2 l+1)} \frac{(Z \alpha)^{4} m_{\mathrm{r}}^{3} c^{2}}{2 n^{3} m_{\mathrm{N}}^{2}}+\cdots, \tag{10}
\end{align*}
$$

where $l$ is the nonrelativistic orbital angular momentum quantum number, $\kappa$ is the angular-momentum-parity quantum number $\kappa=(-1)^{j-l+1 / 2}\left(j+\frac{1}{2}\right), \quad M=m_{\mathrm{e}}+m_{\mathrm{N}}$, and $m_{\mathrm{r}}=m_{\mathrm{e}} m_{\mathrm{N}} /\left(m_{\mathrm{e}}+m_{\mathrm{N}}\right)$ is the reduced mass.

## b. Relativistic recoil

Relativistic corrections to Eq. (10) associated with motion of the nucleus are considered relativistic-recoil corrections. The leading term, to lowest order in $Z \alpha$ and all orders in $m_{\mathrm{e}} / m_{\mathrm{N}}$, is (Erickson, 1977; Sapirstein and Yennie, 1990)

$$
\begin{align*}
E_{\mathrm{S}}= & \frac{m_{\mathrm{r}}^{3}}{m_{\mathrm{e}}^{2} m_{\mathrm{N}}} \frac{(Z \alpha)^{5}}{\pi n^{3}} m_{\mathrm{e}} c^{2}\left\{\frac{1}{3} \delta_{l 0} \ln (Z \alpha)^{-2}-\frac{8}{3} \ln k_{0}(n, l)\right. \\
& -\frac{1}{9} \delta_{l 0}-\frac{7}{3} a_{n}-\frac{2}{m_{\mathrm{N}}^{2}-m_{\mathrm{e}}^{2}} \delta_{l 0} \\
& \left.\times\left[m_{\mathrm{N}}^{2} \ln \left(\frac{m_{\mathrm{e}}}{m_{\mathrm{r}}}\right)-m_{\mathrm{e}}^{2} \ln \left(\frac{m_{\mathrm{N}}}{m_{\mathrm{r}}}\right)\right]\right\} \tag{11}
\end{align*}
$$

where

$$
\begin{equation*}
a_{n}=-2\left[\ln \left(\frac{2}{n}\right)+\sum_{i=1}^{n} \frac{1}{i}+1-\frac{1}{2 n}\right] \delta_{l 0}+\frac{1-\delta_{l 0}}{l(l+1)(2 l+1)} \tag{12}
\end{equation*}
$$

To lowest order in the mass ratio, higher-order corrections in $Z \alpha$ have been extensively investigated; the contribution of the next two orders in $Z \alpha$ is

$$
\begin{align*}
E_{\mathrm{R}}= & \frac{m_{\mathrm{e}}}{m_{\mathrm{N}}} \frac{(Z \alpha)^{6}}{n^{3}} m_{\mathrm{e}} c^{2} \\
& \times\left[D_{60}+D_{72} Z \alpha \ln ^{2}(Z \alpha)^{-2}+\cdots\right] \tag{13}
\end{align*}
$$

where for $n \mathrm{~S}_{1 / 2}$ states (Pachucki and Grotch, 1995; Eides and Grotch, 1997c)

$$
\begin{equation*}
D_{60}=4 \ln 2-\frac{7}{2} \tag{14}
\end{equation*}
$$

and (Melnikov and Yelkhovsky, 1999; Pachucki and Karshenboim, 1999)

$$
\begin{equation*}
D_{72}=-\frac{11}{60 \pi} \tag{15}
\end{equation*}
$$

and for states with $l \geqslant 1$ (Golosov et al., 1995; Elkhovski1̆, 1996; Jentschura and Pachucki, 1996)

$$
\begin{equation*}
D_{60}=\left[3-\frac{l(l+1)}{n^{2}}\right] \frac{2}{\left(4 l^{2}-1\right)(2 l+3)} . \tag{16}
\end{equation*}
$$

In Eq. (16) and subsequent discussion, the first subscript on the coefficient of a term refers to the power of $Z \alpha$ and the second subscript to the power of $\ln (Z \alpha)^{-2}$. The relativistic recoil correction used in the 2006 adjustment is based on Eqs. (11)-(16). The estimated uncertainty for S states is taken to be $10 \%$ of Eq. (13), and for states with $l \geqslant 1$ it is taken to be $1 \%$ of that equation.

Numerical values for the complete contribution of Eq. (13) to all orders in $Z \alpha$ have been obtained by Shabaev et al. (1998). Although the difference between the allorders calculation and the truncated power series for $S$ states is about three times their quoted uncertainty, the two results are consistent within the uncertainty assigned here. The covariances of the theoretical values are calculated by assuming that the uncertainties are predominately due to uncalculated terms proportional to $\left(m_{\mathrm{e}} / m_{\mathrm{N}}\right) / n^{3}$.

## c. Nuclear polarization

Interactions between the atomic electron and the nucleus which involve excited states of the nucleus give rise to nuclear polarization corrections. For hydrogen, we use the result (Khriplovich and Sen'kov, 2000)

$$
\begin{equation*}
E_{\mathrm{P}}(\mathrm{H})=-0.070(13) h \frac{\delta_{l 0}}{n^{3}} \mathrm{kHz} \tag{17}
\end{equation*}
$$

For deuterium, the sum of the proton polarizability, the neutron polarizability (Khriplovich and Sen'kov, 1998), and the dominant nuclear structure polarizability (Friar and Payne, 1997a) gives

$$
\begin{equation*}
E_{\mathrm{P}}(\mathrm{D})=-21.37(8) h \frac{\delta_{l 0}}{n^{3}} \mathrm{kHz} \tag{18}
\end{equation*}
$$

We assume that this effect is negligible in states of higher $l$.

## d. Self energy

The one-photon electron self energy is given by

$$
\begin{equation*}
E_{\mathrm{SE}}^{(2)}=\frac{\alpha}{\pi} \frac{(Z \alpha)^{4}}{n^{3}} F(Z \alpha) m_{\mathrm{e}} c^{2} \tag{19}
\end{equation*}
$$

where

$$
\begin{align*}
F(Z \alpha)= & A_{41} \ln (Z \alpha)^{-2}+A_{40}+A_{50}(Z \alpha) \\
& +A_{62}(Z \alpha)^{2} \ln ^{2}(Z \alpha)^{-2}+A_{61}(Z \alpha)^{2} \ln (Z \alpha)^{-2} \\
& +G_{\mathrm{SE}}(Z \alpha)(Z \alpha)^{2} \tag{20}
\end{align*}
$$

From Erickson and Yennie (1965) and earlier papers cited therein,

TABLE VII. Bethe logarithms $\ln k_{0}(n, l)$ relevant to the determination of $R_{\infty}$.

| $n$ | S | P | D |
| :--- | :---: | :---: | :---: |
| 1 | 2.984128556 |  |  |
| 2 | 2.811769893 | -0.030016709 |  |
| 3 | 2.767663612 |  |  |
| 4 | 2.749811840 | -0.041954895 | -0.006740939 |
| 6 | 2.735664207 |  | -0.008147204 |
| 8 | 2.730267261 |  | -0.008785043 |
| 12 |  | -0.009342954 |  |

$$
\begin{align*}
A_{41}= & \frac{4}{3} \delta_{l 0}, \\
A_{40}= & -\frac{4}{3} \ln k_{0}(n, l)+\frac{10}{9} \delta_{l 0}-\frac{1}{2 \kappa(2 l+1)}\left(1-\delta_{l 0}\right), \\
A_{50}= & \left(\frac{139}{32}-2 \ln 2\right) \pi \delta_{l 0}, \\
A_{62}= & -\delta_{l 0}, \\
A_{61}= & {\left[4\left(1+\frac{1}{2}+\cdots+\frac{1}{n}\right)+\frac{28}{3} \ln 2-4 \ln n-\frac{601}{180}\right.} \\
& \left.-\frac{77}{45 n^{2}}\right] \delta_{l 0}+\left(1-\frac{1}{n^{2}}\right)\left(\frac{2}{15}+\frac{1}{3} \delta_{j 1 / 2}\right) \delta_{l 1} \\
& +\frac{96 n^{2}-32 l(l+1)}{3 n^{2}(2 l-1)(2 l)(2 l+1)(2 l+2)(2 l+3)}\left(1-\delta_{l 0}\right) . \tag{21}
\end{align*}
$$

The Bethe logarithms $\ln k_{0}(n, l)$ in Eq. (21) are given in Table VII (Drake and Swainson, 1990).

The function $G_{\text {SE }}(Z \alpha)$ in Eq. (20) is the higher-order contribution (in $Z \alpha$ ) to the self energy, and the values for $G_{\text {SE }}(\alpha)$ that we use here are listed in Table VIII. For S and P states with $n \leqslant 4$, the values in the table are based on direct numerical evaluations by Jentschura and Mohr $(2004,2005)$ and Jentschura et al. $(1999,2001)$. The values of $G_{\text {SE }}(\alpha)$ for the 6 S and 8 S states are based on the low- $Z$ limit of this function $G_{\text {SE }}(0)=A_{60}$ (Jentschura,

Czarnecki, and Pachucki, 2005) together with extrapolations of complete numerical calculation results of $F(Z \alpha)$ [see Eq. (20)] at higher $Z$ (Kotochigova and Mohr, 2006). The values of $G_{\mathrm{SE}}(\alpha)$ for D states are from Jentschura, Kotochigova, Le Bigot, et al. (2005).

The dominant effect of the finite mass of the nucleus on the self energy correction is taken into account by multiplying each term of $F(Z \alpha)$ by the reduced-mass factor $\left(m_{\mathrm{r}} / m_{\mathrm{e}}\right)^{3}$, except that the magnetic moment term $-1 /[2 \kappa(2 l+1)]$ in $A_{40}$ is multiplied instead by the factor $\left(m_{\mathrm{r}} / m_{\mathrm{e}}\right)^{2}$. In addition, the argument $(Z \alpha)^{-2}$ of the logarithms is replaced by $\left(m_{\mathrm{e}} / m_{\mathrm{r}}\right)(Z \alpha)^{-2}$ (Sapirstein and Yennie, 1990).

The uncertainty of the self energy contribution to a given level arises entirely from the uncertainty of $G_{\text {SE }}(\alpha)$ listed in Table VIII and is taken to be entirely of type $u_{n}$.
e. Vacuum polarization

The second-order vacuum-polarization level shift is

$$
\begin{equation*}
E_{\mathrm{VP}}^{(2)}=\frac{\alpha}{\pi} \frac{(Z \alpha)^{4}}{n^{3}} H(Z \alpha) m_{\mathrm{e}} c^{2} \tag{22}
\end{equation*}
$$

where the function $H(Z \alpha)$ is divided into the part corresponding to the Uehling potential, denoted here by $H^{(1)}(Z \alpha)$, and the higher-order remainder $H^{(\mathrm{R})}(Z \alpha)$, where

$$
\begin{align*}
H^{(1)}(Z \alpha)= & V_{40}+V_{50}(Z \alpha)+V_{61}(Z \alpha)^{2} \ln (Z \alpha)^{-2} \\
& +G_{\mathrm{VP}}^{(1)}(Z \alpha)(Z \alpha)^{2},  \tag{23}\\
H^{(\mathrm{R})}(Z \alpha)= & G_{\mathrm{VP}}^{(\mathrm{R})}(Z \alpha)(Z \alpha)^{2}, \tag{24}
\end{align*}
$$

with

$$
\begin{align*}
& V_{40}=-\frac{4}{15} \delta_{l 0} \\
& V_{50}=\frac{5}{48} \pi \delta_{l 0} \\
& V_{61}=-\frac{2}{15} \delta_{l 0} . \tag{25}
\end{align*}
$$

The part $G_{\mathrm{VP}}^{(1)}(Z \alpha)$ arises from the Uehling potential with values given in Table IX (Mohr, 1982; Kotochigova et al., 2002). The higher-order remainder $G_{\mathrm{VP}}^{(\mathrm{R})}(Z \alpha)$ has been considered by Wichmann and Kroll, and the leading terms in powers of $Z \alpha$ are (Wichmann and Kroll, 1956; Mohr, 1975, 1983)

TABLE VIII. Values of the function $G_{\text {SE }}(\alpha)$.

| $n$ | $\mathrm{~S}_{1 / 2}$ | $\mathrm{P}_{1 / 2}$ | $\mathrm{P}_{3 / 2}$ | $\mathrm{D}_{3 / 2}$ | $\mathrm{D}_{5 / 2}$ |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 1 | $-30.290240(20)$ |  |  |  |  |
| 2 | $-31.185150(90)$ | $-0.97350(20)$ | $-0.48650(20)$ |  |  |
| 3 | $-31.04770(90)$ |  |  |  | $0.03163(22)$ |
| 4 | $-30.9120(40)$ | $-1.1640(20)$ | $-0.6090(20)$ |  | $0.03417(26)$ |
| 6 | $-30.711(47)$ |  |  | $0.007940(90)$ | $0.03484(22)$ |
| 8 | $-30.606(47)$ |  |  | $0.0080(20)$ | $0.0350(30)$ |
| 12 |  |  |  |  |  |

TABLE IX. Values of the function $G_{V \mathrm{~V}}^{(1)}(\alpha)$.

| $n$ | $\mathrm{~S}_{1 / 2}$ | $\mathrm{P}_{1 / 2}$ | $\mathrm{P}_{3 / 2}$ | $\mathrm{D}_{3 / 2}$ | $\mathrm{D}_{5 / 2}$ |
| :--- | :---: | :---: | :---: | :---: | :---: |
| 1 | -0.618724 |  |  |  |  |
| 2 | -0.808872 | -0.064006 | -0.014132 |  |  |
| 3 | -0.814530 |  |  |  | -0.000000 |
| 4 | -0.806579 | -0.080007 | -0.017666 |  | -0.000000 |
| 6 | -0.791450 |  |  | -0.000000 | -0.000000 |
| 8 | -0.781197 |  | -0.000000 | -0.000000 |  |
| 12 |  |  |  |  |  |

$$
\begin{align*}
G_{\mathrm{VP}}^{(\mathrm{R})}(Z \alpha)= & \left(\frac{19}{45}-\frac{\pi^{2}}{27}\right) \delta_{l 0}+\left(\frac{1}{16}-\frac{31 \pi^{2}}{2880}\right) \pi(Z \alpha) \delta_{l 0} \\
& +\cdots \tag{26}
\end{align*}
$$

Higher-order terms omitted from Eq. (26) are negligible.
In a manner similar to that for the self energy, the leading effect of the finite mass of the nucleus is taken into account by multiplying Eq. (22) by the factor $\left(m_{\mathrm{r}} / m_{\mathrm{e}}\right)^{3}$ and including a multiplicative factor of $\left(m_{\mathrm{e}} / m_{\mathrm{r}}\right)$ in the argument of the logarithm in Eq. (23).

There is also a second-order vacuum polarization level shift due to the creation of virtual particle pairs other than the $\mathrm{e}^{-} \mathrm{e}^{+}$pair. The predominant contribution for $n S$ states arises from $\mu^{+} \mu^{-}$, with the leading term being (Eides and Shelyuto, 1995; Karshenboim, 1995)

$$
\begin{equation*}
E_{\mu \mathrm{VP}}^{(2)}=\frac{\alpha}{\pi} \frac{(Z \alpha)^{4}}{n^{3}}\left(-\frac{4}{15}\right)\left(\frac{m_{\mathrm{e}}}{m_{\mu}}\right)^{2}\left(\frac{m_{\mathrm{r}}}{m_{\mathrm{e}}}\right)^{3} m_{\mathrm{e}} c^{2} \tag{27}
\end{equation*}
$$

The next-order term in the contribution of muon vacuum polarization to $n S$ states is of relative order $Z \alpha m_{\mathrm{e}} / m_{\mu}$ and is therefore negligible. The analogous contribution $E_{\tau \vee \mathrm{P}}^{(2)}$ from $\tau^{+} \tau^{-}(-18 \mathrm{~Hz}$ for the 1 S state $)$ is also negligible at the level of uncertainty of current interest.
For the hadronic vacuum polarization contribution, we use the result given by Friar et al. (1999) that utilizes all available $\mathrm{e}^{+} \mathrm{e}^{-}$scattering data,

$$
\begin{equation*}
E_{\mathrm{had} \mathrm{VP}}^{(2)}=0.671(15) E_{\mu \mathrm{VP}}^{(2)}, \tag{28}
\end{equation*}
$$

where the uncertainty is of type $u_{0}$.
The muonic and hadronic vacuum polarization contributions are negligible for P and D states.

## f. Two-photon corrections

Corrections from two virtual photons have been partially calculated as a power series in $Z \alpha$,

$$
\begin{equation*}
E^{(4)}=\left(\frac{\alpha}{\pi}\right)^{2} \frac{(Z \alpha)^{4}}{n^{3}} m_{\mathrm{e}} c^{2} F^{(4)}(Z \alpha) \tag{29}
\end{equation*}
$$

where

$$
\begin{align*}
F^{(4)}(Z \alpha)= & B_{40}+B_{50}(Z \alpha)+B_{63}(Z \alpha)^{2} \ln ^{3}(Z \alpha)^{-2} \\
& +B_{62}(Z \alpha)^{2} \ln ^{2}(Z \alpha)^{-2}+B_{61}(Z \alpha)^{2} \ln (Z \alpha)^{-2} \\
& +B_{60}(Z \alpha)^{2}+\cdots . \tag{30}
\end{align*}
$$

The leading term $B_{40}$ is well known,

$$
\begin{align*}
B_{40}= & {\left[\frac{3 \pi^{2}}{2} \ln 2-\frac{10 \pi^{2}}{27}-\frac{2179}{648}-\frac{9}{4} \zeta(3)\right] \delta_{l 0} } \\
& +\left[\frac{\pi^{2} \ln 2}{2}-\frac{\pi^{2}}{12}-\frac{197}{144}-\frac{3 \zeta(3)}{4}\right] \frac{1-\delta_{l 0}}{\kappa(2 l+1)} . \tag{31}
\end{align*}
$$

The second term is (Pachucki, 1993a, 1994; Eides and Shelyuto, 1995; Eides et al., 1997)

$$
\begin{equation*}
B_{50}=-21.5561(31) \delta_{l 0} \tag{32}
\end{equation*}
$$

and the next coefficient is (Karshenboim, 1993; Manohar and Stewart, 2000; Yerokhin, 2000; Pachucki, 2001)

$$
\begin{equation*}
B_{63}=-\frac{8}{27} \delta_{l 0} \tag{33}
\end{equation*}
$$

For S states, the coefficient $B_{62}$ is given by

$$
\begin{equation*}
B_{62}=\frac{16}{9}\left[\frac{71}{60}-\ln 2+\gamma+\psi(n)-\ln n-\frac{1}{n}+\frac{1}{4 n^{2}}\right] \tag{34}
\end{equation*}
$$

where $\gamma=0.577 \ldots$ is Euler's constant and $\psi$ is the psi function (Abramowitz and Stegun, 1965). The difference $B_{62}(1)-B_{62}(n)$ was calculated by Karshenboim (1996) and confirmed by Pachucki (2001), who also calculated the $n$-independent additive constant. For P states, the calculated value is (Karshenboim, 1996)

$$
\begin{equation*}
B_{62}=\frac{4}{27} \frac{n^{2}-1}{n^{2}} \tag{35}
\end{equation*}
$$

This result has been confirmed by Jentschura and Nándori (2002), who also showed that for D and higher angular momentum states $B_{62}=0$.

Recent work has led to new results for $B_{61}$ and higherorder coefficients. In the paper of Jentschura, Czarnecki, and Pachucki (2005), an additional state-independent contribution to the coefficient $B_{61}$ for S states is given, which differs slightly ( $2 \%$ ) from the earlier result of Pa chucki (2001) quoted in CODATA 2002. The revised coefficient for $S$ states is

TABLE X. Values of $N$ used in the 2006 adjustment.

| $n$ | $N(n \mathrm{~S})$ | $N(n \mathrm{P})$ |
| :--- | :---: | :---: |
| 1 | $17.85567203(1)$ |  |
| 2 | $12.03214158(1)$ | $0.003300635(1)$ |
| 3 | $10.449809(1)$ |  |
| 4 | $9.722413(1)$ | $-0.000394332(1)$ |
| 6 | $9.031832(1)$ |  |
| 8 | $8.697639(1)$ |  |

$$
\begin{align*}
B_{61}= & \frac{413581}{64800}+\frac{4 N(n S)}{3}+\frac{2027 \pi^{2}}{864}-\frac{616 \ln 2}{135} \\
& -\frac{2 \pi^{2} \ln 2}{3}+\frac{40 \ln ^{2} 2}{9}+\zeta(3)+\left(\frac{304}{135}-\frac{32 \ln 2}{9}\right) \\
& \times\left[\frac{3}{4}+\gamma+\psi(n)-\ln n-\frac{1}{n}+\frac{1}{4 n^{2}}\right] \tag{36}
\end{align*}
$$

where $\zeta$ is the Riemann zeta function (Abramowitz and Stegun, 1965). The coefficients $N(n \mathrm{~S})$ are listed in Table X . The state-dependent part $B_{61}(n \mathrm{~S})-B_{61}(1 \mathrm{~S})$ was confirmed by Jentschura, Czarnecki, and Pachucki (2005) in their Eqs. (4.26) and (6.3). For higher-l states, $B_{61}$ has been calculated by Jentschura, Czarnecki, and Pachucki (2005); for P states,

$$
\begin{align*}
& B_{61}\left(n \mathrm{P}_{1 / 2}\right)=\frac{4}{3} N(n \mathrm{P})+\frac{n^{2}-1}{n^{2}}\left(\frac{166}{405}-\frac{8}{27} \ln 2\right),  \tag{37}\\
& B_{61}\left(n \mathrm{P}_{3 / 2}\right)=\frac{4}{3} N(n \mathrm{P})+\frac{n^{2}-1}{n^{2}}\left(\frac{31}{405}-\frac{8}{27} \ln 2\right), \tag{38}
\end{align*}
$$

and for D states,

$$
\begin{equation*}
B_{61}(n \mathrm{D})=0 . \tag{39}
\end{equation*}
$$

The coefficient $B_{61}$ also vanishes for states with $l>2$. The necessary values of $N(n \mathrm{P})$ are given in Eq. (17) of Jentschura (2003) and are listed in Table X.
The next term is $B_{60}$, and recent work has also been done for this contribution. For S states, the state dependence is considered first, and is given by Czarnecki et al. (2005) and Jentschura, Czarnecki, and Pachucki (2005),

$$
\begin{equation*}
B_{60}(n \mathrm{~S})-B_{60}(1 \mathrm{~S})=b_{\mathrm{L}}(n \mathrm{~S})-b_{\mathrm{L}}(1 \mathrm{~S})+A(n), \tag{40}
\end{equation*}
$$

where

$$
\begin{aligned}
A(n)= & \left(\frac{38}{45}-\frac{4}{3} \ln 2\right)[N(n S)-N(1 S)]-\frac{337043}{129600} \\
& -\frac{94261}{21600 n}+\frac{902609}{129600 n^{2}}+\left(\frac{4}{3}-\frac{16}{9 n}+\frac{4}{9 n^{2}}\right) \ln ^{2} 2 \\
& +\left(-\frac{76}{45}+\frac{304}{135 n}-\frac{76}{135 n^{2}}\right) \ln 2+\left(-\frac{53}{15}+\frac{35}{2 n}\right. \\
& \left.-\frac{419}{30 n^{2}}\right) \zeta(2) \ln 2+\left(\frac{28003}{10800}-\frac{11}{2 n}\right.
\end{aligned}
$$

TABLE XI. Values of $b_{\mathrm{L}}, B_{60}$, and $\Delta B_{71}$ used in the 2006 adjustment. See the text for an explanation of the uncertainty (33.7).

| $n$ | $b_{\mathrm{L}}(n \mathrm{~S})$ | $B_{60}(n \mathrm{~S})$ | $\Delta B_{71}(n \mathrm{~S})$ |
| :--- | :---: | :---: | :--- |
| 1 | $-81.4(0.3)$ | $-95.3(0.3)(33.7)$ |  |
| 2 | $-66.6(0.3)$ | $-80.2(0.3)(33.7)$ | $16(8)$ |
| 3 | $-63.5(0.6)$ | $-77.0(0.6)(33.7)$ | $22(11)$ |
| 4 | $-61.8(0.8)$ | $-75.3(0.8)(33.7)$ | $25(12)$ |
| 6 | $-59.8(0.8)$ | $-73.3(0.8)(33.7)$ | $28(14)$ |
| 8 | $-58.8(2.0)$ | $-72.3(2.0)(33.7)$ | $29(15)$ |

$$
\begin{align*}
& \left.+\frac{31397}{10800 n^{2}}\right) \zeta(2)+\left(\frac{53}{60}-\frac{35}{8 n}+\frac{419}{120 n^{2}}\right) \zeta(3) \\
& +\left(\frac{37793}{10800}+\frac{16}{9} \ln ^{2} 2-\frac{304}{135} \ln 2+8 \zeta(2) \ln 2\right. \\
& \left.-\frac{13}{3} \zeta(2)-2 \zeta(3)\right)[\gamma+\psi(n)-\ln n] \tag{41}
\end{align*}
$$

The term $A(n)$ makes a small contribution in the range 0.3 to 0.4 for the states under consideration.

The two-loop Bethe logarithms $b_{\mathrm{L}}$ in Eq. (40) are listed in Table XI. The values for $n=1$ to 6 are from Jentschura (2004) and Pachucki and Jentschura (2003), and the value at $n=8$ is obtained by extrapolation of the calculated values from $n=4$ to $6 \quad\left[b_{\mathrm{L}}(5 \mathrm{~S})\right.$ $=-60.6(8)]$ with a function of the form

$$
\begin{equation*}
b_{\mathrm{L}}(n \mathrm{~S})=a+\frac{b}{n}+\frac{c}{n(n+1)}, \tag{42}
\end{equation*}
$$

which yields

$$
\begin{equation*}
b_{\mathrm{L}}(n \mathrm{~S})=-55.8-\frac{24}{n} . \tag{43}
\end{equation*}
$$

It happens that the fit gives $c=0$. An estimate for $B_{60}$ given by

$$
\begin{equation*}
B_{60}(n \mathrm{~S})=b_{\mathrm{L}}(n \mathrm{~S})+\frac{10}{9} N(n \mathrm{~S})+\cdots \tag{44}
\end{equation*}
$$

was derived by Pachucki (2001). The dots represent uncalculated contributions at the relative level of $15 \%$ ( $\mathrm{Pa}-$ chucki and Jentschura, 2003). Equation (44) gives $B_{60}(1 \mathrm{~S})=-61.6(9.2)$. However, more recently Yerokhin et al. (2003, 2005a, 2005b, 2007) have calculated the 1 S-state two-loop self energy correction for $Z \geqslant 10$. This is expected to give the main contribution to the higherorder two-loop correction. Their results extrapolated to $Z=1$ yield a value for the contribution of all terms of order $B_{60}$ or higher of $-127 \times(1 \pm 0.3)$, which corresponds to a value of roughly $B_{60}=-129(39)$, assuming a linear extrapolation from $Z=1$ to 0 . This differs by about a factor of 2 from the result given by Eq. (44). In view of this difference between the two calculations, for the 2006 adjustment, we use the average of the two values with an uncertainty that is half the difference, which gives

$$
\begin{equation*}
B_{60}(1 \mathrm{~S})=-95.3(0.3)(33.7) \tag{45}
\end{equation*}
$$

In Eq. (45), the first number in parentheses is the statedependent uncertainty $u_{n}\left(B_{60}\right)$ associated with the twoloop Bethe logarithm, and the second number in parentheses is the state-independent uncertainty $u_{0}\left(B_{60}\right)$ that is common to all S-state values of $B_{60}$. Values of $B_{60}$ for all relevant $S$ states are given in Table XI. For higher- $l$ states, $B_{60}$ has not been calculated, so we take it to be zero, with uncertainties $u_{n}\left[B_{60}(n \mathrm{P})\right]=5.0$ and $u_{n}\left[B_{60}(n \mathrm{D})\right]=1.0$. We assume that these uncertainties account for higher-order P - and D -state uncertainties as well. For S states, higher-order terms have been estimated by Jentschura, Czarnecki, and Pachucki (2005) with an effective potential model. They find that the next term has a coefficient of $B_{72}$ and is state independent. We thus assume that the uncertainty $u_{0}\left[B_{60}(n S)\right]$ is sufficient to account for the uncertainty due to omitting such a term and higher-order state-independent terms. In addition, they find an estimate for the state dependence of the next term, given by

$$
\begin{align*}
\Delta B_{71}(n \mathrm{~S})= & B_{71}(n \mathrm{~S})-B_{71}(1 \mathrm{~S}) \\
= & \pi\left(\frac{427}{36}-\frac{16}{3} \ln 2\right) \\
& \times\left[\frac{3}{4}-\frac{1}{n}+\frac{1}{4 n^{2}}+\gamma+\psi(n)-\ln n\right] \tag{46}
\end{align*}
$$

with a relative uncertainty of $50 \%$. We include this additional term, which is listed in Table XI, along with the estimated uncertainty $u_{n}\left(B_{71}\right)=B_{71} / 2$.
The disagreement of the analytic and numerical calculations results in an uncertainty of the two-photon contribution that is larger than the estimated uncertainty used in the 2002 adjustment. As a result, the uncertainties of the recommended values of the Rydberg constant and proton and deuteron radii are slightly larger in the 2006 adjustment, although the 2002 and 2006 recommended values are consistent with each other. On the other hand, the uncertainty of the 2 P state fine structure is reduced as a result of the new analytic calculations.

As in the case of the order $\alpha$ self energy and vacuumpolarization contributions, the dominant effect of the finite mass of the nucleus is taken into account by multiplying each term of the two-photon contribution by the reduced-mass factor $\left(m_{\mathrm{r}} / m_{\mathrm{e}}\right)^{3}$, except that the magnetic moment term, the second line of Eq. (31), is instead multiplied by the factor $\left(m_{\mathrm{r}} / m_{\mathrm{e}}\right)^{2}$. In addition, the argument $(Z \alpha)^{-2}$ of the logarithms is replaced by $\left(m_{\mathrm{e}} / m_{\mathrm{r}}\right)(Z \alpha)^{-2}$.

## g. Three-photon corrections

The leading contribution from three virtual photons is expected to have the form

$$
\begin{equation*}
E^{(6)}=\left(\frac{\alpha}{\pi}\right)^{3} \frac{(Z \alpha)^{4}}{n^{3}} m_{\mathrm{e}} c^{2}\left[C_{40}+C_{50}(Z \alpha)+\cdots\right] \tag{47}
\end{equation*}
$$

in analogy with Eq. (29) for two photons. The leading term $C_{40}$ is (Baikov and Broadhurst, 1995; Eides and

Grotch, 1995a; Laporta and Remiddi, 1996; Melnikov and van Ritbergen, 2000)

$$
\begin{align*}
C_{40}= & {\left[-\frac{568 \mathrm{a}_{4}}{9}+\frac{85 \zeta(5)}{24}-\frac{121 \pi^{2} \zeta(3)}{72}-\frac{84071 \zeta(3)}{2304}\right.} \\
& -\frac{71 \ln ^{4} 2}{27}-\frac{239 \pi^{2} \ln ^{2} 2}{135}+\frac{4787 \pi^{2} \ln 2}{108} \\
& \left.+\frac{1591 \pi^{4}}{3240}-\frac{252251 \pi^{2}}{9720}+\frac{679441}{93312}\right] \delta_{l 0} \\
& +\left[-\frac{100 a_{4}}{3}+\frac{215 \zeta(5)}{24}-\frac{83 \pi^{2} \zeta(3)}{72}-\frac{139 \zeta(3)}{18}\right. \\
& -\frac{25 \ln ^{4} 2}{18}+\frac{25 \pi^{2} \ln ^{2} 2}{18}+\frac{298 \pi^{2} \ln 2}{9}+\frac{239 \pi^{4}}{2160} \\
& \left.-\frac{17101 \pi^{2}}{810}-\frac{28259}{5184}\right] \frac{1-\delta_{l 0}}{\kappa(2 l+1)} \tag{48}
\end{align*}
$$

where $a_{4}=\sum_{n=1}^{\infty} 1 /\left(2^{n} n^{4}\right)=0.517479061 \ldots$. Higher-order terms have not been calculated, although partial results have been obtained (Eides and Shelyuto, 2007). An uncertainty is assigned by taking $u_{0}\left(C_{50}\right)=30 \delta_{l 0}$ and $u_{n}\left(C_{63}\right)=1$, where $C_{63}$ is defined by the usual convention. The dominant effect of the finite mass of the nucleus is taken into account by multiplying the term proportional to $\delta_{l 0}$ by the reduced-mass factor $\left(m_{\mathrm{r}} / m_{\mathrm{e}}\right)^{3}$ and the term proportional to $1 / \kappa(2 l+1)$, the magnetic moment term, by the factor $\left(m_{\mathrm{r}} / m_{\mathrm{e}}\right)^{2}$.

The contribution from four photons is expected to be of order

$$
\begin{equation*}
\left(\frac{\alpha}{\pi}\right)^{4} \frac{(Z \alpha)^{4}}{n^{3}} m_{\mathrm{e}} c^{2} \tag{49}
\end{equation*}
$$

which is about 10 Hz for the 1 S state and is negligible at the level of uncertainty of current interest.

## h. Finite nuclear size

At low $Z$, the leading contribution due to the finite size of the nucleus is

$$
\begin{equation*}
E_{\mathrm{NS}}^{(0)}=\mathcal{E}_{\mathrm{NS}} \delta_{l 0}, \tag{50}
\end{equation*}
$$

with

$$
\begin{equation*}
\mathcal{E}_{\mathrm{NS}}=\frac{2}{3}\left(\frac{m_{\mathrm{r}}}{m_{\mathrm{e}}}\right)^{3} \frac{(Z \alpha)^{2}}{n^{3}} m_{\mathrm{e}} c^{2}\left(\frac{Z \alpha R_{\mathrm{N}}}{\chi_{\mathrm{C}}}\right)^{2}, \tag{51}
\end{equation*}
$$

where $R_{\mathrm{N}}$ is the bound-state root-mean-square (rms) charge radius of the nucleus and $\lambda_{\mathrm{C}}$ is the Compton wavelength of the electron divided by $2 \pi$. The leading higher-order contributions have been examined by Friar (1979b), Friar and Payne (1997b), and Karshenboim (1997) [see also Borisoglebsky and Trofimenko (1979) and Mohr (1983)]. The expressions that we employ to evaluate the nuclear size correction are the same as those discussed in more detail in CODATA-98.

For S states, the leading and next-order corrections are given by

$$
\begin{align*}
E_{\mathrm{NS}}= & \mathcal{E}_{\mathrm{NS}}\left\{1-C_{\eta} \frac{m_{\mathrm{r}}}{m_{\mathrm{e}}} \frac{R_{\mathrm{N}}}{\lambda_{\mathrm{C}}} Z \alpha-\left[\ln \left(\frac{m_{\mathrm{r}}}{m_{\mathrm{e}}} \frac{R_{\mathrm{N}}}{\lambda_{\mathrm{C}}} \frac{Z \alpha}{n}\right)\right.\right. \\
& \left.\left.+\psi(n)+\gamma-\frac{(5 n+9)(n-1)}{4 n^{2}}-C_{\theta}\right](Z \alpha)^{2}\right\}, \tag{52}
\end{align*}
$$

where $C_{\eta}$ and $C_{\theta}$ are constants that depend on the details of the assumed charge distribution in the nucleus. The values used here are $C_{\eta}=1.7(1)$ and $C_{\theta}=0.47(4)$ for hydrogen or $C_{\eta}=2.0(1)$ and $C_{\theta}=0.38(4)$ for deuterium.
For the $\mathrm{P}_{1 / 2}$ states in hydrogen, the leading term is

$$
\begin{equation*}
E_{\mathrm{NS}}=\mathcal{E}_{\mathrm{NS}} \frac{(Z \alpha)^{2}\left(n^{2}-1\right)}{4 n^{2}} \tag{53}
\end{equation*}
$$

For $\mathrm{P}_{3 / 2}$ states and D states, the nuclear-size contribution is negligible.

## i. Nuclear-size correction to self energy and vacuum polarization

For the self energy, the additional contribution due to the finite size of the nucleus is (Pachucki, 1993b; Eides and Grotch, 1997b; Milstein et al., 2002, 2003a)

$$
\begin{equation*}
E_{\mathrm{NSE}}=\left(4 \ln 2-\frac{23}{4}\right) \alpha(Z \alpha) \mathcal{E}_{\mathrm{NS}} \delta_{l 0}, \tag{54}
\end{equation*}
$$

and for the vacuum polarization it is (Friar, 1979a, 1981; Hylton, 1985; Eides and Grotch, 1997b)

$$
\begin{equation*}
E_{\mathrm{NVP}}=\frac{3}{4} \alpha(Z \alpha) \mathcal{E}_{\mathrm{NS}} \delta_{l 0} . \tag{55}
\end{equation*}
$$

For the self energy term, higher-order size corrections for S states (Milstein et al., 2002) and size corrections for P states have been calculated (Jentschura, 2003; Milstein et al., 2003b), but these corrections are negligible for the current work, and are not included. The D-state corrections are assumed to be negligible.

## j. Radiative-recoil corrections

The dominant effect of nuclear motion on the self energy and vacuum polarization has been taken into account by including appropriate reduced-mass factors. The additional contributions beyond this prescription are termed radiative-recoil effects with leading terms given by

$$
\begin{align*}
E_{\mathrm{RR}}= & \frac{m_{\mathrm{r}}^{3}}{m_{\mathrm{e}}^{2} m_{\mathrm{N}}} \frac{\alpha(Z \alpha)^{5}}{\pi^{2} n^{3}} m_{\mathrm{e}} c^{2} \delta_{l 0}\left[6 \zeta(3)-2 \pi^{2} \ln 2+\frac{35 \pi^{2}}{36}\right. \\
& \left.-\frac{448}{27}+\frac{2}{3} \pi(Z \alpha) \ln ^{2}(Z \alpha)^{-2}+\cdots\right] . \tag{56}
\end{align*}
$$

The constant term in Eq. (56) is the sum of the analytic result for the electron-line contribution (Czarnecki and Melnikov, 2001; Eides et al., 2001a) and the vacuumpolarization contribution (Eides and Grotch, 1995b; Pachucki, 1995). This term agrees with the numerical value (Pachucki, 1995) used in CODATA-98. The log-squared term has been calculated by Pachucki and Karshenboim (1999) and by Melnikov and Yelkhovsky (1999).

For the uncertainty, we take a term of order $(Z \alpha) \ln (Z \alpha)^{-2}$ relative to the square brackets in Eq. (56)
with numerical coefficients 10 for $u_{0}$ and 1 for $u_{n}$. These coefficients are roughly what one would expect for the higher-order uncalculated terms. For higher- $l$ states in the present evaluation, we assume that the uncertainties of the two- and three-photon corrections are much larger than the uncertainty of the radiative-recoil correction. Thus, we assign no uncertainty for the radiativerecoil correction for P and D states.

## k. Nucleus self energy

An additional contribution due to the self energy of the nucleus has been given by Pachucki (1995),

$$
\begin{align*}
E_{\mathrm{SEN}}= & \frac{4 Z^{2} \alpha(Z \alpha)^{4}}{3 \pi n^{3}} \frac{m_{\mathrm{r}}^{3}}{m_{\mathrm{N}}^{2}} c^{2}\left[\ln \left(\frac{m_{\mathrm{N}}}{m_{\mathrm{r}}(Z \alpha)^{2}}\right) \delta_{l 0}\right. \\
& \left.-\ln k_{0}(n, l)\right] . \tag{57}
\end{align*}
$$

This correction has also been examined by Eides et al. (2001b), who consider how it is modified by the effect of structure of the proton. The structure effect would lead to an additional model-dependent constant in the square brackets in Eq. (57).
To evaluate the nucleus self energy correction, we use Eq. (57) and assign an uncertainty $u_{0}$ that corresponds to an additive constant of 0.5 in the square brackets for S states. For P and D states, the correction is small and its uncertainty, compared to other uncertainties, is negligible.

## l. Total energy and uncertainty

The total energy $E_{n L j}^{X}$ of a particular level (where L $=\mathrm{S}, \mathrm{P}, \ldots$ and $X=\mathrm{H}, \mathrm{D}$ ) is the sum of the various contributions listed above plus an additive correction $\delta_{n \mathrm{~L} j}^{X}$ that accounts for the uncertainty in the theoretical expression for $E_{n \mathrm{Lj}}^{X}$. Our theoretical estimate for the value of $\delta_{n \mathrm{~L} j}^{X}$ for a particular level is zero with a standard uncertainty of $u\left(\delta_{n \mathrm{~L} j}^{X}\right)$ equal to the square root of the sum of the squares of the individual uncertainties of the contributions; as they are defined above, the contributions to the energy of a given level are independent. (Components of uncertainty associated with the fundamental constants are not included here, because they are determined by the least-squares adjustment itself.) Thus, we have for the square of the uncertainty, or variance, of a particular level

$$
\begin{equation*}
u^{2}\left(\delta_{n \mathrm{~L} j}^{X}\right)=\sum_{i} \frac{u_{0 i}^{2}(X \mathrm{~L} j)+u_{n i}^{2}(X L j)}{n^{6}}, \tag{58}
\end{equation*}
$$

where the individual values $u_{0 i}(X \mathrm{~L} j) / n^{3}$ and $u_{n i}(X \mathrm{~L} j) / n^{3}$ are the components of uncertainty from each of the contributions, labeled by $i$, discussed above. [The factors of $1 / n^{3}$ are isolated so that $u_{0 i}(X L j)$ is explicitly independent of $n$.]
The covariance of any two $\delta$ s follows from Eq. (F7) of Appendix F of CODATA-98. For a given isotope $X$, we have

$$
\begin{equation*}
u\left(\delta_{n_{1} L j}^{X}, \delta_{n_{2} \mathrm{~L} j}^{X}\right)=\sum_{i} \frac{u_{0 i}^{2}(X \mathrm{~L} j)}{\left(n_{1} n_{2}\right)^{3}} \tag{59}
\end{equation*}
$$

which follows from the fact that $u\left(u_{0 i}, u_{n i}\right)=0$ and $u\left(u_{n_{1} i}, u_{n_{2} i}\right)=0$ for $n_{1} \neq n_{2}$. We also set

$$
\begin{equation*}
u\left(\delta_{n_{1} \mathrm{~L}_{1} j_{1}}^{X}, \delta_{n_{2} \mathrm{~L}_{2} j_{2}}^{X}\right)=0 \tag{60}
\end{equation*}
$$

if $\mathrm{L}_{1} \neq \mathrm{L}_{2}$ or $j_{1} \neq j_{2}$.
For covariances between $\delta$ s for hydrogen and deuterium, we have for states of the same $n$

$$
\begin{equation*}
u\left(\delta_{n \mathrm{~L} j}^{\mathrm{H}}, \delta_{n \mathrm{~L} j}^{\mathrm{D}}\right)=\sum_{i=i_{\mathrm{c}}} \frac{u_{0 i}(\mathrm{HL} j) u_{0 i}(\mathrm{DL} j)+u_{n i}(\mathrm{HL} j) u_{n i}(\mathrm{DL} j)}{n^{6}}, \tag{61}
\end{equation*}
$$

and for $n_{1} \neq n_{2}$

$$
\begin{equation*}
u\left(\delta_{n_{1} \mathrm{~L} j}^{\mathrm{H}}, \delta_{n_{2} \mathrm{~L} j}^{\mathrm{D}}\right)=\sum_{i=i_{\mathrm{c}}} \frac{u_{0 i}(\mathrm{HL} j) u_{0 i}(\mathrm{DL} j)}{\left(n_{1} n_{2}\right)^{3}} \tag{62}
\end{equation*}
$$

where the summation is over the uncertainties common to hydrogen and deuterium. In most cases, the uncertainties can in fact be viewed as common except for a known multiplicative factor that contains all of the mass dependence. We assume

$$
\begin{equation*}
u\left(\delta_{n_{1} \mathrm{~L}_{1} j_{1}}^{\mathrm{H}}, \delta_{n_{2} \mathrm{~L}_{2} j_{2}}^{\mathrm{D}}\right)=0 \tag{63}
\end{equation*}
$$

if $\mathrm{L}_{1} \neq \mathrm{L}_{2}$ or $j_{1} \neq j_{2}$.
The values of $u\left(\delta_{n \mathrm{~L} j}^{X}\right)$ of interest for the 2006 adjustment are given in Table XXVIII of Sec. XII, and the non-negligible covariances of the $\delta$ s are given in the form of correlation coefficients in Table XXIX of that section. These coefficients are as large as 0.9999 .

Since the transitions between levels are measured in frequency units $(\mathrm{Hz})$, in order to apply the above equations for the energy level contributions we divide the theoretical expression for the energy difference $\Delta E$ of the transition by the Planck constant $h$ to convert it to a frequency. Further, since we take the Rydberg constant $R_{\infty}=\alpha^{2} m_{\mathrm{e}} c / 2 h$ (expressed in $\mathrm{m}^{-1}$ ) rather than the electron mass $m_{\mathrm{e}}$ to be an adjusted constant, we replace the group of constants $\alpha^{2} m_{\mathrm{e}} c^{2} / 2 h$ in $\Delta E / h$ by $c R_{\infty}$.

## $m$. Transition frequencies between levels with $n=2$

As an indication of the consistency of the theory summarized above and the experimental data, we list below values of the transition frequencies between levels with $n=2$ in hydrogen. These results are based on values of the constants obtained in a variation of the 2006 leastsquares adjustment in which the measurements of the directly related transitions (items $A 38, A 39.1$, and $A 39.2$ in Table XXVIII) are not included, and the weakly dependent constants $A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), A_{\mathrm{r}}(\mathrm{d})$, and $\alpha$ are assigned their 2006 adjusted values. The results are

$$
\nu_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)=1057843.9(2.5) \mathrm{kHz}
$$

$$
\left[2.3 \times 10^{-6}\right]
$$

$$
\begin{align*}
\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-2 \mathrm{P}_{3 / 2}\right)= & 9911197.6(2.5) \mathrm{kHz} \\
& {\left[2.5 \times 10^{-7}\right] } \\
\nu_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}-2 \mathrm{P}_{3 / 2}\right)= & 10969041.475(99) \mathrm{kHz} \\
& {\left[9.0 \times 10^{-9}\right] } \tag{64}
\end{align*}
$$

which agree well with the relevant experimental results of Table XXVIII. Although the first two values in Eq. (64) have changed only slightly from the results of the 2002 adjustment, the third value, the fine-structure splitting, has an uncertainty that is almost an order of magnitude smaller than the 2002 value, due mainly to improvements in the theory of the two-photon correction.

A value of the fine-structure constant $\alpha$ can be obtained from data on the hydrogen and deuterium transitions. This is done by running a variation of the 2006 least-squares adjustment that includes all the transition frequency data in Table XXVIII and the 2006 adjusted values of $A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p})$, and $A_{\mathrm{r}}(\mathrm{d})$. The resulting value is

$$
\begin{equation*}
\alpha^{-1}=137.036002(48) \quad\left[3.5 \times 10^{-7}\right] \tag{65}
\end{equation*}
$$

which is consistent with the 2006 recommended value, although substantially less accurate. This result is included in Table XXXIV.

## 2. Experiments on hydrogen and deuterium

Table XII summarizes the transition frequency data relevant to the determination of $R_{\infty}$. With the exception of the first entry, which is the most recent result for the $1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}$ transition frequency in hydrogen from the group at the Max-Planck-Institute für Quantenoptik (MPQ), Garching, Germany, all these data are the same as those used in the 2002 adjustment. Since these data are reviewed in CODATA-98 or CODATA-02, they are not discussed here. For a brief discussion of data not included in Table XII, see Sec. II.B. 3 of CODATA-02.

The new MPQ result,

$$
\begin{align*}
\nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)= & 2466061413187.074(34) \mathrm{kHz} \\
& {\left[1.4 \times 10^{-14}\right] } \tag{66}
\end{align*}
$$

was obtained in the course of an experiment to search for a temporal variation of the fine-structure constant $\alpha$ (Fischer et al., 2004a, 2004b; Hänsch et al., 2005; Udem, 2006). It is consistent with, but has a somewhat smaller uncertainty than, the previous result from the MPQ group, $\nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ $=2466061413187.103(46) \mathrm{kHz}\left[1.9 \times 10^{-14}\right]$ (Niering et al., 2000), which was the value used in the 2002 adjustment. The improvements that led to the reduction in uncertainty include a more stable external reference cavity for locking the 486 nm cw dye laser, thereby reducing its linewidth; an upgraded vacuum system that lowered the background gas pressure in the interaction region, thereby reducing the background gas pressure shift and its associated uncertainty; and a significantly reduced within-day Type A (i.e., statistical) uncertainty due to the narrower laser linewidth and better signal-tonoise ratio.

TABLE XII. Summary of measured transition frequencies $\nu$ considered in the present work for the determination of the Rydberg constant $R_{\infty}$ ( H is hydrogen and D is deuterium).

| Authors | Laboratory | Frequency interval(s) | Reported value ( $\nu / \mathrm{kHz}$ ) | Rel. stand. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: |
| Fischer et al., 2004a, 2004b | MPQ | $\nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 2466061413 187.074(34) | $1.4 \times 10^{-14}$ |
| Weitz et al., 1995 | MPQ | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{~S}_{1 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 4797 338(10) | $2.1 \times 10^{-6}$ |
|  |  | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{D}_{5 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | $6490144(24)$ | $3.7 \times 10^{-6}$ |
|  |  | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{~S}_{1 / 2}\right)-\frac{1}{4} \nu_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 4801 693(20) | $4.2 \times 10^{-6}$ |
|  |  | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{D}_{5 / 2}\right)-\frac{1}{4} \nu_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 6494841 (41) | $6.3 \times 10^{-6}$ |
| Huber et al., 1998 | MPQ | $\nu_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)-\nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 670994 334.64(15) | $2.2 \times 10^{-10}$ |
| de Beauvoir et al., 1997 | LKB/SYRTE | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{~S}_{1 / 2}\right)$ | $770649350012.0(8.6)$ | $1.1 \times 10^{-11}$ |
|  |  | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{D}_{3 / 2}\right)$ | $770649504450.0(8.3)$ | $1.1 \times 10^{-11}$ |
|  |  | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{D}_{5 / 2}\right)$ | 770649561 584.2(6.4) | $8.3 \times 10^{-12}$ |
|  |  | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{~S}_{1 / 2}\right)$ | 770859041 245.7(6.9) | $8.9 \times 10^{-12}$ |
|  |  | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{D}_{3 / 2}\right)$ | $770859195701.8(6.3)$ | $8.2 \times 10^{-12}$ |
|  |  | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{D}_{5 / 2}\right)$ | 770859252 849.5(5.9) | $7.7 \times 10^{-12}$ |
| Schwob et al., 1999, 2001 | LKB/SYRTE | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-12 \mathrm{D}_{3 / 2}\right)$ | 799191710 472.7(9.4) | $1.2 \times 10^{-11}$ |
|  |  | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-12 \mathrm{D}_{5 / 2}\right)$ | 799191727 403.7(7.0) | $8.7 \times 10^{-12}$ |
|  |  | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-12 \mathrm{D}_{3 / 2}\right)$ | 799409168 038.0(8.6) | $1.1 \times 10^{-11}$ |
|  |  | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-12 \mathrm{D}_{5 / 2}\right)$ | 799409184 966.8(6.8) | $8.5 \times 10^{-12}$ |
| Bourzeix et al., 1996 | LKB | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-6 \mathrm{~S}_{1 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-3 \mathrm{~S}_{1 / 2}\right)$ | 4197 604(21) | $4.9 \times 10^{-6}$ |
|  |  | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-6 \mathrm{D}_{5 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-3 \mathrm{~S}_{1 / 2}\right)$ | 4699 099(10) | $2.2 \times 10^{-6}$ |
| Berkeland et al., 1995 | Yale | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{P}_{1 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | $4664269(15)$ | $3.2 \times 10^{-6}$ |
|  |  | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{P}_{3 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 6035 373(10) | $1.7 \times 10^{-6}$ |
| Hagley and Pipkin, 1994 | Harvard | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-2 \mathrm{P}_{3 / 2}\right)$ | 9911 200(12) | $1.2 \times 10^{-6}$ |
| Lundeen and Pipkin, 1986 | Harvard | $\nu_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 1057 845.0(9.0) | $8.5 \times 10^{-6}$ |
| Newton et al., 1979 | U. Sussex | $\nu_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 1057 862(20) | $1.9 \times 10^{-5}$ |

The MPQ result in Eq. (66) and Table XII for $\nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ was provided by Udem (2006) of the MPQ group. It follows from the measured value $\quad \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)=2466061102474.851(34) \mathrm{kHz}$ $\left[1.4 \times 10^{-14}\right] \quad$ obtained for the $\left(1 \mathrm{~S}, F=1, m_{F}= \pm 1\right)$ $\rightarrow\left(2 \mathrm{~S}, F^{\prime}=1, m_{F}^{\prime}= \pm 1\right)$ transition frequency (Fischer et al., 2004a, 2004b; Hänsch et al., 2005) by using the well known 1S and 2S hyperfine splittings (Ramsey, 1990; Kolachevsky et al., 2004) to convert it to the frequency corresponding to the hyperfine centroid.

## 3. Nuclear radii

The theoretical expressions for the finite nuclear size correction to the energy levels of hydrogen H and deuterium D (see Sec. IV.A.1.h) are functions of the boundstate nuclear rms charge radius of the proton $R_{\mathrm{p}}$ and of the deuteron $R_{\mathrm{d}}$. These values are treated as variables in the adjustment, so the transition frequency data, together with theory, determine values for the radii. The radii are also determined by elastic electron-proton scattering data in the case of $R_{\mathrm{p}}$ and from elastic electrondeuteron scattering data in the case of $R_{\mathrm{d}}$. These independently determined values are used as additional information on the radii. There have been no new results during the last four years and thus we take as input
data for these two radii the values used in the 2002 adjustment,

$$
\begin{align*}
& R_{\mathrm{p}}=0.895(18) \mathrm{fm},  \tag{67}\\
& R_{\mathrm{d}}=2.130(10) \mathrm{fm} . \tag{68}
\end{align*}
$$

The result for $R_{\mathrm{p}}$ is due to Sick (2003) [see also Sick (2007)]. The result for $R_{\mathrm{d}}$ is that given in Sec. III.B. 7 of CODATA-98 based on the analysis of Sick and Trautmann (1998).

An experiment currently underway to measure the Lamb shift in muonic hydrogen may eventually provide a significantly improved value of $R_{\mathrm{p}}$ and hence an improved value of $R_{\infty}$ (Nebel et al., 2007).

## B. Antiprotonic helium transition frequencies and $A_{\mathrm{r}}(\mathrm{e})$

The antiprotonic helium atom is a three-body system consisting of a ${ }^{4} \mathrm{He}$ or ${ }^{3} \mathrm{He}$ nucleus, an antiproton, and an electron, denoted by $\overline{\mathrm{p}} \mathrm{He}^{+}$. Even though the Bohr radius for the antiproton in the field of the nucleus is about 1836 times smaller than the electron Bohr radius, in the highly excited states studied experimentally, the average orbital radius of the antiproton is comparable to the electron Bohr radius, giving rise to relatively long-

TABLE XIII. Summary of data related to the determination of $A_{\mathrm{r}}(\mathrm{e})$ from measurements on antiprotonic helium.

| Transition <br> $(n, l) \rightarrow\left(n^{\prime}, l^{\prime}\right)$ | Experimental <br> value (MHz) | Calculated <br> value $(\mathrm{MHz})$ | $a$ <br> $\left(2 c R_{\infty}\right)$ | $b$ <br> $\left(2 c R_{\infty}\right)$ |
| :--- | :---: | ---: | ---: | ---: |
| $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}:(32,31) \rightarrow(31,30)$ | $1132609209(15)$ | $1132609223.50(82)$ | 0.2179 | 0.0437 |
| $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}:(35,33) \rightarrow(34,32)$ | $804633059.0(8.2)$ | $804633058.0(1.0)$ | 0.1792 | 0.0360 |
| $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}:(36,34) \rightarrow(35,33)$ | $717474004(10)$ | $717474001.1(1.2)$ | 0.1691 | 0.0340 |
| $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}:(37,34) \rightarrow(36,33)$ | $636878139.4(7.7)$ | $636878151.7(1.1)$ | 0.1581 | 0.0317 |
| $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}:(39,35) \rightarrow(38,34)$ | $501948751.6(4.4)$ | $501948755.4(1.2)$ | 0.1376 | 0.0276 |
| $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}:(40,35) \rightarrow(39,34)$ | $445608557.6(6.3)$ | $445608569.3(1.3)$ | 0.1261 | 0.0253 |
| $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}:(37,35) \rightarrow(38,34)$ | $412885132.2(3.9)$ | $412885132.8(1.8)$ | -0.1640 | -0.0329 |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(32,31) \rightarrow(31,30)$ | $1043128608(13)$ | $1043128579.70(91)$ | 0.2098 | 0.0524 |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(34,32) \rightarrow(33,31)$ | $822809190(12)$ | $822809170.9(1.1)$ | 0.1841 | 0.0460 |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(36,33) \rightarrow(35,32)$ | $646180434(12)$ | $646180408.2(1.2)$ | 0.1618 | 0.0405 |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(38,34) \rightarrow(37,33)$ | $505222295.7(8.2)$ | $505222280.9(1.1)$ | 0.1398 | 0.0350 |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(36,34) \rightarrow(37,33)$ | $414147507.8(4.0)$ | $414147509.3(1.8)$ | -0.1664 | -0.0416 |

lived states. Also, for the high- $l$ states studied, because of the vanishingly small overlap of the antiproton wave function with the helium nucleus, strong interactions between the antiproton and the nucleus are negligible.
One of the goals of antiprotonic helium experiments is to measure the antiproton-electron mass ratio. However, since we assume that $C P T$ is a valid symmetry, for the purpose of the least-squares adjustment we take the masses of the antiproton and proton to be equal and use the data to determine the proton-electron mass ratio. Since the proton relative atomic mass is known more accurately than the electron relative atomic mass from other experiments, the mass ratio yields information primarily on the electron relative atomic mass. Other experiments have demonstrated the equality of the charge-to-mass ratio of p and $\overline{\mathrm{p}}$ to within 9 parts in $10^{11}$; see Gabrielse (2006).

## 1. Theory relevant to antiprotonic helium

Calculations of transition frequencies of antiprotonic helium have been done by Kino et al. (2003) and by Korobov (2003, 2005). The uncertainties of calculations by Korobov (2005) are of the order of 1 MHz to 2 MHz , while the uncertainties and scatter relative to the experimental values of the results of Kino et al. (2003) are substantially larger, so we use the results of Korobov (2005) in the 2006 adjustment. [See also the remarks in Hayano (2007) concerning the theory.]
The dominant contribution to the energy levels is the nonrelativistic solution of the Schrödinger equation for the three-body system together with relativistic and radiative corrections treated as perturbations. The nonrelativistic levels are resonances, because the states can decay by the Auger effect in which the electron is ejected. Korobov (2005) calculated the nonrelativistic energy by using one of two formalisms, depending on whether the Auger rate is small or large. In the case in which the rate is small, the Feshbach formalism is used with an optical
potential. The optical potential is omitted in the calculation of higher-order relativistic and radiative corrections. For broad resonances with a higher Auger rate, the nonrelativistic energies are calculated with the complex coordinate rotation method. In checking the convergence of the nonrelativistic levels, attention was paid to the convergence of the expectation value of the the delta function operators used in the evaluation of the relativistic and radiative corrections.
Korobov (2005) evaluated the relativistic and radiative corrections as perturbations to the nonrelativistic levels, including relativistic corrections of order $\alpha^{2} R_{\infty}$, anomalous magnetic moment corrections of order $\alpha^{3} R_{\infty}$ and higher, one-loop self energy and vacuumpolarization corrections of order $\alpha^{3} R_{\infty}$, and higher-order one-loop and leading two-loop corrections of order $\alpha^{4} R_{\infty}$. Higher-order relativistic corrections of order $\alpha^{4} R_{\infty}$ and radiative corrections of order $\alpha^{5} R_{\infty}$ were estimated with effective operators. The uncertainty estimates account for uncalculated terms of order $\alpha^{5} \ln \alpha R_{\infty}$.
Transition frequencies obtained by Korobov (2005, 2006) using the CODATA-02 values of the relevant constants are listed in Table XIII under the column header "calculated value." We denote these values of the frequencies by $\nu_{\bar{p} H e}^{(0)}\left(n, l: n^{\prime}, l^{\prime}\right)$, where He is either ${ }^{3} \mathrm{He}^{+}$or ${ }^{4} \mathrm{He}^{+}$. Also calculated are the leading-order changes in the theoretical values of the transition frequencies as a function of the relative changes in the mass ratios $A_{\mathrm{r}}(\overline{\mathrm{p}}) / A_{\mathrm{r}}(\mathrm{e})$ and $A_{\mathrm{r}}(N) / A_{\mathrm{r}}(\overline{\mathrm{p}})$, where $N$ is either ${ }^{3} \mathrm{He}^{2+}$ or ${ }^{4} \mathrm{He}^{2+}$. If we denote the transition frequencies as functions of these mass ratios by $\nu_{\overline{\mathrm{p}} \mathrm{He}}\left(n, l: n^{\prime}, l^{\prime}\right)$, then the changes can be written as

$$
\begin{equation*}
a_{\overline{\mathrm{p} H e}}\left(n, l: n^{\prime}, l^{\prime}\right)=\left(\frac{A_{\mathrm{r}}(\overline{\mathrm{p}})}{A_{\mathrm{r}}(\mathrm{e})}\right)^{(0)} \frac{\partial \nu_{\overline{\mathrm{p} H e}}\left(n, l: n^{\prime}, l^{\prime}\right)}{\partial\left(\frac{A_{\mathrm{r}}(\overline{\mathrm{p}})}{A_{\mathrm{r}}(\mathrm{e})}\right)}, \tag{69}
\end{equation*}
$$

$$
\begin{equation*}
b_{\overline{\mathrm{p} H e}}\left(n, l: n^{\prime}, l^{\prime}\right)=\left(\frac{A_{\mathrm{r}}(\mathrm{He})}{A_{\mathrm{r}}(\overline{\mathrm{p}})}\right)^{(0)} \frac{\partial \nu_{\overline{\mathrm{p} H e}}\left(n, l: n^{\prime}, l^{\prime}\right)}{\partial\left(\frac{A_{\mathrm{r}}(N)}{A_{\mathrm{r}}(\overline{\mathrm{p}})}\right)} . \tag{70}
\end{equation*}
$$

Values of these derivatives, in units of $2 c R_{\infty}$, are listed in Table XIII in the columns with the headers " $a$ " and " $b$," respectively. The zero-order frequencies and the derivatives are used in the expression

$$
\begin{align*}
\nu_{\overline{\mathrm{p} H e}}\left(n, l: n^{\prime}, l^{\prime}\right)= & \nu_{\overline{\mathrm{p} H e}}^{(0)}\left(n, l: n^{\prime}, l^{\prime}\right)+a_{\overline{\mathrm{p} H e}}\left(n, l: n^{\prime}, l^{\prime}\right) \\
& \times\left[\left(\frac{A_{\mathrm{r}}(\mathrm{e})}{A_{\mathrm{r}}(\overline{\mathrm{p}})}\right)^{(0)}\left(\frac{A_{\mathrm{r}}(\overline{\mathrm{p}})}{A_{\mathrm{r}}(\mathrm{e})}\right)-1\right] \\
& +b_{\overline{\mathrm{p} H e}}\left(n, l: n^{\prime}, l^{\prime}\right)\left[\left(\frac{A_{\mathrm{r}}(\overline{\mathrm{p}})}{A_{\mathrm{r}}(N)}\right)^{(0)}\right. \\
& \left.\times\left(\frac{A_{\mathrm{r}}(N)}{A_{\mathrm{r}}(\overline{\mathrm{p}})}\right)-1\right]+\cdots, \tag{71}
\end{align*}
$$

which provides a first-order approximation to the transition frequencies as a function of changes to the mass ratios. This expression is used to incorporate the experimental data and calculations for the antiprotonic system as a function of the mass ratios into the least-squares adjustment. It should be noted that even though the mass ratios are the independent variables in Eq. (71) and the relative atomic masses $A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p})$, and $A_{\mathrm{r}}(N)$ are the adjusted constants in the 2006 least-squares adjustment, the primary effect of including these data in the adjustment is on the electron relative atomic mass, because independent data in the adjustment provide values of the proton and helium nuclei relative atomic masses with significantly smaller uncertainties.

The uncertainties in the theoretical expressions for the transition frequencies are included in the adjustment as additive constants $\delta_{\overline{\mathrm{p}} \mathrm{He}}\left(n, l: n^{\prime}, l^{\prime}\right)$. Values for the theoretical uncertainties and covariances used in the adjustment are given in Sec. XII, Tables XXXII and XXXIII, respectively (Korobov, 2006).

## 2. Experiments on antiprotonic helium

Experimental work on antiprotonic helium began in the early 1990s and it continues to be an active field of research; a comprehensive review through 2000 has been given by Yamazaki et al. (2002) and a concise review through 2006 by Hayano (2007). The first measurements of $\overline{\mathrm{p}} \mathrm{He}^{+}$transition frequencies at CERN with $u_{\mathrm{r}}$ $<10^{-6}$ were reported in 2001 (Hori et al., 2001), improved results were reported in 2003 (Hori et al., 2003), and transition frequencies with uncertainties sufficiently small that they can, together with the theory of the transitions, provide a competitive value of $A_{\mathrm{r}}(\mathrm{e})$, were reported in 2006 (Hori et al., 2006).
The 12 transition frequencies-seven for ${ }^{4} \mathrm{He}$ and five for ${ }^{3} \mathrm{He}$ given by Hori et al. (2006) -which we take as input data in the 2006 adjustment, are listed in column 2 of Table XIII with the corresponding transitions indicated in column 1. To reduce rounding errors, an additional digit for both the frequencies and their uncertain-
ties as provided by Hori (2006) have been included. All 12 frequencies are correlated; their correlation coefficients, based on detailed uncertainty budgets for each, also provided by Hori (2006), are given in Table XXXIII in Sec XII.

In the current version of the experiment, 5.3 MeV antiprotons from the CERN Antiproton Decelerator (AD) are decelerated using a radio-frequency quadrupole decelerator (RFQD) to energies in the range 10 keV to 120 keV controlled by a dc potential bias on the RFQD's electrodes. The decelerated antiprotons, about $30 \%$ of the antiprotons entering the RFQD, are then diverted to a low-pressure cryogenic helium gas target at 10 K by an achromatic momentum analyzer, the purpose of which is to eliminate the large background that the remaining $70 \%$ of undecelerated antiprotons would have produced.

About $3 \%$ of the $\overline{\mathrm{p}}$ stopped in the target form $\overline{\mathrm{p}} \mathrm{He}^{+}$, in which a $\bar{p}$ with large principle quantum number ( $n$ $\approx 38$ ) and angular momentum quantum number ( $l \approx n$ ) circulates in a localized, nearly circular orbit around the $\mathrm{He}^{2+}$ nucleus while the electron occupies the distributed 1S state. These $\bar{p}$ energy levels are metastable with lifetimes of several microseconds and de-excite radiatively. There are also short-lived $\bar{p}$ states with similar values of $n$ and $l$ but with lifetimes on the order of 10 ns and which de-excite by Auger transitions to form $\overline{\mathrm{p}} \mathrm{He}^{2+}$ hydrogenlike ions. These undergo Stark collisions, which cause the rapid annihilation of the $\bar{p}$ in the helium nucleus. The annihilation rate versus time elapsed since $\overline{\mathrm{p}} \mathrm{He}^{+}$formation, or delayed annihilation time spectrum (DATS), is measured using Cherenkov counters.

With the exception of the $(36,34) \rightarrow(35,33)$ transition frequency, all frequencies given in Table XIII were obtained by stimulating transitions from the $\overline{\mathrm{p}} \mathrm{He}^{+}$metastable states with values of $n$ and $l$ indicated in column one on the left-hand side of the arrow to the short-lived, Auger-decaying states with values of $n$ and $l$ indicated on the right-hand side of the arrow.

The megawatt-scale light intensities needed to induce the $\overline{\mathrm{p}} \mathrm{He}^{+}$transitions, which cover the wavelength range 265 nm to 726 nm , can only be provided by a pulsed laser. Frequency and linewidth fluctuations and frequency calibration problems associated with such lasers were overcome by starting with a cw "seed" laser beam of frequency $\nu_{\mathrm{cw}}$, known with $u_{\mathrm{r}}<4 \times 10^{-10}$ through its stabilization by an optical frequency comb, and then amplifying the intensity of the laser beam by a factor of $10^{6}$ in a cw pulse amplifier consisting of three dye cells pumped by a pulsed Nd:YAG laser. The 1 W seed laser beam with wavelength in the range 574 nm to 673 nm was obtained from a pumped cw dye laser, and the 1 W seed laser beam with wavelength in the range 723 nm to 941 nm was obtained from a pumped cw Ti:sapphire laser. The shorter wavelengths ( 265 nm to 471 nm ) for inducing transitions were obtained by frequency doubling the amplifier output at 575 nm and 729 nm to 941 nm or by frequency tripling its 794 nm output. The frequency of the seed laser beam $\nu_{\mathrm{cw}}$, and thus the fre-
quency $\nu_{\mathrm{pl}}$ of the pulse amplified beam, was scanned over a range of $\pm 4 \mathrm{GHz}$ around the $\overline{\mathrm{p}} \mathrm{He}^{+}$transition frequency by changing the repetition frequency $f_{\text {rep }}$ of the frequency comb.

The resonance curve for a transition was obtained by plotting the area under the resulting DATS peak versus $\nu_{\mathrm{pl}}$. Because of the approximate 400 MHz Doppler broadening of the resonance due to the 10 K thermal motion of the $\overline{\mathrm{p}} \mathrm{He}^{+}$atoms, a rather sophisticated theoretical line shape that takes into account many factors must be used to obtain the desired transition frequency.
Two other effects of major importance are the socalled chirp effect and linear shifts in the transition frequencies due to collisions between the $\overline{\mathrm{p}} \mathrm{He}^{+}$and background helium atoms. The frequency $\nu_{\mathrm{pl}}$ can deviate from $\nu_{\mathrm{cw}}$ due to sudden changes in the index of refraction of the dye in the cells of the amplifier. This chirp, which can be expressed as $\Delta \nu_{\mathrm{c}}(t)=\nu_{\mathrm{pl}}(t)-\nu_{\mathrm{cw}}$, can shift the measured $\overline{\mathrm{p}} \mathrm{He}^{+}$frequencies from their actual values. Hori et al. (2006) eliminated this effect by measuring $\Delta \nu_{\mathrm{c}}(t)$ in real time and applying a frequency shift to the seed laser, thereby canceling the dye-cell chirp. This effect is the predominant contributor to the correlations among the 12 transitions (Hori, 2006). The collisional shift was eliminated by measuring the frequencies of ten transitions in helium gas targets with helium atom densities $\rho$ in the range $2 \times 10^{18} / \mathrm{cm}^{3}$ to $3 \times 10^{21} / \mathrm{cm}^{3}$ to determine $d \nu / d \rho$. The in vacuo $(\rho=0)$ values were obtained by applying a suitable correction in the range -14 MHz to 1 MHz to the initially measured frequencies obtained at $\rho \approx 2 \times 10^{18} / \mathrm{cm}^{3}$.

In contrast to the other 11 transition frequencies in Table XIII, which were obtained by inducing a transition from a long-lived, metastable state to a short-lived, Auger-decaying state, the $(36,34) \rightarrow(35,33)$ transition frequency was obtained by inducing a transition from the $(36,34)$ metastable state to the $(35,33)$ metastable state using three different lasers. This was done by first depopulating at time $t_{1}$ the $(35,33)$ metastable state by inducing the $(35,33) \rightarrow(34,32)$ metastable state to short-lived-state transition, then at time $t_{2}$ inducing the $(36,34) \rightarrow(35,33)$ transition using the cw pulse-amplified laser, and then at time $t_{3}$ again inducing the $(35,33)$ $\rightarrow(34,32)$ transition. The resonance curve for the $(36,34) \rightarrow(35,33)$ transition was obtained from the DATS peak resulting from this last induced transition.

The 4 MHz to 15 MHz standard uncertainties of the transition frequencies in Table XIII arise from the resonance line-shape fit ( 3 MHz to 13 MHz , statistical or Type A), not completely eliminating the chirp effect (2 MHz to 4 MHz , nonstatistical or Type B), collisional shifts ( 0.1 MHz to 2 MHz , Type B ), and frequency doubling or tripling ( 1 MHz to 2 MHz , Type B ).

## 3. Values of $\boldsymbol{A}_{\mathrm{r}}(\mathrm{e})$ inferred from antiprotonic helium

From the theory of the 12 antiprotonic transition frequencies discussed in Sec IV.B.1, the 2006 recommended values of the relative atomic masses of the proton, alpha
particle ( $\alpha$, nucleus of the ${ }^{4} \mathrm{He}$ atom), and the helion (h, nucleus of the ${ }^{3} \mathrm{He}$ atom), $A_{\mathrm{r}}(\mathrm{p}), A_{\mathrm{r}}(\alpha)$, and $A_{\mathrm{r}}(\mathrm{h})$, respectively, together with the 12 experimental values for these frequencies given in Table XIII, we find the following three values for $A_{\mathrm{r}}(\mathrm{e})$ from the seven $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}$frequencies alone, from the five $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}$frequencies alone, and from the 12 frequencies together:

$$
\begin{align*}
& A_{\mathrm{r}}(\mathrm{e})=0.0005485799103(12) \quad\left[2.1 \times 10^{-9}\right]  \tag{72}\\
& A_{\mathrm{r}}(\mathrm{e})=0.0005485799053(15) \quad\left[2.7 \times 10^{-9}\right]  \tag{73}\\
& A_{\mathrm{r}}(\mathrm{e})=0.00054857990881(91) \quad\left[1.7 \times 10^{-9}\right] . \tag{74}
\end{align*}
$$

The separate inferred values from the $\bar{p}^{4} \mathrm{He}^{+}$and $\bar{p}^{3} \mathrm{He}^{+}$ frequencies differ somewhat, but the value from all 12 frequencies not only agrees with the three other available results for $A_{\mathrm{r}}(\mathrm{e})$ (see Table XXXVI, Sec. XII.A), but has a competitive level of uncertainty as well.

## C. Hyperfine structure and fine structure

## 1. Hyperfine structure

Because the ground-state hyperfine transition frequencies $\Delta \nu_{\mathrm{H}}, \Delta \nu_{\mathrm{Mu}}$, and $\Delta \nu_{\mathrm{Ps}}$ of the comparatively simple atoms hydrogen, muonium, and positronium, respectively, are proportional to $\alpha^{2} R_{\infty} c$, in principle a value of $\alpha$ can be obtained by equating an experimental value of one of these transition frequencies to its presumed readily calculable theoretical expression. However, currently only measurements of $\Delta \nu_{\mathrm{Mu}}$ and the theory of the muonium hyperfine structure have sufficiently small uncertainties to provide a useful result for the 2006 adjustment, and even in this case the result is not a competitive value of $\alpha$, but rather the most accurate value of the electron-muon mass ratio $m_{\mathrm{e}} / m_{\mu}$. Indeed, we discuss the relevant experiments and theory in Sec. VI.B.

Although the ground-state hyperfine transition frequency of hydrogen has long been of interest as a potential source of an accurate value of $\alpha$ because it is experimentally known with $u_{\mathrm{r}} \approx 10^{-12}$ (Ramsey, 1990), the relative uncertainty of the theory is still of the order of $10^{-6}$. Thus, $\Delta \nu_{\mathrm{H}}$ cannot yet provide a competitive value of the fine-structure constant. At present, the main sources of uncertainty in the theory arise from the internal structure of the proton, namely (i) the electric charge and magnetization densities of the proton, which are taken into account by calculating the proton's so-called Zemach radius; and (ii) the polarizability of the proton (that is, protonic excited states). For details of the progress made over the last four years in reducing the uncertainties from both sources, see Carlson (2007), Pachucki (2007), Sick (2007), and references therein. Because the muon is a structureless pointlike particle, the theory of $\Delta \nu_{\mathrm{Mu}}$ is free from such uncertainties.

It is also not yet possible to obtain a useful value of $\alpha$ from $\Delta \nu_{\text {Ps }}$ since the most accurate experimental result has $u_{\mathrm{r}}=3.6 \times 10^{-6}$ (Ritter et al., 1984). The theoretical un-
certainty of $\Delta \nu_{\text {Ps }}$ is not significantly smaller and may in fact be larger (Adkins et al., 2002; Penin, 2004).

## 2. Fine structure

As in the case of hyperfine splittings, fine-structure transition frequencies are proportional to $\alpha^{2} R_{\infty} c$ and could be used to deduce a value of $\alpha$. Some data related to the fine structure of hydrogen and deuterium are discussed in Sec. IV.A. 2 in connection with the Rydberg constant. They are included in the adjustment because of their influence on the adjusted value of $R_{\infty}$. However, the value of $\alpha$ that can be derived from these data is not competitive; see Eq. (65). See also Sec. III.B. 3 of CODATA-02 for a discussion of why earlier fine structure-related results in H and D are not considered.

Because the transition frequencies corresponding to the differences in energy of the three $2{ }^{3} \mathrm{P}$ levels of ${ }^{4} \mathrm{He}$ can be both measured and calculated with reasonable accuracy, the fine structure of ${ }^{4} \mathrm{He}$ has long been viewed as a potential source of a reliable value of $\alpha$. The three frequencies of interest are $\nu_{01} \approx 29.6 \mathrm{GHz}$, $\nu_{12} \approx 2.29 \mathrm{GHz}$, and $\nu_{02} \approx 31.9 \mathrm{GHz}$, which correspond to the intervals $2{ }^{3} \mathrm{P}_{1}-2{ }^{3} \mathrm{P}_{0}, 2{ }^{3} \mathrm{P}_{2}-2{ }^{3} \mathrm{P}_{1}$, and $2{ }^{3} \mathrm{P}_{2}-2{ }^{3} \mathrm{P}_{0}$, respectively. The value with the smallest uncertainty for any of these frequencies was obtained at Harvard (Zelevinsky et al., 2005),

$$
\begin{equation*}
\nu_{01}=29616951.66(70) \mathrm{kHz} \quad\left[2.4 \times 10^{-8}\right] \tag{75}
\end{equation*}
$$

It is consistent with the value of $\nu_{01}$ reported by George et al. (2001) with $u_{\mathrm{r}}=3.0 \times 10^{-8}$, and that reported by Giusfredi et al. (2005) with $u_{\mathrm{r}}=3.4 \times 10^{-8}$. If the theoretical expression for $\nu_{01}$ were exactly known, the weighted mean of the three results would yield a value of $\alpha$ with $u_{\mathrm{r}} \approx 8 \times 10^{-9}$.

However, as discussed in CODATA-02, the theory of the $2{ }^{3} \mathrm{P}_{J}$ transition frequencies is far from satisfactory. First, different calculations disagree, and because of the considerable complexity of the calculations and the history of their evolution, there is general agreement that results that have not been confirmed by independent evaluation should be taken as tentative. Second, there are significant disagreements between theory and experiment. Recently, Pachucki (2006) has advanced the theory by calculating the complete contribution to the $2{ }^{3} \mathrm{P}_{J}$ fine-structure levels of order $m \alpha^{7}$ (or $\alpha^{5} \mathrm{Ry}$ ), with the final theoretical result for $\nu_{01}$ being

$$
\begin{equation*}
\nu_{01}=29616943.01(17) \mathrm{kHz} \quad\left[5.7 \times 10^{-9}\right] \tag{76}
\end{equation*}
$$

This value disagrees with the experimental value given in Eq. (75) as well as with the theoretical value $\nu_{01}$ $=29616946.42(18) \mathrm{kHz}\left[6.1 \times 10^{-9}\right]$ given by Drake (2002), which also disagrees with the experimental value. These disagreements suggest that there is a problem with theory and/or experiment that must be resolved before a meaningful value of $\alpha$ can be obtained from the helium fine structure (Pachucki, 2006). Therefore, as in the 2002 adjustment, we do not include ${ }^{4} \mathrm{He}$ finestructure data in the 2006 adjustment.

## V. MAGNETIC MOMENT ANOMALIES AND $\boldsymbol{g}$-FACTORS

In this section, theory and experiment for the magnetic moment anomalies of the free electron and muon and the bound-state $g$-factor of the electron in hydrogenic carbon $\left({ }^{12} \mathrm{C}^{5+}\right)$ and in hydrogenic oxygen $\left({ }^{16} \mathrm{O}^{7+}\right)$ are reviewed.

The magnetic moment of any of the three charged leptons $\ell=\mathrm{e}, \mu, \tau$ is written as

$$
\begin{equation*}
\mu_{\ell}=g_{\ell} \frac{e}{2 m_{\ell}} \boldsymbol{s} \tag{77}
\end{equation*}
$$

where $g_{\ell}$ is the $g$-factor of the particle, $m_{\ell}$ is its mass, and $s$ is its spin. In Eq. (77), $e$ is the elementary charge and is positive. For the negatively charged leptons $\ell^{-}, g_{\ell}$ is negative, and for the corresponding antiparticles $\ell^{+}, g_{\ell}$ is positive. $C P T$ invariance implies that the masses and absolute values of the $g$-factors are the same for each particle-antiparticle pair. These leptons have eigenvalues of spin projection $s_{z}= \pm \hbar / 2$, and it is conventional to write, based on Eq. (77),

$$
\begin{equation*}
\mu_{\ell}=\frac{g_{\ell}}{2} \frac{e \hbar}{2 m_{\ell}} \tag{78}
\end{equation*}
$$

where in the case of the electron, $\mu_{\mathrm{B}}=e \hbar / 2 m_{\mathrm{e}}$ is the Bohr magneton.

The free lepton magnetic moment anomaly $a_{\ell}$ is defined as

$$
\begin{equation*}
\left|g_{\ell}\right|=2\left(1+a_{\ell}\right) \tag{79}
\end{equation*}
$$

where $g_{\mathrm{D}}=-2$ is the value predicted by the free-electron Dirac equation. The theoretical expression for $a_{\ell}$ may be written as

$$
\begin{equation*}
a_{\ell}(\mathrm{th})=a_{\ell}(\mathrm{QED})+a_{\ell}(\text { weak })+a_{\ell}(\mathrm{had}) \tag{80}
\end{equation*}
$$

where the terms denoted by QED, weak, and had account for the purely quantum electrodynamic, predominantly electroweak, and predominantly hadronic (that is, strong interaction) contributions to $a_{\ell}$, respectively.

The QED contribution may be written as (Kinoshita et al., 1990)

$$
\begin{align*}
a_{\ell}(\mathrm{QED})= & A_{1}+A_{2}\left(m_{\ell} / m_{\ell^{\prime}}\right)+A_{2}\left(m_{\ell} / m_{\ell^{\prime \prime}}\right) \\
& +A_{3}\left(m_{\ell} / m_{\ell^{\prime}}, m_{\ell} / m_{\ell^{\prime \prime}}\right) \tag{81}
\end{align*}
$$

where for the electron, $\left(\ell, \ell^{\prime}, \ell^{\prime \prime}\right)=(e, \mu, \tau)$ and for the muon, $\left(\ell, \ell^{\prime}, \ell^{\prime \prime}\right)=(\mu, \mathrm{e}, \tau)$. The anomaly for the $\tau$, which is poorly known experimentally (Yao et al., 2006), is not considered here. For recent work on the theory of $a_{\tau}$, see Eidelman and Passera (2007). In Eq. (81), the term $A_{1}$ is mass independent, and the mass dependence of $A_{2}$ and $A_{3}$ arises from vacuum polarization loops with lepton $\ell^{\prime}, \ell^{\prime \prime}$, or both. Each of the four terms on the righthand side of Eq. (81) can be expressed as a power series in the fine-structure constant $\alpha$,

$$
\begin{align*}
A_{i}= & A_{i}^{(2)}\left(\frac{\alpha}{\pi}\right)+A_{i}^{(4)}\left(\frac{\alpha}{\pi}\right)^{2}+A_{i}^{(6)}\left(\frac{\alpha}{\pi}\right)^{3}+A_{i}^{(8)}\left(\frac{\alpha}{\pi}\right)^{4} \\
& +A_{i}^{(10)}\left(\frac{\alpha}{\pi}\right)^{5}+\cdots \tag{82}
\end{align*}
$$

where $A_{2}^{(2)}=A_{3}^{(2)}=A_{3}^{(4)}=0$. Coefficients proportional to $(\alpha / \pi)^{n}$ are of order $e^{2 n}$ and are referred to as $2 n$ th-order coefficients.

For the mass-independent term $A_{1}$, the second-order coefficient is known exactly, and the fourth- and sixthorder coefficients are known analytically in terms of readily evaluated functions,

$$
\begin{align*}
& A_{1}^{(2)}=\frac{1}{2}  \tag{83}\\
& A_{1}^{(4)}=-0.328478965579 \ldots,  \tag{84}\\
& A_{1}^{(6)}=1.181241456 \ldots \tag{85}
\end{align*}
$$

A total of 891 Feynman diagrams give rise to the mass-independent eighth-order coefficient $A_{1}^{(8)}$, and only a few of these are known analytically. However, in an effort that has its origins in the 1960s, Kinoshita and collaborators have calculated $A_{1}^{(8)}$ numerically; the result of this ongoing project that is used in the 2006 adjustment is (Kinoshita and Nio, 2006a; Gabrielse et al., 2006, 2007)

$$
\begin{equation*}
A_{1}^{(8)}=-1.7283(35) \tag{86}
\end{equation*}
$$

Work was done in the evaluation and checking of this coefficient in an effort to obtain a reliable quantitative result. A subset of 373 diagrams containing closed electron loops was verified by more than one independent formulation. The remaining 518 diagrams with no closed electron loops were formulated in only one way. As a check on this set, extensive cross checking was performed on the renormalization terms both among themselves and with lower-order diagrams that are known exactly (Kinoshita and Nio, 2006a) [see also Gabrielse et al. $(2006,2007)]$. For the final numerical integrations, an adaptive-iterative Monte Carlo routine was used. A time-consuming part of the work was checking for round-off error in the integration.
The 0.0035 standard uncertainty of $A_{1}^{(8)}$ contributes a standard uncertainty to $a_{\mathrm{e}}(\mathrm{th})$ of $0.88 \times 10^{-10} a_{\mathrm{e}}$, which is smaller than the uncertainty due to uncalculated higherorder contributions. Independent work is in progress on analytic calculations of eighth-order integrals. See, for example, Laporta (2001), Laporta et al. (2004), and Mastrolia and Remiddi (2001).

Little is known about the tenth-order coefficient $A_{1}^{(10)}$ and higher-order coefficients, although Kinoshita et al. (2006) are starting the numerical evaluation of the 12672 Feynman diagrams for this coefficient. To evaluate the contribution to the uncertainty of $a_{\mathrm{e}}($ th) due to lack of knowledge of $A_{1}^{(10)}$, we follow CODATA-98 to obtain $A_{1}^{(10)}=0.0(3.7)$. The 3.7 standard uncertainty of $A_{1}^{(10)}$ contributes a standard uncertainty component to $a_{\mathrm{e}}(\mathrm{th})$ of $2.2 \times 10^{-10} a_{\mathrm{e}}$; the uncertainty contributions to
$a_{\mathrm{e}}(\mathrm{th})$ from all other higher-order coefficients, which should be significantly smaller, are assumed to be negligible.

The 2006 least-squares adjustment was carried out using the theoretical results given above, including the value of $A_{1}^{(8)}$ given in Eq. (86). Well after the deadline for new data and the recommended values from the adjustment were made public (Mohr et al., 2007), it was discovered by Aoyama et al. (2007) that 2 of the 47 integrals representing 518 QED diagrams that had not previously been confirmed independently required a corrected treatment of infrared divergences. The revised value they give is

$$
\begin{equation*}
A_{1}^{(8)}=-1.9144(35) \tag{87}
\end{equation*}
$$

although the new calculation is still tentative (Aoyama et $a l ., 2007$ ). This result would lead to the value

$$
\begin{equation*}
\alpha^{-1}=137.035999070(98) \quad\left[7.1 \times 10^{-10}\right] \tag{88}
\end{equation*}
$$

for the inverse fine-structure constant derived from the electron anomaly using the Harvard measurement result for $a_{\mathrm{e}}$ (Gabrielse et al., 2006, 2007). This number is shifted down from the previous result by $641 \times 10^{-9}$ and its uncertainty is increased from (96) to (98) (see Sec. V.A.3), but it is still consistent with the values obtained from recoil experiments (see Table XXVI). If this result for $A_{1}^{(8)}$ had been used in the 2006 adjustment, the recommended value of the inverse fine-structure constant would differ by a similar, although slightly smaller, amount. The effect on the muon anomaly theory is completely negligible.

The mass independent term $A_{1}$ contributes equally to the free electron and muon anomalies and the boundelectron $g$-factors. The mass-dependent terms are different for the electron and muon and are considered separately in the following. For the bound-electron $g$-factor, there are bound-state corrections in addition to the freeelectron value of the $g$-factor, as discussed below.

## A. Electron magnetic moment anomaly $a_{\mathrm{e}}$ and the fine-structure constant $\alpha$

The combination of theory and experiment for the electron magnetic moment anomaly yields the value for the fine-structure constant $\alpha$ with the smallest estimated uncertainty (see Table XIV).

## 1. Theory of $a_{\mathrm{e}}$

The mass-dependent coefficients of interest and corresponding contributions to the theoretical value of the anomaly $a_{\mathrm{e}}(\mathrm{th})$, based on the 2006 recommended values of the mass ratios, are

$$
\begin{align*}
A_{2}^{(4)}\left(m_{\mathrm{e}} / m_{\mu}\right) & =5.19738678(26) \times 10^{-7} \\
& \rightarrow 24.182 \times 10^{-10} a_{\mathrm{e}},  \tag{89}\\
A_{2}^{(4)}\left(m_{\mathrm{e}} / m_{\tau}\right) & =1.83763(60) \times 10^{-9} \\
& \rightarrow 0.085 \times 10^{-10} a_{\mathrm{e}}, \tag{90}
\end{align*}
$$

TABLE XIV. Summary of data related to magnetic moments of the electron and muon and inferred values of the fine-structure constant. (The source data and not the inferred values given here are used in the adjustment.)

| Quantity | Value | Relative standard <br> uncertainty $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :--- | :---: | :---: | :---: | :--- |
| $a_{\mathrm{e}}$ | $1.1596521883(42) \times 10^{-3}$ | $3.7 \times 10^{-9}$ | UWash-87 | V.A.2.a (102) |
| $\alpha^{-1}\left(a_{\mathrm{e}}\right)$ | $137.03599883(50)$ | $3.7 \times 10^{-9}$ |  | V.A.3 (104) |
| $a_{\mathrm{e}}$ | $1.15965218085(76) \times 10^{-3}$ | $6.6 \times 10^{-10}$ | HarvU-06 | V.A.2.b (103) |
| $\alpha^{-1}\left(a_{\mathrm{e}}\right)$ | $137.035999711(96)$ | $7.0 \times 10^{-10}$ |  | V.A.3 (105) |
| $\bar{R}$ |  |  |  |  |
| $a_{\mu}$ | $1.16592093(63) \times 10^{-3}$ | $5.4 \times 10^{-7}$ | BNL-06 | V.B.2 (128) |
| $\alpha^{-1}(\bar{R})$ | $137.03567(26)$ | $1.9 \times 10^{-7}$ |  | V.B.2 (129) |

$$
\begin{align*}
A_{2}^{(6)}\left(m_{\mathrm{e}} / m_{\mu}\right) & =-7.37394172(27) \times 10^{-6} \\
& \rightarrow-0.797 \times 10^{-10} a_{\mathrm{e}},  \tag{91}\\
A_{2}^{(6)}\left(m_{\mathrm{e}} / m_{\tau}\right) & =-6.5819(19) \times 10^{-8} \\
& \rightarrow-0.007 \times 10^{-10} a_{\mathrm{e}}, \tag{92}
\end{align*}
$$

where the standard uncertainties of the coefficients are due to the uncertainties of the mass ratios but are negligible for $a_{\mathrm{e}}(\mathrm{th})$. The contributions from $A_{3}^{(6)}\left(m_{\mathrm{e}} / m_{\mu}\right.$, $m_{\mathrm{e}} / m_{\tau}$ ) and all higher-order mass-dependent terms are negligible as well.

The value for $A_{2}^{(6)}\left(m_{\mathrm{e}} / m_{\mu}\right)$ in Eq. (91) has been updated from the value in CODATA-02 and is in agreement with the result of Passera (2007) based on a calculation to all orders in the mass ratio. The change is given by the term

$$
\begin{align*}
& \frac{17 x^{6} \zeta(3)}{36}-\frac{4381 x^{6} \ln ^{2} x}{30240}+\frac{24761 x^{6} \ln x}{158760} \\
& -\frac{13 \pi^{2} x^{6}}{1344}-\frac{1840256147 x^{6}}{3556224000} \tag{93}
\end{align*}
$$

where $x=m_{\mathrm{e}} / m_{\mu}$, which was not included in CODATA02. The earlier result was based on Eq. (4) of Laporta and Remiddi (1993), which only included terms to order $x^{4}$. The additional term was kindly provided by Laporta and Remiddi (2006).

For the electroweak contribution, we have

$$
\begin{align*}
a_{\mathrm{e}}(\text { weak }) & =0.02973(52) \times 10^{-12} \\
& =0.2564(45) \times 10^{-10} a_{\mathrm{e}} \tag{94}
\end{align*}
$$

as calculated in CODATA-98 but with the current values of $G_{\mathrm{F}}$ and $\sin ^{2} \theta_{\mathrm{W}}$ (see Sec. XI.B).

The hadronic contribution is

$$
\begin{equation*}
a_{\mathrm{e}}(\mathrm{had})=1.682(20) \times 10^{-12}=1.450(17) \times 10^{-9} a_{\mathrm{e}} . \tag{95}
\end{equation*}
$$

It is the sum of the following three contributions: $a_{\mathrm{e}}^{(4)}(\mathrm{had})=1.875(18) \times 10^{-12}$ obtained by Davier and

Höcker (1998), $a_{\mathrm{e}}^{(6 a)}(\mathrm{had})=-0.225(5) \times 10^{-12}$ given by Krause (1997), and $a_{\mathrm{e}}^{(\gamma \gamma)}(\mathrm{had})=0.0318(58) \times 10^{-12}$ calculated by multiplying the corresponding result for the muon given in Sec. V.B. 1 by the factor $\left(m_{\mathrm{e}} / m_{\mu}\right)^{2}$, since $a_{\mathrm{e}}^{(\gamma \gamma)}(\mathrm{had})$ is assumed to vary approximately as the square of the mass.

Because the dependence on $\alpha$ of any contribution other than $a_{\mathrm{e}}(\mathrm{QED})$ is negligible, the anomaly as a function of $\alpha$ is given by combining terms that have like powers of $\alpha / \pi$ to yield

$$
\begin{equation*}
a_{\mathrm{e}}(\mathrm{th})=a_{\mathrm{e}}(\mathrm{QED})+a_{\mathrm{e}}(\text { weak })+a_{\mathrm{e}}(\mathrm{had}) \tag{96}
\end{equation*}
$$

where

$$
\begin{align*}
a_{\mathrm{e}}(\mathrm{QED})= & C_{\mathrm{e}}^{(2)}\left(\frac{\alpha}{\pi}\right)+C_{\mathrm{e}}^{(4)}\left(\frac{\alpha}{\pi}\right)^{2}+C_{\mathrm{e}}^{(6)}\left(\frac{\alpha}{\pi}\right)^{3} \\
& +C_{\mathrm{e}}^{(8)}\left(\frac{\alpha}{\pi}\right)^{4}+C_{\mathrm{e}}^{(10)}\left(\frac{\alpha}{\pi}\right)^{5}+\cdots \tag{97}
\end{align*}
$$

with

$$
\begin{align*}
& C_{\mathrm{e}}^{(2)}=0.5, \\
& C_{\mathrm{e}}^{(4)}=-0.32847844400, \\
& C_{\mathrm{e}}^{(6)}=1.181234017, \\
& C_{\mathrm{e}}^{(8)}=-1.7283(35), \\
& C_{\mathrm{e}}^{(10)}=0.0(3.7) \tag{98}
\end{align*}
$$

and $a_{\mathrm{e}}$ (weak) and $a_{\mathrm{e}}($ had $)$ are as given in Eqs. (94) and (95).

The standard uncertainty of $a_{\mathrm{e}}(\mathrm{th})$ from the uncertainties of the terms listed above, other than that due to $\alpha$, is

$$
\begin{equation*}
u\left[a_{\mathrm{e}}(\mathrm{th})\right]=0.27 \times 10^{-12}=2.4 \times 10^{-10} a_{\mathrm{e}} \tag{99}
\end{equation*}
$$

and is dominated by the uncertainty of the coefficient $C_{\mathrm{e}}^{(10)}$.
For the purpose of the least-squares calculations carried out in Sec. XII.B, we define an additive correction
$\delta_{\mathrm{e}}$ to $a_{\mathrm{e}}(\mathrm{th})$ to account for the lack of exact knowledge of $a_{\mathrm{e}}(\mathrm{th})$, and hence the complete theoretical expression for the electron anomaly is

$$
\begin{equation*}
a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)=a_{\mathrm{e}}(\mathrm{th})+\delta_{\mathrm{e}} . \tag{100}
\end{equation*}
$$

Our theoretical estimate of $\delta_{\mathrm{e}}$ is zero and its standard uncertainty is $u\left[a_{\mathrm{e}}(\mathrm{th})\right]$,

$$
\begin{equation*}
\delta_{\mathrm{e}}=0.00(27) \times 10^{-12} \tag{101}
\end{equation*}
$$

## 2. Measurements of $\boldsymbol{a}_{\mathrm{e}}$

a. University of Washington

The classic series of measurements of the electron and positron anomalies carried out at the University of Washington by Van Dyck et al. (1987) yields the value

$$
\begin{equation*}
a_{\mathrm{e}}=1.1596521883(42) \times 10^{-3} \quad\left[3.7 \times 10^{-9}\right] \tag{102}
\end{equation*}
$$

as discussed in CODATA-98. This result assumes that $C P T$ invariance holds for the electron-positron system.

## b. Harvard University

A new determination of the electron anomaly using a cylindrical Penning trap has been carried out by Odom et al. (2006) at Harvard University, yielding the value

$$
\begin{equation*}
a_{\mathrm{e}}=1.15965218085(76) \times 10^{-3} \quad\left[6.6 \times 10^{-10}\right] \tag{103}
\end{equation*}
$$

which has an uncertainty that is nearly six times smaller than that of the University of Washington result.

As in the University of Washington experiment, the anomaly is obtained in essence from the relation $a_{\text {e }}$ $=f_{\mathrm{a}} / f_{\mathrm{c}}$ by determining, in the same magnetic flux density $B$ (about 5 T ), the anomaly difference frequency $f_{\mathrm{a}}=f_{\mathrm{s}}$ $-f_{\mathrm{c}}$ and cyclotron frequency $f_{\mathrm{c}}=e B / 2 \pi m_{\mathrm{e}}$, where $f_{\mathrm{s}}$ $=g_{\mathrm{e}} \mu_{\mathrm{B}} B / h$ is the electron spin-flip (often called precession) frequency. The marked improvement achieved by the Harvard group, the culmination of a 20 year effort, is due in large part to the use of a cylindrical Penning trap with a resonant cavity that interacts with the trapped electron in a readily calculable way, and through its high $Q$ resonances, significantly increases the lifetime of the electron in its lowest few energy states by inhibiting the decay of these states through spontaneous emission. Further, cooling the trap and its vacuum enclosure to 100 mK by means of a dilution refrigerator eliminates blackbody radiation that could excite the electron from these states.

The frequencies $f_{\mathrm{a}}$ and $f_{\mathrm{c}}$ are determined by applying quantum-jump spectroscopy (QJS) to transitions between the lowest spin $\left(m_{\mathrm{s}}= \pm 1 / 2\right)$ and cyclotron ( $n$ $=0,1,2$ ) quantum states of the electron in the trap. (In QJS, the quantum jumps per attempt to drive them are measured as a function of drive frequency.) The transitions are induced by applying a signal of frequency $\approx f_{\mathrm{a}}$ to trap electrodes or by transmitting microwaves of frequency $\approx f_{\mathrm{c}}$ into the trap cavity. A change in the cyclotron or spin state of the electron is reflected in a shift in $\bar{\nu}_{z}$, the self-excited axial oscillation of the electron. (The
trap axis and $B$ are in the $z$ direction.) This oscillation induces a signal in a resonant circuit that is amplified and fed back to the trap to drive the oscillation. Saturated nickel rings surrounding the trap produce a small magnetic bottle that provides quantum nondemolition couplings of the spin and cyclotron energies to $\bar{\nu}_{z}$. Failure to resolve the cyclotron energy levels would result in an increase of uncertainty due to the leading relativistic correction $\delta / f_{\mathrm{c}} \equiv h f_{\mathrm{c}} / m c^{2} \approx 10^{-9}$.

Another unique feature of the Harvard experiment is that the effect of the trap cavity modes on $f_{c}$, and hence on the measured value of $a_{\mathrm{e}}$, are directly observed for the first time. The modes are quantitatively identified as the familiar transverse electric (TE) and transverse magnetic (TM) modes by observing the response of a cloud of electrons to an axial parametric drive, and, based on the work of Brown and Gabrielse (1986), the range of possible shifts of $f_{\mathrm{c}}$ for a cylindrical cavity with a $Q$ $>500$ as used in the Harvard experiment can be readily calculated. Two measurements of $a_{\mathrm{e}}$ were made: one, which resulted in the value of $a_{\mathrm{e}}$ given in Eq. (103), was at a value of $B$ for which $f_{\mathrm{c}}=149 \mathrm{GHz}$, far from modes that couple to the cyclotron motion; the other was at 146.8 GHz , close to mode $\mathrm{TE}_{127}$. Within the calibration and identification uncertainties for the mode frequencies, good agreement was found between the measured and predicted difference in the two values. Indeed, their weighted mean gives a value of $a_{\mathrm{e}}$ that is larger than the value in Eq. (103) by only the fractional amount 0.5 $\times 10^{-10}$, with $u_{\mathrm{r}}$ slightly reduced to $6.5 \times 10^{-10}$.

The largest component of uncertainty, $5.2 \times 10^{-10}$, in the $6.6 \times 10^{-10} u_{\mathrm{r}}$ of the Harvard result for $a_{\mathrm{e}}$ arises from fitting the resonance line shapes for $f_{\mathrm{a}}$ and $f_{\mathrm{c}}$ obtained from the quantum jump spectroscopy data. It is based on the consistency of three different methods of extracting these frequencies from the line shapes. The method that yielded the best fits and which was used to obtain the reported value of $a_{\mathrm{e}}$ weights each drive frequency, spin flip or cyclotron, by the number of quantum jumps it produces, and then uses the weighted average of the resulting spin flip and cyclotron frequencies in the final calculation of $a_{\mathrm{e}}$. Although the cavity shifts are well characterized, they account for the second largest fractional uncertainty component, $3.4 \times 10^{-10}$. The statistical (Type A) component, which is the next largest, is only $1.5 \times 10^{-10}$.

## 3. Values of $\boldsymbol{\alpha}$ inferred from $\boldsymbol{a}_{\mathrm{e}}$

Equating the theoretical expression with the two experimental values of $a_{\mathrm{e}}$ given in Eqs. (102) and (103) yields

$$
\begin{equation*}
\alpha^{-1}\left(a_{\mathrm{e}}\right)=137.03599883(50) \quad\left[3.7 \times 10^{-9}\right] \tag{104}
\end{equation*}
$$

from the University of Washington result and

$$
\begin{equation*}
\alpha^{-1}\left(a_{\mathrm{e}}\right)=137.035999711(96) \quad\left[7.0 \times 10^{-10}\right] \tag{105}
\end{equation*}
$$

from the Harvard University result. The contribution of the uncertainty in $a_{\mathrm{e}}(\mathrm{th})$ to the relative uncertainty of
either of these results is $2.4 \times 10^{-10}$. The value in Eq. (105) has the smallest uncertainty of any value of alpha currently available. Both values are included in Table XIV.

## B. Muon magnetic moment anomaly $a_{\mu}$

Comparison of theory and experiment for the muon magnetic moment anomaly gives a test of the theory of the hadronic contributions, with the possibility of revealing physics beyond the standard model.

## 1. Theory of $a_{\mu}$

The current theory of $a_{\mu}$ has been throughly reviewed in a number of recent publications; see, for example, Davier et al. (2006), Jegerlehner (2007), Melnikov and Vainshtein (2006), Miller et al. (2007), and Passera (2005).

The relevant mass-dependent terms and corresponding contributions to $a_{\mu}($ th $)$, based on the 2006 recommended values of the mass ratios, are

$$
\begin{align*}
A_{2}^{(4)}\left(m_{\mu} / m_{\mathrm{e}}\right) & =1.0942583088(82) \\
& \rightarrow 506386.4561(38) \times 10^{-8} a_{\mu},  \tag{106}\\
A_{2}^{(4)}\left(m_{\mu} / m_{\tau}\right) & =0.000078064(25) \\
& \rightarrow 36.126(12) \times 10^{-8} a_{\mu},  \tag{107}\\
A_{2}^{(6)}\left(m_{\mu} / m_{\mathrm{e}}\right) & =22.86837997(19) \\
& \rightarrow 24581.76616(20) \times 10^{-8} a_{\mu},  \tag{108}\\
A_{2}^{(6)}\left(m_{\mu} / m_{\tau}\right) & =0.00036051(21) \\
& \rightarrow 0.38752(22) \times 10^{-8} a_{\mu},  \tag{109}\\
A_{2}^{(8)}\left(m_{\mu} / m_{\mathrm{e}}\right) & =132.6823(72) \rightarrow 331.288(18) \times 10^{-8} a_{\mu}, \tag{110}
\end{align*}
$$

$$
\begin{equation*}
A_{2}^{(10)}\left(m_{\mu} / m_{\mathrm{e}}\right)=663(20) \tag{111}
\end{equation*}
$$

$$
\begin{equation*}
\rightarrow 3.85(12) \times 10^{-8} a_{\mu} \tag{112}
\end{equation*}
$$

$$
A_{3}^{(6)}\left(m_{\mu} / m_{\mathrm{e}}, m_{\mu} / m_{\tau}\right)=0.00052766(17)
$$

$$
\begin{equation*}
\rightarrow 0.56720(18) \times 10^{-8} a_{\mu} \tag{113}
\end{equation*}
$$

$$
A_{3}^{(8)}\left(m_{\mu} / m_{\mathrm{e}}, m_{\mu} / m_{\tau}\right)=0.037594(83)
$$

$$
\begin{equation*}
\rightarrow 0.09387(21) \times 10^{-8} a_{\mu} \tag{114}
\end{equation*}
$$

These contributions and their uncertainties, as well as the values (including their uncertainties) of $a_{\mu}$ (weak) and $a_{\mu}(\mathrm{had})$ given below, should be compared with the $54 \times 10^{-8} a_{\mu}$ standard uncertainty of the experimental value of $a_{\mu}$ from Brookhaven National Laboratory (BNL) (see next section).

Some of the above terms reflect the results of recent calculations. The value of $A_{2}^{(6)}\left(m_{\mu} / m_{\tau}\right)$ in Eq. (109) includes an additional contribution as discussed in
connection with Eq. (91). The terms $A_{2}^{(8)}\left(m_{\mu} / m_{\mathrm{e}}\right)$ and $A_{3}^{(8)}\left(m_{\mu} / m_{\mathrm{e}}, m_{\mu} / m_{\tau}\right)$ have been updated by Kinoshita and Nio (2004), with the resulting value for $A_{2}^{(8)}\left(m_{\mu} / m_{\mathrm{e}}\right)$ in Eq. (110) differing from the previous value of 127.50 (41) due to the elimination of various problems with earlier calculations, and the resulting value for $A_{3}^{(8)}\left(m_{\mu} / m_{\mathrm{e}}, m_{\mu} / m_{\tau}\right)$ in Eq. (114) differing from the previous value of $0.079(3)$, because diagrams that were thought to be negligible do in fact contribute to the result. Further, the value for $A_{2}^{(10)}\left(m_{\mu} / m_{\mathrm{e}}\right)$ in Eq. (111) from Kinoshita and Nio (2006b) replaces the previous value, 930 (170). These authors believed that their result, obtained from the numerical evaluation of all integrals from 17 key subsets of Feynman diagrams, accounts for the leading contributions to $A_{2}^{(10)}\left(m_{\mu} / m_{\mathrm{e}}\right)$, and the work of Kataev (2006), based on the so-called renormalization-group-inspired scheme-invariant approach, supports this view.

The electroweak contribution to $a_{\mu}(\mathrm{th})$ is taken to be

$$
\begin{equation*}
a_{\mu}(\text { weak })=154(2) \times 10^{-11} \tag{115}
\end{equation*}
$$

as given by Czarnecki et al. $(2003,2006)$. This value was used in the 2002 adjustment and is discussed in CODATA-02.

The hadronic contribution to $a_{\mu}($ th $)$ may be written as

$$
\begin{equation*}
a_{\mu}(\mathrm{had})=a_{\mu}^{(4)}(\mathrm{had})+a_{\mu}^{(6 a)}(\mathrm{had})+a_{\mu}^{(\gamma \gamma)}(\mathrm{had})+\cdots \tag{116}
\end{equation*}
$$

where $a_{\mu}^{(4)}($ had $)$ and $a_{\mu}^{(6 a)}($ had $)$ arise from hadronic vacuum polarization and are of order $(\alpha / \pi)^{2}$ and $(\alpha / \pi)^{3}$, respectively; and $a_{\mu}^{(\gamma \gamma)}(\mathrm{had})$, which arises from hadronic light-by-light vacuum polarization, is also of order $(\alpha / \pi)^{3}$.

Values of $a_{\mu}^{(4)}$ (had) are obtained from calculations that evaluate dispersion integrals over measured cross sections for the scattering of $\mathrm{e}^{+} \mathrm{e}^{-}$into hadrons. In addition, in some such calculations, data on decays of the $\tau$ into hadrons are used to replace the $\mathrm{e}^{+} \mathrm{e}^{-}$data in certain parts of the calculation. In the 2002 adjustment, results from both types of calculation were averaged to obtain a value that would be representative of both approaches.

There have been improvements in the calculations that use only $\mathrm{e}^{+} \mathrm{e}^{-}$data with the addition of new data from the detectors CMD-2 at Novosibirsk, KLOE at Frascati, BaBar at the Stanford Linear Accelerator Center, and corrected data from the detector SND at Novosibirsk (Davier, 2007; Hagiwara et al., 2007; Jegerlehner, 2007). However, there is a persistent disagreement between the results that include the $\tau$ decay data and those that use only $\mathrm{e}^{+} \mathrm{e}^{-}$data. In view of the improvements in the results based solely on $\mathrm{e}^{+} \mathrm{e}^{-}$data and the unresolved questions concerning the assumptions required to incorporate the $\tau$ data into the analysis (Melnikov and Vainshtein, 2006; Davier, 2006; Davier et al., 2006), we use in the 2006 adjustment results based solely on $\mathrm{e}^{+} \mathrm{e}^{-}$data. The value employed is

$$
\begin{equation*}
a_{\mu}^{(4)}(\mathrm{had})=690(21) \times 10^{-10} \tag{117}
\end{equation*}
$$

which is the unweighted mean of the values $a_{\mu}^{(4)}(\mathrm{had})$ $=689.4(4.6) \times 10^{-10} \quad$ (Hagiwara et al., 2007) and $a_{\mu}^{(4)}(\mathrm{had})=690.9(4.4) \times 10^{-10}$ (Davier, 2007). The uncertainty assigned to the value of $a_{\mu}^{(4)}$ (had), as expressed in Eq. (117), is essentially the difference between the values that include $\tau$ data and those that do not. In particular, the result that includes $\tau$ data that we use to estimate the uncertainty is $711.0(5.8) \times 10^{-11}$ from Davier et al. (2003); the value of $a_{\mu}^{(4)}$ (had) used in the 2002 adjustment was based in part on this result. Although there is the smaller value $701.8(5.8) \times 10^{-11}$ from Trocóniz and Ynduráin (2005), we use only the larger value in order to obtain an uncertainty that covers the possibility of physics beyond the standard model not included in the calculation of $a_{\mu}(\mathrm{th})$. Other, mostly older results for $a_{\mu}^{(4)}(\mathrm{had})$, but which in general agree with the two values we have averaged, are summarized in Table III of Jegerlehner (2007).

For the second term in Eq. (116), we employ the value

$$
\begin{equation*}
a_{\mu}^{(6 a)}(\mathrm{had})=-97.90(95) \times 10^{-11} \tag{118}
\end{equation*}
$$

calculated by Hagiwara et al. (2004), which was also used in the 2002 adjustment.
The light-by-light contribution in Eq. (116) has been calculated by Melnikov and Vainshtein (2004, 2006), who obtain the value

$$
\begin{equation*}
a_{\mu}^{(\gamma \gamma)}(\mathrm{had})=136(25) \times 10^{-11} \tag{119}
\end{equation*}
$$

It is somewhat larger than earlier results, because it includes short distance constraints imposed by quantum chromodynamics (QCD) that were not included in the previous calculations. It is consistent with the $95 \%$ confidence limit upper bound of $159 \times 10^{-11}$ for $a_{\mu}^{(\gamma \gamma)}(\mathrm{had})$ obtained by Erler and Sánchez (2006), the value $110(40) \times 10^{-11}$ proposed by Bijnens and Prades (2007), and the value $125(35) \times 10^{-11}$ suggested by Davier and Marciano (2004).

The total hadronic contribution is

$$
\begin{equation*}
a_{\mu}(\mathrm{had})=694(21) \times 10^{-10}=595(18) \times 10^{-7} a_{\mu} \tag{120}
\end{equation*}
$$

Combining terms in $a_{\mu}(\mathrm{QED})$ that have like powers of $\alpha / \pi$, we summarize the theory of $a_{\mu}$ as follows:

$$
\begin{equation*}
a_{\mu}(\mathrm{th})=a_{\mu}(\mathrm{QED})+a_{\mu}(\text { weak })+a_{\mu}(\mathrm{had}) \tag{121}
\end{equation*}
$$

where

$$
\begin{align*}
a_{\mu}(\mathrm{QED})= & C_{\mu}^{(2)}\left(\frac{\alpha}{\pi}\right)+C_{\mu}^{(4)}\left(\frac{\alpha}{\pi}\right)^{2}+C_{\mu}^{(6)}\left(\frac{\alpha}{\pi}\right)^{3} \\
& +C_{\mu}^{(8)}\left(\frac{\alpha}{\pi}\right)^{4}+C_{\mu}^{(10)}\left(\frac{\alpha}{\pi}\right)^{5}+\cdots \tag{122}
\end{align*}
$$

with

$$
\begin{aligned}
& C_{\mu}^{(2)}=0.5, \\
& C_{\mu}^{(4)}=0.765857408(27), \\
& C_{\mu}^{(6)}=24.05050959(42),
\end{aligned}
$$

$$
\begin{align*}
& C_{\mu}^{(8)}=130.9916(80), \\
& C_{\mu}^{(10)}=663(20), \tag{123}
\end{align*}
$$

and $a_{\mu}$ (weak) and $a_{\mu}$ (had) are as given in Eqs. (115) and (120). The standard uncertainty of $a_{\mu}($ th $)$ from the uncertainties of the terms listed above, other than that due to $\alpha$, is

$$
\begin{equation*}
u\left[a_{\mu}(\mathrm{th})\right]=2.1 \times 10^{-9}=1.8 \times 10^{-6} a_{\mu} \tag{124}
\end{equation*}
$$

and is primarily due to the uncertainty of $a_{\mu}$ (had).
For the purpose of the least-squares calculations carried out in Sec. XII.B, we define an additive correction $\delta_{\mu}$ to $a_{\mu}($ th $)$ to account for the lack of exact knowledge of $a_{\mu}($ th $)$, and hence the complete theoretical expression for the muon anomaly is

$$
\begin{equation*}
a_{\mu}\left(\alpha, \delta_{\mu}\right)=a_{\mu}(\mathrm{th})+\delta_{\mu} \tag{125}
\end{equation*}
$$

Our theoretical estimate of $\delta_{\mu}$ is zero and its standard uncertainty is $u\left[a_{\mu}(\right.$ th $\left.)\right]$,

$$
\begin{equation*}
\delta_{\mu}=0.0(2.1) \times 10^{-9} . \tag{126}
\end{equation*}
$$

Although $a_{\mu}(\mathrm{th})$ and $a_{\mathrm{e}}(\mathrm{th})$ have some common components of uncertainty, the covariance of $\delta_{\mu}$ and $\delta_{\mathrm{e}}$ is negligible.

## 2. Measurement of $\boldsymbol{a}_{\mu}$ : Brookhaven

Experiment E821 at Brookhaven National Laboratory (BNL), Upton, New York, was initiated by the Muon $g-2$ Collaboration in the early 1980s with the goal of measuring $a_{\mu}$ with a significantly smaller uncertainty than $u_{\mathrm{r}}=7.2 \times 10^{-6}$. This is the uncertainty achieved in the third $g-2$ experiment carried out at the European Organization for Nuclear Research (CERN), Geneva, Switzerland, in the mid-1970s using both positive and negative muons and which was the culmination of nearly 20 years of effort (Bailey et al., 1979).

The basic principle of the experimental determination of $a_{\mu}$ is similar to that used to determine $a_{\mathrm{e}}$ and involves measuring the anomaly difference frequency $f_{\mathrm{a}}=f_{\mathrm{s}}-f_{\mathrm{c}}$, where $f_{\mathrm{s}}=\left|g_{\mu}\right|\left(e \hbar / 2 m_{\mu}\right) B / h$ is the muon spin-flip (often called precession) frequency in the applied magnetic flux density $B$ and $f_{\mathrm{c}}=e B / 2 \pi m_{\mu}$ is the corresponding muon cyclotron frequency. However, instead of eliminating $B$ by measuring $f_{\mathrm{c}}$ as is done for the electron, $B$ is determined from proton nuclear magnetic resonance (NMR) measurements. As a consequence, the value of $\mu_{\mu} / \mu_{\mathrm{p}}$ is required to deduce the value of $a_{\mu}$ from the data. The relevant equation is

$$
\begin{equation*}
a_{\mu}=\frac{\bar{R}}{\left|\mu_{\mu} / \mu_{\mathrm{p}}\right|-\bar{R}}, \tag{127}
\end{equation*}
$$

where $\bar{R}=f_{\mathrm{a}} / \bar{f}_{\mathrm{p}}$ and $\bar{f}_{\mathrm{p}}$ is the free proton NMR frequency corresponding to the average flux density seen by the muons in their orbits in the muon storage ring used in the experiment. (Of course, in the corresponding experiment for the electron, a Penning trap is employed rather than a storage ring.)

The BNL $a_{\mu}$ experiment was discussed in both CODATA-98 and CODATA-02. In the 1998 adjustment, the CERN final result for $\bar{R}$ with $u_{\mathrm{r}}=7.2 \times 10^{-6}$, and the first BNL result for $\bar{R}$, obtained from the 1997 engineering run using positive muons and with $u_{\mathrm{r}}=13$ $\times 10^{-6}$, were taken as input data. By the time of the 2002 adjustment, the BNL experiment had progressed to the point where the CERN result was no longer competitive, and the input datum used was the BNL mean value of $\bar{R}$ with $u_{\mathrm{r}}=6.7 \times 10^{-7}$ obtained from the 1998,1999 , and 2000 runs using $\mu^{+}$. The final run of the BNL E821 experiment was carried out in 2001 with $\mu^{-}$and achieved an uncertainty for $\bar{R}$ of $u_{\mathrm{r}}=7.0 \times 10^{-7}$, but the result only became available in early 2004, well after the closing date of the 2002 adjustment.

Based on the data obtained in all five runs and assuming $C P T$ invariance, an assumption justified by the consistency of the values of $\bar{R}$ obtained from either $\mu^{+}$or $\mu^{-}$, the final report on the E821 experiment gives as the final value of $\frac{p}{R}$ (Bennett et al., 2006) [see also (Miller et al., 2007)]:

$$
\begin{equation*}
\bar{R}=0.0037072064(20) \quad\left[5.4 \times 10^{-7}\right] \tag{128}
\end{equation*}
$$

which we take as an input datum in the 2006 adjustment. A new BNL experiment to obtain a value of $\bar{R}$ with a smaller uncertainty is under discussion (Hertzog 2007).

The experimental value of $a_{\mu}$ implied by this value of $\bar{R}$ is, from Eq. (127) and the 2006 recommended value of $\mu_{\mu} / \mu_{\mathrm{p}}$ (the uncertainty of which is inconsequential in this application),

$$
\begin{equation*}
a_{\mu}(\exp )=1.16592093(63) \times 10^{-3} \quad\left[5.4 \times 10^{-7}\right] \tag{129}
\end{equation*}
$$

Further, with the aid of Eq. (217) in Sec. VI.B, Eq. (127) can be written as

$$
\begin{equation*}
\bar{R}=-\frac{a_{\mu}\left(\alpha, \delta_{\mu}\right)}{1+a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)} \frac{m_{\mathrm{e}}}{m_{\mu}} \frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}}, \tag{130}
\end{equation*}
$$

where we have used $g_{\mathrm{e}}=-2\left(1+a_{\mathrm{e}}\right)$ and $g_{\mu}=-2\left(1+a_{\mu}\right)$ and replaced $a_{\mathrm{e}}$ and $a_{\mu}$ with their complete theoretical expressions $a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)$ and $a_{\mu}\left(\alpha, \delta_{\mu}\right)$, discussed in Secs. V.A. 1 and V.B.1, respectively. Equation (130) is, in fact, the observational equation for the input datum $\bar{R}$.
a. Theoretical value of $a_{\mu}$ and inferred value of $\alpha$

Evaluation of the theoretical expression for $a_{\mu}$ in Eq. (121) with the 2006 recommended value of $\alpha$, the uncertainty of which is negligible in this context, yields

$$
\begin{equation*}
a_{\mu}(\mathrm{th})=1.1659181(21) \times 10^{-3} \quad\left[1.8 \times 10^{-6}\right] \tag{131}
\end{equation*}
$$

which may be compared to the value in Eq. (129) deduced from the BNL result for $\bar{R}$ given in Eq. (128). The experimental value exceeds the theoretical value by $1.3 u_{\text {diff }}$, where $u_{\text {diff }}$ is the standard uncertainty of the difference. It should be recognized, however, that this agreement is a consequence of the comparatively large uncertainty we have assigned to $a_{\mu}^{(4)}$ (had) [see Eq. (120)].

If the result for $a_{\mu}^{(4)}(\mathrm{had})$ that includes tau data were ignored and the uncertainty of $a_{\mu}^{(4)}$ (had) were based on the estimated uncertainties of the calculated values using only $\mathrm{e}^{+} \mathrm{e}^{-}$data, then the experimental value would exceed the theoretical value by $3.5 u_{\text {diff }}$. This inconsistency is well known to the high-energy physics community and is of considerable interest because it may be an indication of "new physics" beyond the Standard Model, such as supersymmetry (Stöckinger, 2007).

One might ask, why include the theoretical value for $a_{\mu}$ in the 2006 adjustment given its current problems? By retaining the theoretical expression with an increased uncertainty, we ensure that the 2006 recommended value of $a_{\mu}$ reflects, albeit with a comparatively small weight, the existence of the theoretical value.

The consistency between theory and experiment may also be examined by considering the value of $\alpha$ obtained by equating the theoretical expression for $a_{\mu}$ with the BNL experimental value, as done for $a_{\mathrm{e}}$ in Sec. V.A.3. The result is

$$
\begin{equation*}
\alpha^{-1}=137.03567(26) \quad\left[1.9 \times 10^{-6}\right] \tag{132}
\end{equation*}
$$

which is the value included in Table XIV.

## C. Bound electron $g$-factor in ${ }^{12} \mathrm{C}^{5+}$ and in ${ }^{16} \mathrm{O}^{7+}$ and $A_{\mathrm{r}}(\mathrm{e})$

Precise measurements and theoretical calculations for the $g$-factor of the electron in hydrogenic ${ }^{12} \mathrm{C}$ and in hydrogenic ${ }^{16} \mathrm{O}$ lead to values of $A_{\mathrm{r}}(\mathrm{e})$ that contribute to the determination of the 2006 recommended value of this important constant.

For a ground-state hydrogenic ion ${ }^{A} X^{(Z-1)+}$ with mass number $A$, atomic number (proton number) $Z$, nuclear spin quantum number $i=0$, and $g$-factor $g_{\mathrm{e}}-\left({ }^{A} X^{(Z-1)+}\right)$ in an applied magnetic flux density $B$, the ratio of the electron's spin-flip (often called precession) frequency $f_{\mathrm{s}}$ $=\left|g_{\mathrm{e}^{-}}\left({ }^{A} X^{(Z-1)+}\right)\right|\left(e \hbar / 2 m_{\mathrm{e}}\right) B / h$ to the cyclotron frequency of the ion $f_{\mathrm{c}}=(Z-1) e B / 2 \pi m\left({ }^{A} X^{(Z-1)+}\right)$ in the same magnetic flux density is

$$
\begin{equation*}
\frac{f_{\mathrm{s}}\left({ }^{A} X^{(Z-1)+}\right)}{f_{\mathrm{c}}\left({ }^{A} X^{(Z-1)+}\right)}=-\frac{g_{\mathrm{e}}\left({ }^{A} X^{(Z-1)+}\right)}{2(Z-1)} \frac{A_{\mathrm{r}}\left({ }^{A} X^{(Z-1)+}\right)}{A_{\mathrm{r}}(\mathrm{e})}, \tag{133}
\end{equation*}
$$

where $A_{\mathrm{r}}(X)$ is the relative atomic mass of particle $X$. If the frequency ratio $f_{\mathrm{s}} / f_{\mathrm{c}}$ is determined experimentally with high accuracy, and $A_{\mathrm{r}}\left({ }^{A} X^{(Z-1)+}\right)$ of the ion is also accurately known, then this expression can be used to determine an accurate value of $A_{\mathrm{r}}(\mathrm{e})$, assuming the bound-state electron $g$-factor can be calculated from QED theory with sufficient accuracy; or the $g$-factor can be determined if $A_{\mathrm{r}}(\mathrm{e})$ is accurately known from another experiment. In fact, a broad program involving a number of European laboratories has been underway since the mid-1990s to measure the frequency ratio and calculate the $g$-factor for different ions, most notably (to date) ${ }^{12} \mathrm{C}^{5+}$ and ${ }^{16} \mathrm{O}^{7+}$. The measurements themselves are being performed at the Gesellschaft für Schwerionenforschung, Darmstadt, Germany (GSI) by GSI and University of Mainz researchers, and we discuss the ex-

TABLE XV. Theoretical contributions and total for the $g$-factor of the electron in hydrogenic carbon 12 based on the 2006 recommended values of the constants.

| Contribution | Value | Source |
| :--- | ---: | :---: |
| Dirac $g_{\mathrm{D}}$ | $-1.998721354402(2)$ | Eq. (138) |
| $\Delta g_{\mathrm{SE}}^{(2)}$ | $-0.002323672426(4)$ | Eq. (146) |
| $\Delta g_{\mathrm{VP}}^{(2)}$ | $0.000000008512(1)$ | Eq. (152) |
| $\Delta g^{(4)}$ | $0.000003545677(25)$ | Eq. (156) |
| $\Delta g^{(6)}$ | -0.000000029618 | Eq. (158) |
| $\Delta g^{(8)}$ | 0.000000000101 | Eq. (159) |
| $\Delta g^{(10)}$ | $0.000000000000(1)$ | Eq. (160) |
| $\Delta g_{\text {rec }}$ | $-0.000000087639(10)$ | Eqs. (161)-(163) |
| $\Delta g_{\text {ns }}$ | $-0.000000000408(1)$ | Eq. (165) |
| $g_{\mathrm{e}-}{ }^{\left({ }^{12} \mathrm{C}^{5+}\right)}$ | $-2.001041590203(28)$ | Eq. (166) |

perimental determinations of $f_{\mathrm{s}} / f_{\mathrm{c}}$ for ${ }^{12} \mathrm{C}^{5+}$ and ${ }^{16} \mathrm{O}^{7+}$ at GSI in Secs. V.C.2.a and V.C.2.b. The theoretical expressions for the bound-electron $g$-factors of these two ions are reviewed in the next section.

## 1. Theory of the bound electron $g$-factor

In this section, we consider an electron in the 1 S state of hydrogenlike carbon 12 or oxygen 16 within the framework of bound-state QED. The measured quantity is the transition frequency between the two Zeeman levels of the atom in an externally applied magnetic field.
The energy of a free electron with spin projection $s_{z}$ in a magnetic flux density $B$ in the $z$ direction is

$$
\begin{equation*}
E=-\boldsymbol{\mu} \cdot \boldsymbol{B}=-g_{\mathrm{e}} \frac{e}{2 m_{\mathrm{e}}} s_{z} B, \tag{134}
\end{equation*}
$$

and hence the spin-flip energy difference is

$$
\begin{equation*}
\Delta E=-g_{\mathrm{e}}-\mu_{\mathrm{B}} B . \tag{135}
\end{equation*}
$$

(In keeping with the definition of the $g$-factor in Sec. V, the quantity $g_{\mathrm{e}^{-}}$is negative.) The analogous expression for ions with no nuclear spin is

$$
\begin{equation*}
\Delta E_{\mathrm{b}}(X)=-g_{\mathrm{e}^{-}}(X) \mu_{\mathrm{B}} B, \tag{136}
\end{equation*}
$$

which defines the bound-state electron $g$-factor, and where $X$ is either ${ }^{12} \mathrm{C}^{5+}$ or ${ }^{16} \mathrm{O}^{7+}$.
The theoretical expression for $g_{\mathrm{e}}-(X)$ is written as

$$
\begin{equation*}
g_{\mathrm{e}^{-}}(X)=g_{\mathrm{D}}+\Delta g_{\mathrm{rad}}+\Delta g_{\mathrm{rec}}+\Delta g_{\mathrm{ns}}+\cdots \tag{137}
\end{equation*}
$$

where the individual terms are the Dirac value, the radiative corrections, the recoil corrections, and the nuclear size corrections, respectively. These theoretical contributions are discussed in the following paragraphs; numerical results based on the 2006 recommended values are summarized in Tables XV and XVI. In the 2006 adjustment, $\alpha$ in the expression for $g_{\mathrm{D}}$ is treated as a variable, but the constants in the rest of the calculation of the $g$-factors are taken as fixed quantities.
Breit (1928) obtained the exact value

TABLE XVI. Theoretical contributions and total for the $g$-factor of the electron in hydrogenic oxygen 16 based on the 2006 recommended values of the constants.

| Contribution | Value | Source |
| :--- | :---: | :---: |
| Dirac $g_{\mathrm{D}}$ | -1.99772600308 | Eq. $(138)$ |
| $\Delta g_{\mathrm{SE}}^{(2)}$ | $-0.00232444212(1)$ | Eq. $(146)$ |
| $\Delta g_{\mathrm{VP}}^{(2)}$ | 0.00000002638 | Eq. $(152)$ |
| $\Delta g^{(4)}$ | $0.00000354654(11)$ | Eq. $(156)$ |
| $\Delta g^{(6)}$ | -0.00000002963 | Eq. $(158)$ |
| $\Delta g^{(8)}$ | 0.00000000010 | Eq. $(159)$ |
| $\Delta g^{(10)}$ | 0.00000000000 | Eq. $(160)$ |
| $\Delta g_{\text {rec }}$ | $-0.00000011702(1)$ | Eqs. $(161)-(163)$ |
| $\Delta g_{\text {ns }}$ | $-0.00000000156(1)$ | Eq. $(165)$ |
| $g_{\mathrm{e}^{-}\left({ }^{16} \mathrm{O}^{7+}\right)}$ | $-2.00004702038(11)$ | Eq. $(166)$ |

$$
\begin{align*}
g_{\mathrm{D}} & =-\frac{2}{3}\left[1+2 \sqrt{1-(Z \alpha)^{2}}\right] \\
& =-2\left[1-\frac{1}{3}(Z \alpha)^{2}-\frac{1}{12}(Z \alpha)^{4}-\frac{1}{24}(Z \alpha)^{6}+\cdots\right] \tag{138}
\end{align*}
$$

from the Dirac equation for an electron in the field of a fixed-point charge of magnitude $Z e$, where the only uncertainty is that due to the uncertainty in $\alpha$.
The radiative corrections may be written as

$$
\begin{equation*}
\Delta g_{\mathrm{rad}}=-2\left[C_{\mathrm{e}}^{(2)}(Z \alpha)\left(\frac{\alpha}{\pi}\right)+C_{\mathrm{e}}^{(4)}(Z \alpha)\left(\frac{\alpha}{\pi}\right)^{2}+\cdots\right] \tag{139}
\end{equation*}
$$

where the coefficients $C_{\mathrm{e}}^{(2 n)}(Z \alpha)$, corresponding to $n$ virtual photons, are slowly varying functions of $Z \alpha$. These coefficients are defined in direct analogy with the corresponding coefficients for the free electron $C_{\mathrm{e}}^{(2 n)}$ given in Eq. (98) so that

$$
\begin{equation*}
\lim _{Z \alpha \rightarrow 0} C_{\mathrm{e}}^{(2 n)}(Z \alpha)=C_{\mathrm{e}}^{(2 n)} \tag{140}
\end{equation*}
$$

The first two terms of the coefficient $C_{\mathrm{e}}^{(2)}(Z \alpha)$ have been known for some time (Faustov, 1970; Grotch, 1970; Close and Osborn, 1971). Recently, Pachucki et al. (2004), Pachucki, Czarnecki, Jentschura, et al. (2005), and Pachucki, Jentschura, and Yerokhin (2005) calculated additional terms with the result

$$
\begin{align*}
C_{\mathrm{e}, \mathrm{SE}}^{(2)}(Z \alpha)= & \frac{1}{2}\left\{1+\frac{(Z \alpha)^{2}}{6}+(Z \alpha)^{4}\left[\frac{32}{9} \ln (Z \alpha)^{-2}\right.\right. \\
& \left.+\frac{247}{216}-\frac{8}{9} \ln k_{0}-\frac{8}{3} \ln k_{3}\right] \\
& \left.+(Z \alpha)^{5} R_{\mathrm{SE}}(Z \alpha)\right\}, \tag{141}
\end{align*}
$$

where

$$
\begin{align*}
& \ln k_{0}=2.984128556,  \tag{142}\\
& \ln k_{3}=3.272806545, \tag{143}
\end{align*}
$$

$$
\begin{equation*}
R_{\mathrm{SE}}(6 \alpha)=22.160(10), \tag{144}
\end{equation*}
$$

$$
\begin{equation*}
R_{\mathrm{SE}}(8 \alpha)=21.859(4) \tag{145}
\end{equation*}
$$

The quantity $\ln k_{0}$ is the Bethe logarithm for the 1 S state (see Table VII) and $\ln k_{3}$ is a generalization of the Bethe logarithm relevant to the $g$-factor calculation. The remainder function $R_{\mathrm{SE}}(Z \alpha)$ was obtained by Pachucki et al. (2004) and Pachucki, Jentschura, and Yerokhin, (2005) by extrapolation of numerical calculations of the self energy for $Z>8$ by Yerokhin et al. (2002) using Eq. (141) to remove the lower-order terms. For $Z=6$ and 8 , this yields

$$
\begin{align*}
& C_{\mathrm{e}, \mathrm{SE}}^{(2)}(6 \alpha)=0.50018360665(80) \\
& C_{\mathrm{e}, \mathrm{SE}}^{(2)}(8 \alpha)=0.5003492887(14) \tag{146}
\end{align*}
$$

The lowest-order vacuum-polarization correction consists of a wave-function correction and a potential correction. The wave-function correction has been calculated numerically by Beier et al. (2000) with the result (in our notation)

$$
\begin{align*}
& C_{\mathrm{e}, \mathrm{VPwf}}^{(2)}(6 \alpha)=-0.0000018403431(43) \\
& C_{\mathrm{e}, \mathrm{VPwf}}^{(2)}(8 \alpha)=-0.000005712028(26) \tag{147}
\end{align*}
$$

Each of these values is the sum of the Uehling potential contribution and the higher-order Wichmann-Kroll contribution, which were calculated separately with the uncertainties added linearly, as done by Beier et al. (2000). The values in Eq. (147) are consistent with an evaluation of the correction in powers of $Z \alpha$ (Karshenboim, 2000; Karshenboim et al., 2001a, 2001b). For the potential correction, Beier et al. (2000) found that the Uehling potential contribution is zero and calculated the WichmannKroll contribution numerically over a wide range of $Z$ (Beier, 2000). An extrapolation of the numerical values from higher $Z$, taken together with the analytic result of Karshenboim and Milstein (2002),

$$
\begin{equation*}
C_{\mathrm{e}, \mathrm{VPp}}^{(2)}(Z \alpha)=\frac{7 \pi}{432}(Z \alpha)^{5}+\cdots \tag{148}
\end{equation*}
$$

for the lowest-order Wichmann-Kroll contribution, yields

$$
\begin{align*}
& C_{\mathrm{e}, \mathrm{VPp}}^{(2)}(6 \alpha)=0.000000007 \text { 9595(69) } \\
& C_{\mathrm{e}, \mathrm{VPp}}^{(2)}(8 \alpha)=0.000000033235(29) \tag{149}
\end{align*}
$$

More recently, Lee et al. (2005) obtained the result

$$
\begin{align*}
& C_{\mathrm{e}, V \mathrm{Vp}}^{(2)}(6 \alpha)=0.000000008201(11) \\
& C_{\mathrm{e}, V \mathrm{Vp}}^{(2)}(8 \alpha)=0.00000003423(11) \tag{150}
\end{align*}
$$

The values in Eqs. (149) and (150) disagree somewhat, so in the present analysis we use a value that is an unweighted average of the two, with half the difference for the uncertainty. These average values are

$$
\begin{align*}
& C_{\mathrm{e}, \mathrm{VPp}}^{(2)}(6 \alpha)=0.00000000808(12) \\
& C_{\mathrm{e}, \mathrm{VPp}}^{(2)}(8 \alpha)=0.00000003373(50) \tag{151}
\end{align*}
$$

The total one-photon vacuum polarization coefficients are given by the sum of Eqs. (147) and (151),

$$
\begin{align*}
C_{\mathrm{e}, \mathrm{VP}}^{(2)}(6 \alpha) & =C_{\mathrm{e}, \mathrm{VPwf}}^{(2)}(6 \alpha)+C_{\mathrm{e}, \mathrm{VPp}}^{(2)}(6 \alpha) \\
& =-0.00000183226(12), \\
C_{\mathrm{e}, \mathrm{VP}}^{(2)}(8 \alpha) & =C_{\mathrm{e}, \mathrm{VPwf}}^{(2)}(8 \alpha)+C_{\mathrm{e}, \mathrm{VPp}}^{(2)}(8 \alpha) \\
& =-0.00000567830(50) . \tag{152}
\end{align*}
$$

The total for the one-photon coefficient $C_{\mathrm{e}}^{(2)}(Z \alpha)$, given by the sum of Eqs. (146) and (152), is

$$
\begin{align*}
C_{\mathrm{e}}^{(2)}(6 \alpha) & =C_{\mathrm{e}, \mathrm{SE}}^{(2)}(6 \alpha)+C_{\mathrm{e}, \mathrm{VP}}^{(2)}(6 \alpha) \\
& =0.50018177438(81) \\
C_{\mathrm{e}}^{(2)}(8 \alpha) & =C_{\mathrm{e}, \mathrm{SE}}^{(2)}(8 \alpha)+C_{\mathrm{e}, \mathrm{VP}}^{(2)}(8 \alpha) \\
& =0.5003436104(14) \tag{153}
\end{align*}
$$

and the total one-photon contribution $\Delta g^{(2)}$ to the $g$-factor is

$$
\begin{array}{rlr}
\Delta g^{(2)} & =-2 C_{\mathrm{e}}^{(2)}(Z \alpha)\left(\frac{\alpha}{\pi}\right) \\
& =-0.002323663914(4) & \text { for } Z=6 \\
& =-0.002324415746(7) & \text { for } Z=8 \tag{154}
\end{array}
$$

The separate one-photon self energy and vacuum polarization contributions to the $g$-factor are given in Tables XV and XVI.

Calculations by Eides and Grotch (1997a) using the Bargmann-Michel-Telegdi equation and by Czarnecki et al. (2001) using an effective potential approach yield

$$
\begin{equation*}
C_{\mathrm{e}}^{(2 n)}(Z \alpha)=C_{\mathrm{e}}^{(2 n)}\left[1+\frac{(Z \alpha)^{2}}{6}+\cdots\right] \tag{155}
\end{equation*}
$$

as the leading binding correction to the free-electron coefficients $C_{\mathrm{e}}^{(2 n)}$ for any order $n$. For $C_{\mathrm{e}}^{(2)}(Z \alpha)$, this correction was known for some time. For higher-order terms, it provides the leading binding effect.

The two-loop contribution of relative order $(Z \alpha)^{4}$ was recently calculated by Jentschura, et al. (2006) and Pachucki, Czarnecki, Jentschura, et al. (2005) for any S state. Their result for the ground-state correction is

$$
\begin{align*}
C_{\mathrm{e}}^{(4)}(Z \alpha)= & C_{\mathrm{e}}^{(4)}\left(1+\frac{(Z \alpha)^{2}}{6}\right)+(Z \alpha)^{4}\left[\frac{14}{9} \ln (Z \alpha)^{-2}\right. \\
& +\frac{991343}{155520}-\frac{2}{9} \ln k_{0}-\frac{4}{3} \ln k_{3}+\frac{679 \pi^{2}}{12960} \\
& \left.-\frac{1441 \pi^{2}}{720} \ln 2+\frac{1441}{480} \zeta(3)\right]+\mathcal{O}(Z \alpha)^{5} \\
= & -0.3285778(23) \quad \text { for } Z=6 \\
= & -0.3286578(97) \quad \text { for } Z=8 \tag{156}
\end{align*}
$$

where $\ln k_{0}$ and $\ln k_{3}$ are given in Eqs. (142) and (143). The uncertainty due to uncalculated terms is estimated by assuming that unknown higher-order terms, of order $(Z \alpha)^{5}$ or higher for two loops, are comparable to the
higher-order one-loop terms scaled by the free-electron coefficients in each case, with an extra factor of 2 included (Pachucki, Czarnecki, Jentschura, et al., 2005),

$$
\begin{equation*}
u\left[C_{\mathrm{e}}^{(4)}(Z \alpha)\right]=2\left|(Z \alpha)^{5} C_{\mathrm{e}}^{(4)} R_{\mathrm{SE}}(Z \alpha)\right| \tag{157}
\end{equation*}
$$

The three- and four-photon terms are calculated with the leading binding correction included,

$$
\begin{align*}
C_{\mathrm{e}}^{(6)}(Z \alpha) & =C_{\mathrm{e}}^{(6)}\left[1+\frac{(Z \alpha)^{2}}{6}+\cdots\right] \\
& =1.181611 \ldots \quad \text { for } Z=6 \\
& =1.181905 \ldots \quad \text { for } Z=8 \tag{158}
\end{align*}
$$

where $C_{\mathrm{e}}^{(6)}=1.181234 \ldots$, and

$$
\begin{align*}
C_{\mathrm{e}}^{(8)}(Z \alpha) & =C_{\mathrm{e}}^{(8)}\left[1+\frac{(Z \alpha)^{2}}{6}+\cdots\right] \\
& =-1.7289(35) \ldots \quad \text { for } Z=6 \\
& =-1.7293(35) \ldots \quad \text { for } Z=8 \tag{159}
\end{align*}
$$

where $C_{\mathrm{e}}^{(8)}=-1.7283(35) \quad$ (Kinoshita and Nio, 2006a). This value would shift somewhat if the more recent tentative value $C_{\mathrm{e}}^{(8)}=-1.9144(35)$ (Aoyama et al., 2007) were used (see Sec. V). An uncertainty estimate

$$
\begin{equation*}
C_{\mathrm{e}}^{(10)}(Z \alpha) \approx C_{\mathrm{e}}^{(10)}=0.0(3.7) \tag{160}
\end{equation*}
$$

is included for the five-loop correction.
The recoil correction to the bound-state $g$-factor associated with the finite mass of the nucleus is denoted by $\Delta g_{\text {rec }}$, which we write here as the sum $\Delta g_{\text {rec }}^{(0)}+\Delta g_{\text {rec }}^{(2)}$ corresponding to terms that are zero and first order in $\alpha / \pi$, respectively. For $\Delta g_{\text {rec }}^{(0)}$, we have

$$
\begin{align*}
\Delta g_{\mathrm{rec}}^{(0)}= & \left\{-(Z \alpha)^{2}+\frac{(Z \alpha)^{4}}{3\left[1+\sqrt{1-(Z \alpha)^{2}}\right]^{2}}\right. \\
& \left.-(Z \alpha)^{5} P(Z \alpha)\right\} \frac{m_{\mathrm{e}}}{m_{\mathrm{N}}}+\mathcal{O}\left(\frac{m_{\mathrm{e}}}{m_{\mathrm{N}}}\right)^{2} \\
= & -0.00000008771(1) \ldots \quad \text { for } Z=6 \\
= & -0.00000011711(1) \ldots \quad \text { for } Z=8 \tag{161}
\end{align*}
$$

where $m_{\mathrm{N}}$ is the mass of the nucleus. The mass ratios, obtained from the 2006 adjustment, are $m_{\mathrm{e}} / m\left({ }^{12} \mathrm{C}^{6+}\right)$ $=0.0000457275 \ldots$ and $m_{\mathrm{e}} / m\left({ }^{16} \mathrm{O}^{8+}\right)=0.0000343065 \ldots$. The recoil terms are the same as in CODATA-02, and references to the original calculations are given there. An additional term of the order of the mass ratio squared is included as

$$
\begin{equation*}
S_{Z}(Z \alpha)^{2}\left(\frac{m_{\mathrm{e}}}{m_{\mathrm{N}}}\right)^{2} \tag{162}
\end{equation*}
$$

where $S_{Z}$ is taken to be the average of the disagreeing values $1+Z$, obtained by Eides (2002) and Eides and Grotch (1997a), and $Z / 3$ obtained by Martynenko and Faustov $(2001,2002)$ for this term. The uncertainty in $S_{Z}$ is taken to be half the difference of the two values.

For $\Delta g_{\text {rec }}^{(2)}$, we have

$$
\begin{align*}
\Delta g_{\text {rec }}^{(2)} & =\frac{\alpha}{\pi} \frac{(Z \alpha)^{2}}{3} \frac{m_{\mathrm{e}}}{m_{\mathrm{N}}}+\cdots \\
& =0.00000000006 \ldots \quad \text { for } Z=6 \\
& =0.00000000009 \ldots \quad \text { for } Z=8 . \tag{163}
\end{align*}
$$

There is a small correction to the bound-state $g$-factor due to the finite size of the nucleus, of order

$$
\begin{equation*}
\Delta g_{\mathrm{ns}}=-\frac{8}{3}(Z \alpha)^{4}\left(\frac{R_{\mathrm{N}}}{\chi_{\mathrm{C}}}\right)^{2}+\cdots \tag{164}
\end{equation*}
$$

where $R_{\mathrm{N}}$ is the bound-state nuclear rms charge radius and $\lambda_{\mathrm{C}}$ is the Compton wavelength of the electron divided by $2 \pi$. This term is calculated as in CODATA-02 (Glazov and Shabaev, 2002) with updated values for the nuclear radii, $R_{\mathrm{N}}=2.4703(22) \mathrm{fm}$ and $R_{\mathrm{N}}=2.7013(55) \mathrm{fm}$, from the compilation of Angeli (2004) for ${ }^{12} \mathrm{C}$ and ${ }^{16} \mathrm{O}$, respectively. This yields the correction

$$
\begin{align*}
& \Delta g_{\mathrm{ns}}=-0.000000000408(1) \quad \text { for }{ }^{12} \mathrm{C} \\
& \Delta g_{\mathrm{ns}}=-0.00000000156(1) \quad \text { for }{ }^{16} \mathrm{O} \tag{165}
\end{align*}
$$

The theoretical value for the $g$-factor of the electron in hydrogenic carbon 12 or oxygen 16 is the sum of the individual contributions discussed above and summarized in Tables XV and XVI,

$$
\begin{align*}
& g_{\mathrm{e}^{-}}\left({ }^{12} \mathrm{C}^{5+}\right)=-2.001041590203(28), \\
& g_{\mathrm{e}^{-}}\left({ }^{16} \mathrm{O}^{7+}\right)=-2.00004702038(11) \tag{166}
\end{align*}
$$

For the purpose of the least-squares calculations carried out in Sec. XII.B, we define $g_{\mathrm{C}}($ th $)$ to be the sum of $g_{\mathrm{D}}$ as given in Eq. (138), the term $-2(\alpha / \pi) C_{\mathrm{e}}^{(2)}$, and the numerical values of the remaining terms in Eq. (137) as given in Table XV, where the standard uncertainty of these latter terms is

$$
\begin{equation*}
u\left[g_{\mathrm{C}}(\mathrm{th})\right]=0.3 \times 10^{-10}=1.4 \times 10^{-11}\left|g_{\mathrm{C}}(\mathrm{th})\right| \tag{167}
\end{equation*}
$$

The uncertainty in $g_{\mathrm{C}}($ th) due to the uncertainty in $\alpha$ enters the adjustment primarily through the functional dependence of $g_{\mathrm{D}}$ and the term $-2(\alpha / \pi) C_{\mathrm{e}}^{(2)}$ on $\alpha$. Therefore, this particular component of uncertainty is not explicitly included in $u\left[g_{C}(\right.$ th $\left.)\right]$. To take the uncertainty $u\left[g_{\mathrm{C}}(\right.$ th $\left.)\right]$ into account, we employ as the theoretical expression for the $g$-factor

$$
\begin{equation*}
g_{\mathrm{C}}\left(\alpha, \delta_{\mathrm{C}}\right)=g_{\mathrm{C}}(\text { th })+\delta_{\mathrm{C}}, \tag{168}
\end{equation*}
$$

where the input value of the additive correction $\delta_{\mathrm{C}}$ is taken to be zero and its standard uncertainty is $u\left[g_{\mathrm{C}}(\mathrm{th})\right]$,

$$
\begin{equation*}
\delta_{\mathrm{C}}=0.00(27) \times 10^{-10} \tag{169}
\end{equation*}
$$

Analogous considerations apply for the $g$-factor in oxygen,

$$
\begin{align*}
& u\left[g_{\mathrm{O}}(\mathrm{th})\right]=1.1 \times 10^{-10}=5.3 \times 10^{-11} \mid g_{\mathrm{O}}(\text { th }) \mid,  \tag{170}\\
& g_{\mathrm{O}}\left(\alpha, \delta_{\mathrm{O}}\right)=g_{\mathrm{O}}(\text { th })+\delta_{\mathrm{O}} \tag{171}
\end{align*}
$$

TABLE XVII. Summary of experimental data on the electron bound-state $g$-factor in hydrogenic carbon and oxygen and inferred values of the relative atomic mass of the electron.

| Input datum | Value | Relative standard <br> uncertainty $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :--- | :--- | :---: | :---: | :---: |
| $f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right) / f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)$ | $4376.2104989(23)$ | $5.2 \times 10^{-10}$ | GSI-02 | V.C.2.a (175) |
| $A_{\mathrm{r}}(\mathrm{e})$ | $0.00054857990932(29)$ | $5.2 \times 10^{-10}$ |  | V.C.2.a (177) |
| $f_{\mathrm{s}}\left({ }^{16} \mathrm{O}^{7+}\right) / f_{\mathrm{c}}\left({ }^{16} \mathrm{O}^{7+}\right)$ | $4164.3761837(32)$ | $7.6 \times 10^{-10}$ | GSI-02 | V.C.2.b (178) |
| $A_{\mathrm{r}}(\mathrm{e})$ | $0.00054857990958(42)$ | $7.6 \times 10^{-10}$ |  | V.C.2.b (181) |

$$
\begin{equation*}
\delta_{\mathrm{O}}=0.0(1.1) \times 10^{-10} \tag{172}
\end{equation*}
$$

Since the uncertainties of the theoretical values of the carbon and oxygen $g$-factors arise primarily from the same sources, the quantities $\delta_{\mathrm{C}}$ and $\delta_{\mathrm{O}}$ are highly correlated. Their covariance is

$$
\begin{equation*}
u\left(\delta_{\mathrm{C}}, \delta_{\mathrm{O}}\right)=27 \times 10^{-22} \tag{173}
\end{equation*}
$$

which corresponds to a correlation coefficient of $r\left(\delta_{\mathrm{C}}, \delta_{\mathrm{O}}\right)=0.92$.

The theoretical value of the ratio of the two $g$-factors, which is relevant to the comparison to experiment in Sec. V.C.2.c, is

$$
\begin{equation*}
\frac{g_{\mathrm{e}^{-}}\left({ }^{12} \mathrm{C}^{5+}\right)}{g_{\mathrm{e}^{-}\left({ }^{16} \mathrm{O}^{7+}\right)}=1.000497273218(41), ~, ~, ~} \tag{174}
\end{equation*}
$$

where the covariance, including the contribution from the uncertainty in $\alpha$ for this case, is taken into account.

## 2. Measurements of $g_{\mathrm{e}}\left({ }^{12} \mathrm{C}^{5+}\right)$ and $g_{\mathrm{e}}\left({ }^{16} \mathrm{O}^{7+}\right)$

The experimental data on the electron bound-state $g$-factor in hydrogenic carbon and oxygen and the inferred values of $A_{\mathrm{r}}(\mathrm{e})$ are summarized in Table XVII.

## a. Experiment on $g_{\mathrm{e}}\left({ }^{12} \mathrm{C}^{5+}\right)$ and inferred value of $A_{\mathrm{r}}(\mathrm{e})$

The accurate determination of the frequency ratio $f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right) / f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)$ at GSI based on the double Penningtrap technique was discussed in CODATA-02. [See also the recent concise review by Werth et al. (2006).] Since the result used as an input datum in the 2002 adjustment is unchanged, we take it as an input datum in the 2006 adjustment as well (Beier et al., 2002; Häffner et al., 2003; Werth, 2003),

$$
\begin{equation*}
\frac{f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right)}{f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)}=4376.2104989(23) \tag{175}
\end{equation*}
$$

From Eqs. (133) and (4), we have

$$
\begin{align*}
\frac{f_{\mathrm{s}}\left(12 \mathrm{C}^{5+}\right)}{f_{\mathrm{c}}\left(12 \mathrm{C}^{5+}\right)}= & -\frac{g_{\mathrm{e}}\left({ }^{12} \mathrm{C}^{5+}\right)}{10 A_{\mathrm{r}}(\mathrm{e})}\left[12-5 A_{\mathrm{r}}(\mathrm{e})\right. \\
& \left.+\frac{\left.E_{\mathrm{b}}\left({ }^{12} \mathrm{C}\right)-E_{\mathrm{b}}{ }^{12} \mathrm{C}^{5+}\right)}{m_{\mathrm{u}} c^{2}}\right] \tag{176}
\end{align*}
$$

which is the basis for the observational equation for the ${ }^{12} \mathrm{C}^{5+}$ frequency-ratio input datum.

Evaluation of this expression using the result for $f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right) / f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)$ in Eq. (175), the theoretical result for $g_{e^{-}}\left({ }^{12} \mathrm{C}^{5+}\right)$ in Table XV, and the relevant binding energies in Table IV of CODATA-02, yields

$$
\begin{equation*}
A_{\mathrm{r}}(\mathrm{e})=0.00054857990932(29) \quad\left[5.2 \times 10^{-10}\right] \tag{177}
\end{equation*}
$$

This value is consistent with that from antiprotonic helium given in Eq. (74) and that from the University of Washington given in Eq. (5), but has about a factor of 3 to 4 smaller uncertainty.

## b. Experiment on $g_{\mathrm{e}}\left({ }^{16} \mathrm{O}^{7+}\right)$ and inferred value of $A_{\mathrm{r}}(\mathrm{e})$

The double Penning-trap determination of the frequency ratio $f_{\mathrm{s}}\left({ }^{16} \mathrm{C}^{7+}\right) / f_{\mathrm{c}}\left({ }^{16} \mathrm{C}^{7+}\right)$ at GSI was also discussed in CODATA-02, but the value used as an input datum was not quite final (Verdú et al., 2002, 2003; Werth, 2003). A slightly different value for the ratio was given in the final report of the measurement (Verdú et al., 2004), which is the value we take as the input datum in the 2006 adjustment but modified slightly as follows based on information provided by Verdú (2006): (i) an unrounded instead of a rounded value for the correction due to extrapolating the axial temperature $T_{z}$ to 0 K was added to the uncorrected ratio $(-0.0000047$ in place of -0.000005 ); and (ii) a more detailed uncertainty budget was employed to evaluate the uncertainty of the ratio. The resulting value is

$$
\begin{equation*}
\frac{f_{\mathrm{s}}\left(1^{16} \mathrm{O}^{7+}\right)}{f_{\mathrm{c}}\left({ }^{16} \mathrm{O}^{7+}\right)}=4164.3761837(32) \tag{178}
\end{equation*}
$$

In analogy with what was done above with the ratio $f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right) / f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)$, from Eqs. (133) and (4) we have

$$
\begin{equation*}
\frac{f_{\mathrm{s}}\left({ }^{16} \mathrm{O}^{7+}\right)}{f_{\mathrm{c}}\left({ }^{16} \mathrm{O}^{7+}\right)}=-\frac{g_{\mathrm{e}^{-}}\left({ }^{16} \mathrm{O}^{7+}\right)}{14 A_{\mathrm{r}}(\mathrm{e})} A_{\mathrm{r}}\left({ }^{16} \mathrm{O}^{7+}\right) \tag{179}
\end{equation*}
$$

with

$$
\begin{align*}
A_{\mathrm{r}}\left({ }^{16} \mathrm{O}\right)= & A_{\mathrm{r}}\left({ }^{(16} \mathrm{O}^{7+}\right)+7 A_{\mathrm{r}}(\mathrm{e}) \\
& -\frac{E_{\mathrm{b}}\left({ }^{16} \mathrm{O}\right)-E_{\mathrm{b}}\left({ }^{16} \mathrm{O}^{7+}\right)}{m_{\mathrm{u}} c^{2}}, \tag{180}
\end{align*}
$$

which are the basis for the observational equations for the oxygen frequency ratio and $A_{\mathrm{r}}\left({ }^{16} \mathrm{O}\right)$, respectively. The first expression, evaluated using the result for $f_{\mathrm{s}}\left({ }^{16} \mathrm{C}^{7+}\right) / f_{\mathrm{c}}\left({ }^{16} \mathrm{C}^{7+}\right)$ in Eq. (178) and the theoretical result
for $g_{e^{-}}\left({ }^{16} \mathrm{C}^{7+}\right)$ in Table XVI, in combination with the second expression, evaluated using the value of $A_{\mathrm{r}}\left({ }^{16} \mathrm{O}\right)$ in Table IV and the relevant binding energies in Table IV of CODATA-02, yields

$$
\begin{equation*}
A_{\mathrm{r}}(\mathrm{e})=0.00054857990958(42) \quad\left[7.6 \times 10^{-10}\right] \tag{181}
\end{equation*}
$$

It is consistent with both the University of Washington value in Eq. (5) and the value in Eq. (177) obtained from $f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right) / f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)$.
c. Relations between $g_{\mathrm{e}}\left({ }^{12} \mathrm{C}^{5+}\right)$ and $g_{\mathrm{e}}\left({ }^{16} \mathrm{O}^{7+}\right)$

It should be noted that the GSI frequency ratios for ${ }^{12} \mathrm{C}^{5+}$ and ${ }^{16} \mathrm{O}^{7+}$ are correlated. Based on the detailed uncertainty budgets of the two results (Werth, 2003; Verdú, 2006), we find the correlation coefficient to be

$$
\begin{equation*}
r\left[\frac{\left.f_{\mathrm{s}}{ }^{12} \mathrm{C}^{5+}\right)}{f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)}, \frac{f_{\mathrm{s}}\left({ }^{16} \mathrm{O}^{7+}\right)}{f_{\mathrm{c}}\left({ }^{16} \mathrm{O}^{7+}\right)}\right]=0.082 \tag{182}
\end{equation*}
$$

Finally, as a consistency test, it is of interest to compare the experimental and theoretical values of the ratio of $g_{\mathrm{e}^{-}}\left({ }^{12} \mathrm{C}^{5+}\right)$ to $g_{\mathrm{e}^{-}}\left({ }^{16} \mathrm{C}^{7+}\right)$ (Karshenboim and Ivanov, 2002). The main reason is that the experimental value of the ratio is only weakly dependent on the value of $A_{\mathrm{r}}(\mathrm{e})$. The theoretical value of the ratio is given in Eq. (174) and takes into account the covariance of the two theoretical values. The experimental value of the ratio can be obtained by combining Eqs. (175), (176), (178)-(180), and (182), and using the 2006 recommended value for $A_{\mathrm{r}}(\mathrm{e})$. Because of the weak dependence of the experimental ratio on $A_{\mathrm{r}}(\mathrm{e})$, the value used is not at all critical. The result is

$$
\begin{equation*}
\frac{g_{\mathrm{e}}\left({ }^{12} \mathrm{C}^{5+}\right)}{g_{\mathrm{e}^{-}}\left({ }^{16} \mathrm{O}^{7+}\right)}=1.00049727368(89) \quad\left[8.9 \times 10^{-10}\right] \tag{183}
\end{equation*}
$$

in agreement with the theoretical value.

## VI. MAGNETIC MOMENT RATIOS AND THE MUONELECTRON MASS RATIO

Magnetic moment ratios and the muon-electron mass ratio are determined by experiments on bound states of relevant particles. The free electron and muon magnetic moments are discussed in Sec. V and the theory of the $g$-factor of an electron bound in an atom with no nuclear spin is considered in Sec. V.C.1.
For nucleons or nuclei with spin $\boldsymbol{I}$, the magnetic moment can be written as

$$
\begin{equation*}
\boldsymbol{\mu}=g \frac{e}{2 m_{\mathrm{p}}} \boldsymbol{I} \tag{184}
\end{equation*}
$$

or

$$
\begin{equation*}
\mu=g \mu_{\mathrm{N}} i \tag{185}
\end{equation*}
$$

In Eq. (185), $\mu_{\mathrm{N}}=e \hbar / 2 m_{\mathrm{p}}$ is the nuclear magneton, defined in analogy with the Bohr magneton, and $i$ is the
spin quantum number of the nucleus defined by $\boldsymbol{I}^{2}=i(i$ $+1) \hbar^{2}$ and $I_{z}=-i \hbar, \ldots,(i-1) \hbar$, $i \hbar$, where $I_{z}$ is the spin projection. However, in some publications, moments of nucleons are expressed in terms of the Bohr magneton with a corresponding change in the definition of the $g$-factor.

For atoms with a nonzero nuclear spin, bound-state $g$-factors are defined by considering the contribution to the Hamiltonian from the interaction of the atom with an applied magnetic flux density $B$. For example, for hydrogen, in the framework of the Pauli approximation, we have

$$
\begin{align*}
\mathcal{H} & =\beta(\mathrm{H}) \boldsymbol{\mu}_{\mathrm{e}^{-}} \cdot \boldsymbol{\mu}_{\mathrm{p}}-\boldsymbol{\mu}_{\mathrm{e}^{-}}(\mathrm{H}) \cdot \boldsymbol{B}-\boldsymbol{\mu}_{\mathrm{p}}(\mathrm{H}) \cdot \boldsymbol{B} \\
& =\frac{2 \pi}{\hbar} \Delta \nu_{\mathrm{H}^{\boldsymbol{s}}} \cdot \boldsymbol{I}-g_{\mathrm{e}^{-}}(\mathrm{H}) \frac{\mu_{\mathrm{B}}}{\hbar} \boldsymbol{s} \cdot \boldsymbol{B}-g_{\mathrm{p}}(\mathrm{H}) \frac{\mu_{\mathrm{N}}}{\hbar} \boldsymbol{I} \cdot \boldsymbol{B}, \tag{186}
\end{align*}
$$

where $\beta(\mathrm{H})$ characterizes the strength of the hyperfine interaction, $\Delta \nu_{\mathrm{H}}$ is the ground-state hyperfine frequency, $\boldsymbol{s}$ is the spin of the electron, and $\boldsymbol{I}$ is the spin of the nucleus, that is, the proton. Equation (186), or its analog for other combinations of particles, serves to define the corresponding bound-state $g$-factors, which are $g_{\mathrm{e}^{-}}(\mathrm{H})$ and $g_{\mathrm{p}}(\mathrm{H})$ in this case.

## A. Magnetic moment ratios

A number of magnetic moment ratios are of interest for the 2006 adjustment. The results of measurements and the inferred values of various quantities are summarized in Sec. VI.A.2, and the measurement results themselves are also summarized in Table XIX.

The inferred moment ratios depend on the relevant theoretical binding corrections that relate the $g$-factor measured in the bound state to the corresponding freeparticle $g$-factor. To use the results of these experiments in the 2006 adjustment, we employ theoretical expressions that give predictions for the moments and $g$-factors of the bound particles in terms of free-particle moments and $g$-factors as well as adjusted constants; this is discussed in the following section. However, in a number of cases, the differences between the bound-state and freestate values are sufficiently small that the adjusted constants can be taken as exactly known.

## 1. Theoretical ratios of atomic bound-particle to free-particle $g$-factors

Theoretical $g$-factor-related quantities used in the 2006 adjustment are the ratio of the $g$-factor of the electron in the ground state of hydrogen to that of the free electron $g_{\mathrm{e}^{-}}(\mathrm{H}) / g_{\mathrm{e}^{-}}$; the ratio of the $g$-factor of the proton in hydrogen to that of the free proton $g_{\mathrm{p}}(\mathrm{H}) / g_{\mathrm{p}}$; the analogous ratios for the electron and deuteron in deuterium, $g_{\mathrm{e}^{-}}(\mathrm{D}) / g_{\mathrm{e}^{-}}$and $g_{\mathrm{d}}(\mathrm{D}) / g_{\mathrm{d}}$, respectively; and the analogous ratios for the electron and positive muon in muonium, $g_{\mathrm{e}^{-}}(\mathrm{Mu}) / g_{\mathrm{e}^{-}}$and $g_{\mu^{+}}(\mathrm{Mu}) / g_{\mu^{+}}$, respectively.

These ratios and the references for the relevant calculations are discussed in CODATA-98 and CODATA-02; only a summary of the results is included here.

For the electron in hydrogen, we have

$$
\begin{align*}
\frac{g_{\mathrm{e}^{-}}(\mathrm{H})}{g_{\mathrm{e}^{-}}}= & 1-\frac{1}{3}(Z \alpha)^{2}-\frac{1}{12}(Z \alpha)^{4}+\frac{1}{4}(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right) \\
& +\frac{1}{2}(Z \alpha)^{2} \frac{m_{\mathrm{e}}}{m_{\mathrm{p}}}+\frac{1}{2}\left(A_{1}^{(4)}-\frac{1}{4}\right)(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right)^{2} \\
& -\frac{5}{12}(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right) \frac{m_{\mathrm{e}}}{m_{\mathrm{p}}}+\cdots \tag{187}
\end{align*}
$$

where $A_{1}^{(4)}$ is given in Eq. (84). For the proton in hydrogen, we have

$$
\begin{align*}
\frac{g_{\mathrm{p}}(\mathrm{H})}{g_{\mathrm{p}}}= & 1-\frac{1}{3} \alpha(Z \alpha)-\frac{97}{108} \alpha(Z \alpha)^{3} \\
& +\frac{1}{6} \alpha(Z \alpha) \frac{m_{\mathrm{e}}}{m_{\mathrm{p}}} \frac{3+4 a_{\mathrm{p}}}{1+a_{\mathrm{p}}}+\cdots \tag{188}
\end{align*}
$$

where the proton magnetic moment anomaly $a_{\mathrm{p}}$ is defined by

$$
\begin{equation*}
a_{\mathrm{p}}=\frac{\mu_{\mathrm{p}}}{\left(e \hbar / 2 m_{\mathrm{p}}\right)}-1 \approx 1.793 \tag{189}
\end{equation*}
$$

For deuterium, similar expressions apply for the electron,

$$
\begin{align*}
\frac{g_{\mathrm{e}^{-}}(\mathrm{D})}{g_{\mathrm{e}^{-}}}= & 1-\frac{1}{3}(Z \alpha)^{2}-\frac{1}{12}(Z \alpha)^{4}+\frac{1}{4}(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right) \\
& +\frac{1}{2}(Z \alpha)^{2} \frac{m_{\mathrm{e}}}{m_{\mathrm{d}}}+\frac{1}{2}\left(A_{1}^{(4)}-\frac{1}{4}\right)(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right)^{2} \\
& -\frac{5}{12}(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right) \frac{m_{\mathrm{e}}}{m_{\mathrm{d}}}+\cdots \tag{190}
\end{align*}
$$

and deuteron,

$$
\begin{align*}
\frac{g_{\mathrm{d}}(\mathrm{D})}{g_{\mathrm{d}}}= & 1-\frac{1}{3} \alpha(Z \alpha)-\frac{97}{108} \alpha(Z \alpha)^{3} \\
& +\frac{1}{6} \alpha(Z \alpha) \frac{m_{\mathrm{e}}}{m_{\mathrm{d}}} \frac{3+4 a_{\mathrm{d}}}{1+a_{\mathrm{d}}}+\cdots \tag{191}
\end{align*}
$$

where the deuteron magnetic moment anomaly $a_{\mathrm{d}}$ is defined by

$$
\begin{equation*}
a_{\mathrm{d}}=\frac{\mu_{\mathrm{d}}}{\left(e \hbar / m_{\mathrm{d}}\right)}-1 \approx-0.143 \tag{192}
\end{equation*}
$$

In the case of Mu , some additional higher-order terms are included because of the larger mass ratio. For the electron in muonium, we have

TABLE XVIII. Theoretical values for various bound-particle to free-particle $g$-factor ratios relevant to the 2006 adjustment based on the 2006 recommended values of the constants.

| Ratio | Value |
| :--- | :---: |
| $g_{\mathrm{e}^{-}}(\mathrm{H}) / g_{\mathrm{e}^{-}}$ | $1-17.7054 \times 10^{-6}$ |
| $g_{\mathrm{p}}(\mathrm{H}) / g_{\mathrm{p}}$ | $1-17.7354 \times 10^{-6}$ |
| $g_{\mathrm{e}^{-}}(\mathrm{D}) / g_{\mathrm{e}^{-}}$ | $1-17.7126 \times 10^{-6}$ |
| $g_{\mathrm{d}}(\mathrm{D}) / g_{\mathrm{d}}$ | $1-17.7461 \times 10^{-6}$ |
| $g_{\mathrm{e}^{-}}(\mathrm{Mu}) / g_{\mathrm{e}^{-}}$ | $1-17.5926 \times 10^{-6}$ |
| $g_{\mu^{+}}(\mathrm{Mu}) / g_{\mu^{+}}$ | $1-17.6254 \times 10^{-6}$ |

$$
\begin{align*}
\frac{g_{\mathrm{e}^{-}}(\mathrm{Mu})}{g_{\mathrm{e}^{-}}}= & 1-\frac{1}{3}(Z \alpha)^{2}-\frac{1}{12}(Z \alpha)^{4}+\frac{1}{4}(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right) \\
& +\frac{1}{2}(Z \alpha)^{2} \frac{m_{\mathrm{e}}}{m_{\mu}}+\frac{1}{2}\left(A_{1}^{(4)}-\frac{1}{4}\right)(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right)^{2} \\
& -\frac{5}{12}(Z \alpha)^{2}\left(\frac{\alpha}{\pi}\right) \frac{m_{\mathrm{e}}}{m_{\mu}}-\frac{1}{2}(1+Z) \\
& \times(Z \alpha)^{2}\left(\frac{m_{\mathrm{e}}}{m_{\mu}}\right)^{2}+\cdots \tag{193}
\end{align*}
$$

and for the muon in muonium, the ratio is

$$
\begin{align*}
\frac{g_{\mu^{+}}(\mathrm{Mu})}{g_{\mu^{+}}}= & 1-\frac{1}{3} \alpha(Z \alpha)-\frac{97}{108} \alpha(Z \alpha)^{3}+\frac{1}{2} \alpha(Z \alpha) \frac{m_{\mathrm{e}}}{m_{\mu}} \\
& +\frac{1}{12} \alpha(Z \alpha)\left(\frac{\alpha}{\pi}\right) \frac{m_{\mathrm{e}}}{m_{\mu}}-\frac{1}{2}(1+Z) \alpha(Z \alpha) \\
& \times\left(\frac{m_{\mathrm{e}}}{m_{\mu}}\right)^{2}+\cdots \tag{194}
\end{align*}
$$

The numerical values of the corrections in Eqs. (187)-(194), based on the 2006 adjusted values of the relevant constants, are listed in Table XVIII. Uncertainties are negligible at the level of uncertainty of the relevant experiments.

## 2. Ratio measurements

## a. Electron to proton magnetic moment ratio $\mu_{\mathrm{e}} / \mu_{\mathrm{p}}$

The ratio $\mu_{\mathrm{e}} / \mu_{\mathrm{p}}$ is obtained from measurements of the ratio of the magnetic moment of the electron to the magnetic moment of the proton in the 1 S state of hydrogen $\mu_{\mathrm{e}^{-}}(\mathrm{H}) / \mu_{\mathrm{p}}(\mathrm{H})$. We use the value obtained by Winkler et al. (1972) at MIT,

$$
\begin{equation*}
\frac{\mu_{\mathrm{e}^{-}}(\mathrm{H})}{\mu_{\mathrm{p}}(\mathrm{H})}=-658.2107058(66) \quad\left[1.0 \times 10^{-8}\right] \tag{195}
\end{equation*}
$$

where a minor typographical error in the original publication has been corrected (Kleppner, 1997). The freeparticle ratio $\mu_{\mathrm{e}} / \mu_{\mathrm{p}}$ follows from the bound-particle ratio and the relation

$$
\begin{align*}
\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}} & =\frac{g_{\mathrm{p}}(\mathrm{H})}{g_{\mathrm{p}}}\left(\frac{g_{\mathrm{e}^{-}}(\mathrm{H})}{g_{\mathrm{e}^{-}}}\right)^{-1} \frac{\mu_{\mathrm{e}^{-}}(\mathrm{H})}{\mu_{\mathrm{p}}(\mathrm{H})} \\
& =-658.2106860(66) \quad\left[1.0 \times 10^{-8}\right], \tag{196}
\end{align*}
$$

where the bound-state $g$-factor ratios are given in Table XVIII.

## b. Deuteron to electron magnetic moment ratio $\mu_{\mathrm{d}} / \mu_{\mathrm{e}}$

From measurements of the ratio $\mu_{\mathrm{d}}(\mathrm{D}) / \mu_{\mathrm{e}}-(\mathrm{D})$ in the 1S state of deuterium, Phillips et al. (1984) at MIT obtained

$$
\begin{equation*}
\frac{\mu_{\mathrm{d}}(\mathrm{D})}{\mu_{\mathrm{e}^{-}}(\mathrm{D})}=-4.664345392(50) \times 10^{-4} \quad\left[1.1 \times 10^{-8}\right] \tag{197}
\end{equation*}
$$

Although this result has not been published, as in the 1998 and 2002 adjustments, we include it as an input datum, because the method has been described in detail by Winkler et al. (1972) in connection with their measurement of $\mu_{e^{-}}(\mathrm{H}) / \mu_{\mathrm{p}}(\mathrm{H})$. The free-particle ratio is given by

$$
\begin{align*}
\frac{\mu_{\mathrm{d}}}{\mu_{\mathrm{e}^{-}}} & =\frac{g_{\mathrm{e}^{-}}(\mathrm{D})}{g_{\mathrm{e}^{-}}}\left(\frac{g_{\mathrm{d}}(\mathrm{D})}{g_{\mathrm{d}}}\right)^{-1} \frac{\mu_{\mathrm{d}}(\mathrm{D})}{\mu_{\mathrm{e}^{-}}(\mathrm{D})} \\
& =-4.664345548(50) \times 10^{-4} \quad\left[1.1 \times 10^{-8}\right] \tag{198}
\end{align*}
$$

with the bound-state $g$-factor ratios given in Table XVIII.
c. Proton to deuteron and triton to proton magnetic moment
ratios $\mu_{\mathrm{p}} / \mu_{\mathrm{d}}$ and $\mu_{\mathrm{t}} / \mu_{\mathrm{p}}$

The ratios $\mu_{\mathrm{p}} / \mu_{\mathrm{d}}$ and $\mu_{\mathrm{t}} / \mu_{\mathrm{p}}$ can be determined by NMR measurements on the HD molecule (bound state of hydrogen and deuterium) and the HT molecule (bound state of hydrogen and tritium, ${ }^{3} \mathrm{H}$ ), respectively. The relevant expressions are (see CODATA-98)

$$
\begin{align*}
& \frac{\mu_{\mathrm{p}}(\mathrm{HD})}{\mu_{\mathrm{d}}(\mathrm{HD})}=\left[1+\sigma_{\mathrm{d}}(\mathrm{HD})-\sigma_{\mathrm{p}}(\mathrm{HD})\right] \frac{\mu_{\mathrm{p}}}{\mu_{\mathrm{d}}},  \tag{199}\\
& \frac{\mu_{\mathrm{t}}(\mathrm{HT})}{\mu_{\mathrm{p}}(\mathrm{HT})}=\left[1-\sigma_{\mathrm{t}}(\mathrm{HT})+\sigma_{\mathrm{p}}(\mathrm{HT})\right] \frac{\mu_{\mathrm{t}}}{\mu_{\mathrm{p}}} \tag{200}
\end{align*}
$$

where $\mu_{\mathrm{p}}(\mathrm{HD})$ and $\mu_{\mathrm{d}}(\mathrm{HD})$ are the proton and deuteron magnetic moments in HD , respectively, and $\sigma_{\mathrm{p}}(\mathrm{HD})$ and $\sigma_{\mathrm{d}}(\mathrm{HD})$ are the corresponding nuclear magnetic shielding corrections. Similarly, $\mu_{\mathrm{t}}(\mathrm{HT})$ and $\mu_{\mathrm{p}}(\mathrm{HT})$ are the triton (nucleus of tritium) and proton magnetic moments in HT, respectively, and $\sigma_{\mathrm{t}}(\mathrm{HT})$ and $\sigma_{\mathrm{p}}(\mathrm{HT})$ are the corresponding nuclear magnetic shielding corrections. [Note that $\mu($ bound $)=(1-\sigma) \mu($ free $)$ and the nuclear magnetic shielding corrections are small.]

The determination of $\mu_{\mathrm{d}} / \mu_{\mathrm{p}}$ from NMR measurements on HD by Wimett (1953) and by a Russian group in St. Petersburg (Neronov et al., 1975; Gorshkov et al., 1989) was discussed in CODATA-98. However, for rea-
sons given there, mainly the lack of sufficient information to assign a reliable uncertainty to the reported values of $\mu_{\mathrm{d}}(\mathrm{HD}) / \mu_{\mathrm{p}}(\mathrm{HD})$ and also to the nuclear magnetic shielding correction difference $\sigma_{\mathrm{d}}(\mathrm{HD})-\sigma_{\mathrm{p}}(\mathrm{HD})$, the results were not used in the 1998 or 2002 adjustments. Further, since neither of these adjustments addressed quantities related to the triton, the determination of $\mu_{\mathrm{t}} / \mu_{\mathrm{p}}$ from measurements on HT by the Russian group (Neronov and Barzakh, 1977) was not considered in either of these adjustments. It may be recalled that a systematic error related to the use of separate inductance coils for the proton and deuteron NMR resonances in the measurements of Neronov et al. (1975) was eliminated in the HT measurements of Neronov and Barzakh (1977) as well as in the HD measurements of Gorshkov et al. (1989).

Recently, a member of the earlier St. Petersburg group together with one or more Russian colleagues published the following results based in part on new measurements and reexamination of relevant theory (Neronov and Karshenboim, 2003; Karshenboim et al., 2005),

$$
\begin{align*}
& \frac{\mu_{\mathrm{p}}(\mathrm{HD})}{\mu_{\mathrm{d}}(\mathrm{HD})}=3.257199531(29) \quad\left[8.9 \times 10^{-9}\right]  \tag{201}\\
& \frac{\mu_{\mathrm{t}}(\mathrm{HT})}{\mu_{\mathrm{p}}(\mathrm{HT})}=1.066639887(10) \quad\left[9.4 \times 10^{-9}\right]  \tag{202}\\
& \sigma_{\mathrm{dp}} \equiv \sigma_{\mathrm{d}}(\mathrm{HD})-\sigma_{\mathrm{p}}(\mathrm{HD})=15(2) \times 10^{-9}  \tag{203}\\
& \sigma_{\mathrm{tp}} \equiv \sigma_{\mathrm{t}}(\mathrm{HT})-\sigma_{\mathrm{p}}(\mathrm{HT})=20(3) \times 10^{-9} \tag{204}
\end{align*}
$$

which together with Eqs. (199) and (200) yield

$$
\begin{align*}
& \frac{\mu_{\mathrm{p}}}{\mu_{\mathrm{d}}}=3.257199482(30) \quad\left[9.1 \times 10^{-9}\right],  \tag{205}\\
& \frac{\mu_{\mathrm{t}}}{\mu_{\mathrm{p}}}=1.066639908(10) \quad\left[9.8 \times 10^{-9}\right] . \tag{206}
\end{align*}
$$

The purpose of the new work (Neronov and Karshenboim, 2003; Karshenboim et al., 2005) was (i) to check whether rotating the NMR sample and using a highpressure gas as the sample ( 60 to 130 atmospheres), which was the case in most of the older Russian experiments, influenced the results and to report a value of $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$ with a reliable uncertainty; and (ii) to reexamine the theoretical values of the nuclear magnetic shielding correction differences $\sigma_{\mathrm{dp}}$ and $\sigma_{\mathrm{tp}}$ and their uncertainties. It was also anticipated that based on this new work, a value of $\mu_{\mathrm{t}}(\mathrm{HT}) / \mu_{\mathrm{p}}(\mathrm{HT})$ with a reliable uncertainty could be obtained from the highly precise measurements of Neronov and Barzakh (1977). However, Gorshkov et al. (1989), as part of their experiment to determine $\mu_{\mathrm{d}} / \mu_{\mathrm{p}}$, compared the result from a 100 atmosphere HD rotating sample with a 100 atmosphere HD nonrotating sample and found no statistically significant difference.

TABLE XIX. Summary of data for magnetic moment ratios of various bound particles.

| Quantity | Value | Relative standard <br> uncertainty $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :--- | :--- | :---: | :--- | ---: |
| $\mu_{\mathrm{e}}-(\mathrm{H}) / \mu_{\mathrm{p}}(\mathrm{H})$ | $-658.2107058(66)$ | $1.0 \times 10^{-8}$ | MIT-72 | VI.A.2.a (195) |
| $\mu_{\mathrm{d}}(\mathrm{D}) / \mu_{\mathrm{e}}-(\mathrm{D})$ | $-4.664345392(50) \times 10^{-4}$ | $1.1 \times 10^{-8}$ | MIT-84 | VI.A.2.b (197) |
| $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$ | $3.257199531(29)$ | $8.9 \times 10^{-9}$ | StPtrsb-03 | VI.A.2.c (201) |
| $\sigma_{\mathrm{dp}}$ | $15(2) \times 10^{-9}$ |  | StPtrsb-03 | VI.A.2.c (203) |
| $\mu_{\mathrm{t}}(\mathrm{HT}) / \mu_{\mathrm{p}}(\mathrm{HT})$ | $1.066639887(10)$ | $9.4 \times 10^{-9}$ | StPtrsb-03 | VI.A.2.c (202) |
| $\sigma_{\mathrm{tp}}$ | $20(3) \times 10^{-9}$ |  | StPtrsb-03 | VI.A.2.c (204) |
| $\mu_{\mathrm{e}}-(\mathrm{H}) / \mu_{\mathrm{p}}^{\prime}$ | $-658.2159430(72)$ | $1.1 \times 10^{-8}$ | MIT-77 | VI.A.2.d (209) |
| $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime}$ | $-0.7617861313(33)$ | $4.3 \times 10^{-9}$ | NPL-93 | VI.A.2.e (211) |
| $\mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime}$ | $-0.68499694(16)$ | $2.4 \times 10^{-7}$ | ILL-79 | VI.A.2.f (212) |

To test the effect of sample rotation and sample pressure, Neronov and Karshenboim (2003) performed measurements using a commercial NMR spectrometer operating at a magnetic flux density of about 7 T and a nonrotating 10 atmosphere HD gas sample. Because of the relatively low pressure, the NMR signals were comparatively weak and a measurement time of 1 h was required. To simplify the measurements, the frequency of the proton NMR signal from HD was determined relative to the frequency of the more easily measured proton NMR signal from acetone, $\left(\mathrm{CH}_{3}\right)_{2} \mathrm{CO}$. Similarly, the frequency of the deuteron NMR signal from HD was determined relative to the frequency of the more easily measured deuteron NMR signal from deuterated acetone, $\left(\mathrm{CD}_{3}\right)_{2} \mathrm{CO}$. A number of tests involving the measurement of the hyperfine interaction constant in the case of the proton triplet NMR spectrum, and the isotopic shift in the case of the deuteron, where the deuteron HD doublet NMR spectrum was compared with the singlet spectrum of $D_{2}$, were carried out to investigate the reliability of the new data. The results of the tests were in good agreement with the older results obtained with sample rotation and high gas pressure.

The more recent result for $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$ reported by Karshenboim et al. (2005), which was obtained with the same NMR spectrometer employed by Neronov and Karshenboim (2003) but with a 20 atmosphere nonrotating gas sample, agrees with the 10 atmosphere nonrotating sample result of the latter researchers and is interpreted by Karshenboim et al. (2005) as confirming the 2003 result. Although the values of $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$ reported by the Russian researchers in 2005, 2003, and 1989 agree, the 2003 result as given in Eq. (201) and Table XIX, the uncertainty of which is dominated by the proton NMR line fitting procedure, is taken as the input datum in the 2006 adjustment because of the attention paid to possible systematic effects, both experimental and theoretical.
Based on their HD measurements and related analysis, especially the fact that sample pressure and rotation do not appear to be a problem at the current level of uncertainty, Neronov and Karshenboim (2003) concluded that the result for $\mu_{\mathrm{t}}(\mathrm{HT}) / \mu_{\mathrm{p}}(\mathrm{HT})$ reported by

Neronov and Barzakh (1977) is reliable but that it should be assigned about the same relative uncertainty as their result for $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$. We therefore include as an input datum in the 2006 adjustment the result for $\mu_{\mathrm{t}}(\mathrm{HT}) / \mu_{\mathrm{p}}(\mathrm{HT})$ given in Eq. (202) and Table XIX.

Without reliable theoretically calculated values for the shielding correction differences $\sigma_{\mathrm{dp}}$ and $\sigma_{\mathrm{tp}}$, reliable experimental values for the ratios $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$ and $\mu_{\mathrm{t}}(\mathrm{HT}) / \mu_{\mathrm{p}}(\mathrm{HT})$ are of little use. Although Neronov and Barzakh (1977) give theoretical estimates of these quantities based on their own calculations, they do not discuss the uncertainties of their estimates. To address this issue, Neronov and Karshenboim (2003) examined the calculations and concluded that a reasonable estimate of the relative uncertainty is $15 \%$. This leads to the values for $\sigma_{\mathrm{dp}}$ and $\sigma_{\mathrm{tp}}$ in Eqs. (203) and (204) and Table XIX, which we also take as input data for the 2006 adjustment. [For simplicity, we use StPtrsb-03 as the identifier in Table XIX for $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD}), \mu_{\mathrm{t}}(\mathrm{HT}) / \mu_{\mathrm{p}}(\mathrm{HT}), \sigma_{\mathrm{dp}}$, and $\sigma_{\mathrm{tp}}$, because they are directly or indirectly a consequence of the work of Neronov and Karshenboim (2003).]

The equations for the measured moment ratios $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$ and $\mu_{\mathrm{t}}(\mathrm{HT}) / \mu_{\mathrm{p}}(\mathrm{HT})$ in terms of the adjusted constants $\mu_{\mathrm{e}^{-}} / \mu_{\mathrm{p}}, \mu_{\mathrm{d}} / \mu_{\mathrm{e}^{-}}, \mu_{\mathrm{t}} / \mu_{\mathrm{p}}, \sigma_{\mathrm{dp}}$, and $\sigma_{\mathrm{tp}}$ are, from Eqs. (199) and (200),

$$
\begin{align*}
& \frac{\mu_{\mathrm{p}}(\mathrm{HD})}{\mu_{\mathrm{d}}(\mathrm{HD})}=\left[1+\sigma_{\mathrm{dp}}\right]\left(\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}}\right)^{-1}\left(\frac{\mu_{\mathrm{d}}}{\mu_{\mathrm{e}^{-}}}\right)^{-1},  \tag{207}\\
& \frac{\mu_{\mathrm{t}}(\mathrm{HT})}{\mu_{\mathrm{p}}(\mathrm{HT})}=\left[1-\sigma_{\mathrm{tp}}\right] \frac{\mu_{\mathrm{t}}}{\mu_{\mathrm{p}}} \tag{208}
\end{align*}
$$

## d. Electron to shielded proton magnetic moment ratio $\mu_{\mathrm{e}} / \mu_{\mathrm{p}}^{\prime}$

Based on the measurement of the ratio of the electron moment in the 1 S state of hydrogen to the shielded proton moment at $34.7^{\circ} \mathrm{C}$ by Phillips et al. (1977) at MIT, and temperature-dependence measurements of the shielded proton moment by Petley and Donaldson
(1984) at the National Physical Laboratory (NPL), Teddington, UK, we have

$$
\begin{equation*}
\frac{\mu_{\mathrm{e}}-(\mathrm{H})}{\mu_{\mathrm{p}}^{\prime}}=-658.2159430(72) \quad\left[1.1 \times 10^{-8}\right], \tag{209}
\end{equation*}
$$

where the prime indicates that the protons are in a spherical sample of pure $\mathrm{H}_{2} \mathrm{O}$ at $25^{\circ} \mathrm{C}$ surrounded by vacuum. Hence

$$
\begin{align*}
\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}^{\prime}} & =\left(\frac{g_{\mathrm{e}^{-}}(\mathrm{H})}{g_{\mathrm{e}^{-}}}\right)^{-1} \frac{\mu_{\mathrm{e}^{-}}(\mathrm{H})}{\mu_{p}^{\prime}} \\
& =-658.2275971(72) \quad\left[1.1 \times 10^{-8}\right] \tag{210}
\end{align*}
$$

where the bound-state $g$-factor ratio is given in Table XVIII. Support for the MIT result in Eq. (210) from measurements at NPL on the helion (see the following section) is discussed in CODATA-02.
e. Shielded helion to shielded proton magnetic moment ratio $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime}$

The ratio of the magnetic moment of the helion $h$, the nucleus of the ${ }^{3} \mathrm{He}$ atom, to the magnetic moment of the proton in $\mathrm{H}_{2} \mathrm{O}$ was determined in a high-accuracy experiment at NPL (Flowers et al., 1993) with the result

$$
\begin{equation*}
\frac{\mu_{\mathrm{h}}^{\prime}}{\mu_{\mathrm{p}}^{\prime}}=-0.7617861313(33) \quad\left[4.3 \times 10^{-9}\right] \tag{211}
\end{equation*}
$$

The prime on the symbol for the helion moment indicates that the helion is not free, but is bound in a helium atom. Although the exact shape and temperature of the gaseous ${ }^{3} \mathrm{He}$ sample are unimportant, we assume that it is spherical, at $25^{\circ} \mathrm{C}$, and surrounded by vacuum.

## f. Neutron to shielded proton magnetic moment ratio $\mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime}$

Based on a measurement carried out at the Institut Max von Laue-Paul Langevin (ILL) in Grenoble, France (Greene et al., 1977 1979), we have

$$
\begin{equation*}
\frac{\mu_{\mathrm{n}}}{\mu_{\mathrm{p}}^{\prime}}=-0.68499694(16) \quad\left[2.4 \times 10^{-7}\right] \tag{212}
\end{equation*}
$$

The observational equations for the measured values of $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime}$ and $\mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime}$ are simply

$$
\begin{align*}
& \mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime}=\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime}  \tag{213}\\
& \mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime}=\mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime}, \tag{214}
\end{align*}
$$

while the observational equations for the measured values of $\mu_{\mathrm{e}^{-}}(\mathrm{H}) / \mu_{\mathrm{p}}(\mathrm{H}), \mu_{\mathrm{d}}(\mathrm{D}) / \mu_{e^{-}}(\mathrm{D})$, and $\mu_{\mathrm{e}^{-}}(\mathrm{H}) / \mu_{\mathrm{p}}^{\prime}$ follow directly from Eqs. (196), (198), and (210), respectively.

## B. Muonium transition frequencies, the muon-proton magnetic moment ratio $\mu_{\mu} / \mu_{\mathrm{p}}$, and muon-electron mass ratio $m_{\mu} / m_{\mathrm{e}}$

Measurements of transition frequencies between Zeeman energy levels in muonium (the $\mu^{+} \mathrm{e}^{-}$atom) yield values of $\mu_{\mu} / \mu_{\mathrm{p}}$ and the muonium ground-state hyper-
fine splitting $\Delta \nu_{\mathrm{Mu}}$ that depend only weakly on theory. The relevant expression for the magnetic moment ratio is

$$
\begin{equation*}
\frac{\mu_{\mu^{+}}}{\mu_{\mathrm{p}}}=\frac{\Delta \nu_{\mathrm{Mu}}^{2}-\nu^{2}\left(f_{\mathrm{p}}\right)+2 s_{\mathrm{e}} f_{\mathrm{p}} \nu\left(f_{\mathrm{p}}\right)}{4 s_{\mathrm{e}} f_{\mathrm{p}}^{2}-2 f_{\mathrm{p}} \nu\left(f_{\mathrm{p}}\right)}\left(\frac{g_{\mu^{+}}(\mathrm{Mu})}{g_{\mu^{+}}}\right)^{-1} \tag{215}
\end{equation*}
$$

where $\Delta \nu_{\mathrm{Mu}}$ and $\nu\left(f_{\mathrm{p}}\right)$ are the sum and difference of two measured transition frequencies, $f_{\mathrm{p}}$ is the free proton NMR reference frequency corresponding to the magnetic flux density used in the experiment, $g_{\mu^{+}}(\mathrm{Mu}) / g_{\mu^{+}}$is the bound-state correction for the muon in muonium given in Table XVIII, and

$$
\begin{equation*}
s_{\mathrm{e}}=\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}} \frac{g_{\mathrm{e}^{-}}(\mathrm{Mu})}{g_{\mathrm{e}^{-}}} \tag{216}
\end{equation*}
$$

where $g_{\mathrm{e}^{-}}(\mathrm{Mu}) / g_{\mathrm{e}^{-}}$is the bound-state correction for the electron in muonium given in the same table.

The muon to electron mass ratio $m_{\mu} / m_{\mathrm{e}}$ and the muon to proton magnetic moment ratio $\mu_{\mu} / \mu_{\mathrm{p}}$ are related by

$$
\begin{equation*}
\frac{m_{\mu}}{m_{\mathrm{e}}}=\left(\frac{\mu_{\mathrm{e}}}{\mu_{\mathrm{p}}}\right)\left(\frac{\mu_{\mu}}{\mu_{\mathrm{p}}}\right)^{-1}\left(\frac{g_{\mu}}{g_{\mathrm{e}}}\right) . \tag{217}
\end{equation*}
$$

The theoretical expression for the hyperfine splitting $\Delta \nu_{\mathrm{Mu}}(\mathrm{th})$ is discussed in the following section and may be written as

$$
\begin{align*}
\Delta \nu_{\mathrm{Mu}}(\mathrm{th}) & =\frac{16}{3} c R_{\infty} \alpha^{2} \frac{m_{\mathrm{e}}}{m_{\mu}}\left(1+\frac{m_{\mathrm{e}}}{m_{\mu}}\right)^{-3} \mathcal{F}\left(\alpha, m_{\mathrm{e}} / m_{\mu}\right) \\
& =\Delta \nu_{\mathrm{F}} \mathcal{F}\left(\alpha, m_{\mathrm{e}} / m_{\mu}\right) \tag{218}
\end{align*}
$$

where the function $\mathcal{F}$ depends weakly on $\alpha$ and $m_{\mathrm{e}} / m_{\mu}$. By equating this expression to an experimental value of $\Delta \nu_{\mathrm{Mu}}$, one can calculate a value of $\alpha$ from a given value of $m_{\mu} / m_{\mathrm{e}}$ or one can calculate a value of $m_{\mu} / m_{\mathrm{e}}$ from a given value of $\alpha$.

## 1. Theory of the muonium ground-state hyperfine splitting

This section gives a brief summary of the present theory of $\Delta \nu_{\mathrm{Mu}}$, the ground-state hyperfine splitting of muonium ( $\mu^{+} \mathrm{e}^{-}$atom). There has been essentially no change in the theory since the 2002 adjustment. Although complete results of the relevant calculations are given here, references to the original literature included in CODATA-98 or CODATA-02 are generally not repeated.

The hyperfine splitting is given by the Fermi formula,

$$
\begin{equation*}
\Delta \nu_{\mathrm{F}}=\frac{16}{3} c R_{\infty} Z^{3} \alpha^{2} \frac{m_{\mathrm{e}}}{m_{\mu}}\left[1+\frac{m_{\mathrm{e}}}{m_{\mu}}\right]^{-3} . \tag{219}
\end{equation*}
$$

Some of the following theoretical expressions correspond to a muon with charge $Z e$ rather than $e$ in order to identify the source of the terms. The theoretical value of the hyperfine splitting is given by

$$
\begin{align*}
\Delta \nu_{\mathrm{Mu}}(\text { th })= & \Delta \nu_{\mathrm{D}}+\Delta \nu_{\mathrm{rad}}+\Delta \nu_{\mathrm{rec}}+\Delta \nu_{\mathrm{r}-\mathrm{r}}+\Delta \nu_{\text {weak }} \\
& +\Delta \nu_{\mathrm{had}} \tag{220}
\end{align*}
$$

where the terms labeled D, rad, rec, r-r, weak, and had
account for the Dirac (relativistic), radiative, recoil, radiative-recoil, electroweak, and hadronic (strong interaction) contributions to the hyperfine splitting, respectively.

The contribution $\Delta \nu_{\mathrm{D}}$, given by the Dirac equation, is

$$
\begin{equation*}
\Delta \nu_{\mathrm{D}}=\Delta \nu_{\mathrm{F}}\left(1+a_{\mu}\right)\left[1+\frac{3}{2}(Z \alpha)^{2}+\frac{17}{8}(Z \alpha)^{4}+\cdots\right] \tag{221}
\end{equation*}
$$

where $a_{\mu}$ is the muon magnetic moment anomaly.
The radiative corrections are written as

$$
\begin{align*}
\Delta \nu_{\mathrm{rad}}= & \Delta \nu_{\mathrm{F}}\left(1+a_{\mu}\right)\left[D^{(2)}(Z \alpha)\left(\frac{\alpha}{\pi}\right)+D^{(4)}(Z \alpha)\left(\frac{\alpha}{\pi}\right)^{2}\right. \\
& \left.+D^{(6)}(Z \alpha)\left(\frac{\alpha}{\pi}\right)^{3}+\cdots\right] \tag{222}
\end{align*}
$$

where the functions $D^{(2 n)}(Z \alpha)$ are contributions associated with $n$ virtual photons. The leading term is

$$
\begin{align*}
D^{(2)}(Z \alpha)= & A_{1}^{(2)}+\left(\ln 2-\frac{5}{2}\right) \pi Z \alpha+\left[-\frac{2}{3} \ln ^{2}(Z \alpha)^{-2}\right. \\
& \left.+\left(\frac{281}{360}-\frac{8}{3} \ln 2\right) \ln (Z \alpha)^{-2}+16.9037 \ldots\right] \\
& \times(Z \alpha)^{2}+\left[\left(\frac{5}{2} \ln 2-\frac{547}{96}\right) \ln (Z \alpha)^{-2}\right] \\
& \times \pi(Z \alpha)^{3}+G(Z \alpha)(Z \alpha)^{3} \tag{223}
\end{align*}
$$

where $A_{1}^{(2)}=\frac{1}{2}$, as in Eq. (83). The function $G(Z \alpha)$ accounts for all higher-order contributions in powers of $Z \alpha$, and can be divided into parts that correspond to the self energy or vacuum polarization, $G(Z \alpha)=G_{\text {SE }}(Z \alpha)$ $+G_{\mathrm{VP}}(Z \alpha)$. We adopt the value

$$
\begin{equation*}
G_{\mathrm{SE}}(\alpha)=-14(2), \tag{224}
\end{equation*}
$$

which is the simple mean and standard deviation of the following three values: $G_{\mathrm{SE}}(\alpha)=-12.0(2.0)$ from Blundell et al. (1997), $G_{\mathrm{SE}}(0)=-15.9(1.6)$ from Nio (2001, 2002), and $G_{\mathrm{SE}}(\alpha)=-14.3(1.1)$ from Yerokhin and Shabaev (2001). The vacuum polarization part $G_{\mathrm{VP}}(Z \alpha)$ has been calculated to several orders of $Z \alpha$ by Karshenboim et al. (1999, 2000). Their expression yields

$$
\begin{equation*}
G_{\mathrm{VP}}(\alpha)=7.227(9) \tag{225}
\end{equation*}
$$

For $D^{(4)}(Z \alpha)$, as in CODATA-02, we have

$$
\begin{align*}
D^{(4)}(Z \alpha)= & A_{1}^{(4)}+0.7717(4) \pi Z \alpha+\left[-\frac{1}{3} \ln ^{2}(Z \alpha)^{-2}\right. \\
& \left.-0.6390 \ldots \times \ln (Z \alpha)^{-2}+10(2.5)\right](Z \alpha)^{2} \\
& +\cdots, \tag{226}
\end{align*}
$$

where $A_{1}^{(4)}$ is given in Eq. (84).
Finally,

$$
\begin{equation*}
D^{(6)}(Z \alpha)=A_{1}^{(6)}+\cdots, \tag{227}
\end{equation*}
$$

where only the leading contribution $A_{1}^{(6)}$ as given in Eq. (85) is known. Higher-order functions $D^{(2 n)}(Z \alpha)$ with $n$ $>3$ are expected to be negligible.
The recoil contribution is given by

$$
\begin{align*}
\Delta \nu_{\mathrm{rec}}= & \Delta \nu_{\mathrm{F}} \frac{m_{\mathrm{e}}}{m_{\mu}}\left(-\frac{3}{1-\left(m_{\mathrm{e}} / m_{\mu}\right)^{2}} \ln \left(\frac{m_{\mu}}{m_{\mathrm{e}}}\right) \frac{Z \alpha}{\pi}\right. \\
& +\frac{1}{\left(1+m_{\mathrm{e}} / m_{\mu}\right)^{2}}\left\{\ln (Z \alpha)^{-2}-8 \ln 2+\frac{65}{18}\right. \\
& +\left[\frac{9}{2 \pi^{2}} \ln ^{2}\left(\frac{m_{\mu}}{m_{\mathrm{e}}}\right)+\left(\frac{27}{2 \pi^{2}}-1\right) \ln \left(\frac{m_{\mu}}{m_{\mathrm{e}}}\right)+\frac{93}{4 \pi^{2}}\right. \\
& \left.\left.+\frac{33 \zeta(3)}{\pi^{2}}-\frac{13}{12}-12 \ln 2\right] \frac{m_{\mathrm{e}}}{m_{\mu}}\right\}(Z \alpha)^{2} \\
& +\left\{-\frac{3}{2} \ln \left(\frac{m_{\mu}}{m_{\mathrm{e}}}\right) \ln (Z \alpha)^{-2}-\frac{1}{6} \ln ^{2}(Z \alpha)^{-2}\right. \\
& \left.\left.+\left(\frac{101}{18}-10 \ln 2\right) \ln (Z \alpha)^{-2}+40(10)\right\} \frac{(Z \alpha)^{3}}{\pi}\right) \\
& +\cdots, \tag{228}
\end{align*}
$$

as discussed in CODATA-02
The radiative-recoil contribution is

$$
\begin{aligned}
\Delta \nu_{\mathrm{r}-\mathrm{r}}= & \nu_{\mathrm{F}}\left(\frac{\alpha}{\pi}\right)^{2} \frac{m_{\mathrm{e}}}{m_{\mu}}\left\{\left[-2 \ln ^{2}\left(\frac{m_{\mu}}{m_{\mathrm{e}}}\right)+\frac{13}{12} \ln \left(\frac{m_{\mu}}{m_{\mathrm{e}}}\right)\right.\right. \\
& \left.+\frac{21}{2} \zeta(3)+\frac{\pi^{2}}{6}+\frac{35}{9}\right]+\left[\frac{4}{3} \ln ^{2} \alpha^{-2}+\left(\frac{16}{3} \ln 2\right.\right. \\
& \left.\left.-\frac{341}{180}\right) \ln \alpha^{-2}-40(10)\right] \pi \alpha+\left[-\frac{4}{3} \ln ^{3}\left(\frac{m_{\mu}}{m_{\mathrm{e}}}\right)\right. \\
& \left.\left.+\frac{4}{3} \ln ^{2}\left(\frac{m_{\mu}}{m_{\mathrm{e}}}\right)\right] \frac{\alpha}{\pi}\right\}-\nu_{\mathrm{F}} \alpha^{2}\left(\frac{m_{\mathrm{e}}}{m_{\mu}}\right)^{2}\left(6 \ln 2+\frac{13}{6}\right)
\end{aligned}
$$

$$
\begin{equation*}
+\cdots \tag{229}
\end{equation*}
$$

where, for simplicity, the explicit dependence on $Z$ is not shown.

The electroweak contribution due to the exchange of a $Z^{0}$ boson is (Eides, 1996)

$$
\begin{equation*}
\Delta \nu_{\text {weak }}=-65 \mathrm{~Hz} \tag{230}
\end{equation*}
$$

For the hadronic vacuum polarization contribution, we use the result of Eidelman et al. (2002),

$$
\begin{equation*}
\Delta \nu_{\mathrm{had}}=236(4) \mathrm{Hz} \tag{231}
\end{equation*}
$$

which takes into account experimental data on the cross section for $\mathrm{e}^{-} \mathrm{e}^{+} \rightarrow \pi^{+} \pi^{-}$and on the $\phi$ meson leptonic width. The leading hadronic contribution is 231.2(2.9) Hz and the next-order term is $5(2) \mathrm{Hz}$, giving a total of $236(4) \mathrm{Hz}$. The pion and kaon contributions to the hadronic correction have been considered within a chiral unitary approach and found to be in general agreement with (but have a three times larger uncertainty than) the corresponding contributions given in earlier studies using data from $\mathrm{e}^{+}-\mathrm{e}^{-}$scattering (Palomar, 2003).

TABLE XX. Summary of data related to the hyperfine splitting in muonium and inferred values of $\mu_{\mu} / \mu_{\mathrm{p}}, m_{\mu} / m_{\mathrm{e}}$, and $\alpha$ from the combined 1982 and 1999 LAMPF data.

| Quantity | Value | Relative standard <br> uncertainty $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :--- | :--- | :---: | :--- | :--- |
| $\Delta \nu_{\mathrm{Mu}}$ | $4463302.88(16) \mathrm{kHz}$ | $3.6 \times 10^{-8}$ | LAMPF-82 | VI.B.2.a (236) |
| $\nu\left(f_{\mathrm{p}}\right)$ | $627994.77(14) \mathrm{kHz}$ | $2.2 \times 10^{-7}$ | LAMPF-82 | VI.B.2.a (237) |
|  |  |  |  |  |
| $\Delta \nu_{\mathrm{Mu}}$ | $4463302765(53) \mathrm{Hz}$ | $1.2 \times 10^{-8}$ | LAMPF-99 | VI.B.2.b (239) |
| $\nu\left(f_{\mathrm{p}}\right)$ | $668223166(57) \mathrm{Hz}$ | $8.6 \times 10^{-8}$ | LAMPF-99 | VI.B.2.b (240) |
|  |  |  |  |  |
| $\mu_{\mu} / \mu_{\mathrm{p}}$ | $3.18334524(37)$ | $1.2 \times 10^{-7}$ | LAMPF | VI.B.2.c (242) |
| $m_{\mu} / m_{\mathrm{e}}$ | $206.768276(24)$ | $1.2 \times 10^{-7}$ | LAMPF | VI.B.2.c (243) |
| $\alpha^{-1}$ | $137.0360017(80)$ | $5.8 \times 10^{-8}$ | LAMPF | VI.B.2.c (244) |

The standard uncertainty of $\Delta \nu_{\mathrm{Mu}}($ th) was discussed in Appendix E of CODATA-02. The four principal sources of uncertainty are the terms $\Delta \nu_{\mathrm{rad}}, \Delta \nu_{\mathrm{rec}}, \Delta \nu_{\mathrm{r}-\mathrm{r}}$, and $\Delta \nu_{\mathrm{had}}$ in Eq. (220). Included in the 67 Hz uncertainty of $\Delta \nu_{\mathrm{r}-\mathrm{r}}$ is a 41 Hz component, based on the partial calculations of Eides et al. $(2002,2003)$ and Li et al. (1993), to account for a possible uncalculated radiative-recoil contribution of order $\Delta \nu_{\mathrm{F}}\left(m_{\mathrm{e}} / m_{\mu}\right)(\alpha / \pi)^{3} \ln \left(m_{\mu} / m_{\mathrm{e}}\right)$ and nonlogarithmic terms. Since the completion of the 2002 adjustment, results of additional partial calculations have been published that would lead to a small reduction in the 41 Hz estimate (Eides and Shelyuto, 2003, 2004, 2007; Eides et al., 2004, 2005). However, because the calculations are not yet complete and the decrease of the 101 Hz total uncertainty assigned to $\Delta \nu_{\mathrm{Mu}}($ th) for the 2002 adjustment would only be a few percent, the Task Group decided to retain the 101 Hz uncertainty for the 2006 adjustment.

We thus have for the standard uncertainty of the theoretical expression for the muonium hyperfine splitting $\Delta \nu_{\mathrm{Mu}}($ th $)$

$$
\begin{equation*}
u\left[\Delta \nu_{\mathrm{Mu}}(\mathrm{th})\right]=101 \mathrm{~Hz} \quad\left[2.3 \times 10^{-8}\right] \tag{232}
\end{equation*}
$$

For the least-squares calculations, we use as the theoretical expression for the hyperfine splitting,

$$
\begin{equation*}
\Delta \nu_{\mathrm{Mu}}\left(R_{\infty}, \alpha, \frac{m_{\mathrm{e}}}{m_{\mu}}, \delta_{\mu}, \delta_{\mathrm{Mu}}\right)=\Delta \nu_{\mathrm{Mu}}(\mathrm{th})+\delta_{\mathrm{Mu}}, \tag{233}
\end{equation*}
$$

where $\delta_{\mathrm{Mu}}$ is assigned a priori the value

$$
\begin{equation*}
\delta_{\mathrm{Mu}}=0(101) \mathrm{Hz} \tag{234}
\end{equation*}
$$

in order to account for the uncertainty of the theoretical expression.

The theory summarized above predicts

$$
\begin{equation*}
\Delta \nu_{\mathrm{Mu}}=4463302881(272) \mathrm{Hz} \quad\left[6.1 \times 10^{-8}\right] \tag{235}
\end{equation*}
$$

based on values of the constants obtained from a variation of the 2006 least-squares adjustment that omits as input data the two LAMPF measured values of $\Delta \nu_{\mathrm{Mu}}$ discussed in the following section.

The main source of uncertainty in this value is the mass ratio $m_{\mathrm{e}} / m_{\mu}$ that appears in the theoretical expression as an overall factor. See the text following Eq.
(D14) of Appendix D of CODATA-98 for an explanation of why the relative uncertainty of the predicted value of $\Delta \nu_{\mathrm{Mu}}$ in Eq. (235) is smaller than the relative uncertainty of the electron-muon mass ratio as given in Eq. (243) of Sec. VI.B.2.c.

## 2. Measurements of muonium transition frequencies and values of $\mu_{\mu} / \mu_{\mathrm{p}}$ and $\boldsymbol{m}_{\mu} / \boldsymbol{m}_{\mathrm{e}}$

The two most precise determinations of muonium Zeeman transition frequencies were carried out at the Clinton P. Anderson Meson Physics Facility at Los Alamos (LAMPF), USA, and reviewed in CODATA-98. The following three sections and Table XX give the key results.

## a. LAMPF 1982

The results obtained by Mariam (1981) and Mariam et al. (1982), which we take as input data in the current adjustment as in the two previous adjustments, may be summarized as follows:

$$
\begin{align*}
& \Delta \nu_{\mathrm{Mu}}=4463302.88(16) \mathrm{kHz} \quad\left[3.6 \times 10^{-8}\right],  \tag{236}\\
& \nu\left(f_{\mathrm{p}}\right)=627994.77(14) \mathrm{kHz} \quad\left[2.2 \times 10^{-7}\right],  \tag{237}\\
& r\left[\Delta \nu_{\mathrm{Mu}}, \nu\left(f_{\mathrm{p}}\right)\right]=0.23, \tag{238}
\end{align*}
$$

where $f_{\mathrm{p}}$ is very nearly 57.972993 MHz , corresponding to the flux density of about 1.3616 T used in the experiment, and $r\left[\Delta \nu_{\mathrm{Mu}}, \nu\left(f_{\mathrm{p}}\right)\right]$ is the correlation coefficient of $\Delta \nu_{\mathrm{Mu}}$ and $\nu\left(f_{\mathrm{p}}\right)$.

## b. LAMPF 1999

The results obtained by Liu et al. (1999), which we also take as input data in the current adjustment as in the 1998 and 2002 adjustments, may be summarized as follows:

$$
\begin{gather*}
\Delta \nu_{\mathrm{Mu}}=4463302765(53) \mathrm{Hz} \quad\left[1.2 \times 10^{-8}\right],  \tag{239}\\
\nu\left(f_{\mathrm{p}}\right)=668223166(57) \mathrm{Hz} \quad\left[8.6 \times 10^{-8}\right], \tag{240}
\end{gather*}
$$

$$
\begin{equation*}
r\left[\Delta \nu_{\mathrm{Mu}}, \nu\left(f_{\mathrm{p}}\right)\right]=0.19 \tag{241}
\end{equation*}
$$

where $f_{\mathrm{p}}$ is exactly 72.320000 MHz , corresponding to the flux density of approximately 1.7 T used in the experiment, and $r\left[\Delta \nu_{\mathrm{Mu}}, \nu\left(f_{\mathrm{p}}\right)\right]$ is the correlation coefficient of $\Delta \nu_{\mathrm{Mu}}$ and $\nu\left(f_{\mathrm{p}}\right)$.

## c. Combined LAMPF results

By carrying out a least-squares adjustment using only the LAMPF 1982 and LAMPF 1999 data, the 2006 recommended values of the quantities $R_{\infty}, \mu_{\mathrm{e}} / \mu_{\mathrm{p}}, g_{\mathrm{e}}$, and $g_{\mu}$, together with Eqs. (215)-(218), we obtain

$$
\begin{align*}
& \frac{\mu_{\mu^{+}}}{\mu_{\mathrm{p}}}=3.18334524(37) \quad\left[1.2 \times 10^{-7}\right]  \tag{242}\\
& \frac{m_{\mu}}{m_{\mathrm{e}}}=206.768276(24) \quad\left[1.2 \times 10^{-7}\right]  \tag{243}\\
& \alpha^{-1}=137.0360017(80) \quad\left[5.8 \times 10^{-8}\right] \tag{244}
\end{align*}
$$

where this value of $\alpha$ may be called the muonium value of the fine-structure constant and denoted as $\alpha^{-1}\left(\Delta \nu_{\mathrm{Mu}}\right)$.
It is noteworthy that the uncertainty of the value of the mass ratio $m_{\mu} / m_{\mathrm{e}}$ given in Eq. (243) is about four times the uncertainty of the 2006 recommended value. The reason is that taken together, the experimental value of and theoretical expression for the hyperfine splitting essentially determine only the value of the product $\alpha^{2} m_{\mu} / m_{\mathrm{e}}$, as is evident from Eq. (218). In the full adjustment, the value of $\alpha$ is determined by other data with an uncertainty significantly smaller than that of the value in Eq. (244), which in turn determines the value of $m_{\mu} / m_{\mathrm{e}}$ with a smaller uncertainty than that of Eq. (243).

## VII. ELECTRICAL MEASUREMENTS

This section is devoted to a discussion of quantities that require electrical measurements of the most basic kind for their determination: the gyromagnetic ratios of the shielded proton and helion, the von Klitzing constant $R_{\mathrm{K}}$, the Josephson constant $K_{\mathrm{J}}$, the product $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, and the Faraday constant. However, some results we discuss were taken as input data for the 2002 adjustment but were not included in the final least-squares adjustment from which the 2002 recommended values were obtained, mainly because of their comparatively large uncertainties and hence low weight. Nevertheless, we take them as potential input data in the 2006 adjustment because they provide information on the overall consistency of the available data and tests of the exactness of the relations $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$. The lone exception is the low-field measurement of the gyromagnetic ratio of the helion reported by Tarbeev et al. (1989). Because of its large uncertainty and strong disagreement with many other data, we no longer consider it-see CODATA-02.

## A. Shielded gyromagnetic ratios $\boldsymbol{\gamma}^{\prime}$, the fine-structure constant $\alpha$, and the Planck constant $h$

The gyromagnetic ratio $\gamma$ of a bound particle of spin quantum number $i$ and magnetic moment $\mu$ is given by

$$
\begin{equation*}
\gamma=\frac{2 \pi f}{B}=\frac{\omega}{B}=\frac{|\mu|}{i \hbar}, \tag{245}
\end{equation*}
$$

where $f$ is the precession (that is, spin-flip) frequency and $\omega$ is the angular precession frequency of the particle in the magnetic flux density $B$. The SI unit of $\gamma$ is $\mathrm{s}^{-1}$ $\mathrm{T}^{-1}=\mathrm{C} \mathrm{kg}^{-1}=\mathrm{Askg}^{-1}$. In this section, we summarize measurements of the gyromagnetic ratio of the shielded proton

$$
\begin{equation*}
\gamma_{\mathrm{p}}^{\prime}=\frac{2 \mu_{\mathrm{p}}^{\prime}}{\hbar} \tag{246}
\end{equation*}
$$

and of the shielded helion

$$
\begin{equation*}
\gamma_{\mathrm{h}}^{\prime}=\frac{2\left|\mu_{\mathrm{h}}^{\prime}\right|}{\hbar} \tag{247}
\end{equation*}
$$

where, as in previous sections that dealt with magneticmoment ratios involving these particles, protons are those in a spherical sample of pure $\mathrm{H}_{2} \mathrm{O}$ at $25^{\circ} \mathrm{C}$ surrounded by vacuum; and helions are those in a spherical sample of low-pressure, pure ${ }^{3} \mathrm{He}$ gas at $25^{\circ} \mathrm{C}$ surrounded by vacuum.

As discussed in CODATA-98, two methods are used to determine the shielded gyromagnetic ratio $\gamma^{\prime}$ of a particle: the low- and high-field methods. In either case, the measured current $I$ in the experiment can be expressed in terms of the product $K_{\mathrm{J}} R_{\mathrm{K}}$, but $B$ depends on $I$ differently in the two cases. In essence, the low-field experiments determine $\gamma^{\prime} / K_{\mathrm{J}} R_{\mathrm{K}}$ and the high-field experiments determine $\gamma^{\prime} K_{\mathrm{J}} R_{\mathrm{K}}$. This leads to

$$
\begin{align*}
& \gamma^{\prime}=\Gamma_{90}^{\prime}(\mathrm{lo}) \frac{K_{\mathrm{J}} R_{\mathrm{K}}}{K_{\mathrm{J}-90} R_{\mathrm{K}-90}},  \tag{248}\\
& \gamma^{\prime}=\Gamma_{90}^{\prime}(\mathrm{hi}) \frac{K_{\mathrm{J}-90} R_{\mathrm{K}-90}}{K_{\mathrm{J}} R_{\mathrm{K}}} \tag{249}
\end{align*}
$$

where $\Gamma_{90}^{\prime}(\mathrm{lo})$ and $\Gamma_{90}^{\prime}(\mathrm{hi})$ are the experimental values of $\gamma^{\prime}$ in SI units that would result from the low- and highfield experiments if $K_{\mathrm{J}}$ and $R_{\mathrm{K}}$ had the exactly known conventional values of $K_{\mathrm{J}-90}$ and $R_{\mathrm{K}-90}$, respectively. The quantities $\Gamma_{90}^{\prime}(\mathrm{lo})$ and $\Gamma_{90}^{\prime}(\mathrm{hi})$ are the input data used in the adjustment, but the observational equations take into account the fact that $K_{\mathrm{J}-90} \neq K_{\mathrm{J}}$ and $R_{\mathrm{K}-90} \neq R_{\mathrm{K}}$.

Accurate values of $\Gamma_{90}^{\prime}(\mathrm{lo})$ and $\Gamma_{90}^{\prime}(\mathrm{hi})$ for the proton and helion are of potential importance because they provide information on the values of $\alpha$ and $h$. Assuming the validity of the relations $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$, the following expressions apply to the four available proton results and one available helion result:

$$
\begin{equation*}
\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})=\frac{K_{\mathrm{J}-90} R_{\mathrm{K}-90} g_{\mathrm{e}^{-}}}{4 \mu_{0} R_{\infty}} \frac{\mu_{\mathrm{p}}^{\prime}}{\mu_{\mathrm{e}^{-}}} \alpha^{3}, \tag{250}
\end{equation*}
$$

TABLE XXI. Summary of data related to shielded gyromagnetic ratios of the proton and helion, and inferred values of $\alpha$ and $h$.

| Quantity | Value | Relative standard uncertainty $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :---: | :---: | :---: | :---: | :---: |
| $\begin{gathered} \Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo}) \\ \alpha^{-1} \end{gathered}$ | $2.67515405(30) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $1.1 \times 10^{-7}$ | NIST-89 | VII.A.1.a (253) |
|  | $137.0359879(51)$ | $3.7 \times 10^{-8}$ |  | VII.A.1.a (254) |
| $\begin{gathered} \Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo}) \\ \alpha^{-1} \end{gathered}$ | $2.6751530(18) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $6.6 \times 10^{-7}$ | NIM-95 | VII.A.1.b (255) |
|  | $137.036006(30)$ | $2.2 \times 10^{-7}$ |  | VII.A.1.b (256) |
| $\begin{gathered} \Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo}) \\ \alpha^{-1} \end{gathered}$ | $2.03789537(37) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $1.8 \times 10^{-7}$ | KR/VN-98 | VII.A.1.c (257) |
|  | 137.035 9852(82) | $6.0 \times 10^{-8}$ |  | VII.A.1.c (258) |
| $\underset{\mathrm{p}-90}{\prime}(\mathrm{hi})$ | $2.6751525(43) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $1.6 \times 10^{-6}$ | NIM-95 | VII.A.2.a (259) |
|  | $6.626071(11) \times 10^{-34} \mathrm{~J} \mathrm{~s}$ | $1.6 \times 10^{-6}$ |  | VII.A.2.a (261) |
| $\underset{\mathrm{p}-90}{\Gamma_{h}^{\prime}}(\mathrm{hi})$ | $2.6751518(27) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $1.0 \times 10^{-6}$ | NPL-79 | VII.A.2.b (262) |
|  | $6.6260729(67) \times 10^{-34} \mathrm{~J} \mathrm{~s}$ | $1.0 \times 10^{-6}$ |  | VII.A.2.b (263) |

$$
\begin{align*}
& \Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})=-\frac{K_{\mathrm{J}-90} R_{\mathrm{K}-90} g_{\mathrm{e}^{-}}}{4 \mu_{0} R_{\infty}} \frac{\mu_{\mathrm{h}}^{\prime}}{\mu_{\mathrm{e}^{-}}} \alpha^{3},  \tag{251}\\
& \Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})=\frac{c \alpha^{2} g_{\mathrm{e}^{-}}}{2 K_{\mathrm{J}-90} R_{\mathrm{K}-90} R_{\infty}} \frac{\mu_{\mathrm{p}}^{\prime}}{\mu_{\mathrm{e}^{-}}^{\prime}} \frac{1}{h} \tag{252}
\end{align*}
$$

Since the five experiments, including necessary corrections, were discussed fully in CODATA-98, only a brief summary is given in the following sections. The five results, together with the value of $\alpha$ inferred from each low-field measurement and the value of $h$ inferred from each high-field measurement, are collected in Table XXI.

## 1. Low-field measurements

A number of national metrology institutes have long histories of measuring the gyromagnetic ratio of the shielded proton, motivated, in part, by their need to monitor the stability of their practical unit of current based on groups of standard cells and standard resistors. This was prior to the development of the Josephson and quantum Hall effects for the realization of practical electric units.

## a. NIST: Low field

The most recent National Institute of Standards and Technology (NIST), Gaithersburg, USA, low-field measurement was reported by Williams et al. (1989). Their result is

$$
\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})=2.67515405(30) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}
$$

$$
\begin{equation*}
\left[1.1 \times 10^{-7}\right] \tag{253}
\end{equation*}
$$

where $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ is related to $\gamma_{\mathrm{p}}^{\prime}$ by Eq. (248).

The value of $\alpha$ that may be inferred from this result follows from Eq. (250). Using the 2006 recommended values for other relevant quantities, the uncertainties of which are significantly smaller than the uncertainty of the NIST result (statements that also apply to the following four similar calculations), we obtain

$$
\begin{equation*}
\alpha^{-1}=137.0359879(51) \quad\left[3.7 \times 10^{-8}\right], \tag{254}
\end{equation*}
$$

where the relative uncertainty is about one-third the relative uncertainty of the NIST value of $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ because of the cube-root dependence of $\alpha$ on $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$.

## b. NIM: Low field

The latest low-field proton gyromagnetic ratio experiment carried out by researchers at the National Institute of Metrology (NIM), Beijing, PRC, yielded (Liu et al., 1995)

$$
\begin{align*}
\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})= & 2.6751530(18) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1} \\
& {\left[6.6 \times 10^{-7}\right] . } \tag{255}
\end{align*}
$$

Based on Eq. (250), the inferred value of $\alpha$ from the NIM result is

$$
\begin{equation*}
\alpha^{-1}=137.036006(30) \quad\left[2.2 \times 10^{-7}\right] \tag{256}
\end{equation*}
$$

## c. KRISS/VNIIM: Low field

The determination of $\gamma_{\mathrm{h}}^{\prime}$ at the Korea Research Institute of Standards and Science (KRISS), Taedok Science Town, Republic of Korea, was carried out in a collaborative effort with researchers from the Mendeleyev AllRussian Research Institute for Metrology (VNIIM), St. Petersburg, Russian Federation (Kim et al., 1995; Shifrin et al., 1998a, 1998b, 1999; Park et al., 1999). The result of this work can be expressed as

TABLE XXII. Summary of data related to the von Klitzing constant $R_{\mathrm{K}}$ and inferred values of $\alpha$.

| Quantity | Value | Relative standard <br> uncertainty $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :---: | :---: | :---: | :---: | :---: |
| $R_{\mathrm{K}}$ | $25812.80831(62) \Omega$ | $2.4 \times 10^{-8}$ | NIST-97 | VII.B.1 (265) |
| $\alpha^{-1}$ | $137.0360037(33)$ | $2.4 \times 10^{-8}$ |  | VII.B.1 (266) |
| $R_{\mathrm{K}}$ | $25812.8071(11) \Omega$ | $4.4 \times 10^{-8}$ | NMI-97 | VII.B.2 (267) |
| $\alpha^{-1}$ | $137.0359973(61)$ | $4.4 \times 10^{-8}$ |  | VII.B.2 (268) |
|  |  |  |  |  |
| $R_{\mathrm{K}}$ | $25812.8092(14) \Omega$ | $5.4 \times 10^{-8}$ | NPL-88 | VII.B.3 (269) |
| $\alpha^{-1}$ | $137.0360083(73)$ | $5.4 \times 10^{-8}$ |  | VII.B.3 (270) |
|  |  | $1.3 \times 10^{-7}$ |  | NIM-95 |
| $R_{\mathrm{K}}$ | $25812.8084(34) \Omega$ | $1.3 \times 10^{-7}$ |  | VII.B.4 (271) |
| $\alpha^{-1}$ | $137.036004(18)$ |  |  | VII.B.4 (272) |
| $R_{\mathrm{K}}$ | $25812.8081(14) \Omega$ | $137.0360023(73)$ | $5.3 \times 10^{-8}$ | LNE-01 |
| $\alpha^{-1}$ |  |  | VII.B.5 (273) |  |

$$
\begin{align*}
\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})= & 2.03789537(37) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1} \\
& {\left[1.8 \times 10^{-7}\right], } \tag{257}
\end{align*}
$$

and the value of $\alpha$ that may be inferred from it through Eq. (251) is

$$
\begin{equation*}
\alpha^{-1}=137.0359852(82) \quad\left[6.0 \times 10^{-8}\right] . \tag{258}
\end{equation*}
$$

## 2. High-field measurements

## a. NIM: High field

The latest high-field proton gyromagnetic ratio experiment at NIM yielded (Liu et al., 1995)

$$
\begin{align*}
\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})= & 2.6751525(43) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1} \\
& {\left[1.6 \times 10^{-6}\right] } \tag{259}
\end{align*}
$$

where $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})$ is related to $\gamma_{\mathrm{p}}^{\prime}$ by Eq. (249). Its correlation coefficient with the NIM low-field result in Eq. (255) is

$$
\begin{equation*}
r(\mathrm{lo}, \mathrm{hi})=-0.014 \tag{260}
\end{equation*}
$$

Based on Eq. (252), the value of $h$ that may be inferred from the NIM high-field result is

$$
\begin{equation*}
h=6.626071(11) \times 10^{-34} \mathrm{~J} \mathrm{~s} \quad\left[1.6 \times 10^{-6}\right] \tag{261}
\end{equation*}
$$

## b. NPL: High field

The most accurate high-field $\gamma_{\mathrm{p}}^{\prime}$ experiment was carried out at NPL by Kibble and Hunt (1979), with the result

$$
\begin{align*}
\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})= & 2.6751518(27) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1} \\
& {\left[1.0 \times 10^{-6}\right] . } \tag{262}
\end{align*}
$$

This leads to the inferred value

$$
\begin{equation*}
h=6.6260729(67) \times 10^{-34} \mathrm{~J} \mathrm{~s} \quad\left[1.0 \times 10^{-6}\right] \tag{263}
\end{equation*}
$$

based on Eq. (252).

## B. von Klitzing constant $R_{\mathrm{K}}$ and $\alpha$

Since the quantum Hall effect, the von Klitzing constant $R_{\mathrm{K}}$ associated with it, and available determinations of $R_{\mathrm{K}}$ are fully discussed in CODATA-98 and CODATA-02, we only outline the main points here.

The quantity $R_{\mathrm{K}}$ is measured by comparing a quantized Hall resistance $R_{\mathrm{H}}(i)=R_{\mathrm{K}} / i$, where $i$ is an integer, to a resistance $R$ whose value is known in terms of the SI unit of resistance $\Omega$. In practice, the latter quantity, the ratio $R / \Omega$, is determined by means of a calculable cross capacitor, a device based on a theorem in electrostatics from the 1950s (Thompson and Lampard, 1956; Lampard, 1957). The theorem allows one to construct a cylindrical capacitor, generally called a ThompsonLampard calculable capacitor (Thompson, 1959), whose capacitance, to high accuracy, depends only on its length.

As indicated in Sec. II, if one assumes the validity of the relation $R_{\mathrm{K}}=h / e^{2}$, then $R_{\mathrm{K}}$ and the fine-structure constant $\alpha$ are related by

$$
\begin{equation*}
\alpha=\mu_{0} c / 2 R_{\mathrm{K}} \tag{264}
\end{equation*}
$$

Hence, the relative uncertainty of the value of $\alpha$ that may be inferred from a particular experimental value of $R_{\mathrm{K}}$ is the same as the relative uncertainty of that value.

The values of $R_{\mathrm{K}}$ we take as input data in the 2006 adjustment and the corresponding inferred values of $\alpha$ are given in the following sections and summarized in Table XXII.

## 1. NIST: Calculable capacitor

The result obtained at NIST is (Jeffery et al., 1997) [see also Jeffery et al. (1998)]

$$
\begin{align*}
R_{\mathrm{K}} & =25812.8\left[1+0.322(24) \times 10^{-6}\right] \Omega \\
& =25812.80831(62) \Omega \quad\left[2.4 \times 10^{-8}\right] \tag{265}
\end{align*}
$$

and is viewed as superseding the NIST result reported by Cage et al. (1989). Work by Jeffery et al. (1999) provides additional support for the uncertainty budget of the NIST calculable capacitor.
The value of $\alpha$ that may be inferred from the NIST value of $R_{\mathrm{K}}$ is, from Eq. (264),

$$
\begin{equation*}
\alpha^{-1}=137.0360037(33) \quad\left[2.4 \times 10^{-8}\right] \tag{266}
\end{equation*}
$$

## 2. NMI: Calculable capacitor

Based on measurements carried out at the National Metrology Institute (NMI), Lindfield, Australia, from December 1994 to April 1995, and a complete reassessment of uncertainties associated with their calculable capacitor and associated apparatus, Small et al. (1997) reported the result

$$
\begin{align*}
R_{\mathrm{K}} & =R_{\mathrm{K}-90}\left[1+0.4(4.4) \times 10^{-8}\right] \\
& =25812.8071(11) \Omega \quad\left[4.4 \times 10^{-8}\right] . \tag{267}
\end{align*}
$$

The value of $\alpha$ it implies is

$$
\begin{equation*}
\alpha^{-1}=137.0359973(61) \quad\left[4.4 \times 10^{-8}\right] . \tag{268}
\end{equation*}
$$

Because of problems associated with the 1989 NMI value of $R_{\mathrm{K}}$, only the result reported in 1997 is used in the 2006 adjustment, as was the case in the 1998 and 2002 adjustments.

## 3. NPL: Calculable capacitor

The NPL calculable capacitor is similar in design to those of NIST and NMI. The result for $R_{\mathrm{K}}$ reported by Hartland et al. (1988) is

$$
\begin{align*}
R_{\mathrm{K}} & =25812.8\left[1+0.356(54) \times 10^{-6}\right] \Omega \\
& =25812.8092(14) \Omega \quad\left[5.4 \times 10^{-8}\right], \tag{269}
\end{align*}
$$

and the value of $\alpha$ that one may infer from it is

$$
\begin{equation*}
\alpha^{-1}=137.0360083(73) \quad\left[5.4 \times 10^{-8}\right] \tag{270}
\end{equation*}
$$

## 4. NIM: Calculable capacitor

The NIM calculable cross capacitor differs markedly from the version used at NIST, NMI, and NPL. The four bars (electrodes) that comprise the capacitor are horizontal rather than vertical and the length that determines its known capacitance is fixed rather than variable. The NIM result for $R_{\mathrm{K}}$, as reported by Zhang et al. (1995) is

$$
\begin{equation*}
R_{\mathrm{K}}=25812.8084(34) \Omega \quad\left[1.3 \times 10^{-7}\right] \tag{271}
\end{equation*}
$$

which implies

$$
\begin{equation*}
\alpha^{-1}=137.036004(18) \quad\left[1.3 \times 10^{-7}\right] \tag{272}
\end{equation*}
$$

## 5. LNE: Calculable capacitor

The value of $R_{\mathrm{K}}$ obtained at the Laboratoire National d'Essais (LNE), Trappes, France, is (Trapon et al., 2001, 2003)

$$
\begin{equation*}
R_{\mathrm{K}}=25812.8081(14) \Omega \quad\left[5.3 \times 10^{-8}\right] \tag{273}
\end{equation*}
$$

which implies

$$
\begin{equation*}
\alpha^{-1}=137.0360023(73) \quad\left[5.3 \times 10^{-8}\right] \tag{274}
\end{equation*}
$$

The LNE Thompson-Lampard calculable capacitor is unique among all calculable capacitors in that it consists of five horizontal bars arranged at the corners of a regular pentagon.

## C. Josephson constant $K_{\mathrm{J}}$ and $\boldsymbol{h}$

Again, since the Josephson effect, the Josephson constant $K_{\mathrm{J}}$ associated with it, and the available determinations of $K_{\mathrm{J}}$ are fully discussed in CODATA-98 and CODATA-02, we only outline the main points here.

The quantity $K_{\mathrm{J}}$ is measured by comparing a Josephson voltage $U_{\mathrm{J}}(n)=n f / K_{\mathrm{J}}$ to a high voltage $U$ whose value is known in terms of the SI unit of voltage V. Here $n$ is an integer and $f$ is the frequency of the microwave radiation applied to the Josephson device. In practice, the latter quantity, the ratio $U / \mathrm{V}$, is determined by counterbalancing an electrostatic force arising from the voltage $U$ with a known gravitational force.

A measurement of $K_{\mathrm{J}}$ can also provide a value of $h$. If, as discussed in Sec. II, we assume the validity of the relation $K_{\mathrm{J}}=2 e / h$ and recall that $\alpha=e^{2} / 4 \pi \epsilon_{0} \hbar c$ $=\mu_{0} c e^{2} / 2 h$, we have

$$
\begin{equation*}
h=\frac{8 \alpha}{\mu_{0} c K_{\mathrm{J}}^{2}} . \tag{275}
\end{equation*}
$$

Since $u_{\mathrm{r}}$ of the fine-structure constant is significantly smaller than $u_{\mathrm{r}}$ of the measured values of $K_{\mathrm{J}}, u_{\mathrm{r}}$ of $h$ derived from Eq. (275) will be essentially twice the $u_{\mathrm{r}}$ of $K_{\mathrm{J}}$.

The values of $K_{\mathrm{J}}$ we take as input data in the 2006 adjustment, and the corresponding inferred values of $h$, are given in the following two sections and summarized in Table XXIII. Also summarized in that table are the measured values of the product $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ and the quantity $\mathcal{F}_{90}$ related to the Faraday constant $F$, together with their corresponding inferred values of $h$. These results are discussed in Secs. VII.D and VII.E.

## 1. NMI: $\mathbf{H g}$ electrometer

The determination of $K_{\mathrm{J}}$ at NMI, carried out using an apparatus called a liquid-mercury electrometer, yielded the result (Clothier et al., 1989)

$$
\begin{align*}
K_{\mathrm{J}} & =483594\left[1+8.087(269) \times 10^{-6}\right] \mathrm{GHz} / \mathrm{V} \\
& =483597.91(13) \mathrm{GHz} / \mathrm{V} \quad\left[2.7 \times 10^{-7}\right] \tag{276}
\end{align*}
$$

Equation (275), the NMI value of $K_{\mathrm{J}}$, and the 2006 recommended value of $\alpha$, which has a much smaller $u_{\mathrm{r}}$,

TABLE XXIII. Summary of data related to the Josephson constant $K_{\mathrm{J}}$, the product $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, the Faraday constant $F$, and inferred values of $h$.

| Quantity | Value | Relative standard <br> uncertainty $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :---: | :---: | :---: | :---: | :---: |
| $K_{\mathrm{J}}$ | $483597.91(13) \mathrm{GHz} \mathrm{V}^{-1}$ | $2.7 \times 10^{-7}$ | NMI-89 | VII.C.1 (276) |
| $h$ | $6.6260684(36) \times 10^{-34} \mathrm{~J} \mathrm{~s}$ | $5.4 \times 10^{-7}$ |  | VII.C.1 (277) |
|  |  |  |  |  |
| $K_{\mathrm{J}}$ | $483597.96(15) \mathrm{GHz} \mathrm{V}^{-1}$ | $3.1 \times 10^{-7}$ | PTB-91 | VII.C.2 (278) |
| $h$ | $6.6260670(42) \times 10^{-34} \mathrm{~J} \mathrm{~s}$ | $6.3 \times 10^{-7}$ |  | VII.C.2 (279) |
|  |  |  |  |  |
| $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $6.0367625(12) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1}$ | $2.0 \times 10^{-7}$ | NPL-90 | VII.D.1 (281) |
| $h$ | $6.6260682(13) \times 10^{-34} \mathrm{~J} \mathrm{~s}^{2}$ | $2.0 \times 10^{-7}$ |  | VII.D.1 (282) |
| $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $6.03676185(53) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1}$ | $8.7 \times 10^{-8}$ | NIST-98 | VII.D.2.a (283) |
| $h$ | $6.62606891(58) \times 10^{-34} \mathrm{~J} \mathrm{~s}$ | $8.7 \times 10^{-8}$ |  | VII.D.2.a (284) |
| $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $6.03676185(22) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1}$ | $3.6 \times 10^{-8}$ | NIST-07 | VII.D.2.b (287) |
| $h$ | $6.62606891(24) \times 10^{-34} \mathrm{~J} \mathrm{~s}$ | $3.6 \times 10^{-8}$ |  | VII.D.2.b (288) |
| $\mathcal{F}_{90}$ | $96485.39(13) \mathrm{C} \mathrm{mol}{ }^{-1}$ |  |  |  |
| $h$ | $6.6260657(88) \times 10^{-34} \mathrm{~J} \mathrm{~s}$ | $1.3 \times 10^{-6}$ | NIST-80 | VII.E.1 (295) |

yields an inferred value for the Planck constant of

$$
\begin{equation*}
h=6.6260684(36) \times 10^{-34} \mathrm{~J} \mathrm{~s} \quad\left[5.4 \times 10^{-7}\right] \tag{277}
\end{equation*}
$$

## 2. PTB: Capacitor voltage balance

The determination of $K_{\mathrm{J}}$ at PTB was carried out using a voltage balance consisting of two coaxial cylindrical electrodes (Sienknecht and Funck, 1985, 1986; Funck and Sienknecht, 1991). Taking into account the correction associated with the reference capacitor used in the PTB experiment as described in CODATA-98, the result of the PTB determination is

$$
\begin{equation*}
K_{\mathrm{J}}=483597.96(15) \mathrm{GHz} / \mathrm{V} \quad\left[3.1 \times 10^{-7}\right] \tag{278}
\end{equation*}
$$

from which we infer, using Eq. (275),

$$
\begin{equation*}
h=6.6260670(42) \times 10^{-34} \mathrm{~J} \mathrm{~s} \quad\left[6.3 \times 10^{-7}\right] \tag{279}
\end{equation*}
$$

## D. Product $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ and $h$

A value for the product $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ is important to determine the Planck constant $h$, because if one assumes that the relations $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ are valid, then

$$
\begin{equation*}
h=\frac{4}{K_{\mathrm{J}}^{2} R_{\mathrm{K}}} . \tag{280}
\end{equation*}
$$

The product $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ is determined by comparing electrical power known in terms of a Josephson voltage and quantized Hall resistance to the equivalent mechanical power known in the SI unit $\mathrm{W}=\mathrm{m}^{2} \mathrm{~kg} \mathrm{~s}^{-3}$. Comparison is carried out using an apparatus known as a moving-coil watt balance first proposed by Kibble (1975) at NPL. To date,
two laboratories, NPL and NIST, have determined $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ using this method.

## 1. NPL: Watt balance

Shortly after Kibble's original proposal in 1975, Kibble and Robinson (1977) carried out a feasibility study of the idea based on experience with the NPL apparatus that was used to determine $\gamma_{\mathrm{p}}^{\prime}$ by the highfield method (Kibble and Hunt, 1979). The work continued and led to a result with an uncertainty of about 2 parts in $10^{7}$ (Kibble et al., 1990). This result, discussed in CODATA-98 and which was taken as an input datum in the 1998 and 2002 adjustments, and which we also take as an input datum in the 2006 adjustment, may be expressed as

$$
\begin{align*}
K_{\mathrm{J}}^{2} R_{\mathrm{K}} & =K_{\mathrm{J}-\mathrm{NPL}}^{2} R_{\mathrm{K}-\mathrm{NPL}}\left[1+16.14(20) \times 10^{-6}\right] \\
& =6.0367625(12) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1} \quad\left[2.0 \times 10^{-7}\right] \tag{281}
\end{align*}
$$

where $\quad K_{\mathrm{J}-\mathrm{NPL}}=483594 \mathrm{GHz} / \mathrm{V}$ and $R_{\mathrm{K}-\mathrm{NPL}}$ $=25812.8092 \Omega$. The value of $h$ that may be inferred from the NPL result is, according to Eq. (280),

$$
\begin{equation*}
h=6.6260682(13) \times 10^{-34} \mathrm{~J} \mathrm{~s} \quad\left[2.0 \times 10^{-7}\right] \tag{282}
\end{equation*}
$$

Based on the experience gained in this experiment, NPL researchers designed and constructed what is essentially a completely new apparatus, called the NPL Mark II watt balance, that could possibly achieve a result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ with an uncertainty of a few parts in $10^{8}$ (Robinson and Kibble, 1997; Kibble and Robinson, 2003). Although the balance itself employs the same bal-
ance beam as the previous NPL watt balance, little else from that experiment is retained in the new experiment.

Over 1000 measurements in vacuum were carried out with the MK II between January 2000 and November 2001. Many were made in an effort to identify the cause of an observed fractional change in the value of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ of about $3 \times 10^{-7}$ that occurred in mid-April 2000 (Robinson and Kibble, 2002). A change in the alignment of the apparatus was suspected of contributing to the shift.

Significant improvements were subsequently made in the experiment, and recently, based on measurements carried out from October 2006 to March 2007, the initial result from MK II, $h=6.62607095(44) \mathrm{J} \mathrm{s}\left[6.6 \times 10^{-8}\right]$, was reported by Robinson and Kibble (2007) assuming the validity of Eq. (280). Although this result became available much too late to be considered for the 2006 adjustment, we do note that it lies between the value of $h$ inferred from the 2007 NIST result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ discussed in Sec. VII.D.2.b, and that inferred from the measurement of the molar volume of silicon $V_{\mathrm{m}}(\mathrm{Si})$ discussed in Sec. VIII.B. The NPL work is continuing and a result with a smaller uncertainty is anticipated (Robinson and Kibble, 2007).

## 2. NIST: Watt balance

## a. 1998 measurement

Work on a moving-coil watt balance at NIST began shortly after Kibble made his 1975 proposal. A first result with $u_{\mathrm{r}}=1.3 \times 10^{-6}$ was reported by NIST researchers in 1989 (Cage et al., 1989). Significant improvements were then made to the apparatus and the final result from this phase of the NIST effort was reported in 1998 by Williams et al. (1998),

$$
\begin{align*}
K_{\mathrm{J}}^{2} R_{\mathrm{K}} & =K_{\mathrm{J}-90}^{2} R_{\mathrm{K}-90}\left[1-8(87) \times 10^{-9}\right] \\
& =6.03676185(53) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1} \quad\left[8.7 \times 10^{-8}\right] . \tag{283}
\end{align*}
$$

A lengthy paper giving the details of the NIST 1998 watt balance experiment was published in 2005 by Steiner, Newell, and Williams (2005). This was the NIST result taken as an input datum in the 1998 and 2002 adjustments; although the 1989 result was consistent with that of 1998 , its uncertainty was about 15 times larger. The value of $h$ implied by the 1998 NIST result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ is

$$
\begin{equation*}
h=6.62606891(58) \times 10^{-34} \mathrm{~J} \text { s } \quad\left[8.7 \times 10^{-8}\right] \tag{284}
\end{equation*}
$$

## b. 2007 measurement

Based on the lessons learned in the decade-long effort with a watt balance operating in air that led to their 1998 result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, the NIST watt-balance researchers initiated a new program with the goal of measuring $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ with $u_{\mathrm{r}} \approx 10^{-8}$. The experiment was completely disassembled, and renovations to the research facility were made to improve vibration isolation, reduce electromagnetic interference, and incorporate a multilayer temperature control system. A new watt balance with major
changes and improvements was constructed with little remaining of the earlier apparatus except the superconducting magnet used to generate the required radial magnetic flux density and the wheel used as the balance.

The most notable change in the experiment is that in the new apparatus the entire balance mechanism and moving coil are in vacuum, which eliminates the uncertainties of the corrections in the previous experiment for the index of refraction of air in the laser position measurements ( $u_{\mathrm{r}}=43 \times 10^{-9}$ ) and for the buoyancy force exerted on the mass standard $\left(u_{\mathrm{r}}=23 \times 10^{-9}\right)$. Alignment uncertainties were reduced by over a factor of 4 by (i) incorporating a more comprehensive understanding of all degrees of freedom involving the moving coil; and (ii) the application of precise alignment techniques for all degrees of freedom involving the moving coil, the superconducting magnet, and the velocity measuring interferometers. Hysteresis effects were reduced by a factor of 4 by using a diamondlike carbon-coated knife edge and flat (Schwarz et al., 2001), employing a hysteresis erasure procedure, and reducing the balance deflections during mass exchanges with improved control systems. A programmable Josephson array voltage standard (Benz et al., 1997) was connected directly to the experiment, eliminating two voltage transfers required in the old experiment and reducing the voltage traceability uncertainty by a factor of 15 .

A total of 6023 individual values of $W_{90} / \mathrm{W}$ were obtained over the 2-year period from March 2003 to February 2005 as part of the effort to develop and improve the new experiment. (The results are converted to the notation used here by the relation $W_{90} / \mathrm{W}$ $=K_{\mathrm{J}-90}^{2} R_{\mathrm{K}-90} / K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ discussed in CODATA-98.) The initial result from that work was reported by Steiner, Williams, Newell, et al. (2005),

$$
\begin{align*}
K_{\mathrm{J}}^{2} R_{\mathrm{K}} & =K_{\mathrm{J}-90}^{2} R_{\mathrm{K}-90}\left[1-24(52) \times 10^{-9}\right] \\
& =6.03676175(31) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1} \quad\left[5.2 \times 10^{-8}\right] \tag{285}
\end{align*}
$$

which yields a value for the Planck constant of

$$
\begin{equation*}
h=6.62606901(34) \times 10^{-34} \mathrm{~J} \mathrm{~s} \quad\left[5.2 \times 10^{-8}\right] \tag{286}
\end{equation*}
$$

This result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ was obtained from data spanning the final 7 months of the 2-year period. It is based on the weighted mean of $48 W_{90} / \mathrm{W}$ measurement sets using a Au mass standard and 174 sets using a PtIr mass standard, where a typical measurement set consists of 12 to 15 individual values of $W_{90} / \mathrm{W}$. The 2005 NIST result is consistent with the 1998 NIST result, but its uncertainty has been reduced by a factor of 1.7.

Following this initial effort with the new apparatus, further improvements were made to it in order to reduce the uncertainties from various systematic effects, the most notable reductions being in the determination of the local acceleration due to gravity $g$ (a factor of 2.5), the effect of balance wheel surface roughness (a factor of 10), and the effect of the magnetic susceptibility of the mass standard (a factor of 1.6). An improved result was
then obtained based on 2183 values of $W_{90} / \mathrm{W}$ recorded in 134 measurement sets from January 2006 to June 2006. Due to a wear problem with the gold mass standard, only a PtIr mass standard was used in these measurements. The result, first reported in 2006 and subsequently published in 2007 by Steiner et al. (2007), is

$$
\begin{align*}
K_{\mathrm{J}}^{2} R_{\mathrm{K}} & =K_{\mathrm{J}-90}^{2} R_{\mathrm{K}-90}\left[1-8(36) \times 10^{-9}\right] \\
& =6.03676185(22) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1} \quad\left[3.6 \times 10^{-8}\right] . \tag{287}
\end{align*}
$$

The value of $h$ that may be inferred from this value of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ is

$$
\begin{equation*}
h=6.62606891(24) \times 10^{-34} \mathrm{~J} \text { s } \quad\left[3.6 \times 10^{-8}\right] \tag{288}
\end{equation*}
$$

The 2007 NIST result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ is consistent with and has an uncertainty smaller by a factor of 1.4 than the uncertainty of the 2005 NIST result. However, because the two results are from essentially the same experiment and hence are highly correlated, we take only the 2007 result as an input datum in the 2006 adjustment.

On the other hand, the experiment on which the NIST 2007 result is based is only slightly dependent on the experiment on which the NIST 1998 result is based, as can be seen from the above discussions. Thus, in keeping with our practice in similar cases, most notably the 1982 and 1999 LAMPF measurements of muonium Zeeman transition frequencies (see Sec. VI.B.2), we also take the NIST 1998 result in Eq. (283) as an input datum in the 2006 adjustment. But to ensure that we do not give undue weight to the NIST work, an analysis of the uncertainty budgets of the 1998 and 2007 NIST results was performed to determine the level of correlation. Of the relative uncertainty components listed in Table II of Williams et al. (1998) and in Table 2 of Steiner, Williams, Newell, et al. (2005) but as updated in Table 1 of Steiner et al. (2007), the largest common relative uncertainty components were from the magnetic flux profile fit due to the use of the same analysis routine $\left(16 \times 10^{-9}\right)$; leakage resistance and electrical grounding since the same current supply was used in both experiments (10 $\times 10^{-9}$ ) ; and the determination of the local gravitational acceleration $g$ due to the use of the same absolute gravimeter $\left(7 \times 10^{-9}\right)$. The correlation coefficient was thus determined to be

$$
\begin{equation*}
r\left(K_{\mathrm{J}}^{2} R_{\mathrm{K}}-98, K_{\mathrm{J}}^{2} R_{\mathrm{K}}-07\right)=0.14 \tag{289}
\end{equation*}
$$

which we take into account in our calculations as appropriate.

## 3. Other values

Although there is no competitive published value of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ other than those from NPL and NIST discussed above, it is worth noting that at least three additional laboratories have watt-balance experiments in progress: the Swiss Federal Office of Metrology and Accreditation (METAS), Bern-Wabern, Switzerland, the LNE, and the

BIPM. Descriptions of these efforts may be found in Beer et al. (2003), Genevès et al. (2005), and Picard et al. (2007), respectively.

## 4. Inferred value of $\boldsymbol{K}_{\mathbf{J}}$

It is of interest to note that a value of $K_{\mathrm{J}}$ with an uncertainty significantly smaller than those of the directly measured values discussed in Sec. VII.C can be obtained from directly measured watt-balance values of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, together with directly measured calculablecapacitor values of $R_{\mathrm{K}}$, without assuming the validity of the relations $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$. The relevant expression is simply $K_{\mathrm{J}}=\left[\left(K_{\mathrm{J}}^{2} R_{\mathrm{K}}\right)_{\mathrm{W}} /\left(R_{\mathrm{K}}\right)_{\mathrm{C}}\right]^{1 / 2}$, where $\left(K_{\mathrm{J}}^{2} R_{\mathrm{K}}\right)_{\mathrm{W}}$ is from the watt balance and $\left(R_{\mathrm{K}}\right)_{\mathrm{C}}$ is from the calculable capacitor.

Using the weighted mean of the three watt-balance results for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ discussed in this section and the weighted mean of the five calculable-capacitor results for $R_{\mathrm{K}}$ discussed in Sec. VII.B, we have

$$
\begin{align*}
K_{\mathrm{J}} & =K_{\mathrm{J}-90}\left[1-2.8(1.9) \times 10^{-8}\right] \\
& =483597.8865(94) \mathrm{GHz} / \mathrm{V} \quad\left[1.9 \times 10^{-8}\right] \tag{290}
\end{align*}
$$

which is consistent with the two directly measured values but has an uncertainty that is smaller by more than an order of magnitude. This result is implicitly included in the least-squares adjustment, even though the explicit value for $K_{\mathrm{J}}$ obtained here is not used as an input datum.

## E. Faraday constant $\boldsymbol{F}$ and $\boldsymbol{h}$

The Faraday constant $F$ is equal to the Avogadro constant $N_{\mathrm{A}}$ times the elementary charge $e, F=N_{\mathrm{A}} e$; its SI unit is coulomb per mol, $\mathrm{C} \mathrm{mol}{ }^{-1}=\mathrm{As} \mathrm{mol}^{-1}$. It determines the amount of substance $n(X)$ of an entity $X$ that is deposited or dissolved during electrolysis by the passage of a quantity of electricity, or charge, $Q=I t$, due to the flow of a current $I$ in a time $t$. In particular, the Faraday constant $F$ is related to the molar mass $M(X)$ and valence $z$ of entity $X$ by

$$
\begin{equation*}
F=\frac{\operatorname{ItM}(X)}{z m_{\mathrm{d}}(X)}, \tag{291}
\end{equation*}
$$

where $m_{\mathrm{d}}(X)$ is the mass of entity $X$ dissolved as the result of transfer of charge $Q=I t$ during the electrolysis. It follows from the relations $F=N_{\mathrm{A}} e, e^{2}=2 \alpha h / \mu_{0} c$, $m_{\mathrm{e}}=2 R_{\circ} h / c \alpha^{2}$, and $N_{\mathrm{A}}=A_{\mathrm{r}}(\mathrm{e}) M_{\mathrm{u}} / m_{\mathrm{e}}$, where $M_{\mathrm{u}}$ $=10^{-3} \mathrm{~kg} \mathrm{~mol}^{-1}$, that

$$
\begin{equation*}
F=\frac{A_{\mathrm{r}}(\mathrm{e}) M_{\mathrm{u}}}{R_{\infty}}\left(\frac{c}{2 \mu_{0}} \frac{\alpha^{5}}{h}\right)^{1 / 2} . \tag{292}
\end{equation*}
$$

Since, according to Eq. (291), $F$ is proportional to the current $I$, and $I$ is inversely proportional to the product $K_{\mathrm{J}} R_{\mathrm{K}}$ if the current is determined in terms of the Josephson and quantum Hall effects, we may write

$$
\begin{equation*}
\mathcal{F}_{90}=\frac{K_{\mathrm{J}} R_{\mathrm{K}}}{K_{\mathrm{J}-90} R_{\mathrm{K}-90}} \frac{A_{\mathrm{r}}(\mathrm{e}) M_{\mathrm{u}}}{R_{\infty}}\left(\frac{c}{2 \mu_{0}} \frac{\alpha^{5}}{h}\right)^{1 / 2} \tag{293}
\end{equation*}
$$

where $\mathcal{F}_{90}$ is the experimental value of $F$ in SI units that would result from the Faraday experiment if $K_{\mathrm{J}}=K_{\mathrm{J}-90}$ and $R_{\mathrm{K}}=R_{\mathrm{K}-90}$. The quantity $\mathcal{F}_{90}$ is the input datum used in the adjustment, but the observational equation accounts for the fact that $K_{\mathrm{J}-90} \neq K_{\mathrm{J}}$ and $R_{\mathrm{K}-90} \neq R_{\mathrm{K}}$. If one assumes the validity of $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$, then in terms of adjusted constants, Eq. (293) can be written as

$$
\begin{equation*}
\mathcal{F}_{90}=\frac{c M_{\mathrm{u}}}{K_{\mathrm{J}-90} R_{\mathrm{K}-90}} \frac{A_{\mathrm{r}}(\mathrm{e}) \alpha^{2}}{R_{\infty} h} . \tag{294}
\end{equation*}
$$

## 1. NIST: Ag coulometer

There is one high-accuracy experimental value of $\mathcal{F}_{90}$ available, that from NIST (Bower and Davis, 1980). The NIST experiment used a silver dissolution coulometer based on the anodic dissolution by electrolysis of silver, which is monovalent, into a solution of perchloric acid containing a small amount of silver perchlorate. The basic chemical reaction is $\mathrm{Ag} \rightarrow \mathrm{Ag}^{+}+\mathrm{e}^{-}$and occurs at the anode, which in the NIST work was a highly purified silver bar.

As discussed in detail in CODATA-98, the NIST experiment leads to

$$
\begin{equation*}
\mathcal{F}_{90}=96485.39(13) \mathrm{C} \mathrm{~mol}^{-1} \quad\left[1.3 \times 10^{-6}\right] \tag{295}
\end{equation*}
$$

[Note that the new AME2003 values of $A_{\mathrm{r}}\left({ }^{107} \mathrm{Ag}\right)$ and $A_{\mathrm{r}}\left({ }^{109} \mathrm{Ag}\right)$ in Table II do not change this result.]

The value of $h$ that may be inferred from the NIST result, Eq. (294), and the 2006 recommended values for the other quantities is

$$
\begin{equation*}
h=6.6260657(88) \times 10^{-34} \mathrm{~J} \text { s } \quad\left[1.3 \times 10^{-6}\right] \tag{296}
\end{equation*}
$$

where the uncertainties of the other quantities are negligible compared to the uncertainty of $\mathcal{F}_{90}$.

## VIII. MEASUREMENTS INVOLVING SILICON CRYSTALS

Here we discuss experiments relevant to the 2006 adjustment that use highly pure, nearly crystallographically perfect, single crystals of silicon. However, because one such experiment determines the quotient $h / m_{\mathrm{n}}$, where $m_{\mathrm{n}}$ is the mass of the neutron, for convenience and because any experiment that determines the ratio of the Planck constant to the mass of a fundamental particle or atom provides a value of the fine-structure constant $\alpha$, we also discuss in this section two silicon-independent experiments: the 2002 Stanford University, Stanford, California, USA, measurement of $h / m\left({ }^{133} \mathrm{Cs}\right)$ and the 2006 Laboratoire Kastler-Brossel or LKB measurement of $h / m\left({ }^{87} \mathrm{Rb}\right)$.

In this section, W4.2a, NR3, W04, and NR4 are shortened forms of the full crystal designations WASO 4.2a, NRLM3, WASO 04, and NRLM4, respectively, for use in quantity symbols. No distinction is made between different crystals taken from the same ingot. As we use the
current laboratory name to identify a result rather than the laboratory name at the time the measurement was carried out, we have replaced IMGC and NRLM with INRIM and NMIJ.

## A. $\{220\}$ lattice spacing of silicon $\boldsymbol{d}_{\mathbf{2 2 0}}$

A value of the $\{220\}$ lattice spacing of a silicon crystal in meters is relevant to the 2006 adjustment not only because of its role in determining $\alpha$ from $h / m_{\mathrm{n}}$ (see Sec. VIII.D.1), but also because of its role in determining the relative atomic mass of the neutron $A_{\mathrm{r}}(\mathrm{n})$ (see Sec. VIII.C). Further, together with the measured value of the molar volume of silicon $V_{\mathrm{m}}(\mathrm{Si})$, it can provide a competitive value of $h$ (see Sec. VIII.B).

Various aspects of silicon and its crystal plane spacings of interest here are reviewed in CODATA-98 and CODATA-02. [See also the reviews of Becker (2003), Mana (2001), and Becker (2001).] Some points worth noting are that silicon is a cubic crystal with $n=8$ atoms per face-centered-cubic unit cell of edge length (or lattice parameter) $a=543 \mathrm{pm}$ with $d_{220}=a / \sqrt{8}$. The three naturally occurring isotopes of Si are ${ }^{28} \mathrm{Si},{ }^{29} \mathrm{Si}$, and ${ }^{30} \mathrm{Si}$, and the amount-of-substance fractions $x\left({ }^{28} \mathrm{Si}\right), x\left({ }^{29} \mathrm{Si}\right)$, and $x\left({ }^{30} \mathrm{Si}\right)$ of natural silicon are approximately 0.92 , 0.05 , and 0.03 , respectively.

Although the $\{220\}$ lattice spacing of Si is not a fundamental constant in the usual sense, for practical purposes one can consider $a$, and hence $d_{220}$, of an impurityfree, crystallographically perfect or "ideal" silicon crystal under specified conditions, principally of temperature, pressure, and isotopic composition, to be an invariant of nature. The reference temperature and pressure currently adopted are $t_{90}=22.5^{\circ} \mathrm{C}$ and $p=0$ (that is, vacuum), where $t_{90}$ is Celsius temperature on the International Temperature Scale of 1990 (ITS-90) (Preston-Thomas, 1990a, 1990b). However, no reference values for $x\left({ }^{A} \mathrm{Si}\right)$ have yet been adopted, because the variation of $a$ due to the variation of the isotopic composition of the crystals used in high-accuracy experiments is taken to be negligible at the current level of experimental uncertainty in $a$. A much larger effect on $a$ is the impurities that the silicon crystal contains-mainly carbon $(\mathrm{C})$, oxygen $(\mathrm{O})$, and nitrogen ( N ) -and corrections must be applied to convert the $\{220\}$ lattice spacing $d_{220}(X)$ of a real crystal $X$ to the $\{220\}$ lattice spacing $d_{220}$ of an ideal crystal.

Nevertheless, we account for the possible variation in the lattice spacing of different samples taken from the same ingot by including an additional component (or components) of relative standard uncertainty in the uncertainty of any measurement result involving a silicon lattice spacing (or spacings). This additional component is typically $\sqrt{2} \times 10^{-8}$ for each crystal, but it can be larger, for example, $(3 / 2) \sqrt{2} \times 10^{-8}$ in the case of crystal $\mathrm{MO}^{*}$ discussed below, because it is known to contain a comparatively large amount of carbon; see Secs. III.A.c and III.I of CODATA-98 for details. For simplicity, we do

TABLE XXIV. Summary of measurements of the absolute $\{220\}$ lattice spacing of various silicon crystals and inferred values of $d_{220}$. (The latter follows from the data discussed in Sec. VIII.A.2.)

| Quantity | Value | Relative standard <br> uncertainty $u_{r}$ | Identification | Sec. and Eq. |
| :---: | :--- | :---: | :--- | :--- |
| $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ | $192015.563(12) \mathrm{fm}$ | $6.2 \times 10^{-8}$ | PTB-81 | VIII.A.1.a (297) |
| $d_{220}$ | $192015.565(13) \mathrm{fm}$ | $6.5 \times 10^{-8}$ |  |  |
| $d_{220}(\mathrm{NR} 3)$ | $192015.5919(76) \mathrm{fm}$ | $4.0 \times 10^{-8}$ | NMIJ-04 | VIII.A.1.b (298) |
| $d_{220}$ | $192015.5973(84) \mathrm{fm}$ | $4.4 \times 10^{-8}$ |  |  |
| $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ | $192015.5715(33) \mathrm{fm}$ | $1.7 \times 10^{-8}$ | INRIM-07 | VIII.A.1.c (299) |
| $d_{220}$ | $192015.5732(53) \mathrm{fm}$ | $2.8 \times 10^{-8}$ |  |  |
| $d_{220}(\mathrm{MO} *)$ | $192015.5498(51) \mathrm{fm}$ | $2.6 \times 10^{-8}$ | INRIM-07 | VIII.A.1.c (300) |
| $d_{220}$ | $192015.5685(67) \mathrm{fm}$ | $3.5 \times 10^{-8}$ |  |  |
| $h / m_{\mathrm{n}} d_{220}(\mathrm{~W} 04)$ | $2060.267004(84) \mathrm{m} \mathrm{s}$ |  |  |  |
| $d_{220}$ | $192015.5982(79) \mathrm{fm}$ | $4.1 \times 10^{-8}$ | PTB-99 | VIII.D.1 (322) |

not explicitly mention our inclusion of such components in the following discussion.

Further, because of this component and the use of the same samples in different experiments, and because of the existence of other common components of uncertainty in the uncertainty budgets of different experimental results involving silicon crystals, many of the input data discussed in the following sections are correlated. In most cases, we do not explicity give the relevant correlation coefficients in the text; instead Table XXXI in Sec. XII provides all non-negligible correlation coefficients of the input data listed in Table XXX.

## 1. X-ray optical interferometer measurements of $\boldsymbol{d}_{\mathbf{2 2 0}}(X)$

High accuracy measurements of $d_{220}(X)$, where $X$ denotes any one of various crystals, are carried out using a combined x-ray and optical interferometer (XROI) fabricated from a single-crystal of silicon taken from one of several well-characterized single crystal ingots or boules. As discussed in CODATA-98, an XROI is a device that enables x-ray fringes of unknown period $d_{220}(X)$ to be compared with optical fringes of known period by moving one of the crystal plates of the XROI, called the analyzer. Also discussed there are the XROI measurements of $d_{220}(\mathrm{~W} 4.2 \mathrm{a}), d_{220}(\mathrm{MO} *)$, and $d_{220}(\mathrm{NR} 3)$, which were carried out at the PTB in Germany (Becker et al., 1981), the Istituto Nazionale di Ricerca Metrologica, Torino, Italy (INRIM) (Basile et al., 1994), and the National Metrology Institute of Japan (NMIJ), Tsukuba, Japan (Nakayama and Fujimoto, 1997), respectively.

For reasons discussed in CODATA-02 and by Cavagnero et al. (2004a, 2004b), only the NMIJ 1997 result was taken as an input datum in the 2002 adjustment. However, further work, published in an Erratum to that paper, showed that the results obtained at INRIM given in the paper were in error. After the error was discovered, additional work was carried out at INRIM to fully understand and correct it. New results were then reported in 2006 (Becker et al., 2007). Thus, as summarized in Table XXIV and compared in Fig. 1, we take as input
data the four absolute $\{220\}$ lattice spacing values determined in three different laboratories, as discussed in the following three sections. The last value in the table, which is not an XROI result, is discussed in Sec. VIII.D.1.

We point out that not only do we take the $\{220\}$ lattice spacings of the crystals WASO 4.2a, NRLM3, and MO* as adjusted constants, but also the $\{220\}$ lattice spacings of the crystals N, WASO 17, ILL, WASO 04, and NRLM4, because they too were involved in various experiments, including the $d_{220}$ lattice spacing fractional difference measurements discussed in Sec. VIII.A.2. (The inferred values of $d_{220}$ in Table XXIV are based on the data in that section.)


FIG. 1. Inferred values (open circles) of $d_{220}$ from various measurements (solid circles) of $d_{220}(X)$. For comparison, the 2002 and 2006 CODATA recommended values of $d_{220}$ are also shown.

## a. PTB measurement of $d_{220}$ (W4.2a)

The following value, identified as PTB-81 in Table XXIV and Fig. 1, is the original result obtained at PTB as reported by Becker et al. (1981) and discussed in CODATA-98:

$$
\begin{equation*}
d_{220}(\mathrm{~W} 4.2 \mathrm{a})=192015.563(12) \mathrm{fm} \quad\left[6.2 \times 10^{-8}\right] . \tag{297}
\end{equation*}
$$

## b. NMIJ measurement of $d_{220}(\mathrm{NR} 3)$

The following value, identified as NMIJ-04 in Table XXIV and Fig. 1, reflects the NMIJ efforts in the early and mid-1990s as well as the work carried out in the early 2000s:

$$
\begin{equation*}
d_{220}(\mathrm{NR} 3)=192015.5919(76) \mathrm{fm} \quad\left[4.0 \times 10^{-8}\right] . \tag{298}
\end{equation*}
$$

This value, reported by Cavagnero et al. (2004a, 2004b), is the weighted mean of the 1997 NIMJ result of Nakayama and Fujimoto (1997) discussed in CODATA-98 and CODATA-02, and the result from a new series of measurements performed at NMIJ from December 2002 to February 2003 with nearly the same apparatus. One of the principal differences from the earlier experiment was the much improved temperature control system of the room in which the NMIJ XROI was located; the new system provided a temperature stability of about $1 \mathrm{mK} / \mathrm{d}$ and allowed the temperature of the XROI to be set to within 20 mK of $22.5^{\circ} \mathrm{C}$.
The result for $d_{220}($ NR3 $)$ from the 2002-2003 measurements is based on 61 raw data. In each measurement, the phases of the x-ray and optical fringes (optical orders) were compared at the 0th, 100th, and 201st optical orders, and then with the analyzer moving in the reverse direction, at the 201st, 100th, and 0th orders. The $n / m$ ratio was calculated from the phase of the x-ray fringe at the 0 th and 201st orders, where $n$ is the number of x -ray fringes in $m$ optical fringes (optical orders) of period $\lambda / 2$, where $\lambda$ is the wavelength of the laser beam used in the optical interferometer and $d_{220}(\mathrm{NR} 3)=(\lambda / 2) /(n / m)$.
In the new work, the fractional corrections to $d_{220}$ (NR3), in parts in $10^{9}$, total 181(35), the largest by far being the correction 173(33) for laser beam diffraction. The next largest is $5.0(7.1)$ for laser beam alignment. The statistical uncertainty is 33 (Type A).
Before calculating the weighted mean of the new and 1997 results for $d_{220}($ NR3 $)$, Cavagnero et al. (2004a, 2004b) revised the 1997 value based on a reanalysis of the old experiment, taking into account what was learned in the new experiment. The reanalysis resulted not only in a reduction of the statistical uncertainty from (again, in parts in $10^{9}$ ) 50 to 1.8 due to a better understanding of the undulation of $n / m$ values as a function of time, but also in more reliable estimates of the corrections for laser beam diffraction and laser beam alignment. Indeed, the fractional corrections for the revised 1997 NMIJ value of $d_{220}($ NR3 ) total $190(38)$ compared to
the original total of 173 (14), and the final uncertainty of the revised 1997 value is $u_{\mathrm{r}}=3.8 \times 10^{-8}$ compared to $u_{\mathrm{r}}$ $=4.8 \times 10^{-8}$ of the new value.
For completeness, we note that two possible corrections to the NMIJ result have been discussed in the literature. In the Erratum to Cavagnero et al. (2004a, 2004b), it is estimated that a fractional correction to the value of $d_{220}$ (NR3) in Eq. (298) of $-1.3 \times 10^{-8}$ may be required to account for the contamination of the NMIJ laser by a parasitic component of laser radiation as in the case of the INRIM laser discussed in the next section. However, it is not applied, because of its comparatively small size and the fact that no measurements of $d_{220}$ (NR3) have yet been made at NMIJ (or INRIM) with a problem-free laser that confirm the correction, as has been done at INRIM for the crystals WASO 4.2a and MO*.

Fujimoto et al. (2007) estimated, based on a Monte Carlo simulation, that the fractional correction to $d_{220}$ (NR3) labeled Fresnel diffraction in Table I of Nakayama and Fujimoto (1997) and equal to 16.0(8) $\times 10^{-8}$ should be $10(3) \times 10^{-8}$. The change arises from taking into account the misalignment of the interfering beams in the laser interferometer. Because this additional diffraction effect was present in both the 1997 and 2002-2003 measurements but was not considered in the reanalysis of the 1997 result nor in the analysis of the 2002-2003 data, it implies that the weighted mean value for $d_{220}(\mathrm{NR} 3)$ in Eq. (298) should be reduced by this amount and its $u_{\mathrm{r}}$ increased from $4.0 \times 10^{-8}$ to 5.0 $\times 10^{-8}$. However, because the data required for the calculation were not precisely known (they were not logged in the laboratory notebooks because the experimenters were unaware of their importance), the correction is viewed as somewhat conjectural and thus applying it would not be justified (Mana and Massa, 2006).

## c. INRIM measurement of $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ and $d_{220}\left(\mathrm{MO}^{*}\right)$

The following two new INRIM values, with identifier INRIM-07, were reported by Becker et al. (2007):

$$
\begin{align*}
& d_{220}(\mathrm{~W} 4.2 \mathrm{a})=192015.5715(33) \mathrm{fm} \quad\left[1.7 \times 10^{-8}\right],  \tag{299}\\
& d_{220}\left(\mathrm{MO}^{*}\right)=192015.5498(51) \mathrm{fm} \quad\left[2.6 \times 10^{-8}\right] . \tag{300}
\end{align*}
$$

The correlation coefficient of these values is 0.057 , based on the detailed uncertainty budget for $d_{220}\left(\mathrm{MO}^{*}\right)$ in Cavagnero et al. (2004a, 2004b) and the similar uncertainty budget for $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ provided by Fujimoto et al. (2006). Although the 2007 result for $d_{220}\left(\mathrm{MO}^{*}\right)$ of Becker et al. (2007) in Eq. (300) agrees with the 1994 INRIM result of Basile et al. (1994), which was used as an input datum in the 1998 adjustment, because of the many advances incorporated in the new work, we no longer consider the old result.
In addition to the determination described in the previous section of the $\{220\}$ lattice spacing of crystal

NRLM3 carried out at NMIJ in 2002-2003 using the NMIJ NRLM3 x-ray interferometer and associated NMIJ apparatus, Cavagnero et al. (2004a, 2004b) reported the results of measurements carried out at INRIM of the $\{220\}$ lattice spacings of crystals MO* and NRLM3, where in the latter case it was an INRIMNMIJ joint effort that used the NIMJ NRLM3 x-ray interferometer but the INRIM associated apparatus. But as indicated above, both results were subsequently found to be in error: the optical laser beam used to measure the displacement of the x-ray interferometer's analyzer crystal was contaminated by a parasitic component with a frequency that differed by about 1.1 GHz from the frequency assigned the laser beam.

After eliminating the error by replacing the problem laser with a $633 \mathrm{~nm} \mathrm{He}-\mathrm{Ne}$ external-cavity diode laser locked to a ${ }^{127} \mathrm{I}_{2}$ stabilized laser, the INRIM researchers repeated the measurements they had previously carried out with the INRIM MO* x-ray interferometer and with the refurbished PTB WASO 4.2a x-ray interferometer originally used in the PTB experiment that led to the 1981 value of $d_{220}$ (W4.2a) in Eq. (297). The PTB WASO 4.2a x-ray interferometer was refurbished at PTB through remachining, but the result for $d_{220}(\mathrm{~W} 4.2 \mathrm{a}) \mathrm{ob}-$ tained at INRIM with the contaminated laser was not included in Cavagnero et al. (2004a, 2004b). The values of $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ and $d_{220}\left(\mathrm{MO}^{*}\right)$ in Eqs. (299) and (300) resulted from the repeated measurements (Becker et al., 2007).

In principle, based on the experimentally observed shifts in the measured values of $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ and $d_{220}\left(\mathrm{MO}^{*}\right)$ obtained with the malfunctioning laser and the properly functioning laser, the value of $d_{220}(\mathrm{NR} 3)$ obtained in the INRIM-NMIJ joint effort using the malfunctioning laser mentioned above, and the value of $d_{220}$ (WS5C) also obtained with this laser, could be corrected and taken as input data. WS5C is an XROI manufactured by INRIM from a WASO 04 sample, but the value of $d_{220}$ (WS5C) obtained using the contaminated laser was also not included in Cavagnero et al. (2004a, 2004b). However, because of the somewhat erratic history of silicon lattice spacing measurements, the Task Group decided to use only data obtained with a laser known to be functioning properly.
Improvements in the INRIM XROI apparatus since the $1994 d_{220}\left(\mathrm{MO}^{*}\right)$ measurement of Basile et al. (1994) include (i) a new two-axis "tip-tilt" platform for the XROI that is electronically controlled to compensate for parasitic rotations and straightness error of the guiding system that moves the platform; (ii) imaging the x-ray interference pattern formed by the x-ray beam transmitted through the moving analyzer in such a way that detailed information concerning lattice distortion and analyzer pitch can be extracted on line from the analysis of the phases of the x-ray fringes; and (iii) an upgraded computer-aided system for combined interferometer displacement and control, x-ray and optical fringe scanning, signal digitization and sampling, environmental monitoring, and data analysis.

The values of $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ and $d_{220}\left(\mathrm{MO}^{*}\right)$ in Eqs. (299) and (300) are the means of tens of individual values, with each value being the average of about ten data points collected in 1 h measurement cycles during which the analyzer was translated back and forth by 300 optical orders. For the two crystals, respectively, the statistical uncertainties in parts in $10^{9}$ are 3.5 and 11.6, and the various corrections and their uncertainties are laser beam wavelength, $-0.8(4),-0.8(4)$; laser beam diffraction, 12.0(2.2), 12.0(2.2); laser beam alignment, 2.5(3.5), 2.5(3.5); Abbe error, $0.0(2.8), 0.0(3.7)$; trajectory error, $0.0(1.4), 0.0(3.6)$; analyzer temperature, 1.0(5.2), 1.0(7.9); and abberations, $0.0(5.0), 0.0(2.0)$. The total uncertainties are 9.6 and 15.7.

## 2. $\boldsymbol{d}_{220}$ difference measurements

To relate the lattice spacings of crystals used in various experiments, highly accurate measurements are made of the fractional difference $\left[d_{220}(X)\right.$ $-d_{220}($ ref $\left.)\right] / d_{220}($ ref $)$ of the $\{220\}$ lattice spacing of a sample of a single-crystal ingot $X$ and that of a reference crystal "ref." Both NIST and PTB have carried out such measurements, and the fractional differences from these two laboratories that we take as input data in the 2006 adjustment are given in the following two sections and are summarized in Table XXV. For details concerning these measurements, see CODATA-98 and CODATA02.

## a. NIST difference measurements

The following fractional difference involving a crystal denoted simply as N was obtained as part of the NIST effort to measure the wavelengths in meters of the $\mathrm{K} \alpha_{1}$ x-ray lines of $\mathrm{Cu}, \mathrm{Mo}$, and W; see Sec. XI.A:

$$
\begin{equation*}
\frac{d_{220}(\mathrm{~W} 17)-d_{220}(\mathrm{~N})}{d_{220}(\mathrm{~W} 17)}=7(22) \times 10^{-9} \tag{301}
\end{equation*}
$$

The following three fractional differences involving crystals from the four crystals denoted ILL, WASO 17, MO*, and NRLM3 were obtained as part of the NIST effort, discussed in Sec. VIII.C, to determine the relative atomic mass of the neutron $A_{\mathrm{r}}(\mathrm{n})$ (Kessler et al., 1999):

$$
\begin{align*}
& \frac{d_{220}(\mathrm{ILL})-d_{220}(\mathrm{~W} 17)}{d_{220}(\mathrm{ILL})}=-8(22) \times 10^{-9}  \tag{302}\\
& \frac{d_{220}(\mathrm{ILL})-d_{220}\left(\mathrm{MO}^{*}\right)}{d_{220}(\mathrm{ILL})}=86(27) \times 10^{-9}  \tag{303}\\
& \frac{d_{220}(\mathrm{ILL})-d_{220}(\mathrm{NR} 3)}{d_{220}(\mathrm{ILL})}=34(22) \times 10^{-9} \tag{304}
\end{align*}
$$

The following more recent NIST difference measurements, which we also take as input data in the 2006 adjustment, were provided by Kessler (2006) of NIST and are updates of the results reported by Hanke and Kessler (2005):

TABLE XXV. Summary of measurements of the relative $\{220\}$ lattice spacings of silicon crystals.

| Quantity | Value | Identification | Sec. and Eq. |
| :--- | :---: | :---: | :--- |
| $1-d_{220}(\mathrm{~W} 17) / d_{220}(\mathrm{ILL})$ | $-8(22) \times 10^{-9}$ | NIST-99 | VIII.A.2.a (302) |
| $\left.1-d_{220}(\mathrm{MO})^{*}\right) / d_{220}(\mathrm{ILL})$ | $86(27) \times 10^{-9}$ | NIST-99 | VIII.A.2.a (303) |
| $1-d_{220}(\mathrm{NR} 3) / d_{220}(\mathrm{ILL})$ | $34(22) \times 10^{-9}$ | NIST-99 | VIII.A.2.a (304) |
| $1-d_{220}(\mathrm{~N}) / d_{220}(\mathrm{~W} 17)$ | $7(22) \times 10^{-9}$ | NIST-97 | VIII.A.2.a (301) |
| $d_{220}(\mathrm{NR} 3) / d_{220}(\mathrm{~W} 04)-1$ | $-11(21) \times 10^{-9}$ | NIST-06 | VIII.A.2.a (305) |
| $d_{220}(\mathrm{NR} 4) / d_{220}(\mathrm{~W} 04)-1$ | $25(21) \times 10^{-9}$ | NIST-06 | VIII.A.2.a (306) |
| $d_{220}(\mathrm{~W} 17) / d_{220}(\mathrm{~W} 04)-1$ | $11(21) \times 10^{-9}$ | NIST-06 | VIII.A.2.a (307) |
| $d_{220}(\mathrm{~W} 4.2 \mathrm{a}) / d_{220}(\mathrm{~W} 04)-1$ | $-1(21) \times 10^{-9}$ | PTB-98 | VIII.A.2.b (308) |
| $d_{220}(\mathrm{~W} 17) / d_{220}(\mathrm{~W} 04)-1$ | $22(22) \times 10^{-9}$ | PTB-98 | VIII.A.2.b (309) |
| $d_{220}(\mathrm{MO} *) / d_{220}(\mathrm{~W} 04)-1$ | $-103(28) \times 10^{-9}$ | PTB-98 | VIII.A.2.b (310) |
| $d_{220}(\mathrm{NR} 3) / d_{220}(\mathrm{~W} 04)-1$ | $-23(21) \times 10^{-9}$ | PTB-98 | VIII.A.2.b (311) |
| $d_{220} / d_{220}(\mathrm{~W} 04)-1$ | $10(11) \times 10^{-9}$ | PTB-03 | VIII.A.2.b (312) |

$$
\begin{align*}
& \frac{d_{220}(\mathrm{NR} 3)-d_{220}(\mathrm{~W} 04)}{d_{220}(\mathrm{~W} 04)}=-11(21) \times 10^{-9}  \tag{305}\\
& \frac{d_{220}(\mathrm{NR} 4)-d_{220}(\mathrm{~W} 04)}{d_{220}(\mathrm{~W} 04)}=25(21) \times 10^{-9}  \tag{306}\\
& \frac{d_{220}(\mathrm{~W} 17)-d_{220}(\mathrm{~W} 04)}{d_{220}(\mathrm{~W} 04)}=11(21) \times 10^{-9} \tag{307}
\end{align*}
$$

The full designations of the two new crystals involved in these comparisons are WASO 04 and NRLM4. The measurements benefited significantly from the relocation of the NIST lattice comparator to a new laboratory where the temperature varied by only about 5 mK in several weeks compared to the previous laboratory where the temperature varied by about 40 mK in one day (Hanke and Kessler, 2005).

## b. PTB difference measurements

Results for the $\{220\}$ lattice-spacing fractional differences of various crystals that we also take as input data in the 2006 adjustment have been obtained at the PTB (Martin et al., 1998),

$$
\begin{align*}
& \frac{d_{220}(\mathrm{~W} 4.2 \mathrm{a})-d_{220}(\mathrm{~W} 04)}{d_{220}(\mathrm{~W} 04)}=-1(21) \times 10^{-9},  \tag{308}\\
& \frac{d_{220}(\mathrm{~W} 17)-d_{220}(\mathrm{~W} 04)}{d_{220}(\mathrm{~W} 04)}=22(22) \times 10^{-9},  \tag{309}\\
& \frac{d_{220}\left(\mathrm{MO}^{*}\right)-d_{220}(\mathrm{~W} 04)}{d_{220}(\mathrm{~W} 04)}=-103(28) \times 10^{-9},  \tag{310}\\
& \frac{d_{220}(\mathrm{NR} 3)-d_{220}(\mathrm{~W} 04)}{d_{220}(\mathrm{~W} 04)}=-23(21) \times 10^{-9} . \tag{311}
\end{align*}
$$

To relate $d_{220}$ (W04) to the $\{220\}$ lattice spacing $d_{220}$ of an ideal silicon crystal, we take as an input datum

$$
\begin{equation*}
\frac{d_{220}-d_{220}(\mathrm{~W} 04)}{d_{220}(\mathrm{~W} 04)}=10(11) \times 10^{-9} \tag{312}
\end{equation*}
$$

given by Becker et al. (2003), who obtained it by taking into account the known carbon, oxygen, and nitrogen impurities in WASO 04. However, following what was done in the 1998 and 2002 adjustments, we have included an additional component of uncertainty of 1 $\times 10^{-8}$ to account for the possibility that, even after correction for $\mathrm{C}, \mathrm{O}$, and N impurities, the crystal WASO 04, although well characterized as to its purity and crystallographic perfection, does not meet all criteria for an ideal crystal. Indeed, in general, we prefer to use experimentally measured fractional lattice spacing differences rather than differences implied by the $\mathrm{C}, \mathrm{O}$, and N impurity content of the crystals in order to avoid the need to assume that all crystals of interest meet these criteria.

In order to include this fractional difference in the 2002 adjustment, the quantity $d_{220}$ is also taken as an adjusted constant.

## B. Molar volume of silicon $V_{\mathrm{m}}(\mathrm{Si})$ and the Avogadro constant $N_{\text {A }}$

The definition of the molar volume of silicon $V_{\mathrm{m}}(\mathrm{Si})$ and its relationship to the Avogadro constant $N_{\mathrm{A}}$ and Planck constant $h$ as well as other constants is discussed in CODATA-98 and summarized in CODATA-02. In brief we have

$$
\begin{align*}
& m(\mathrm{Si})=\rho(\mathrm{Si}) \frac{a^{3}}{n}  \tag{313}\\
& V_{\mathrm{m}}(\mathrm{Si})=\frac{M(\mathrm{Si})}{\rho(\mathrm{Si})}=\frac{A_{\mathrm{r}}(\mathrm{Si}) M_{\mathrm{u}}}{\rho(\mathrm{Si})},  \tag{314}\\
& N_{\mathrm{A}}=\frac{V_{\mathrm{m}}(\mathrm{Si})}{a^{3} / n}=\frac{A_{\mathrm{r}}(\mathrm{Si}) M_{\mathrm{u}}}{\sqrt{8} d_{220}^{3} \rho(\mathrm{Si})} \tag{315}
\end{align*}
$$

$$
\begin{equation*}
V_{\mathrm{m}}(\mathrm{Si})=\frac{\sqrt{2} c M_{\mathrm{u}} A_{\mathrm{r}}(\mathrm{e}) \alpha^{2} d_{220}^{3}}{R_{\infty} h} \tag{316}
\end{equation*}
$$

which are to be understood in the context of an impurity-free, crystallographically perfect, ideal silicon crystal at the reference conditions $t_{90}=22.5^{\circ} \mathrm{C}$ and $p$ $=0$, and of isotopic composition in the range normally observed for crystals used in high-accuracy experiments. Thus $m(\mathrm{Si}), V_{\mathrm{m}}(\mathrm{Si}), M(\mathrm{Si})$, and $A_{\mathrm{r}}(\mathrm{Si})$ are the mean mass, mean molar volume, mean molar mass, and mean relative atomic mass of the silicon atoms in such a crystal, respectively, and $\rho(\mathrm{Si})$ is the crystal's macroscopic mass density. Equation (316) is the observational equation for a measured value of $V_{\mathrm{m}}(\mathrm{Si})$.

It follows from Eq. (314) that the experimental determination of $V_{\mathrm{m}}(\mathrm{Si})$ requires (i) measurement of the amount-of-substance ratios $n\left({ }^{29} \mathrm{Si}\right) / n\left({ }^{28} \mathrm{Si}\right)$ and $n\left({ }^{30} \mathrm{Si}\right) / n\left({ }^{28} \mathrm{Si}\right)$ of a nearly perfect silicon crystal-and hence amount-of-substance fractions $x\left({ }^{A} \mathrm{Si}\right)$-and then calculation of $A_{\mathrm{r}}(\mathrm{Si})$ from the well-known values of $A_{\mathrm{r}}{ }^{A} \mathrm{Si}$ ); and (ii) measurement of the macroscopic mass density $\rho(\mathrm{Si})$ of the crystal. Determining $N_{\mathrm{A}}$ from Eq. (315) by measuring $V_{\mathrm{m}}(\mathrm{Si})$ in this way and $d_{220}$ using x rays is called the x-ray-crystal-density (XRCD) method.
An extensive international effort has been underway since the early 1990s to determine $N_{\mathrm{A}}$ using this technique with the smallest possible uncertainty. The effort is being coordinated by the Working Group on the Avogadro Constant (WGAC) of the Consultative Committee for Mass and Related Quantities (CCM) of the CIPM. The WGAC, which has representatives from all major research groups working in areas relevant to the determination of $N_{\mathrm{A}}$, is currently chaired by P. Becker of PTB.

As discussed in CODATA-02, the value of $V_{\mathrm{m}}(\mathrm{Si})$ used as an input datum in the 2002 adjustment was provided to the CODATA Task Group by the WGAC and was a consensus value based on independent measurements of $\rho(\mathrm{Si})$ at NMIJ and PTB using a number of different silicon crystals, and measurements of their molar masses $M(\mathrm{Si})$ using isotopic mass spectrometry at the Institute for Reference Materials and Measurements (IRMM), European Commission, Geel, Belgium. This value, identified as $\mathrm{N} / \mathrm{P} / \mathrm{I}-03$ in recognition of the work done by researchers at NMIJ, PTB, and IRMM, is $V_{\mathrm{m}}(\mathrm{Si})=12.0588257(36) \times 10^{-6} \mathrm{~m}^{3} \mathrm{~mol}^{-1} \quad\left[3.0 \times 10^{-7}\right]$. Since then, the data used to obtain it were reanalyzed by the WGAC, resulting in the slightly revised value (Fujii et al., 2005)

$$
\begin{align*}
V_{\mathrm{m}}(\mathrm{Si})= & 12.0588254(34) \times 10^{-6} \mathrm{~m}^{3} \mathrm{~mol}^{-1} \\
& {\left[2.8 \times 10^{-7}\right], } \tag{317}
\end{align*}
$$

which we take as an input datum in the 2006 adjustment and identify as $\mathrm{N} / \mathrm{P} / \mathrm{I}-05$. The slight shift in value and reduction in uncertainty are due to the fact that the effect of nitrogen impurities in the silicon crystals used in the NMIJ measurements was taken into account in the reanalysis (Fujii et al., 2005). Note that the new value of
$A_{\mathrm{r}}\left({ }^{29} \mathrm{Si}\right)$ in Table IV does not change this result.
Based on Eq. (316) and the 2006 recommended values of $A_{\mathrm{r}}(\mathrm{e}), \alpha, d_{220}$, and $R_{\infty}$, the value of $h$ implied by this result is

$$
\begin{equation*}
h=6.6260745(19) \times 10^{-34} \mathrm{~J} \mathrm{~s} \quad\left[2.9 \times 10^{-7}\right] \tag{318}
\end{equation*}
$$

A comparison of this value of $h$ with those in Tables XXI and XXIII shows that it is generally not in good agreement with the most accurate of the other values.

In this regard, two relatively recent publications, the first describing work performed in China (Ding et al., 2005) and the second describing work performed in Switzerland (Reynolds et al., 2006), reported results that, if taken at face value, seem to call into question the uncertainty with which the molar mass of naturally occurring silicon is currently known. [See also Valkiers et al. (2005).] These results highlight the importance of the current WGAC project to measure $V_{\mathrm{m}}(\mathrm{Si})$ using highly enriched silicon crystals with $x\left({ }^{28} \mathrm{Si}\right)>0.99985$ (Becker et al., 2006), which should simplify the determination of the molar mass.

## C. Gamma-ray determination of the neutron relative atomic mass $\boldsymbol{A}_{\mathbf{r}}(\mathbf{n})$

Although the value of $A_{\mathrm{r}}(\mathrm{n})$ listed in Table II is a result of AME2003, it is not used in the 2006 adjustment. Instead, $A_{\mathrm{r}}(\mathrm{n})$ is obtained as discussed in this section in order to ensure that its recommended value is consistent with the best current information on the $\{220\}$ lattice spacing of silicon.

The value of $A_{\mathrm{r}}(\mathrm{n})$ can be obtained by measuring the wavelength of the $2.2 \mathrm{MeV} \gamma$ ray in the reaction $\mathrm{n}+\mathrm{p}$ $\rightarrow \mathrm{d}+\gamma$ in terms of the $d_{220}$ lattice spacing of a particular silicon crystal corrected to the commonly used reference conditions $t_{90}=22.5^{\circ} \mathrm{C}$ and $p=0$. The result for the wavelength-to-lattice spacing ratio, obtained from Bragg-angle measurements carried out in 1995 and 1998 using a flat crystal spectrometer of the GAMS4 diffraction facility at the high-flux reactor of the Institut Max von Laue-Paul Langevin (ILL), Grenoble, France, in a NIST and ILL collaboration, is (Kessler et al., 1999)

$$
\begin{equation*}
\frac{\lambda_{\text {meas }}}{d_{220}(\mathrm{ILL})}=0.00290430246(50) \quad\left[1.7 \times 10^{-7}\right] \tag{319}
\end{equation*}
$$

where $d_{220}$ (ILL) is the $\{220\}$ lattice spacing of the silicon crystals of the ILL GAMS4 spectrometer at $t_{90}$ $=22.5^{\circ} \mathrm{C}$ and $p=0$. Relativistic kinematics of the reaction yields

$$
\begin{equation*}
\frac{\lambda_{\text {meas }}}{d_{220}(\mathrm{ILL})}=\frac{\alpha^{2} A_{\mathrm{r}}(\mathrm{e})}{R_{\infty} d_{220}(\mathrm{ILL})} \frac{A_{\mathrm{r}}(\mathrm{n})+A_{\mathrm{r}}(\mathrm{p})}{\left[A_{\mathrm{r}}(\mathrm{n})+A_{\mathrm{r}}(\mathrm{p})\right]^{2}-A_{\mathrm{r}}^{2}(\mathrm{~d})}, \tag{320}
\end{equation*}
$$

where all seven quantities on the right-hand side are adjusted constants.

Dewey et al. (2006) and Rainville et al. (2005) reported determinations of the wavelengths of the gamma rays

TABLE XXVI. Summary of data related to the quotients $h / m_{\mathrm{n}} d_{220}(\mathrm{~W} 04), h / m(\mathrm{Cs})$, and $h / m(\mathrm{Rb})$, together with inferred values of $\alpha$.

| Quantity | Value | Relative standard <br> uncertainty $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :---: | :---: | :---: | :--- | :--- |
| $h / m_{\mathrm{n}} d_{220}($ W04 $)$ | $2060.267004(84) \mathrm{m} \mathrm{s}^{-1}$ | $4.1 \times 10^{-8}$ | PTB-99 | VIIII.D.1 (322) |
| $\alpha^{-1}$ | $137.0360077(28)$ | $2.1 \times 10^{-8}$ |  | VIII.D.1 (324) |
| $h / m(\mathrm{Cs})$ | $3.002369432(46) \times 10^{-9} \mathrm{~m}^{2} \mathrm{~s}^{-1}$ | $1.5 \times 10^{-8}$ | StanfU-02 | VIII.D.2 (329) |
| $\alpha^{-1}$ | $137.0360000(11)$ | $7.7 \times 10^{-9}$ |  | VIII.D.2 (331) |
| $h / m(\mathrm{Rb})$ | $4.591359287(61) \times 10^{-9} \mathrm{~m}^{2} \mathrm{~s}^{-1}$ | $1.3 \times 10^{-8}$ | LKB-06 | VIII.D.2 (332) |
| $\alpha^{-1}$ | $137.03599883(91)$ | $6.7 \times 10^{-9}$ |  | VIII.D.2 (334) |

emitted in the cascade from the neutron capture state to the ground state in the reactions $\mathrm{n}+{ }^{28} \mathrm{Si} \rightarrow{ }^{29} \mathrm{Si}+2 \gamma, \mathrm{n}$ $+{ }^{32} \mathrm{~S} \rightarrow{ }^{33} \mathrm{Si}+3 \gamma$, and $\mathrm{n}+{ }^{35} \mathrm{Cl} \rightarrow{ }^{36} \mathrm{Cl}+2 \gamma$. The gamma-ray energies are 3.5 MeV and 4.9 MeV for the Si reaction, 5.4 MeV, 2.4 MeV, and 0.8 MeV for the S reaction, and 6.1 MeV, 0.5 MeV , and 2.0 MeV for the Cl reaction. While these data together with the relevant relative atomic masses are potentially an additional source of information on the neutron relative atomic mass, the uncertainties are too large for this purpose; the inferred value of $A_{\mathrm{r}}(\mathrm{n})$ has an uncertainty nearly an order of magnitude larger than that obtained from Eq. (320). Instead, this work is viewed as the most accurate test of $E=m c^{2}$ to date (Rainville et al., 2005).

## D. Quotient of Planck constant and particle mass $h / m(X)$ and

 $\boldsymbol{\alpha}$The relation $R_{\infty}=\alpha^{2} m_{\mathrm{e}} c / 2 h$ leads to

$$
\begin{equation*}
\alpha=\left[\frac{2 R_{\infty}}{c} \frac{A_{\mathrm{r}}(X)}{A_{\mathrm{r}}(\mathrm{e})} \frac{h}{m(X)}\right]^{1 / 2}, \tag{321}
\end{equation*}
$$

where $A_{\mathrm{r}}(X)$ is the relative atomic mass of particle $X$ with mass $m(X)$ and $A_{\mathrm{r}}(\mathrm{e})$ is the relative atomic mass of the electron. Because $c$ is exactly known, $u_{\mathrm{r}}$ of $R_{\infty}$ and $A_{\mathrm{r}}(\mathrm{e})$ are less than $7 \times 10^{-12}$ and $5 \times 10^{-10}$, respectively, and $u_{\mathrm{r}}$ of $A_{\mathrm{r}}(X)$ for many particles and atoms is less than that of $A_{\mathrm{r}}(\mathrm{e})$, Eq. (321) can provide a value of $\alpha$ with a competitive uncertainty if $h / m(X)$ is determined with a sufficiently small uncertainty. Here we discuss the determination of $h / m(X)$ for the neutron n , the ${ }^{133} \mathrm{Cs}$ atom, and the ${ }^{87} \mathrm{Rb}$ atom. The results, including the inferred values of $\alpha$, are summarized in Table XXVI.

## 1. Quotient $h / m_{\mathrm{n}}$

The PTB determination of $h / m_{\mathrm{n}}$ was carried out at the ILL high-flux reactor. The de Broglie relation $p$ $=m_{\mathrm{n}} v=h / \lambda$ was used to determine $h / m_{\mathrm{n}}=\lambda v$ for the neutron by measuring both its de Broglie wavelength $\lambda$ and corresponding velocity $v$. More specifically, the de Broglie wavelength, $\lambda \approx 0.25 \mathrm{~nm}$, of slow neutrons was determined using back reflection from a silicon crystal, and the velocity, $v \approx 1600 \mathrm{~m} / \mathrm{s}$, of the neutrons was de-
termined by a special time-of-flight method. The final result of the experiment is (Krüger et al., 1999)

$$
\begin{align*}
\frac{h}{m_{\mathrm{n}} d_{220}(\mathrm{~W} 04)}= & 2060.267004(84) \mathrm{m} \mathrm{~s}^{-1} \\
& {\left[4.1 \times 10^{-8}\right] } \tag{322}
\end{align*}
$$

where $d_{220}(\mathrm{~W} 04)$ is the $\{220\}$ lattice spacing of the crystal WASO 04 at $t_{90}=22.5^{\circ} \mathrm{C}$ in vacuum. This result is correlated with the PTB fractional lattice-spacing differences given in Eqs. (308)-(311)—the correlation coefficients are about 0.2.

The equation for the PTB result, which follows from Eq. (321), is

$$
\begin{equation*}
\frac{h}{m_{\mathrm{n}} d_{220}(\mathrm{~W} 04)}=\frac{A_{\mathrm{r}}(\mathrm{e})}{A_{\mathrm{r}}(\mathrm{n})} \frac{c \alpha^{2}}{2 R_{\infty} d_{220}(\mathrm{~W} 04)} \tag{323}
\end{equation*}
$$

The value of $\alpha$ that can be inferred from this relation and the PTB value of $h / m_{\mathrm{n}} d_{220}$ (W04), the 2006 recommended values of $R_{\infty}, A_{\mathrm{r}}(\mathrm{e})$, and $A_{\mathrm{r}}(\mathrm{n})$, the NIST and PTB fractional lattice-spacing-differences in Table XXV, and the four XROI values of $d_{220}(X)$ in Table XXIV for crystals WASO 4.2a, NRLM3, and MO*, is

$$
\begin{equation*}
\alpha^{-1}=137.0360077(28) \quad\left[2.1 \times 10^{-8}\right] \tag{324}
\end{equation*}
$$

This value is included in Table XXVI as the first entry; it disagrees with the $\alpha$ values from the two other $h / m$ results.

It is also of interest to calculate the value of $d_{220} \mathrm{im}-$ plied by the PTB result for $h / m_{\mathrm{n}} d_{220}$ (W04). Based on Eq. (323), the 2006 recommended values of $R_{\infty}, A_{\mathrm{r}}(\mathrm{e})$, $A_{\mathrm{r}}(\mathrm{p}), A_{\mathrm{r}}(\mathrm{d}), \alpha$, the NIST and PTB fractional lattice-spacing-differences in Table XXV, and the value of $\lambda_{\text {meas }} / d_{220}$ (ILL) given in Eq. (319), we find

$$
\begin{equation*}
d_{220}=192015.5982(79) \mathrm{fm} \quad\left[4.1 \times 10^{-8}\right] \tag{325}
\end{equation*}
$$

This result is included in Table XXIV as the last entry; it agrees with the NMIJ value, but disagrees with the PTB and INRIM values.

## 2. Quotient $\boldsymbol{h} / \boldsymbol{m}\left({ }^{133} \mathrm{Cs}\right)$

The Stanford University atom interferometry experiment to measure the atomic recoil frequency shift of photons absorbed and emitted by ${ }^{133} \mathrm{Cs}$ atoms, $\Delta \nu_{\mathrm{Cs}}$, in
order to determine the quotient $h / m\left({ }^{133} \mathrm{Cs}\right)$ is described in CODATA-02. As discussed there, the expression applicable to the Stanford experiment is

$$
\begin{equation*}
\frac{h}{m\left({ }^{133} \mathrm{Cs}\right)}=\frac{c^{2} \Delta \nu_{\mathrm{Cs}}}{2 \nu_{\mathrm{eff}}^{2}}, \tag{326}
\end{equation*}
$$

where the frequency $\nu_{\text {eff }}$ corresponds to the sum of the energy difference between the ground-state hyperfine level with $F=3$ and the $6 \mathrm{P}_{1 / 2}$ state $F=3$ hyperfine level and the energy difference between the ground-state hyperfine level with $F=4$ and the same $6 \mathrm{P}_{1 / 2}$ hyperfine level. The result for $\Delta \nu_{\mathrm{Cs}} / 2$ obtained at Stanford is (Wicht et al., 2002)

$$
\begin{equation*}
\Delta \nu_{\mathrm{CS}} / 2=15006.27688(23) \mathrm{Hz} \quad\left[1.5 \times 10^{-8}\right] \tag{327}
\end{equation*}
$$

The Stanford effort included an extensive study of corrections due to possible systematic effects. The largest component of uncertainty by far contributing to the uncertainty of the final result for $\Delta \nu_{\mathrm{Cs}}, u_{\mathrm{r}}=14 \times 10^{-9}$ (Type B), arises from the possible deviation from 1 of the index of refraction of the dilute background gas of cold cesium atoms that move with the signal atoms. This component, estimated experimentally, places a lower limit on the relative uncertainty of the inferred value of $\alpha$ from Eq. (321) of $u_{\mathrm{r}}=7 \times 10^{-9}$. Without it, $u_{\mathrm{r}}$ of $\alpha$ would be about 3 to 4 parts in $10^{9}$.

In the 2002 adjustment, the value $\nu_{\text {eff }}$ $=670231933044(81) \mathrm{kHz}\left[1.2 \times 10^{-10}\right]$, based on the measured frequencies of ${ }^{133} \mathrm{Cs} \mathrm{D}_{1}$-line transitions reported by Udem et al. (1999), was used to obtain the ratio $h / m\left({ }^{133} \mathrm{Cs}\right)$ from the Stanford value of $\Delta \nu_{\mathrm{Cs}} / 2$. Recently, using a femtosecond laser frequency comb and a narrow-linewidth diode laser, and eliminating Doppler shift by orienting the laser beam perpendicular to the ${ }^{133} \mathrm{Cs}$ atomic beam to within $5 \mu \mathrm{rad}$, Gerginov et al. (2006) remeasured the frequencies of the required transitions and obtained a value of $\nu_{\text {eff }}$ that agrees with the value used in 2002 but which has a $u_{\mathrm{r}} 15$ times smaller,

$$
\begin{equation*}
\nu_{\mathrm{eff}}=670231932889.9(4.8) \mathrm{kHz} \quad\left[7.2 \times 10^{-12}\right] . \tag{328}
\end{equation*}
$$

Evaluation of Eq. (326) with this result for $\nu_{\text {eff }}$ and the value of $\Delta \nu_{\mathrm{Cs}} / 2$ in Eq. (327) yields

$$
\begin{align*}
\frac{h}{m\left({ }^{133} \mathrm{Cs}\right)}= & 3.002369432(46) \times 10^{-9} \mathrm{~m}^{2} \mathrm{~s}^{-1} \\
& {\left[1.5 \times 10^{-8}\right], } \tag{329}
\end{align*}
$$

which we take as an input datum in the 2006 adjustment. The observational equation for this datum is, from Eq. (321),

$$
\begin{equation*}
\frac{h}{m\left({ }^{(133} \mathrm{Cs}\right)}=\frac{A_{\mathrm{r}}(\mathrm{e})}{A_{\mathrm{r}}\left({ }^{133} \mathrm{Cs}\right)} \frac{c \alpha^{2}}{2 R_{\infty}} \tag{330}
\end{equation*}
$$

The value of $\alpha$ that may be inferred from this expression, the Stanford result for $h / m\left({ }^{133} \mathrm{Cs}\right)$ in Eq. (329), the 2006 recommended values of $R_{\infty}$ and $A_{\mathrm{r}}(\mathrm{e})$, and the ASME2003 value of $A_{\mathrm{r}}\left({ }^{133} \mathrm{Cs}\right)$ in Table II, the uncertain-
ties of which are inconsequential in this application, is

$$
\begin{equation*}
\alpha^{-1}=137.0360000(11) \quad\left[7.7 \times 10^{-9}\right] \tag{331}
\end{equation*}
$$

where the dominant component of uncertainty arises from the measured value of the recoil frequency shift, in particular, the component of uncertainty due to a possible index of refraction effect.

In this regard, we note that Campbell et al. (2005) have experimentally demonstrated the reality of one aspect of such an effect with a two-pulse light grating interferometer and shown that it can have a significant impact on precision measurements with atom interferometers. However, theoretical calculations based on simulations of the Stanford interferometer by Sarajlic et al. (2006), although incomplete, suggest that the experimentally based uncertainty component $u_{\mathrm{r}}=14 \times 10^{-9}$ assigned by Wicht et al. (2002) to account for this effect is reasonable. We also note that Wicht et al. (2005) have developed an improved theory of momentum transfer when localized atoms and localized optical fields interact. The details of such interactions are relevant to precision atom interferometry. When Wicht et al. (2005) applied the theory to the Stanford experiment to evaluate possible systematic errors arising from wave-front curvature and distortion, as well as the Gouy phase shift of Gaussian beams, they found that such errors do not limit the uncertainty of the value of $\alpha$ that can be obtained from the experiment at the level of a few parts in $10^{9}$, but will play an important role in future precision atominterferometer photon-recoil experiments to measure $\alpha$ with $u_{\mathrm{r}} \approx 5 \times 10^{-10}$, such as is currently underway at Stanford (Müller et al., 2006).

## 3. Quotient $h / m\left({ }^{87} \mathbf{R b}\right)$

In the LKB experiment (Cladé et al., 2006a, 2006b), the quotient $h / m\left({ }^{87} \mathrm{Rb}\right)$, and hence $\alpha$, is determined by accurately measuring the rubidium recoil velocity $v_{\mathrm{r}}$ $=\hbar k / m\left({ }^{87} \mathrm{Rb}\right)$ when a rubidium atom absorbs or emits a photon of wave vector $k=2 \pi / \lambda$, where $\lambda$ is the wavelength of the photon and $\nu=c / \lambda$ is its frequency. The measurements are based on Bloch oscillations in a vertical accelerated optical lattice.

The basic principle of the experiment is to precisely measure the variation of the atomic velocity induced by an accelerated standing wave using velocity selective Raman transitions between two ground-state hyperfine levels. A Raman $\pi$ pulse of two counterpropagating laser beams selects an initial narrow atomic velocity class. After the acceleration process, the final atomic velocity distribution is probed using a second Raman $\pi$ pulse of two counterpropagating laser beams.

The coherent acceleration of the rubidium atoms arises from a succession of stimulated two-photon transitions also using two counterpropagating laser beams. Each transition modifies the atomic velocity by $2 v_{\mathrm{r}}$ leaving the internal state unchanged. The Doppler shift is compensated by linearly sweeping the frequency difference of the two lasers. This acceleration can conveniently be interpreted in terms of Bloch oscillations in
the fundamental energy band of an optical lattice created by the standing wave, because the interference of the two laser beams leads to a periodic light shift of the atomic energy levels and hence to the atoms experiencing a periodic potential (Ben Dahan et al., 1996; Peik et al., 1997).

An atom's momentum evolves by steps of $2 \hbar k$, each one corresponding to a Bloch oscillation. After $N$ oscillations, the optical lattice is adiabatically released and the final velocity distribution, which is the initial distribution shifted by $2 N v_{\mathrm{r}}$, is measured. Due to the high efficiency of Bloch oscillations, for an acceleration of $2000 \mathrm{~m} \mathrm{~s}^{-2}$, 900 recoil momenta can be transferred to a rubidium atom in 3 ms with an efficiency of $99.97 \%$ per recoil.

The atoms are alternately accelerated upwards and downwards by reversing the direction of the Bloch acceleration laser beams, keeping the same delay between the selection and measurement Raman $\pi$ pulses. The resulting differential measurement is independent of gravity. In addition, the contribution of some systematic effects changes sign when the direction of the selection and measuring Raman beams is exchanged. Hence, for each up and down trajectory, the selection and measuring Raman beams are reversed and two velocity spectra are taken. The mean value of these two measurements is free from systematic errors to first order. Thus each determination of $h / m\left({ }^{87} \mathrm{Rb}\right)$ is obtained from four velocity spectra, each requiring 5 min of integration time, two from reversing the Raman beams when the acceleration is in the up direction and two when in the down direction. The Raman and Bloch lasers are stabilized by means of an ultrastable Fabry-Pérot cavity and the frequency of the cavity is checked several times during the 20 min measurement against a well-known two-photon transition in ${ }^{85} \mathrm{Rb}$.

Taking into account a $(-9.2 \pm 4) \times 10^{-10}$ correction to $h / m\left({ }^{87} \mathrm{Rb}\right)$ not included in the value reported by Cladé et al. (2006a) due to a nonzero force gradient arising from a difference in the radius of curvature of the up and down accelerating beams, the result derived from 72 measurements of $h / m\left({ }^{87} \mathrm{Rb}\right)$ acquired over 4 days, which we take as an input datum in the 2006 adjustment, is (Cladé et al., 2006b)

$$
\begin{align*}
\frac{h}{m\left({ }^{87} \mathrm{Rb}\right)}= & 4.591359287(61) \times 10^{-9} \mathrm{~m}^{2} \mathrm{~s}^{-1} \\
& {\left[1.3 \times 10^{-8}\right] } \tag{332}
\end{align*}
$$

where the quoted $u_{\mathrm{r}}$ contains a statistical component from the 72 measurements of $8.8 \times 10^{-9}$.

Cladé et al. (2006b) examined many possible sources of systematic error, both theoretically and experimentally, in this rather complex, sophisticated experiment in order to ensure that their result was correct. These include light shifts, index of refraction effects, and the effect of a gravity gradient, for which the corrections and their uncertainties are in fact comparatively small. More significant are the fractional corrections of $(16.8 \pm 8)$
$\times 10^{-9}$ for wave front curvature and Guoy phase, $(-13.2 \pm 4) \times 10^{-9}$ for second-order Zeeman effect, and $4(4) \times 10^{-9}$ for the alignment of the Raman and Bloch beams. The total of all corrections is given as $10.98(10.0) \times 10^{-9}$.

From Eq. (321), the observational equation for the LKB value of $h / m\left({ }^{87} \mathrm{Rb}\right)$ in Eq. (332) is

$$
\begin{equation*}
\frac{h}{m\left({ }^{87} \mathrm{Rb}\right)}=\frac{A_{\mathrm{r}}(\mathrm{e})}{A_{\mathrm{r}}\left({ }^{87} \mathrm{Rb}\right)} \frac{c \alpha^{2}}{2 R_{\infty}} . \tag{333}
\end{equation*}
$$

Evaluation of this expression with the LKB result and the 2006 recommended values of $R_{\infty}$ and $A_{\mathrm{r}}(\mathrm{e})$, and the value of $\left.A_{\mathrm{r}}{ }^{(87} \mathrm{Rb}\right)$ resulting from the final least-squares adjustment on which the 2006 recommended values are based, all of whose uncertainties are negligible in this context, yields

$$
\begin{equation*}
\alpha^{-1}=137.03599883(91) \quad\left[6.7 \times 10^{-9}\right], \tag{334}
\end{equation*}
$$

which is included in Table XXVI. The uncertainty of this value of $\alpha^{-1}$ is smaller than the uncertainty of any other value except those in Table XIV deduced from the measurement of $a_{\mathrm{e}}$, exceeding the smallest uncertainty of the two values of $\alpha^{-1}\left[a_{\mathrm{e}}\right]$ in that table by a factor of 10 .

## IX. THERMAL PHYSICAL QUANTITIES

The following sections discuss the molar gas constant, Boltzmann constant, and Stefan-Boltzmann constantconstants associated with phenomena in the fields of thermodynamics and/or statistical mechanics.

## A. Molar gas constant $\boldsymbol{R}$

The square of the speed of sound $c_{\mathrm{a}}^{2}(p, T)$ of a real gas at pressure $p$ and thermodynamic temperature $T$ can be written as (Colclough, 1973)

$$
\begin{equation*}
c_{\mathrm{a}}^{2}(p, T)=A_{0}(T)+A_{1}(T) p+A_{2}(T) p^{2}+A_{3}(T) p^{3}+\cdots \tag{335}
\end{equation*}
$$

where $A_{1}(T)$ is the first acoustic virial coefficient, $A_{2}(T)$ is the second, etc. In the limit $p \rightarrow 0$, Eq. (335) yields

$$
\begin{equation*}
c_{\mathrm{a}}^{2}(0, T)=A_{0}(T)=\frac{\gamma_{0} R T}{A_{\mathrm{r}}(X) M_{\mathrm{u}}} \tag{336}
\end{equation*}
$$

where the expression on the right-hand side is the square of the speed of sound for an unbounded ideal gas, and $\gamma_{0}=c_{p} / c_{V}$ is the ratio of the specific-heat capacity of the gas at constant pressure to that at constant volume, $A_{\mathrm{r}}(X)$ is the relative atomic mass of the atoms or molecules of the gas, and $M_{\mathrm{u}}=10^{-3} \mathrm{~kg} \mathrm{~mol}^{-1}$. For a monatomic ideal gas, $\gamma_{0}=5 / 3$.

The 2006 recommended value of $R$, like the 2002 and 1998 values, is based on measurements of the speed of sound in argon carried out in two independent experiments, one done in the 1970s at NPL and the other done in the 1980 s at NIST. Values of $c_{\mathrm{a}}^{2}\left(p, T_{\mathrm{TPW}}\right)$, where $T_{\text {TPW }}=273.16 \mathrm{~K}$ is the triple point of water, were ob-
tained at various pressures and extrapolated to $p=0$ in order to determine $A_{0}\left(T_{\mathrm{TPW}}\right)=c_{\mathrm{a}}^{2}\left(0, T_{\mathrm{TPW}}\right)$, and hence $R$, from

$$
\begin{equation*}
R=\frac{c_{\mathrm{a}}^{2}\left(0, T_{\mathrm{TPW}}\right) A_{\mathrm{r}}(\mathrm{Ar}) M_{\mathrm{u}}}{\gamma_{0} T_{\mathrm{TPW}}} \tag{337}
\end{equation*}
$$

which follows from Eq. (336).
Because the work of both NIST and NPL is reviewed in CODATA-98 and CODATA-02 and nothing has occurred in the past 4 years that would change the values of $R$ implied by their reported values of $c_{\mathrm{a}}^{2}\left(0, T_{\mathrm{TPW}}\right)$, we give only a summary here. Changes in these values due to the new values of $A_{\mathrm{r}}\left({ }^{A} \mathrm{Ar}\right)$ resulting from the 2003 atomic mass evaluation as given in Table II, or the new IUPAC compilation of atomic weights of the elements given by Wieser (2006), are negligible.

Since $R$ cannot be expressed as a function of any other of the 2006 adjusted constants, $R$ itself is taken as an adjusted constant for the NIST and NPL measurements.

## 1. NIST: Speed of sound in argon

In the NIST experiment of Moldover et al. (1988), a spherical acoustic resonator at a temperature $T=T_{\text {TPW }}$ filled with argon was used to determine $c_{\mathrm{a}}^{2}\left(p, T_{\text {TPW }}\right)$. The final NIST result for the molar gas constant is

$$
\begin{equation*}
R=8.314471(15) \mathrm{J} \mathrm{~mol}^{-1} \mathrm{~K}^{-1} \quad\left[1.8 \times 10^{-6}\right] \tag{338}
\end{equation*}
$$

The mercury employed to determine the volume of the spherical resonator was traceable to the mercury whose density was measured by Cook (1961) [see also Cook and Stone (1957)]. The mercury employed in the NMI Hg electrometer determination of $K_{J}$ (see VII.C.1) was also traceable to the same mercury. Consequently, the NIST value of $R$ and the NMI value of $K_{\mathrm{J}}$ are correlated with the non-negligible correlation coefficient 0.068 .

## 2. NPL: Speed of sound in argon

In contrast to the dimensionally fixed resonator used in the NIST experiment, the NPL experiment employed a variable path length fixed-frequency cylindrical acoustic interferometer to measure $c_{\mathrm{a}}^{2}\left(p, T_{\mathrm{TPW}}\right)$. The final NPL result for the molar gas constant is (Colclough et al., 1979)

$$
\begin{equation*}
R=8.314504(70) \mathrm{J} \mathrm{~mol}^{-1} \mathrm{~K}^{-1} \quad\left[8.4 \times 10^{-6}\right] \tag{339}
\end{equation*}
$$

Although both the NIST and NPL values of $R$ are based on the same values of $A_{\mathrm{r}}\left({ }^{40} \mathrm{Ar}\right), A_{\mathrm{r}}\left({ }^{38} \mathrm{Ar}\right)$, and $A_{\mathrm{r}}\left({ }^{36} \mathrm{Ar}\right)$, the uncertainties of these relative atomic masses are sufficiently small that the covariance of the two values of $R$ is negligible.

## 3. Other values

The most important historical values of $R$ have been reviewed by Colclough (1984) [see also Quinn et al. (1976) and CODATA-98]. However, because of the large
uncertainties of these early values, they were not considered for use in the 1986, 1998, or 2002 CODATA adjustments, and we do not consider them for the 2006 adjustment as well.

Also because of its noncompetitive uncertainty ( $u_{\mathrm{r}}$ $=36 \times 10^{-6}$ ), we exclude from consideration in the 2006 adjustment, as in the 2002 adjustment, the value of $R$ obtained from measurements of the speed of sound in argon reported by He and Liu (2002) at the Xián Jiaotong University, Xián, PRC.

## B. Boltzmann constant $\boldsymbol{k}$

The Boltzmann constant is related to the molar gas constant $R$ and other adjusted constants by

$$
\begin{equation*}
k=\frac{2 R_{\infty} h}{c A_{\mathrm{r}}(\mathrm{e}) M_{\mathrm{u}} \alpha^{2}} R=\frac{R}{N_{\mathrm{A}}} . \tag{340}
\end{equation*}
$$

No competitive directly measured value of $k$ was available for the 1998 or 2002 adjustments, and the situation remains unchanged for the present adjustment. Thus, the 2006 recommended value with $u_{\mathrm{r}}=1.7 \times 10^{-6}$ is obtained from this relation, as were the 1998 and 2002 recommended values. However, a number of experiments are currently underway that might lead to competitive values of $k$ (or $R$ ) in the future; see Fellmuth et al. (2006) for a recent review.

Indeed, one such experiment underway at the PTB based on dielectric constant gas thermometry (DCGT) was discussed in both CODATA-98 and CODATA-02, but no experimental result for $A_{\epsilon} / R$, where $A_{\epsilon}$ is the molar polarizability of the ${ }^{4} \mathrm{He}$ atom, other than that considered in these two reports, has been published by the PTB group [see also Fellmuth et al. (2006) and Luther et al. (1996)]. However, the relative uncertainty in the theoretical value of the static electric dipole polarizability of the ground state of the ${ }^{4} \mathrm{He}$ atom, which is required to calculate $k$ from $A_{\epsilon} / R$, has been lowered by more than a factor of 10 to below $2 \times 10^{-7}$ ( Lach et al., 2004). Nevertheless, the change in its value is negligible at the level of uncertainty of the PTB result for $A_{\epsilon} / R$; hence, the value $k=1.38065(4) \times 10^{-23} \mathrm{~J} \mathrm{~K}^{-1}\left[30 \times 10^{-6}\right]$ from the PTB experiment given in CODATA-02 is unchanged.

In addition, preliminary results from two other ongoing experiments, the first at NIST by Schmidt et al. (2007) and the second at the University of Paris by Daussy et al. (2007), have recently been published.

Schmidt et al. (2007) report $R$ $=8.314487(76) \mathrm{J} \mathrm{mol} \mathrm{K}^{-1} \quad\left[9.1 \times 10^{-6}\right]$, obtained from measurements of the index of refraction $n(p, T)$ of ${ }^{4} \mathrm{He}$ gas as a function of $p$ and $T$ by measuring the difference in the resonant frequencies of a quasispherical microwave resonator when filled with ${ }^{4} \mathrm{He}$ at a given pressure and when evacuated (that is, at $p=0$ ). This experiment has some similarities to the PTB DCGT experiment in that it determines the quantity $A_{\epsilon} / R$ and hence $k$. However, in DCGT one measures the difference in capaci-
tance of a capacitor when filled with ${ }^{4} \mathrm{He}$ at a given pressure and at $p=0$, and hence one determines the dielectric constant of the ${ }^{4} \mathrm{He}$ gas rather than its index of refraction. Because ${ }^{4} \mathrm{He}$ is slightly diamagnetic, this means that to obtain $A_{\epsilon} / R$ in the NIST experiment, a value for $A_{\mu} / R$ is required, where $A_{\mu}=4 \pi \chi_{0} / 3$ and $\chi_{0}$ is the diamagnetic susceptibility of a ${ }^{4} \mathrm{He}$ atom.

Daussy et al. (2007) report $k=1.38065(26)$ $\times 10^{-23} \mathrm{~J} \mathrm{~K}^{-1}\left[190 \times 10^{-6}\right]$, obtained from measurements as a function of pressure of the Doppler profile at $T$ $=273.15 \mathrm{~K}$ (the ice point) of a well-isolated rovibrational line in the $\nu_{2}$ band of the ammonium molecule, ${ }^{14} \mathrm{NH}_{3}$, and extrapolation to $p=0$. The experiment actually measures $R=k N_{\mathrm{A}}$, because the mass of the ammonium molecule in kilograms is required but can only be obtained with the requisite accuracy from the molar masses of ${ }^{14} \mathrm{~N}$ and ${ }^{1} \mathrm{H}$, thereby introducing $N_{\mathrm{A}}$.

It is encouraging that the preliminary values of $k$ and $R$ resulting from these three experiments are consistent with the 2006 recommended values.

## C. Stefan-Boltzmann constant $\sigma$

The Stefan-Boltzmann constant is related to $c, h$, and the Boltzmann constant $k$ by

$$
\begin{equation*}
\sigma=\frac{2 \pi^{5} k^{4}}{15 h^{3} c^{2}} \tag{341}
\end{equation*}
$$

which, with the aid of Eq. (340), can be expressed in terms of the molar gas constant and other adjusted constants as

$$
\begin{equation*}
\sigma=\frac{32 \pi^{5} h}{15 c^{6}}\left(\frac{R_{\infty} R}{A_{\mathrm{r}}(\mathrm{e}) M_{\mathrm{u}} \alpha^{2}}\right)^{4} \tag{342}
\end{equation*}
$$

No competitive directly measured value of $\sigma$ was available for the 1998 or 2002 adjustment, and the situation remains unchanged for the 2006 adjustment. Thus, the 2006 recommended value with $u_{\mathrm{r}}=7.0 \times 10^{-6}$ is obtained from this relation, as were the 1998 and 2002 recommended values. For a concise summary of experiments that might provide a competiive value of $\sigma$, see the review by Fellmuth et al. (2006).

## X. NEWTONIAN CONSTANT OF GRAVITATION $G$

Because there is no known quantitative theoretical relationship between the Newtonian constant of gravitation $G$ and other fundamental constants, and because currently available experimental values of $G$ are independent of all other data relevant to the 2006 adjustment, these experimental values contribute only to the determination of the 2006 recommended value of $G$ and can be considered independently from the other data.
The historic difficulty of determining $G$, as demonstrated by the inconsistencies among different measurements, is described in CODATA-86, CODATA-98, and CODATA-02. Although no new competitive independent result for $G$ has become available in the last

4 years, adjustments to two existing results considered in 2002 have been made by researchers involved in the original work. One of the two results that has changed is from the Huazhong University of Science and Technology (HUST) and is now identified as HUST-05; the other is from the University of Zurich (UZur) and is now identified as UZur-06. These revised results are discussed below.

Table XXVII summarizes the various values of $G$ considered here, which are the same as in 2002 with the exception of these two revised results, and Fig. 2 compares them graphically. For reference purposes, both the table and figure include the 2002 and 2006 CODATA recommended values. The result now identified as TR\&D-96 was previously identified as TR\&D-98. The change is because a 1996 reference (Karagioz and Izmailov, 1996) was found that reports the same result as does the 1998 reference (Karagioz et al., 1998).

For simplicity, we write $G$ as a numerical factor multiplying $G_{0}$, where

$$
\begin{equation*}
G_{0}=10^{-11} \mathrm{~m}^{3} \mathrm{~kg}^{-1} \mathrm{~s}^{-2} \tag{343}
\end{equation*}
$$

## A. Updated values

## 1. Huazhong University of Science and Technology

The HUST group, which determines $G$ by the time-of-swing method using a high- $Q$ torsion pendulum with two horizontal, 6.25 kg stainless steel cylindrical source masses labeled A and B positioned on either side of the test mass, has reported a fractional correction of +360 $\times 10^{-6}$ to their original result given by Luo et al. (1999). It arises in part from recently discovered density inhomogeneities in the source masses, the result of which is a displacement of the center of mass of each source mass from its geometrical center (Hu et al., 2005). Using a "weighbridge" with a commercial electronic balance-a method developed by Davis (1995) to locate the center of mass of a test object with micrometers precision- Hu et al. (2005) found that the axial eccentricities $e_{\mathrm{A}}$ and $e_{\mathrm{B}}$ of the two source masses were $(10.3 \pm 2.6) \mu \mathrm{m}$ and ( $6.3 \pm 3.7) \mu \mathrm{m}$, with the result that the equivalent displacements between the test and source masses are larger than the values used by Luo et al. (1999). Assuming a linear axial density distribution, the calculated fractional correction to the previous result is $+210 \times 10^{-6}$ with an additional component of relative standard uncertainty of $78 \times 10^{-6}$ due to the uncertainties of the eccentricities.

The remaining $150 \times 10^{-6}$ portion of the $360 \times 10^{-6}$ fractional correction has also been discussed by Hu et al. (2005) and arises as follows. In the HUST experiment, $G$ is determined by comparing the period of the torsion pendulum with and without the source masses. When the source masses are removed, they are replaced by air. Since the masses of the source masses used by Luo et al. (1999) are the vacuum masses, a correction for the air, first suggested by R. S. Davis and T. J. Quinn of the BIPM, is required. This correction was privately com-

TABLE XXVII. Summary of the measurements of the Newtonian constant of gravitation relevant to the 2006 adjustment together with the 2002 and 2006 CODATA recommended values.

| Item | Source | Identification ${ }^{\text {a }}$ | Method | $\frac{10^{11} G}{\mathrm{~m}^{3} \mathrm{~kg}^{-1} \mathrm{~s}^{-2}}$ | Rel. stand. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | 2002 CODATA adjustment | CODATA-02 |  | $6.6742(10)$ | $1.5 \times 10^{-4}$ |
| $a$. | Karagioz and Izmailov, 1996 | TR\&D-96 | Fiber torsion balance, dynamic mode | 6.6729(5) | $7.5 \times 10^{-5}$ |
| $b$. | Bagley and Luther, 1997 | LANL-97 | Fiber torsion balance, dynamic mode | 6.6740(7) | $1.0 \times 10^{-4}$ |
| c. | Gundlach and Merkowitz, 2000, 2002, | UWash-00 | Fiber torsion balance, dynamic compensation | 6.674 255(92) | $1.4 \times 10^{-5}$ |
| d. | Quinn et al., 2001 | BIPM-01 | Strip torsion balance, compensation mode, static deflection | 6.675 59(27) | $4.0 \times 10^{-5}$ |
| $e$. | Kleinvoß, 2002, Kleinvoß et al., 2002 | UWup-02 | Suspended body, displacement | 6.674 22(98) | $1.5 \times 10^{-4}$ |
| $f$. | Armstrong and Fitzgerald, 2003 | MSL-03 | Strip torsion balance, compensation mode | 6.673 87(27) | $4.0 \times 10^{-5}$ |
| $g$. | Hu et al., 2005 | HUST-05 | Fiber torsion balance, dynamic mode | 6.6723(9) | $1.3 \times 10^{-4}$ |
| $h$. | Schlamminger et al., 2006 | UZur-06 | Stationary body, weight change | 6.674 25(12) | $1.9 \times 10^{-5}$ |
|  | 2006 CODATA adjustment | CODATA-06 |  | 6.674 28(67) | $1.0 \times 10^{-4}$ |

[^1]municated to the Task Group by the HUST researchers in 2003 and included in the HUST value of $G$ used in the 2002 adjustment.

The HUST revised value of $G$, including the additional component of uncertainty due to the measure-


FIG. 2. Values of the Newtonian constant of gravitation $G$.
ment of the eccentricities $e_{\mathrm{A}}$ and $e_{\mathrm{B}}$, is item $g$ in Table XXVII.

## 2. University of Zurich

The University of Zurich result for $G$ discussed in CODATA-02 and used in the 2002 adjustment, $G=6.67407(22) G_{0}\left[3.3 \times 10^{-5}\right]$, was reported by Schlamminger et al. (2002). It was based on the weighted mean of three highly consistent values obtained from three series of measurements carried out at the Paul Scherrer Institute (PSI), Villigen, Switzerland, in 2001 and 2002 and denoted Cu , Ta I, and Ta II. The designation Cu means that the test masses were gold plated copper, and the designation Ta means that they were tantalum. Following the publication of Schlamminger et al. (2002), an extensive reanalysis of the original data was carried out by these authors together with other University of Zurich researchers, the result being the value of $G$ in Table XXVII, item $h$, as given in the detailed final report on the experiment (Schlamminger et al., 2006).

In the University of Zurich approach to determining $G$, a modified commercial single-pan balance is used to measure the change in the difference in weight of two cylindrical test masses when the relative position of two source masses is changed. The quantity measured is the
$800 \mu \mathrm{~g}$ difference signal obtained at many different working points in the balance calibration range using two sets of 16 individual wire weights, allowing an in situ measurement of the balance nonlinearity over the entire 0.2 g balance calibration interval. A more rigorous analysis using a fitting method with Legendre polynomials has now allowed the relative standard uncertainty contribution to $G$ from balance nonlinearity to be reduced from $18 \times 10^{-6}$ to $6.1 \times 10^{-6}$ based on the Cu testmass data. Various problems with the mass handler for the wire weights that did not allow application of the Legendre polynomial fitting procedure occurred during the Ta test-mass measurements, resulting in large systematic errors. Therefore, the researchers decided to include only the Cu data in their final analysis (Schlamminger et al., 2006).

Each source mass consisted of a cylindrical tank filled with $7.5 \times 10^{3} \mathrm{~kg}$ of mercury. Since the mercury represented approximately $94 \%$ of the total mass, special care was taken in determining its mass and density. These measurements were used further to obtain more accurate values for the key tank dimensions and Hg mass. This was done by minimizing a $\chi^{2}$ function that depended on the tank dimensions and the Hg mass and density, and using the dependence of the density on these dimensions and the Hg mass as a constraint. Calculation of the mass integration constant with these improved values reduced the $u_{\mathrm{r}}$ of this critical quantity from $20.6 \times 10^{-6}$ to $6.7 \times 10^{-6}$.

Although the analysis of Schlamminger et al. (2002) assumed a linear temporal drift of the balance zero point, a careful examination by Schlamminger et al. (2006) found that the drift was significantly nonlinear and was influenced by the previous load history of the balance. A series of Legendre polynomials and a sawtooth function, respectively, were therefore used to describe the slow and rapid variations of the observed balance zero point with time.

The 2002 value of $G$ obtained from the Cu data was $6.67403 G_{0}$, consistent with the Ta I and Ta II values of $6.67409 G_{0}$ and $6.67410 G_{0}$ (Schlamminger et al., 2002), whereas the value from the present Cu data analysis is $6.67425(12) G_{0}$, with the $3.3 \times 10^{-5}$ fractional increase due primarily to the application of the nonlinear zeropoint drift correction. A minor contributor to the difference is the inclusion of the very first Cu data set that was omitted in the 2002 analysis due to a large start-up zeropoint drift that is now correctable with the new Legendre polynomial-sawtooth function analysis technique, and the exclusion of a data set that had a temperature stabilization system failure that went undetected by the old data analysis method (Schlamminger, 2007).

## B. Determination of 2006 recommended value of $G$

The overall agreement of the eight values of $G$ in Table XXVII (items $a$ to $h$ ) has improved somewhat since the 2002 adjustment, but the situation is still far from satisfactory. Their weighted mean is $G$
$=6.674275(68) G_{0}$ with $\chi^{2}=38.6$ for degrees of freedom $\nu=N-M=8-1=7$, Birge ratio $R_{\mathrm{B}}=\sqrt{\chi^{2} / \nu}=2.35$, and normalized residuals $r_{i}$ of $-2.75,-0.39,-0.22,4.87,-0.56$, $-1.50,-2.19$, and -0.19 , respectively (see Appendix E of CODATA-98). The BIPM-01 value with $r_{i}=4.87$ is clearly the most problematic. For comparison, the 2002 weighted mean was $G=6.674232(75) G_{0}$ with $\chi^{2}=57.7$ for $\nu=7$ and $R_{\mathrm{B}}=2.87$.

If the BIPM value is deleted, the weighted mean is reduced by 1.3 standard uncertainties to $G$ $=6.674187(70) G_{0}$, and $\chi^{2}=13.3, \nu=6$, and $R_{\mathrm{B}}=1.49$. In this case, the two remaining data with significant normalized residuals are the TR\&D-96 and HUST-05 results with $r_{i}=-2.57$ and -2.10 , respectively. If these two data, which agree with each other, are deleted, the weighted mean is $G=6.674225(71) G_{0}$ with $\chi^{2}=2.0$, $\nu$ $=4, R_{\mathrm{B}}=0.70$, and with all normalized residuals less than 1 except $r_{i}=-1.31$ for datum MSL-03. Finally, if the UWash-00 and UZur-06 data, which have the smallest assigned uncertainties of the initial eight values and which are in excellent agreement with each other, are deleted from the initial group of eight data, the weighted mean of the remaining six data is $G=6.674384$ (167) $G_{0}$ with $\chi^{2}=38.1, \nu=6$, and $R_{\mathrm{B}}=2.76$. The normalized residuals for these six data, TR\&D-96, LANL-97, BIPM01, UWup-02, MSL-03, and HUST-05, are $-2.97,-0.55$, $4.46,-0.17,-1.91$, and -2.32 , respectively.

Finally, if the uncertainties of each of the eight values of $G$ are multiplied by the Birge ratio associated with their weighted mean, $R_{\mathrm{B}}=2.35$, so that $\chi^{2}$ of their weighted mean becomes equal to its expected value of $\nu=7$ and $R_{\mathrm{B}}=1$, the normalized residual of the datum BIPM-01 would still be larger than 2 .

Based on the results of the above calculations, the historical difficulty of determining $G$, the fact that all eight values of $G$ in Table XXVII are credible, and that the two results with the smallest uncertainties, UWash-00 and UZur-06, are highly consistent with one another, the Task Group decided to take as the 2006 CODATA recommended value of $G$ the weighted mean of all data, but with an uncertainty of $0.00067 G_{0}$, corresponding to $u_{\mathrm{r}}=1.0 \times 10^{-4}$,

$$
\begin{equation*}
G=6.67428(67) \times 10^{-11} \mathrm{~m}^{3} \mathrm{~kg}^{-1} \mathrm{~s}^{-2} \quad\left[1.0 \times 10^{-4}\right] \tag{344}
\end{equation*}
$$

This value exceeds the 2002 recommended value by the fractional amount $1.2 \times 10^{-5}$, which is less than one-tenth of the uncertainty $u_{\mathrm{r}}=1.5 \times 10^{-4}$ of the 2002 value. Further, the uncertainty of the 2006 value, $u_{\mathrm{r}}=1.0 \times 10^{-4}$, is two-thirds that of the 2002 value.

In assigning this uncertainty to the 2006 recommended value of $G$, the Task Group recognized that if the uncertainty was smaller than really justified by the data, taking into account the history of measurements of $G$, it might discourage the initiation of new research efforts to determine $G$, if not the continuation of some of the research efforts already underway. Such efforts need to be encouraged in order to provide a more solid and
redundant data set upon which to base future recommended values. On the other hand, if the uncertainty were too large, for example, if the uncertainty of the 2002 recommended value had been retained for the 2006 value, then the recommended value would not have reflected the fact that we now have two data that are in excellent agreement, have $u_{\mathrm{r}}$ less than $2 \times 10^{-5}$, and are the two most accurate values available.

## C. Prospective values

New techniques to measure $G$ using atom interferometry are currently under development in at least two laboratories-the Università de Firenze in Italy and Stanford University in the United States. This comes as no surprise since atom interferometry is also being developed to measure the local acceleration due to gravity $g$ (see Sec. II). Recent proof-of-principle experiments combine two vertically separated atomic clouds forming an atom-interferometer-gravity gradiometer that measures the change in the gravity gradient when a well characterized source mass is displaced. Measuring the change in the gravity gradient allows the rejection of many possible systematic errors. Bertoldi et al. (2006) at the Università de Firenze used a Rb fountain and a fast launch juggling sequence of two atomic clouds to measure $G$ to $1 \%$, obtaining the value $6.64(6) G_{0}$; they hope to reach a final uncertainty of 1 part in $10^{4}$. Fixler et al. (2007) at Stanford used two separate Cs atom interferometer gravimeters to measure $G$ and obtained the value $6.693(34) G_{0}$. The two largest uncertainties from systematic effects were the determination of the initial atom velocity and position. The Stanford researchers also hope to achieve a final uncertainty of 1 part in $10^{4}$. Although neither of these results is significant for the current analysis of $G$, future results could be of considerable interest.

## XI. X-RAY AND ELECTROWEAK QUANTITIES

## A. X-ray units

Historically, units that have been used to express the wavelengths of $x$-ray lines are the copper $K \alpha_{1} x$ unit, symbol $\mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)$, the molybdenum $\mathrm{K} \alpha_{1} \mathrm{x}$ unit, symbol $\mathrm{xu}\left(\operatorname{MoK} \alpha_{1}\right)$, and the ångström star, symbol $\AA$ *. They are defined by assigning an exact, conventional value to the wavelength of the $\mathrm{CuK} \alpha_{1}, \operatorname{MoK} \alpha_{1}$, and $\mathrm{WK} \alpha_{1}$ x-ray lines when each is expressed in its corresponding unit,

$$
\begin{align*}
& \lambda\left(\mathrm{CuK} \alpha_{1}\right)=1537.400 \mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right),  \tag{345}\\
& \lambda\left(\mathrm{MoK} \alpha_{1}\right)=707.831 \mathrm{xu}\left(\mathrm{MoK} \alpha_{1}\right),  \tag{346}\\
& \lambda\left(\mathrm{WK} \alpha_{1}\right)=0.2090100 \AA^{*} . \tag{347}
\end{align*}
$$

The experimental work that determines the best values of these three units was reviewed in CODATA-98, and the relevant data may be summarized as follows:

$$
\begin{align*}
& \frac{\lambda\left(\mathrm{CuK} \alpha_{1}\right)}{d_{220}(\mathrm{~W} 4.2 \mathrm{a})}=0.80232711(24) \quad\left[3.0 \times 10^{-7}\right]  \tag{348}\\
& \frac{\lambda\left(\mathrm{WK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})}=0.108852175(98) \quad\left[9.0 \times 10^{-7}\right]  \tag{349}\\
& \frac{\lambda\left(\mathrm{MoK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})}=0.36940604(19) \quad\left[5.3 \times 10^{-7}\right] \tag{350}
\end{align*}
$$

$$
\begin{equation*}
\frac{\lambda\left(\mathrm{CuK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})}=0.80232804(77) \quad\left[9.6 \times 10^{-7}\right] \tag{351}
\end{equation*}
$$

where $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ and $d_{220}(\mathrm{~N})$ denote the $\{220\}$ lattice spacings, at the standard reference conditions $p=0$ and $t_{90}=22.5^{\circ} \mathrm{C}$, of particular silicon crystals used in the measurements. The result in Eq. (348) is from a collaboration between researchers from Friedrich-Schiller University (FSU), Jena, Germany and the PTB (Härtwig et al., 1991). The lattice spacing $d_{220}(\mathrm{~N})$ is related to crystals of known lattice spacing through Eq. (301).

In order to obtain best values in the least-squares sense for $\mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right), \mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)$, and $\AA^{*}$, we take these units to be adjusted constants. Thus, the observational equations for the data of Eqs. (348)-(351) are

$$
\begin{align*}
& \frac{\lambda\left(\mathrm{CuK} \alpha_{1}\right)}{d_{220}(\mathrm{~W} 4.2 \mathrm{a})}=\frac{1537.400 \mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)}{d_{220}(\mathrm{~W} 4.2 \mathrm{a})}  \tag{352}\\
& \frac{\lambda\left(\mathrm{WK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})}=\frac{0.2090100 \AA^{*}}{d_{220}(\mathrm{~N})}  \tag{353}\\
& \frac{\lambda\left(\mathrm{MoK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})}=\frac{707.831 \mathrm{xu}\left(\mathrm{MoK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})}  \tag{354}\\
& \frac{\lambda\left(\mathrm{CuK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})}=\frac{1537.400 \mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})} \tag{355}
\end{align*}
$$

where $d_{220}(\mathrm{~N})$ is taken to be an adjusted constant and $d_{220}(\mathrm{~W} 17)$ and $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ are adjusted constants as well.

## B. Particle Data Group input

There are a few cases in the 2006 adjustment in which an inexact constant that is used in the analysis of input data is not treated as an adjusted quantity, because the adjustment has a negligible effect on its value. Three such constants, used in the calculation of the theoretical expressions for the electron and muon magnetic moment anomalies $a_{\mathrm{e}}$ and $a_{\mu}$, are the mass of the tau lepton $m_{\tau}$, the Fermi coupling constant $G_{\mathrm{F}}$, and sine squared of the weak mixing angle $\sin ^{2} \theta_{\mathrm{W}}$, and are obtained from the most recent report of the Particle Data Group (Yao et al., 2006),

$$
\begin{equation*}
m_{\tau} c^{2}=1776.99(29) \mathrm{MeV} \quad\left[1.6 \times 10^{-4}\right] \tag{356}
\end{equation*}
$$

$$
\begin{align*}
& \frac{G_{\mathrm{F}}}{(\hbar c)^{3}}=1.16637(1) \times 10^{-5} \mathrm{GeV}^{-2} \quad\left[8.6 \times 10^{-6}\right]  \tag{357}\\
& \sin ^{2} \theta_{\mathrm{W}}=0.22255(56) \quad\left[2.5 \times 10^{-3}\right] \tag{358}
\end{align*}
$$

To facilitate the calculations, the uncertainty of $m_{\tau} c^{2}$ is symmetrized and taken to be 0.29 MeV rather than $+0.29 \mathrm{MeV},-0.26 \mathrm{MeV}$. We use the definition $\sin ^{2} \theta_{\mathrm{W}}$ $=1-\left(m_{\mathrm{W}} / m_{\mathrm{Z}}\right)^{2}$, where $m_{\mathrm{W}}$ and $m_{\mathrm{Z}}$ are, respectively, the masses of the $\mathrm{W}^{ \pm}$and $\mathrm{Z}^{0}$ bosons, because it is employed in the calculation of the electroweak contributions to $a_{\text {e }}$ and $a_{\mu}$ (Czarnecki et al., 1996). The Particle Data Group's recommended value for the mass ratio of these bosons is $m_{\mathrm{W}} / m_{\mathrm{Z}}=0.88173(32)$, which leads to the value of $\sin ^{2} \theta_{\mathrm{W}}$ given above.

## XII. ANALYSIS OF DATA

The previously discussed input data are examined in this section for their mutual compatibility and their potential role in determining the 2006 recommended values of the constants. Based on this analysis, the data are selected for the final least-squares adjustment from which the recommended values are obtained. Because the data on the Newtonian constant of gravitation $G$ are independent of the other data and are analyzed in Sec. X, they are not examined further. The consistency of the input data is evaluated by directly comparing different measurements of the same quantity, and by directly comparing the values of a single fundamental constant inferred from measurements of different quantities. As noted in the outline section of this paper, the inferred value is for comparison purposes only; the datum from which it is obtained, not the inferred value, is the input datum in the adjustment. The potential role of a particular input datum is gauged by carrying out a least-squares adjustment using all initially considered data. A particular measurement of a quantity is included in the final adjustment if its uncertainty is not more than about ten times the uncertainty of the value of that quantity provided by other data in the adjustment. The measure we use is the "self-sensitivity coefficient" of an input datum $S_{\text {c }}$ (see CODATA-98), which must be greater than 0.01 in order for the datum to be included.

The input data are given in Tables XXVIII, XXX, and XXXII and their covariances are given as correlation coefficients in Tables XXIX, XXXI, and XXXIII. The $\delta$ s given in Tables XXVIII, XXX, and XXXII are quantities added to corresponding theoretical expressions to account for the uncertainties of those expressions, as previously discussed (see, for example, Sec. IV.A.1.l). Note that the value of the Rydberg constant $R_{\infty}$ depends only weakly on changes, at the level of the uncertainties, of the data in Tables XXX and XXXII.

## A. Comparison of data

The classic Lamb shift is the only quantity among the Rydberg constant data with more than one measured value, but there are ten different quantities with more than one measured value among the other data. The item numbers given in Tables XXVIII and XXX for the members of such groups of data $(A 39, B 2, B 11, B 28$, $B 31, B 33-B 36, B 38$, and $B 58$ ) have a decimal point with an additional digit to label each member.

In fact, all data for which there is more than one measurement were directly compared in either the 1998 or 2002 adjustments except the following new data: the University of Washington result for $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right)$, item $B 2.2$, the Harvard University result for $a_{\mathrm{e}}$, item $B 11.2$, the NIST watt-balance result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ item $B 36.3$, and the INRIM result for $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$, item B38.2. The two values of $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right)$ agree well-they differ by only $0.5 u_{\text {diff }}$; the two values of $a_{\mathrm{e}}$ are in acceptable agreement-they differ by $1.7 u_{\text {difff }}$; the two values of $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ also agree well-they differ by $0.7 u_{\text {diff }}$; and the three values of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ are highly consistent-their mean and implied value of $h$ are

$$
\begin{align*}
& K_{\mathrm{J}}^{2} R_{\mathrm{K}}=6.03676187(21) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1},  \tag{359}\\
& h=6.62606889(23) \times 10^{-34} \mathrm{~J} \mathrm{~s} \tag{360}
\end{align*}
$$

with $\chi^{2}=0.27$ for $\nu=N-M=2$ degrees of freedom, where $N$ is the number of measurements and $M$ is the number of unknowns, and with Birge ratio $R_{\mathrm{B}}=\sqrt{\chi^{2} / \nu}=0.37$ (see Appendix E of CODATA-98). The normalized residuals for the three values are $0.52,-0.04$, and -0.09 , and their weights in the calculation of the weighted mean are 0.03 , 0.10 , and 0.87 .

Data for quantities with more than one directly measured value used in earlier adjustments are consistent, with the exception of the VNIIM 1989 result for $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$, which is not included in the present adjustment (see Sec. VII). We also note that none of these data has a weight of less than 0.02 in the weighted mean of measurements of the same quantity.

The consistency of measurements of various quantities of different types is shown mainly by comparing the values of the fine-structure constant $\alpha$ or the Planck constant $h$ inferred from the measured values of the quantities. Such inferred values of $\alpha$ and $h$ are given throughout the data review sections, and the results are summarized and discussed further here.

The consistency of a significant fraction of the data of Tables XXVIII and XXX is indicated in Table XXXIV and Figs. $3-5$, which give and graphically compare the values of $\alpha$ inferred from that data. Figures 3 and 4 compare the data that yield values of $\alpha$ with $u_{\mathrm{r}}<10^{-7}$ and $u_{\mathrm{r}}<10^{-8}$, respectively; Fig. 5 also compares the data that yield values of $\alpha$ with $u_{\mathrm{r}}<10^{-7}$, but does so

TABLE XXVIII. Summary of principal input data for the determination of the 2006 recommended value of the Rydberg constant $R_{\infty}$. [The notation for the additive corrections $\delta_{X}\left(n \mathrm{~L}_{j}\right)$ has the same meaning as the notation $\delta_{n \mathrm{~L} j}^{X}$ in Sec. IV.A.1.1.]

| Item number | Input datum | Value | Relative standard uncertainty ${ }^{\text {a }} u_{\mathrm{r}}$ | Identification | Sec. |
| :---: | :---: | :---: | :---: | :---: | :---: |
| A1 | $\delta_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}\right)$ | $0.0(3.7) \mathrm{kHz}$ | $\left[1.1 \times 10^{-12}\right]$ | Theory | IV.A.1.1 |
| A2 | $\delta_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}\right)$ | $0.00(46) \mathrm{kHz}$ | [5.6 $\left.\times 10^{-13}\right]$ | Theory | IV.A.1.1 |
| A3 | $\delta_{\mathrm{H}}\left(3 \mathrm{~S}_{1 / 2}\right)$ | $0.00(14) \mathrm{kHz}$ | $\left[3.7 \times 10^{-13}\right]$ | Theory | IV.A.1.1 |
| A4 | $\delta_{\mathrm{H}}\left(4 \mathrm{~S}_{1 / 2}\right)$ | $0.000(58) \mathrm{kHz}$ | [2.8 $\left.\times 10^{-13}\right]$ | Theory | IV.A.1.1 |
| A5 | $\delta_{\mathrm{H}}\left(6 \mathrm{~S}_{1 / 2}\right)$ | $0.000(20) \mathrm{kHz}$ | [2.1 $\left.\times 10^{-13}\right]$ | Theory | IV.A.1.1 |
| A6 | $\delta_{\mathrm{H}}\left(8 \mathrm{~S}_{1 / 2}\right)$ | $0.0000(82) \mathrm{kHz}$ | [1.6 $\left.\times 10^{-13}\right]$ | Theory | IV.A.1.1 |
| A7 | $\delta_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}\right)$ | $0.000(69) \mathrm{kHz}$ | $\left[8.4 \times 10^{-14}\right]$ | Theory | IV.A.1.1 |
| A8 | $\delta_{\mathrm{H}}\left(4 \mathrm{P}_{1 / 2}\right)$ | $0.0000(87) \mathrm{kHz}$ | $\left[4.2 \times 10^{-14}\right]$ | Theory | IV.A.1.1 |
| A9 | $\delta_{\mathrm{H}}\left(2 \mathrm{P}_{3 / 2}\right)$ | $0.000(69) \mathrm{kHz}$ | $\left[8.4 \times 10^{-14}\right]$ | Theory | IV.A.1.1 |
| A10 | $\delta_{\mathrm{H}}\left(4 \mathrm{P}_{3 / 2}\right)$ | $0.0000(87) \mathrm{kHz}$ | $\left[4.2 \times 10^{-14}\right]$ | Theory | IV.A.1.1 |
| A11 | $\delta_{\mathrm{H}}\left(8 \mathrm{D}_{3 / 2}\right)$ | $0.00000(48) \mathrm{kHz}$ | [9.3 $\left.\times 10^{-15}\right]$ | Theory | IV.A.1.1 |
| A12 | $\delta_{\mathrm{H}}\left(12 \mathrm{D}_{3 / 2}\right)$ | $0.00000(15) \mathrm{kHz}$ | $\left[6.6 \times 10^{-15}\right]$ | Theory | IV.A.1.1 |
| A13 | $\delta_{\mathrm{H}}\left(4 \mathrm{D}_{5 / 2}\right)$ | $0.0000(38) \mathrm{kHz}$ | $\left[1.9 \times 10^{-14}\right]$ | Theory | IV.A.1.1 |
| A14 | $\delta_{\mathrm{H}}\left(6 \mathrm{D}_{5 / 2}\right)$ | $0.0000(11) \mathrm{kHz}$ | $\left[1.2 \times 10^{-14}\right]$ | Theory | IV.A.1.1 |
| A15 | $\delta_{\mathrm{H}}\left(8 \mathrm{D}_{5 / 2}\right)$ | $0.00000(48) \mathrm{kHz}$ | [ $\left.9.3 \times 10^{-15}\right]$ | Theory | IV.A.1.1 |
| A16 | $\delta_{\mathrm{H}}\left(12 \mathrm{D}_{5 / 2}\right)$ | 0.000 00(16) kHz | [7.0 $\left.\times 10^{-15}\right]$ | Theory | IV.A.1.1 |
| A17 | $\delta_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}\right)$ | $0.0(3.6) \mathrm{kHz}$ | $\left[1.1 \times 10^{-12}\right]$ | Theory | IV.A.1.1 |
| A18 | $\delta_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}\right)$ | $0.00(45) \mathrm{kHz}$ | [5.4 $\left.\times 10^{-13}\right]$ | Theory | IV.A.1.1 |
| A19 | $\delta_{\mathrm{D}}\left(4 \mathrm{~S}_{1 / 2}\right)$ | $0.000(56) \mathrm{kHz}$ | $\left[2.7 \times 10^{-13}\right]$ | Theory | IV.A.1.1 |
| A20 | $\delta_{\mathrm{D}}\left(8 \mathrm{~S}_{1 / 2}\right)$ | $0.0000(80) \mathrm{kHz}$ | $\left[1.6 \times 10^{-13}\right]$ | Theory | IV.A.1.1 |
| A21 | $\delta_{\mathrm{D}}\left(8 \mathrm{D}_{3 / 2}\right)$ | $0.00000(48) \mathrm{kHz}$ | [9.3 $\left.\times 10^{-15}\right]$ | Theory | IV.A.1.1 |
| A22 | $\delta_{\mathrm{D}}\left(12 \mathrm{D}_{3 / 2}\right)$ | $0.00000(15) \mathrm{kHz}$ | $\left[6.6 \times 10^{-15}\right]$ | Theory | IV.A.1.1 |
| A23 | $\delta_{\mathrm{D}}\left(4 \mathrm{D}_{5 / 2}\right)$ | $0.0000(38) \mathrm{kHz}$ | $\left[1.9 \times 10^{-14}\right]$ | Theory | IV.A.1.1 |
| A24 | $\delta_{\mathrm{D}}\left(8 \mathrm{D}_{5 / 2}\right)$ | $0.00000(48) \mathrm{kHz}$ | $\left[9.3 \times 10^{-15}\right]$ | Theory | IV.A.1.1 |
| A25 | $\delta_{\mathrm{D}}\left(12 \mathrm{D}_{5 / 2}\right)$ | 0.000 00(16) kHz | [7.0 $\left.\times 10^{-15}\right]$ | Theory | IV.A.1.1 |
| A26 | $\nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | $2466061413187.074(34) \mathrm{kHz}$ | $1.4 \times 10^{-14}$ | MPQ-04 | IV.A. 2 |
| A27 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{~S}_{1 / 2}\right)$ | $770649350012.0(8.6) \mathrm{kHz}$ | $1.1 \times 10^{-11}$ | LK/SY-97 | IV.A. 2 |
| A28 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{D}_{3 / 2}\right)$ | $770649504450.0(8.3) \mathrm{kHz}$ | $1.1 \times 10^{-11}$ | LK/SY-97 | IV.A. 2 |
| A29 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{D}_{5 / 2}\right)$ | 770649561 584.2(6.4) kHz | $8.3 \times 10^{-12}$ | LK/SY-97 | IV.A. 2 |
| A30 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-12 \mathrm{D}_{3 / 2}\right)$ | 799191710 472.7(9.4) kHz | $1.2 \times 10^{-11}$ | LK/SY-98 | IV.A. 2 |
| A31 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-12 \mathrm{D}_{5 / 2}\right)$ | 799191727 403.7(7.0) kHz | $8.7 \times 10^{-12}$ | LK/SY-98 | IV.A. 2 |
| A32 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{~S}_{1 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | $4797338(10) \mathrm{kHz}$ | $2.1 \times 10^{-6}$ | MPQ-95 | IV.A. 2 |
| A33 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{D}_{5 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 6490 144(24) kHz | $3.7 \times 10^{-6}$ | MPQ-95 | IV.A. 2 |
| A34 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-6 \mathrm{~S}_{1 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-3 \mathrm{~S}_{1 / 2}\right)$ | 4197 604(21) kHz | $4.9 \times 10^{-6}$ | LKB-96 | IV.A. 2 |
| A35 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-6 \mathrm{D}_{5 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-3 \mathrm{~S}_{1 / 2}\right)$ | 4699 099(10) kHz | $2.2 \times 10^{-6}$ | LKB-96 | IV.A. 2 |
| A36 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{P}_{1 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 4664 269(15) kHz | $3.2 \times 10^{-6}$ | YaleU-95 | IV.A. 2 |
| A37 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{P}_{3 / 2}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 6035373 (10) kHz | $1.7 \times 10^{-6}$ | YaleU-95 | IV.A. 2 |
| A38 | $\nu_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}-2 \mathrm{P}_{3 / 2}\right)$ | 9911 200(12) kHz | $1.2 \times 10^{-6}$ | HarvU-94 | IV.A. 2 |
| A39.1 | $\nu_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | $1057845.0(9.0) \mathrm{kHz}$ | $8.5 \times 10^{-6}$ | HarvU-86 | IV.A. 2 |
| A39.2 | $\nu_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | $1057862(20) \mathrm{kHz}$ | $1.9 \times 10^{-5}$ | USus-79 | IV.A. 2 |
| A40 | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{~S}_{1 / 2}\right)$ | 770859041 245.7(6.9) kHz | $8.9 \times 10^{-12}$ | LK/SY-97 | IV.A. 2 |
| A41 | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{D}_{3 / 2}\right)$ | $770859195701.8(6.3) \mathrm{kHz}$ | $8.2 \times 10^{-12}$ | LK/SY-97 | IV.A. 2 |
| A42 | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-8 \mathrm{D}_{5 / 2}\right)$ | 770859252 849.5(5.9) kHz | $7.7 \times 10^{-12}$ | LK/SY-97 | IV.A. 2 |
| A43 | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-12 \mathrm{D}_{3 / 2}\right)$ | $799409168038.0(8.6) \mathrm{kHz}$ | $1.1 \times 10^{-11}$ | LK/SY-98 | IV.A. 2 |
| A44 | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-12 \mathrm{D}_{5 / 2}\right)$ | $799409184966.8(6.8) \mathrm{kHz}$ | $8.5 \times 10^{-12}$ | LK/SY-98 | IV.A. 2 |
| A45 | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{~S}_{1 / 2}\right)-\frac{1}{4} \nu_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 4801 693(20) kHz | $4.2 \times 10^{-6}$ | MPQ-95 | IV.A. 2 |
| A46 | $\nu_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}-4 \mathrm{D}_{5 / 2}\right)-\frac{1}{4} \nu_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 6494 841(41) kHz | $6.3 \times 10^{-6}$ | MPQ-95 | IV.A. 2 |

TABLE XXVIII. (Continued.)

| Item number | Input datum | Value | Relative <br> standard uncertainty ${ }^{\text {a }} u_{\mathrm{r}}$ | Identification | Sec. |
| :---: | :---: | :---: | :---: | :---: | :---: |
| A47 | $\nu_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)-\nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)$ | 670994334.64 (15) kHz | $2.2 \times 10^{-10}$ | MPQ-98 | IV.A. 2 |
| A48 | $R_{\text {p }}$ | 0.895 (18) fm | $2.0 \times 10^{-2}$ | Rp-03 | IV.A. 3 |
| A49 | $R_{\text {d }}$ | 2.130(10) fm | $4.7 \times 10^{-3}$ | Rd-98 | IV.A. 3 |

${ }^{\text {a }}$ The values in brackets are relative to the frequency equivalent of the binding energy of the indicated level.

TABLE XXIX. Correlation coefficients $r\left(x_{i}, x_{j}\right) \geqslant 0.0001$ of the input data related to $R_{\infty}$ in Table XXVIII. For simplicity, the two items of data to which a particular correlation coefficient corresponds are identified by their item numbers in Table XXVIII.

| $r(A 1, A 2)=0.9958$ | $r(A 6, A 19)=0.8599$ | $r(A 27, A 28)=0.3478$ | $r(A 30, A 44)=0.1136$ |
| ---: | ---: | :--- | :--- |
| $r(A 1, A 3)=0.9955$ | $r(A 6, A 20)=0.9913$ | $r(A 27, A 29)=0.4532$ | $r(A 31, A 34)=0.0278$ |
| $r(A 1, A 4)=0.9943$ | $r(A 7, A 8)=0.0043$ | $r(A 27, A 30)=0.0899$ | $r(A 31, A 35)=0.0553$ |
| $r(A 1, A 5)=0.8720$ | $r(A 9, A 10)=0.0043$ | $r(A 27, A 31)=0.1206$ | $r(A 31, A 40)=0.1512$ |
| $r(A 1, A 6)=0.8711$ | $r(A 11, A 12)=0.0005$ | $r(A 27, A 34)=0.0225$ | $r(A 31, A 41)=0.1647$ |
| $r(A 1, A 17)=0.9887$ | $r(A 11, A 21)=0.9999$ | $r(A 27, A 35)=0.0448$ | $r(A 31, A 42)=0.1750$ |
| $r(A 1, A 18)=0.9846$ | $r(A 11, A 22)=0.0003$ | $r(A 27, A 40)=0.1225$ | $r(A 31, A 43)=0.1209$ |
| $r(A 1, A 19)=0.9830$ | $r(A 12, A 21)=0.0003$ | $r(A 27, A 41)=0.1335$ | $r(A 31, A 44)=0.1524$ |
| $r(A 1, A 20)=0.8544$ | $r(A 12, A 22)=0.9999$ | $r(A 27, A 42)=0.1419$ | $r(A 32, A 33)=0.1049$ |
| $r(A 2, A 3)=0.9954$ | $r(A 13, A 14)=0.0005$ | $r(A 27, A 43)=0.0980$ | $r(A 32, A 45)=0.2095$ |
| $r(A 2, A 4)=0.9942$ | $r(A 13, A 15)=0.0005$ | $r(A 27, A 44)=0.1235$ | $r(A 32, A 46)=0.0404$ |
| $r(A 2, A 5)=0.8719$ | $r(A 13, A 16)=0.0004$ | $r(A 28, A 29)=0.4696$ | $r(A 33, A 45)=0.0271$ |
| $r(A 2, A 6)=0.8710$ | $r(A 13, A 23)=0.9999$ | $r(A 28, A 30)=0.0934$ | $r(A 33, A 46)=0.0467$ |
| $r(A 2, A 17)=0.9846$ | $r(A 13, A 24)=0.0002$ | $r(A 28, A 31)=0.1253$ | $r(A 34, A 35)=0.1412$ |
| $r(A 2, A 18)=0.9887$ | $r(A 13, A 25)=0.0002$ | $r(A 28, A 34)=0.0234$ | $r(A 34, A 40)=0.0282$ |
| $r(A 2, A 19)=0.9829$ | $r(A 14, A 15)=0.0005$ | $r(A 28, A 35)=0.0466$ | $r(A 34, A 41)=0.0307$ |
| $r(A 2, A 20)=0.8543$ | $r(A 14, A 16)=0.0005$ | $r(A 28, A 40)=0.1273$ | $r(A 34, A 42)=0.0327$ |
| $r(A 3, A 4)=0.9939$ | $r(A 14, A 23)=0.0002$ | $r(A 28, A 41)=0.1387$ | $r(A 34, A 43)=0.0226$ |
| $r(A 3, A 5)=0.8717$ | $r(A 14, A 24)=0.0003$ | $r(A 28, A 42)=0.1475$ | $r(A 34, A 44)=0.0284$ |
| $r(A 3, A 6)=0.8708$ | $r(A 14, A 25)=0.0002$ | $r(A 28, A 43)=0.1019$ | $r(A 35, A 40)=0.0561$ |
| $r(A 3, A 17)=0.9843$ | $r(A 15, A 16)=0.0005$ | $r(A 28, A 44)=0.1284$ | $r(A 35, A 41)=0.0612$ |
| $r(A 3, A 18)=0.9842$ | $r(A 15, A 23)=0.0002$ | $r(A 29, A 30)=0.1209$ | $r(A 35, A 42)=0.0650$ |
| $r(A 3, A 19)=0.9827$ | $r(A 15, A 24)=0.9999$ | $r(A 29, A 31)=0.1622$ | $r(A 35, A 43)=0.0449$ |
| $r(A 3, A 20)=0.8541$ | $r(A 15, A 25)=0.0002$ | $r(A 29, A 34)=0.0303$ | $r(A 35, A 44)=0.0566$ |
| $r(A 4, A 5)=0.8706$ | $r(A 16, A 23)=0.0002$ | $r(A 29, A 35)=0.0602$ | $r(A 36, A 37)=0.0834$ |
| $r(A 4, A 6)=0.8698$ | $r(A 16, A 24)=0.0002$ | $r(A 29, A 40)=0.1648$ | $r(A 40, A 41)=0.5699$ |
| $r(A 4, A 17)=0.9831$ | $r(A 16, A 25)=0.9999$ | $r(A 29, A 41)=0.1795$ | $r(A 40, A 42)=0.6117$ |
| $r(A 4, A 18)=0.9830$ | $r(A 17, A 18)=0.9958$ | $r(A 29, A 42)=0.1908$ | $r(A 40, A 43)=0.1229$ |
| $r(A 4, A 19)=0.9888$ | $r(A 17, A 19)=0.9942$ | $r(A 29, A 43)=0.1319$ | $r(A 40, A 44)=0.1548$ |
| $r(A 4, A 20)=0.8530$ | $r(A 17, A 20)=0.8641$ | $r(A 29, A 44)=0.1662$ | $r(A 41, A 42)=0.6667$ |
| $r(A 5, A 6)=0.7628$ | $r(A 18, A 19)=0.9941$ | $r(A 30, A 31)=0.4750$ | $r(A 41, A 43)=0.1339$ |
| $r(A 5, A 17)=0.8622$ | $r(A 18, A 20)=0.8640$ | $r(A 30, A 34)=0.0207$ | $r(A 41, A 44)=0.1687$ |
| $r(A 5, A 18)=0.8621$ | $r(A 19, A 20)=0.8627$ | $r(A 30, A 35)=0.0412$ | $r(A 42, A 43)=0.1423$ |
| $r(A 5, A 19)=0.8607$ | $r(A 21, A 22)=0.0001$ | $r(A 30, A 40)=0.1127$ | $r(A 42, A 44)=0.1793$ |
| $r(A 5, A 20)=0.7481$ | $r(A 23, A 24)=0.0001$ | $r(A 30, A 41)=0.1228$ | $r(A 43, A 44)=0.5224$ |
| $r(A 6, A 17)=0.8613$ | $r(A 23, A 25)=0.0001$ | $r(A 30, A 42)=0.1305$ | $r(A 45, A 46)=0.0110$ |
| $r(A 6, A 18)=0.8612$ | $r(A 24, A 25)=0.0001$ | $r(A 30, A 43)=0.0901$ |  |

TABLE XXX. Summary of principal input data for the determination of the 2006 recommended values of the fundamental constants ( $R_{\infty}$ and $G$ excepted).

| Item number | Input datum | Value | Relative standard uncertainty ${ }^{\text {a }}$ $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :---: | :---: | :---: | :---: | :---: | :---: |
| B1 | $A_{\mathrm{r}}\left({ }^{1} \mathrm{H}\right)$ | 1.007825032 07(10) | $1.0 \times 10^{-10}$ | AMDC-03 | III.A |
| B2.1 | $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right)$ | 2.014101777 85(36) | $1.8 \times 10^{-10}$ | AMDC-03 | III.A |
| B2.2 | $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right)$ | 2.014101778 040(80) | $4.0 \times 10^{-11}$ | UWash-06 | III.A |
| B3 | $A_{\mathrm{r}}\left({ }^{3} \mathrm{H}\right)$ | 3.016049 2787(25) | $8.3 \times 10^{-10}$ | MSL-06 | III.A |
| B4 | $A_{\mathrm{r}}\left({ }^{3} \mathrm{He}\right)$ | $3.0160293217(26)$ | $8.6 \times 10^{-10}$ | MSL-06 | III.A |
| B5 | $A_{\mathrm{r}}\left({ }^{4} \mathrm{He}\right)$ | $4.002603254131(62)$ | $1.5 \times 10^{-11}$ | UWash-06 | III.A |
| B6 | $A_{\mathrm{r}}\left({ }^{16} \mathrm{O}\right)$ | 15.994914619 57(18) | $1.1 \times 10^{-11}$ | UWash-06 | III.A |
| B7 | $A_{\mathrm{r}}\left({ }^{87} \mathrm{Rb}\right)$ | $86.909180526(12)$ | $1.4 \times 10^{-10}$ | AMDC-03 | III.A |
| $B 8^{\text {b }}$ | $A_{\mathrm{r}}\left({ }^{133} \mathrm{Cs}\right)$ | 132.905451 932(24) | $1.8 \times 10^{-10}$ | AMDC-03 | III.A |
| B9 | $A_{\mathrm{r}}(\mathrm{e})$ | 0.000548579 9111(12) | $2.1 \times 10^{-9}$ | UWash-95 | III.C (5) |
| B10 | $\delta_{\text {e }}$ | $0.00(27) \times 10^{-12}$ | [2.4 $\times 10^{-10}$ ] | Theory | V.A. 1 (101) |
| B11.1 | $a_{\text {e }}$ | $1.1596521883(42) \times 10^{-3}$ | $3.7 \times 10^{-9}$ | UWash-87 | V.A.2.a (102) |
| B11.2 | $a_{\text {e }}$ | $1.15965218085(76) \times 10^{-3}$ | $6.6 \times 10^{-10}$ | HarvU-06 | V.A.2.b (103) |
| B12 | $\delta_{\mu}$ | $0.0(2.1) \times 10^{-9}$ | [1.8 $\times 10^{-6}$ ] | Theory | V.B. 1 (126) |
| B13 | $\bar{R}$ | 0.003707 2064(20) | $5.4 \times 10^{-7}$ | BNL-06 | V.B. 2 (128) |
| B14 | $\delta_{\text {C }}$ | $0.00(27) \times 10^{-10}$ | $\left[1.4 \times 10^{-11}\right]$ | Theory | V.C. 1 (169) |
| B15 | $\delta_{\mathrm{O}}$ | $0.0(1.1) \times 10^{-10}$ | [5.3 $\times 10^{-11}$ ] | Theory | V.C. 1 (172) |
| B16 | $f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right) / f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)$ | $4376.2104989(23)$ | $5.2 \times 10^{-10}$ | GSI-02 | V.C.2.a (175) |
| B17 | $f_{\mathrm{s}}\left({ }^{16} \mathrm{O}^{7+}\right) / f_{\mathrm{c}}\left({ }^{16} \mathrm{O}^{7+}\right)$ | 4164.3761837 (32) | $7.6 \times 10^{-10}$ | GSI-02 | V.C.2.b (178) |
| B18 | $\mu_{\mathrm{e}^{-}}(\mathrm{H}) / \mu_{\mathrm{p}}(\mathrm{H})$ | -658.210 7058(66) | $1.0 \times 10^{-8}$ | MIT-72 | VI.A.2.a (195) |
| B19 | $\mu_{\mathrm{d}}(\mathrm{D}) / \mu_{\mathrm{e}}-(\mathrm{D})$ | -4.664 $345392(50) \times 10^{-4}$ | $1.1 \times 10^{-8}$ | MIT-84 | VI.A.2.b (197) |
| B20 | $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$ | $3.257199531(29)$ | $8.9 \times 10^{-9}$ | StPtrsb-03 | VI.A.2.c (201) |
| B21 | $\sigma_{\text {dp }}$ | $15(2) \times 10^{-9}$ |  | StPtrsb-03 | VI.A.2.c (203) |
| B22 | $\mu_{\mathrm{t}}(\mathrm{HT}) / \mu_{\mathrm{p}}(\mathrm{HT})$ | 1.066639 887(10) | $9.4 \times 10^{-9}$ | StPtrsb-03 | VI.A.2.c (202) |
| B23 | $\sigma_{\text {tp }}$ | $20(3) \times 10^{-9}$ |  | StPtrsb-03 | VI.A.2.c (204) |
| B24 | $\mu_{\mathrm{e}}-(\mathrm{H}) / \mu_{\mathrm{p}}^{\prime}$ | -658.215 9430(72) | $1.1 \times 10^{-8}$ | MIT-77 | VI.A.2.d (209) |
| B25 | $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime}$ | -0.7617861313(33) | $4.3 \times 10^{-9}$ | NPL-93 | VI.A.2.e (211) |
| B26 | $\mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime}$ | -0.684 996 94(16) | $2.4 \times 10^{-7}$ | ILL-79 | VI.A.2.f (212) |
| B27 | $\delta_{\text {Mu }}$ | 0 (101) Hz | [2.3 $\left.\times 10^{-8}\right]$ | Theory | VI.B. 1 (234) |
| B28.1 | $\Delta \nu_{\mathrm{Mu}}$ | 4463302.88 (16) kHz | $3.6 \times 10^{-8}$ | LAMPF-82 | VI.B.2.a (236) |
| B28.2 | $\Delta \nu_{\mathrm{Mu}}$ | $4463302765(53) \mathrm{Hz}$ | $1.2 \times 10^{-8}$ | LAMPF-99 | VI.B.2.b (239) |
| B29 | $\nu(58 \mathrm{MHz})$ | 627 994.77(14) kHz | $2.2 \times 10^{-7}$ | LAMPF-82 | VI.B.2.a (237) |
| B30 | $\nu(72 \mathrm{MHz})$ | 668223 166(57) Hz | $8.6 \times 10^{-8}$ | LAMPF-99 | VI.B.2.b (240) |
| B31.1 ${ }^{\text {b }}$ | $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ | $2.67515405(30) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $1.1 \times 10^{-7}$ | NIST-89 | VII.A.1.a (253) |
| B31.2 ${ }^{\text {b }}$ | $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ | $2.6751530(18) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $6.6 \times 10^{-7}$ | NIM-95 | VII.A.1.b (255) |
| B32 ${ }^{\text {b }}$ | $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ | $2.03789537(37) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $1.8 \times 10^{-7}$ | KR/VN-98 | VII.A.1.c (257) |
| B33.1 ${ }^{\text {b }}$ | $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})$ | $2.6751525(43) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $1.6 \times 10^{-6}$ | NIM-95 | VII.A.2.a (259) |
| B33.2 ${ }^{\text {b }}$ | $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})$ | $2.6751518(27) \times 10^{8} \mathrm{~s}^{-1} \mathrm{~T}^{-1}$ | $1.0 \times 10^{-6}$ | NPL-79 | VII.A.2.b (262) |
| B34.1 ${ }^{\text {b }}$ | $R_{\text {K }}$ | $25812.80831(62) \Omega$ | $2.4 \times 10^{-8}$ | NIST-97 | VII.B. 1 (265) |
| B34.2 ${ }^{\text {b }}$ | $R_{\text {K }}$ | $25812.8071(11) \Omega$ | $4.4 \times 10^{-8}$ | NMI-97 | VII.B. 2 (267) |
| $B 34.3{ }^{\text {b }}$ | $R_{\text {K }}$ | $25812.8092(14) \Omega$ | $5.4 \times 10^{-8}$ | NPL-88 | VII.B. 3 (269) |
| B34.4 ${ }^{\text {b }}$ | $R_{\text {K }}$ | $25812.8084(34) \Omega$ | $1.3 \times 10^{-7}$ | NIM-95 | VII.B. 4 (271) |
| B34.5 ${ }^{\text {b }}$ | $R_{\text {K }}$ | 25 812.8081(14) $\Omega$ | $5.3 \times 10^{-8}$ | LNE-01 | VII.B. 5 (273) |
| B35.1 ${ }^{\text {b }}$ | $K_{\text {J }}$ | 483 597.91(13) GHz V ${ }^{-1}$ | $2.7 \times 10^{-7}$ | NMI-89 | VII.C. 1 (276) |
| $B 35.2{ }^{\text {b }}$ | $K_{\text {J }}$ | $483597.96(15) \mathrm{GHz} \mathrm{V}^{-1}$ | $3.1 \times 10^{-7}$ | PTB-91 | VII.C. 2 (278) |
| $B 36.1{ }^{\text {c }}$ | $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $6.0367625(12) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1}$ | $2.0 \times 10^{-7}$ | NPL-90 | VII.D. 1 (281) |

TABLE XXX. (Continued.)

| Item number | Input datum | Value | Relative <br> standard uncertainty ${ }^{\text {a }}$ $u_{\mathrm{r}}$ | Identification | Sec. and Eq. |
| :---: | :---: | :---: | :---: | :---: | :---: |
| B36.2 ${ }^{\text {c }}$ | $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $6.03676185(53) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1}$ | $8.7 \times 10^{-8}$ | NIST-98 | VII.D.2.a (283) |
| B36.3 ${ }^{\text {c }}$ | $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $6.03676185(22) \times 10^{33} \mathrm{~J}^{-1} \mathrm{~s}^{-1}$ | $3.6 \times 10^{-8}$ | NIST-07 | VII.D.2.b (287) |
| $B 37{ }^{\text {b }}$ | $\mathcal{F}_{90}$ | 96485.39 (13) $\mathrm{C} \mathrm{mol}^{-1}$ | $1.3 \times 10^{-6}$ | NIST-80 | VII.E. 1 (295) |
| B38.1 ${ }^{\text {c }}$ | $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ | $192015.563(12) \mathrm{fm}$ | $6.2 \times 10^{-8}$ | PTB-81 | VIII.A.1.a (297) |
| B38.2 ${ }^{\text {c }}$ | $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ | $192015.5715(33) \mathrm{fm}$ | $1.7 \times 10^{-8}$ | INRIM-07 | VIII.A.1.c (299) |
| $B 39{ }^{\text {c }}$ | $d_{220}(\mathrm{NR} 3)$ | $192015.5919(76) \mathrm{fm}$ | $4.0 \times 10^{-8}$ | NMIJ-04 | VIII.A.1.b (298) |
| $B 40{ }^{\text {c }}$ | $d_{220}\left(\mathrm{MO}^{*}\right)$ | $192015.5498(51) \mathrm{fm}$ | $2.6 \times 10^{-8}$ | INRIM-07 | VIII.A.1.c (300) |
| B41 | $1-d_{220}(\mathrm{~N}) / d_{220}(\mathrm{~W} 17)$ | $7(22) \times 10^{-9}$ |  | NIST-97 | VIII.A.2.a (301) |
| B42 | $1-d_{220}(\mathrm{~W} 17) / d_{220}(\mathrm{ILL})$ | $-8(22) \times 10^{-9}$ |  | NIST-99 | VIII.A.2.a (302) |
| B43 | $1-d_{220}\left(\mathrm{MO}^{*}\right) / d_{220}($ ILL $)$ | $86(27) \times 10^{-9}$ |  | NIST-99 | VIII.A.2.a (303) |
| B44 | $1-d_{220}(\mathrm{NR} 3) / d_{220}(\mathrm{ILL})$ | $34(22) \times 10^{-9}$ |  | NIST-99 | VIII.A.2.a (304) |
| B45 | $d_{220}(\mathrm{NR} 3) / d_{220}(\mathrm{~W} 04)-1$ | $-11(21) \times 10^{-9}$ |  | NIST-06 | VIII.A.2.a (305) |
| B46 | $d_{220}(\mathrm{NR} 4) / d_{220}(\mathrm{~W} 04)-1$ | $25(21) \times 10^{-9}$ |  | NIST-06 | VIII.A.2.a (306) |
| B47 | $d_{220}(\mathrm{~W} 17) / d_{220}(\mathrm{~W} 04)-1$ | $11(21) \times 10^{-9}$ |  | NIST-06 | VIII.A.2.a (307) |
| B48 | $d_{220}(\mathrm{~W} 4.2 \mathrm{a}) / d_{220}(\mathrm{~W} 04)-1$ | $-1(21) \times 10^{-9}$ |  | PTB-98 | VIII.A.2.b (308) |
| B49 | $d_{220}(\mathrm{~W} 17) / d_{220}(\mathrm{~W} 04)-1$ | $22(22) \times 10^{-9}$ |  | PTB-98 | VIII.A.2.b (309) |
| B50 | $d_{220}\left(\mathrm{MO}^{*}\right) / d_{220}(\mathrm{~W} 04)-1$ | $-103(28) \times 10^{-9}$ |  | PTB-98 | VIII.A.2.b (310) |
| B51 | $d_{220}(\mathrm{NR} 3) / d_{220}(\mathrm{~W} 04)-1$ | $-23(21) \times 10^{-9}$ |  | PTB-98 | VIII.A.2.b (311) |
| B52 | $d_{220} / d_{220}(\mathrm{~W} 04)-1$ | $10(11) \times 10^{-9}$ |  | PTB-03 | VIII.A.2.b (312) |
| $B 53{ }^{\text {c }}$ | $V_{\mathrm{m}}(\mathrm{Si})$ | $12.0588254(34) \times 10^{-6} \mathrm{~m}^{3} \mathrm{~mol}^{-1}$ | $2.8 \times 10^{-7}$ | N/P/I-05 | VIII.B (317) |
| B54 | $\lambda_{\text {meas }} / d_{220}($ ILL $)$ | $0.00290430246(50) \mathrm{m} \mathrm{s}^{-1}$ | $1.7 \times 10^{-7}$ | NIST-99 | VIII.C (319) |
| B55 ${ }^{\text {c }}$ | $h / m_{\mathrm{n}} d_{220}$ (W04) | $2060.267004(84) \mathrm{m} \mathrm{s}^{-1}$ | $4.1 \times 10^{-8}$ | PTB-99 | VIII.D. 1 (322) |
| $B 56{ }^{\text {b }}$ | $h / m\left({ }^{133} \mathrm{Cs}\right)$ | $3.002369432(46) \times 10^{-9} \mathrm{~m}^{2} \mathrm{~s}^{-1}$ | $1.5 \times 10^{-8}$ | StanfU-02 | VIII.D. 2 (329) |
| B57 | $h / m\left({ }^{87} \mathrm{Rb}\right)$ | $4.591359287(61) \times 10^{-9} \mathrm{~m}^{2} \mathrm{~s}^{-1}$ | $1.3 \times 10^{-8}$ | LKB-06 | VIII.D. 3 (332) |
| B58.1 | $R$ | 8.314 471(15) $\mathrm{J} \mathrm{mol}^{-1} \mathrm{~K}^{-1}$ | $1.8 \times 10^{-6}$ | NIST-88 | IX.A. 1 (338) |
| B58.2 | $R$ | $8.314504(70) \mathrm{J} \mathrm{mol}^{-1} \mathrm{~K}^{-1}$ | $8.4 \times 10^{-6}$ | NPL-79 | IX.A. 2 (339) |
| B59 | $\lambda\left(\mathrm{CuK} \alpha_{1}\right) / d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ | $0.80232711(24)$ | $3.0 \times 10^{-7}$ | FSU/PTB-91 | XI.A (348) |
| B60 | $\lambda\left(\mathrm{WK} \alpha_{1}\right) / d_{220}(\mathrm{~N})$ | 0.108852 175(98) | $9.0 \times 10^{-7}$ | NIST-79 | XI.A (349) |
| B61 | $\lambda\left(\mathrm{MoK} \alpha_{1}\right) / d_{220}(\mathrm{~N})$ | $0.36940604(19)$ | $5.3 \times 10^{-7}$ | NIST-73 | XI.A (350) |
| B62 | $\lambda\left(\mathrm{CuK} \alpha_{1}\right) / d_{220}(\mathrm{~N})$ | 0.80232804 (77) | $9.6 \times 10^{-7}$ | NIST-73 | XI.A (351) |

${ }^{\mathrm{a}}$ The values in brackets are relative to the quantities $a_{\mathrm{e}}, a_{\mu}, g_{\mathrm{e}}-\left({ }^{12} \mathrm{C}^{5+}\right), g_{\mathrm{e}}\left({ }^{16} \mathrm{O}^{7+}\right)$, or $\Delta \nu_{\mathrm{Mu}}$ as appropriate.
${ }^{\mathrm{b}}$ Datum not included in the final least-squares adjustment that provides the recommended values of the constants.
${ }^{\mathrm{c}}$ Datum included in the final least-squares adjustment with an expanded uncertainty.
through combined values of $\alpha$ obtained from similar experiments. Most of the values of $\alpha$ are in reasonable agreement, implying that most data from which they are obtained are reasonably consistent. There are, however, two important exceptions.
The value of $\alpha$ inferred from the PTB measurement of $h / m_{\mathrm{n}} d_{220}(\mathrm{~W} 04)$, item B55, is based on the mean value $\overline{d_{220}}$ of $d_{220}$ (W04) implied by the four direct \{220\} XROI lattice spacing measurements, items B38.1-B40. It disagrees by about $2.8 u_{\text {diff }}$ with the value of $\alpha$ with the smallest uncertainty, that inferred from the Harvard University measurement of $a_{\mathrm{e}}$. Also, the value of $\alpha$ in-
ferred from the NIST measurement of $\Gamma_{\mathrm{p}-90}^{\prime}(1 \mathrm{o})$ disagrees with the latter by about $2.3 u_{\text {diff }}$. But it is also worth noting that the value $\alpha^{-1}=137.0360000(38)$ [ $2.8 \times 10^{-8}$ ] implied by $h / m_{\mathrm{n}} d_{220}$ (W04) together with item B39 alone, the NMIJ XROI measurement of $d_{220}$ (NR3), agrees well with the Harvard $a_{\mathrm{e}}$ value of $\alpha$. If instead one uses the three other direct XROI lattice spacing measurements, items $B 38.1, B 38.2$, and $B 40$, which agree among themselves, one finds $\alpha^{-1}$ $=137.0360092(28)\left[2.1 \times 10^{-8}\right]$. This value disagrees with $\alpha$ from the Harvard $a_{\mathrm{e}}$ by $3.3 u_{\text {diff }}$.

TABLE XXXI. Non-negligible correlation coefficients $r\left(x_{i}, x_{j}\right)$ of the input data in Table XXX. For simplicity, the two items of data to which a particular correlation coefficient corresponds are identified by their item numbers in Table XXX.

| $r(B 1, B 2.1)=0.073$ | $r(B 38.1, B 38.2)=0.191$ | $r(B 42, B 46)=0.065$ | $r(B 46, B 47)=0.509$ |
| :--- | :--- | :--- | :--- |
| $r(B 2.2, B 5)=0.127$ | $r(B 38.2, B 40)=0.057$ | $r(B 42, B 47)=-0.367$ | $r(B 48, B 49)=0.469$ |
| $r(B 2.2, B 6)=0.089$ | $r(B 41, B 42)=-0.288$ | $r(B 43, B 44)=0.421$ | $r(B 48, B 50)=0.372$ |
| $r(B 5, B 6)=0.181$ | $r(B 41, B 43)=0.096$ | $r(B 43, B 45)=0.053$ | $r(B 48, B 51)=0.502$ |
| $r(B 14, B 15)=0.919$ | $r(B 41, B 44)=0.117$ | $r(B 43, B 46)=0.053$ | $r(B 48, B 55)=0.258$ |
| $r(B 16, B 17)=0.082$ | $r(B 41, B 45)=0.066$ | $r(B 43, B 47)=0.053$ | $r(B 49, B 50)=0.347$ |
| $r(B 28.1, B 29)=0.227$ | $r(B 41, B 46)=0.066$ | $r(B 44, B 45)=-0.367$ | $r(B 49, B 51)=0.469$ |
| $r(B 28.2, B 30)=0.195$ | $r(B 41, B 47)=0.504$ | $r(B 44, B 46)=0.065$ | $r(B 49, B 55)=0.241$ |
| $r(B 31.2, B 33.1)=-0.014$ | $r(B 42, B 43)=0.421$ | $r(B 44, B 47)=0.065$ | $r(B 50, B 51)=0.372$ |
| $r(B 35.1, B 58.1)=0.068$ | $r(B 42, B 44)=0.516$ | $r(B 45, B 46)=0.509$ | $r(B 50, B 55)=0.192$ |
| $r(B 36.2, B 36.3)=0.140$ | $r(B 42, B 45)=0.065$ | $r(B 45, B 47)=0.509$ | $r(B 51, B 55)=0.258$ |

The values of $\alpha$ compared in Fig. 5 follow from Table

$$
\begin{equation*}
\alpha^{-1}[h / m]=137.03599935(69) \quad\left[5.0 \times 10^{-9}\right] \tag{362}
\end{equation*}
$$ XXXIV and are, again in order of increasing uncertainty,

$$
\begin{equation*}
\alpha^{-1}\left[a_{\mathrm{e}}\right]=137.035999683(94) \quad\left[6.9 \times 10^{-10}\right], \quad(361) \quad \alpha^{-1}\left[R_{\mathrm{K}}\right]=137.0360030(25) \quad\left[1.8 \times 10^{-8}\right] \tag{363}
\end{equation*}
$$

TABLE XXXII. Summary of principal input data for the determination of the relative atomic mass of the electron from antiprotonic helium transitions. The numbers in parentheses ( $n, l: n^{\prime}, l^{\prime}$ ) denote the transition $(n, l) \rightarrow\left(n^{\prime}, l^{\prime}\right)$.

| Item number | Input datum | Value | Relative standard uncertainty ${ }^{\text {a }} u_{\mathrm{r}}$ | Identification | Sec. |
| :---: | :---: | :---: | :---: | :---: | :---: |
| C1 | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(32,31: 31,30)$ | 0.00 (82) MHz | [7.3 $\left.\times 10^{-10}\right]$ | JINR-06 | IV.B |
| C2 | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(35,33: 34,32)$ | $0.0(1.0) \mathrm{MHz}$ | $\left[1.3 \times 10^{-9}\right]$ | JINR-06 | IV.B |
| C3 | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(36,34: 35,33)$ | $0.0(1.2) \mathrm{MHz}$ | $\left[1.6 \times 10^{-9}\right]$ | JINR-06 | IV.B |
| C4 | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(39,35: 38,34)$ | $0.0(1.1) \mathrm{MHz}$ | [ $1.8 \times 10^{-9}$ ] | JINR-06 | IV.B |
| C5 | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}{ }^{+}(40,35: 39,34)$ | $0.0(1.2) \mathrm{MHz}$ | [ $2.4 \times 10^{-9}$ ] | JINR-06 | IV.B |
| C6 | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(32,31: 31,30)$ | $0.0(1.3) \mathrm{MHz}$ | [ $2.9 \times 10^{-9}$ ] | JINR-06 | IV.B |
| C7 | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(37,35: 38,34)$ | $0.0(1.8) \mathrm{MHz}$ | [ $4.4 \times 10^{-9}$ ] | JINR-06 | IV.B |
| C8 | $\delta_{\overline{\mathrm{p}}}{ }^{3} \mathrm{He}^{+}(32,31: 31,30)$ | $0.00(91) \mathrm{MHz}$ | [8.7 $\left.\times 10^{-10}\right]$ | JINR-06 | IV.B |
| C9 | $\delta_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(34,32: 33,31)}$ | $0.0(1.1) \mathrm{MHz}$ | $\left[1.4 \times 10^{-9}\right]$ | JINR-06 | IV.B |
| C10 | $\delta_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(36,33: 35,32)}$ | $0.0(1.2) \mathrm{MHz}$ | [ $1.8 \times 10^{-9}$ ] | JINR-06 | IV.B |
| C11 | $\delta_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(38,34: 37,33)}$ | $0.0(1.1) \mathrm{MHz}$ | [ $\left.2.3 \times 10^{-9}\right]$ | JINR-06 | IV.B |
| C12 | $\delta_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(36,34: 37,33)}$ | $0.0(1.8) \mathrm{MHz}$ | $\left[4.4 \times 10^{-9}\right]$ | JINR-06 | IV.B |
| C13 | $\nu_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(32,31: 31,30)$ | 1132609 209(15) MHz | $1.4 \times 10^{-8}$ | CERN-06 | IV.B |
| C14 | $\nu_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(35,33: 34,32)$ | 804633 059.0(8.2) MHz | $1.0 \times 10^{-8}$ | CERN-06 | IV.B |
| C15 | $\nu_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(36,34: 35,33)$ | 717474 004(10) MHz | $1.4 \times 10^{-8}$ | CERN-06 | IV.B |
| C16 | $\nu_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}{ }^{+}(39,35: 38,34)$ | 636878 139.4(7.7) MHz | $1.2 \times 10^{-8}$ | CERN-06 | IV.B |
| C17 | $\nu_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(40,35: 39,34)$ | 501948 751.6(4.4) MHz | $8.8 \times 10^{-9}$ | CERN-06 | IV.B |
| C18 | $\nu_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(32,31: 31,30)$ | 445608557.6 (6.3) MHz | $1.4 \times 10^{-8}$ | CERN-06 | IV.B |
| C19 | $\nu_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(37,35: 38,34)$ | 412885 132.2(3.9) MHz | $9.4 \times 10^{-9}$ | CERN-06 | IV.B |
| C20 | $\nu_{\overline{\mathrm{p}}}{ }^{3} \mathrm{He}^{+}(32,31: 31,30)$ | 1043128 608(13) MHz | $1.3 \times 10^{-8}$ | CERN-06 | IV.B |
| C21 | $\nu_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(34,32: 33,31)}$ | 822809 190(12) MHz | $1.5 \times 10^{-8}$ | CERN-06 | IV.B |
| C22 | $\nu_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(36,33: 34,32)}$ | 646180 434(12) MHz | $1.9 \times 10^{-8}$ | CERN-06 | IV.B |
| C23 | $\nu_{\overline{\mathrm{p}}}{ }^{3} \mathrm{He}^{+}(38,34: 37,33)$ | 505222 295.7(8.2) MHz | $1.6 \times 10^{-8}$ | CERN-06 | IV.B |
| C24 | $\nu_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(36,34: 37,33)}$ | 414147 507.8(4.0) MHz | $9.7 \times 10^{-9}$ | CERN-06 | IV.B |

[^2]TABLE XXXIII. Non-negligible correlation coefficients $r\left(x_{i}, x_{j}\right)$ of the input data in Table XXXII. For simplicity, the two items of data to which a particular correlation coefficient corresponds are identified by their item numbers in Table XXXII.

| $r(C 1, C 2)=0.929$ | $r(C 9, C 10)=0.925$ | $r(C 14, C 23)=0.132$ | $r(C 17, C 24)=0.287$ |
| :--- | :--- | :--- | :--- |
| $r(C 1, C 3)=0.912$ | $r(C 9, C 11)=0.949$ | $r(C 14, C 24)=0.271$ | $r(C 18, C 19)=0.235$ |
| $r(C 1, C 4)=0.936$ | $r(C 9, C 12)=0.978$ | $r(C 15, C 16)=0.223$ | $r(C 18, C 20)=0.107$ |
| $r(C 1, C 5)=0.883$ | $r(C 10, C 11)=0.907$ | $r(C 15, C 17)=0.198$ | $r(C 18, C 21)=0.118$ |
| $r(C 1, C 6)=0.758$ | $r(C 10, C 12)=0.934$ | $r(C 15, C 18)=0.140$ | $r(C 18, C 22)=0.122$ |
| $r(C 1, C 7)=0.957$ | $r(C 11, C 12)=0.959$ | $r(C 15, C 19)=0.223$ | $r(C 18, C 23)=0.112$ |
| $r(C 2, C 3)=0.900$ | $r(C 13, C 14)=0.210$ | $r(C 15, C 20)=0.128$ | $r(C 18, C 24)=0.229$ |
| $r(C 2, C 4)=0.924$ | $r(C 13, C 15)=0.167$ | $r(C 15, C 21)=0.142$ | $r(C 19, C 20)=0.170$ |
| $r(C 2, C 5)=0.872$ | $r(C 13, C 16)=0.224$ | $r(C 15, C 22)=0.141$ | $r(C 19, C 21)=0.188$ |
| $r(C 2, C 6)=0.748$ | $r(C 13, C 17)=0.197$ | $r(C 15, C 23)=0.106$ | $r(C 19, C 22)=0.191$ |
| $r(C 2, C 7)=0.945$ | $r(C 13, C 18)=0.138$ | $r(C 15, C 24)=0.217$ | $r(C 19, C 23)=0.158$ |
| $r(C 3, C 4)=0.907$ | $r(C 13, C 19)=0.222$ | $r(C 16, C 17)=0.268$ | $r(C 19, C 24)=0.324$ |
| $r(C 3, C 5)=0.856$ | $r(C 13, C 20)=0.129$ | $r(C 16, C 18)=0.193$ | $r(C 20, C 21)=0.109$ |
| $r(C 3, C 6)=0.734$ | $r(C 13, C 21)=0.142$ | $r(C 16, C 19)=0.302$ | $r(C 20, C 22)=0.108$ |
| $r(C 3, C 7)=0.927$ | $r(C 13, C 22)=0.141$ | $r(C 16, C 20)=0.172$ | $r(C 20, C 23)=0.081$ |
| $r(C 4, C 5)=0.878$ | $r(C 13, C 23)=0.106$ | $r(C 16, C 21)=0.190$ | $r(C 20, C 24)=0.166$ |
| $r(C 4, C 6)=0.753$ | $r(C 13, C 24)=0.216$ | $r(C 16, C 22)=0.189$ | $r(C 21, C 22)=0.120$ |
| $r(C 4, C 7)=0.952$ | $r(C 14, C 15)=0.209$ | $r(C 16, C 23)=0.144$ | $r(C 21, C 23)=0.090$ |
| $r(C 5, C 6)=0.711$ | $r(C 14, C 16)=0.280$ | $r(C 16, C 24)=0.294$ | $r(C 21, C 24)=0.184$ |
| $r(C 5, C 7)=0.898$ | $r(C 14, C 17)=0.247$ | $r(C 17, C 18)=0.210$ | $r(C 22, C 23)=0.091$ |
| $r(C 6, C 7)=0.770$ | $r(C 14, C 18)=0.174$ | $r(C 17, C 19)=0.295$ | $r(C 22, C 24)=0.186$ |
| $r(C 8, C 9)=0.978$ | $r(C 14, C 19)=0.278$ | $r(C 17, C 20)=0.152$ | $r(C 23, C 24)=0.154$ |
| $r(C 8, C 10)=0.934$ | $r(C 14, C 20)=0.161$ | $r(C 17, C 21)=0.167$ |  |
| $r(C 8, C 11)=0.959$ | $r(C 14, C 21)=0.178$ | $r(C 17, C 22)=0.169$ |  |
| $r(C 8, C 12)=0.988$ | $r(C 14, C 22)=0.177$ | $r(C 17, C 23)=0.141$ |  |

$$
\begin{array}{ll}
\alpha^{-1}\left[h / m_{\mathrm{n}} d_{220}\right]=137.0360077(28) & {\left[2.1 \times 10^{-8}\right],} \\
\alpha^{-1}\left[\Gamma_{\mathrm{p}, \mathrm{~h}-90}^{\prime}(\mathrm{lo})\right]=137.0359875(43) & {\left[3.1 \times 10^{-8}\right],} \tag{365}
\end{array}
$$

$$
\begin{equation*}
\alpha^{-1}\left[\Delta \nu_{\mathrm{Mu}}\right]=137.0360017(80) \quad\left[5.8 \times 10^{-8}\right] . \tag{366}
\end{equation*}
$$

Here $\alpha^{-1}\left[a_{\mathrm{e}}\right]$ is the weighted mean of the two $a_{\mathrm{e}}$ values of $\alpha ; \alpha^{-1}[h / m]$ is the weighted mean of the $h / m\left({ }^{87} \mathrm{Rb}\right)$ and $h / m\left({ }^{133} \mathrm{Cs}\right)$ values; $\alpha^{-1}\left[R_{\mathrm{K}}\right]$ is the weighted mean of the five quantum Hall effect-calculable capacitor values; $\alpha^{-1}\left[h / m_{\mathrm{n}} d_{22}\right]$ is the value given in Table XXXIV and is based on the measurement of $h / m_{\mathrm{n}} d_{220}$ (W04) and the value of $d_{220}$ (W04) inferred from the four XROI determinations of the $\{220\}$ lattice spacing of three different silicon crystals; $\alpha^{-1}\left[\Gamma_{\mathrm{p}, \mathrm{h}-90}^{\prime}(\mathrm{lo})\right]$ is the weighted mean of the two values of $\alpha^{-1}\left[\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})\right]$ and one value of $\alpha^{-1}\left[\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})\right]$; and $\alpha^{-1}\left[\Delta \nu_{\mathrm{Mu}}\right]$ is the value given in Table XXXIV and is based on the 1982 and 1999 measurements at LAMPF on muonium.
Figures 3-5 show that even if all data of Table XXX were retained, the 2006 recommended value of $\alpha$ would be determined to a great extent by $a_{\mathrm{e}}$, and in particular, the Harvard University determination of $a_{\mathrm{e}}$.

The consistency of a significant fraction of the data of Table XXX is indicated in Table XXXV and Figs. 6 and 7 , which give and graphically compare the values of $h$ inferred from those data. Figure 6 compares the data by showing each inferred value of $h$ in the table, while Fig. 7 compares the data through combined values of $h$ from similar experiments. The values of $h$ are in good agreement, implying that the data from which they are obtained are consistent, with one important exception. The value of $h$ inferred from $V_{\mathrm{m}}(\mathrm{Si})$, item B53, disagrees by $2.9 u_{\text {diff }}$ with the value of $h$ from the weighted mean of the three watt-balance values of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ [uncertainty $u_{\mathrm{r}}$ $=3.4 \times 10^{-8}$-see Eq. (360)].
In this regard, it is worth noting that a value of $d_{220}$ of an ideal silicon crystal is required to obtain a value of $h$ from $V_{\mathrm{m}}(\mathrm{Si})$ [see Eq. (316)], and the value used to obtain the inferred value of $h$ given in Eq. (318) and Table XXXV is based on all four XROI lattice spacing measurements, items B38.1-B40, plus the indirect value from $h / m_{\mathrm{n}} d_{220}$ (W04) (see Table XXIV and Fig. 1). However, the NMIJ measurement of $d_{220}(\mathrm{NR} 3)$, item $B 39$, and the indirect value of $d_{220}$ from $h / m_{\mathrm{n}} d_{220}$ (W04) yield values of $h$ from $V_{\mathrm{m}}(\mathrm{Si})$ that are less consistent with the watt-balance mean value than the three other direct XROI lattice spacing measurements, items B38.1, B38.2,

TABLE XXXIV. Comparison of the input data in Table XXX through inferred values of the finestructure constant $\alpha$ in order of increasing standard uncertainty.

| Primary <br> source | Item <br> number | Identification | Sec. and Eq. | $\alpha^{-1}$ | Relative standard <br> uncertainty $u_{\mathrm{r}}$ |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $a_{\mathrm{e}}$ | $B 11.2$ | HarvU-06 | V.A.3 (105) | $137.035999711(96)$ | $7.0 \times 10^{-10}$ |
| $a_{\mathrm{e}}$ | $B 11.1$ | UWash-87 | V.A.3 (104) | $137.03599883(50)$ | $3.7 \times 10^{-9}$ |
| $h / m(\mathrm{Rb})$ | $B 57$ | LKB-06 | VIII.D.3 (334) | $137.03599883(91)$ | $6.7 \times 10^{-9}$ |
| $h / m(\mathrm{Cs})$ | $B 56$ | StanfU-02 | VIII.D.2 (331) | $137.0360000(11)$ | $7.7 \times 10^{-9}$ |
| $h / m_{\mathrm{n}} d_{220}($ W04 $)$ | $B 55$ | PTB-99 |  |  |  |
| $\overline{d_{220}}$ | $B 38.1-B 40$ | Mean | VIII.D.1 (324) | $137.0360077(28)$ | $2.1 \times 10^{-8}$ |
| $R_{\mathrm{K}}$ | $B 34.1$ | NIST-97 | VII.B.1 (266) | $137.0360037(33)$ | $2.4 \times 10^{-8}$ |
| $\Gamma_{\mathrm{p}-90}^{\prime}(l \mathrm{lo})$ | $B 31.1$ | NIST-89 | VII.A.1.a (254) | $137.0359879(51)$ | $3.7 \times 10^{-8}$ |
| $R_{\mathrm{K}}$ | $B 34.2$ | NMI-97 | VII.B.2 (268) | $137.0359973(61)$ | $4.4 \times 10^{-8}$ |
| $R_{\mathrm{K}}$ | $B 34.5$ | LNE-01 | VII.B.5 (274) | $137.0360023(73)$ | $5.3 \times 10^{-8}$ |
| $R_{\mathrm{K}}$ | $B 34.3$ | NPL-88 | VII.B.3 (270) | $137.0360083(73)$ | $5.4 \times 10^{-8}$ |
| $\Delta \nu_{\mathrm{Mu}}$ | $B 28.1, B 28.2$ | LAMPF | VI.B.2.c (244) | $137.0360017(80)$ | $5.8 \times 10^{-8}$ |
| $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ | $B 32$ | KR/VN-98 | VII.A.1.c (258) | $137.0359852(82)$ | $6.0 \times 10^{-8}$ |
| $R_{\mathrm{K}}$ | $B 34.4$ | NIM-95 | VII.B.4 (272) | $137.036004(18)$ | $1.3 \times 10^{-7}$ |
| $\Gamma_{\mathrm{p}-90}^{\prime}(l \mathrm{lo})$ | $B 31.2$ | NIM-95 | VII.A.1.b (256) | $137.036006(30)$ | $2.2 \times 10^{-7}$ |
| $\nu_{\mathrm{H}}, \nu_{\mathrm{D}}$ | $A 26-A 47$ | Various | IV.A.1.m (65) | $137.036002(48)$ | $3.5 \times 10^{-7}$ |
| $\bar{R}$ | $B 13$ | BNL-02 | V.B.2.a (132) | $137.03567(26)$ | $1.9 \times 10^{-6}$ |

and $B 40$, which agree among themselves (a disagreement of about $3.8 u_{\text {diff }}$ compared to $2.5 u_{\text {diff }}$ ). In contrast, the NMIJ measurement of $d_{220}$ (NR3) yields a value of $\alpha$ from $h / m_{\mathrm{n}} d_{220}$ (W04) that is in excellent agreement with the Harvard University value from $a_{\mathrm{e}}$, while the three other lattice spacing measurements yield a value of $\alpha$ in


FIG. 3. Values of the fine-structure constant $\alpha$ with $u_{\mathrm{r}}<10^{-7}$ implied by the input data in Table XXX, in order of decreasing uncertainty from top to bottom, and the 2002 and 2006 CODATA recommended values of $\alpha$. (See Table XXXIV.) Here "mean" indicates the PTB-99 result for $h / m_{\mathrm{n}} d_{220}$ (W04) using the value of $d_{220}(\mathrm{~W} 04)$ implied by the four XROI latticespacing measurements.
poor agreement with alpha from $a_{\mathrm{e}}\left(3.3 u_{\text {diff }}\right)$.
The values of $h$ compared in Fig. 7 follow from Table XXXV and are, again in order of increasing uncertainty,

$$
\begin{align*}
h\left[K_{\mathrm{J}}^{2} R_{\mathrm{K}}\right]= & 6.62606889(23) \times 10^{-34} \mathrm{~J} \mathrm{~s} \\
& {\left[3.4 \times 10^{-8}\right], }  \tag{367}\\
h\left[V_{\mathrm{m}}(\mathrm{Si})\right]= & 6.6260745(19) \times 10^{-34} \mathrm{~J} \mathrm{~s} \\
& {\left[2.9 \times 10^{-7}\right], } \tag{368}
\end{align*}
$$



FIG. 4. Values of the fine-structure constant $\alpha$ with $u_{\mathrm{r}}<10^{-8}$ implied by the input data in Table XXX, in order of decreasing uncertainty from top to bottom. (See Table XXXIV.)


FIG. 5. Values of the fine-structure constant $\alpha$ with $u_{\mathrm{r}}<10^{-7}$ implied by the input data in Table XXX, taken as a weighted mean when more than one measurement of a given type is considered [see Eqs. (361)-(366)], in order of decreasing uncertainty from top to bottom.

$$
\begin{align*}
& h\left[K_{\mathrm{J}}\right]=6.6260678(27) \times 10^{-34} \mathrm{~J} \mathrm{~s} \\
& \quad\left[4.1 \times 10^{-7}\right],  \tag{369}\\
& h\left[\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})\right]=6.6260724(57) \times 10^{-34} \mathrm{~J} \mathrm{~s} \\
& \quad\left[8.6 \times 10^{-7}\right],  \tag{370}\\
& h\left[\mathcal{F}_{90}\right]=6.6260657(88) \times 10^{-34} \mathrm{~J} \mathrm{~s} \\
& \quad\left[1.3 \times 10^{-6}\right] . \tag{371}
\end{align*}
$$

Here $h\left[K_{\mathrm{J}}^{2} R_{\mathrm{K}}\right]$ is the weighted mean of the three values of $h$ from the three watt-balance measurements of $K_{\mathrm{J}}^{2} R_{\mathrm{K}} ; h\left[V_{\mathrm{m}}(\mathrm{Si})\right]$ is the value given in Table XXXV and based on all four XROI $d_{220}$ lattice spacing measurements plus the indirect lattice spacing value from $h / m_{\mathrm{n}} d_{220}$ (W04); $h\left[K_{\mathrm{J}}\right]$ is the weighted mean of the two direct Josephson effect measurements of $K_{\mathrm{J}} ; h\left[\Gamma_{\mathrm{p}-90}(\mathrm{hi})\right]$


FIG. 6. Values of the Planck constant $h$ implied by the input data in Table XXX, in order of decreasing uncertainty from top to bottom. (See Table XXXV.)
is the weighted mean of the two values of $h$ from the two measurements of $\Gamma_{\mathrm{p}-90}(\mathrm{hi})$; and $h\left[\mathcal{F}_{90}\right]$ is the value given in Table XXXV and comes from the silver coulometer measurement of $\mathcal{F}_{90}$. Figures 6 and 7 show that even if all data of Table XXX were retained, the 2006 recommended value of $h$ would be determined to a large extent by $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, and, in particular, the NIST 2007 determination of this quantity.

We conclude our data comparisons by listing in Table XXXVI the four available values of $A_{\mathrm{r}}(\mathrm{e})$. The reasonable agreement of these values shows that the corresponding input data are consistent. The most important of these data are the University of Washington value of $A_{\mathrm{r}}(\mathrm{e}), \delta_{\mathrm{C}}, \delta_{\mathrm{O}}, f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right) / f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right), f_{\mathrm{s}}\left({ }^{12} \mathrm{O}^{7+}\right) / f_{\mathrm{c}}\left({ }^{12} \mathrm{O}^{7+}\right)$, and the antiprotonic helium data, items $B 9, B 14-B 17$, and C1-C24.

In summary, the data comparisons here have identified the following potential problems: (i) the measure-

TABLE XXXV. Comparison of the input data in Table XXX through inferred values of the Planck constant $h$ in order of increasing standard uncertainty.

| Primary <br> source | Item <br> number | Identification | Sec. and Eq. | $h /(\mathrm{J} \mathrm{s})$ | Relative standard <br> uncertainty $u_{\mathrm{r}}$ |
| :--- | :--- | :---: | :--- | :---: | :---: |
| $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $B 36.3$ | NIST-07 | VII.D.2.b (288) | $6.62606891(24) \times 10^{-34}$ | $3.6 \times 10^{-8}$ |
| $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $B 36.2$ | NIST-98 | VII.D.2.a (284) | $6.62606891(58) \times 10^{-34}$ | $8.7 \times 10^{-8}$ |
| $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ | $B 36.1$ | NPL-90 | VII.D.1 (282) | $6.6260682(13) \times 10^{-34}$ | $2.0 \times 10^{-7}$ |
| $V_{\mathrm{m}}(\mathrm{Si})$ | $B 53$ | N/P/I-05 | VIII.B (318) | $6.6260745(19) \times 10^{-34}$ | $2.9 \times 10^{-7}$ |
| $K_{\mathrm{J}}$ | $B 35.1$ | NMI-89 | VII.C.1 (277) | $6.6260684(36) \times 10^{-34}$ | $5.4 \times 10^{-7}$ |
| $K_{\mathrm{J}}$ | $B 35.2$ | PTB-91 | VII.C.2 (279) | $6.6260670(42) \times 10^{-34}$ | $6.3 \times 10^{-7}$ |
| $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})$ | $B 33.2$ | NPL-79 | VII.A.2.b (263) | $6.6260729(67) \times 10^{-34}$ | $1.0 \times 10^{-6}$ |
| $\mathcal{F}_{90}$ | $B 37$ | NIST-80 | VII.E.1 (296) | $6.6260657(88) \times 10^{-34}$ | $1.3 \times 10^{-6}$ |
| $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi})$ | $B 33.1$ | NIM-95 | VII.A.2.a (261) | $6.626071(11) \times 10^{-34}$ | $1.6 \times 10^{-6}$ |

TABLE XXXVI. Values of $A_{\mathrm{r}}(\mathrm{e})$ implied by the input data in Table XXX in order of increasing standard uncertainty.

| Primary <br> source | Item <br> number | Identification | Sec. and Eq. | $A_{\mathrm{r}}(\mathrm{e})$ | Relative standard <br> uncertainty $u_{\mathrm{r}}$ |
| :--- | :--- | :--- | :--- | :--- | :---: |
| $f_{\mathrm{s}}(\mathrm{C}) / f_{\mathrm{c}}(\mathrm{C})$ | $B 16$ | GSI-02 | V.C.2.a (177) | $0.00054857990932(29)$ | $5.2 \times 10^{-10}$ |
| $f_{\mathrm{s}}(\mathrm{O}) / f_{\mathrm{c}}(\mathrm{O})$ | $B 17$ | GSI-02 | V.C.2.b (181) | $0.00054857990958(42)$ | $7.6 \times 10^{-10}$ |
| $\Delta \nu_{\overline{\mathrm{p}}} \mathrm{He}^{+}$ | $C 1-C 24$ | JINR/CERN-06 | IV.B.3 (74) | $0.00054857990881(91)$ | $1.7 \times 10^{-9}$ |
| $A_{\mathrm{r}}(\mathrm{e})$ | $B 9$ | UWash-95 | III.C (5) | $0.0005485799111(12)$ | $2.1 \times 10^{-9}$ |

ment of $V_{\mathrm{m}}(\mathrm{Si})$, item $B 53$, is inconsistent with the wattbalance measurements of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, items B36.1-B36.3, and somewhat inconsistent with the mercury-electrometer and voltage-balance measurements of $K_{\mathrm{J}}$; (ii) the XROI $\{220\}$ lattice spacing values $d_{220}(\mathrm{~W} 4.2 \mathrm{a}), d_{220}(\mathrm{~W} 4.2 \mathrm{a})$, and $d_{220}\left(\mathrm{MO}^{*}\right)$, items $B 38.1, B 38.2$, and $B 40$, are inconsistent with the value of $d_{220}(\mathrm{NR} 3)$, item $B 39$, and the measurement of $h / m_{\mathrm{n}} d_{220}$ (W04), item $B 55$; (iii) the NIST-89 measurement of $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$, item B33.1, is inconsistent with the most accurate data that also determine the value of the fine-structure constant; (iv) although not a problem in the sense of (i)-(iii), there are a number of input data with uncertainties so large that they are unlikely to make a contribution to the determination of the 2006 CODATA recommended values.

Furthermore, we note that some inferred values of $\alpha$ in Table XXXIV and most inferred values of $h$ in Table XXXV depend on either one or both of the relations $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$. The question of whether relaxing the assumption that these relations are exact would reduce or possibly even eliminate some of the observed inconsistencies, considered in Appendix F of CODATA02 , is addressed below. This study indeed confirms the Josephson and quantum Hall effect relations.

## B. Multivariate analysis of data

The multivariate analysis of the data is based on the fact that measured quantities can be expressed as theoretical functions of fundamental constants. These expressions, or observational equations, are written in terms of a particular independent subset of the constants whose members are called adjusted constants. The goal of the analysis is to find the values of the adjusted constants that predict values for the measured data that best agree with the data themselves in the least-squares sense (see Appendix E of CODATA-98).

The symbol $\doteq$ is used to indicate that an observed value of an input datum of the particular type shown on the left-hand side is ideally given by the function of the adjusted constants on the right-hand side; however, the two sides are not necessarily equal, because the equation is one of an overdetermined set relating the data to the adjusted constants. The best estimate of a quantity is given by its observational equation evaluated with the least-squares estimated values of the adjusted constants on which it depends.

In essence, we follow the least-squares approach of Aitken (1934) [see also Sheppard (1912)], who treated the case in which the input data are correlated. The 150 input data of Tables XXVIII, XXX, and XXXII are of 135 distinct types and are expressed as functions of the 79 adjusted constants listed in Tables XXXVII, XXXIX, and XLI. The observational equations that relate the input data to the adjusted constants are given in Tables XXXVIII, XL, and XLII.

Note that the various binding energies $E_{\mathrm{b}}(X) / m_{\mathrm{u}} c^{2}$ in Table XL, such as in the equation for item $B 1$, are treated as fixed quantities with negligible uncertainties. Similarly, the bound-state $g$-factor ratios in this table, such as in the equation for item $B 18$, are treated in the same way. Further, the frequency $f_{\mathrm{p}}$ is not an adjusted constant but is included in the equation for items $B 29$ and $B 30$ to indicate that they are functions of $f_{\mathrm{p}}$. Finally, the observational equation for items $B 29$ and $B 30$, based on Eqs. (215)-(217) of Sec. VI.B, includes the functions $a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)$ and $a_{\mu}\left(\alpha, \delta_{\mu}\right)$ as well as the theoretical expression for input data of Type $B 28, \Delta \nu_{\mathrm{Mu}}$. The latter expression is discussed in Sec. VI.B. 1 and is a function of $R_{\infty}, \alpha$, $m_{\mathrm{e}} / m_{\mu}, a_{\mu}\left(\alpha, \delta_{\mu}\right)$, and $\delta_{\mathrm{Mu}}$.

## 1. Summary of adjustments

A number of adjustments were carried out to gauge the compatibility of the input data in Tables XXVIII, XXX, and XXXII (together with their covariances in Tables XXIX, XXXI, and XXXIII) and to assess their influence on the values of the adjusted constants. The results of 11 of these are given in Tables XLIII and XLV and are discussed here. Because the adjusted value of the Rydberg constant $R_{\infty}$ is essentially the same for all six adjustments summarized in Table XLIII and equal to that of adjustment 4 of Table XLV, the value of $R_{\infty}$ is not listed in Table XLIII. It should also be noted that adjustment 4 of all three tables is the same adjustment.

Adjustment 1. This initial adjustment includes all input data, four of which have normalized residuals $r_{i}$ with absolute magnitudes significantly greater than 2 ; the values of $r_{i}$ for these four data resulting from adjustments 1-6 are given in Table XLIV. Consistent with the previous discussion, the four most inconsistent items are the molar volume of silicon $V_{\mathrm{m}}(\mathrm{Si})$, the quotient $h / m_{\mathrm{n}} d_{220}$ (W04), the XROI measurement of the $\{220\}$ lattice spacing $d_{220}($ NR3 $)$, and the NIST- 89 value of

TABLE XXXVII. The 28 adjusted constants (variables) used in the least-squares multivariate analysis of the Rydbergconstant data given in Table XXVIII. These adjusted constants appear as arguments of the functions on the right-hand side of the observational equations of Table XXXVIII. The notation for hydrogenic energy levels $E_{X}\left(n \mathrm{~L}_{j}\right)$ and for additive corrections $\delta_{X}\left(n \mathrm{~L}_{j}\right)$ have the same meaning as the notations $E_{n \mathrm{~L} j}^{X}$ and $\delta_{n \mathrm{~L} j}^{X}$ in Sec. IV.A.1.1.

| Adjusted constant | Symbol |
| :--- | :--- |
| Rydberg constant | $R_{\infty}$ |
| bound-state proton rms charge radius | $R_{\mathrm{p}}$ |
| bound-state deuteron rms charge radius | $R_{\mathrm{d}}$ |
| additive correction to $E_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(3 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(3 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(4 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(4 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(6 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(6 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(8 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(8 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(2 \mathrm{P}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(4 \mathrm{P}_{1 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(4 \mathrm{P}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(2 \mathrm{P}_{3 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(2 \mathrm{P}_{3 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(4 \mathrm{P}_{3 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(4 \mathrm{P}_{3 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(8 \mathrm{D}_{3 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(8 \mathrm{D}_{3 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(12 \mathrm{D}_{3 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(12 \mathrm{D}_{3 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(4 \mathrm{D}_{5 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(4 \mathrm{D}_{5 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(6 \mathrm{D}_{5 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(6 \mathrm{D}_{5 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(8 \mathrm{D}_{5 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(8 \mathrm{D}_{5 / 2}\right)$ |
| additive correction to $E_{\mathrm{H}}\left(12 \mathrm{D}_{5 / 2}\right) / h$ | $\delta_{\mathrm{H}}\left(12 \mathrm{D}_{5 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(4 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(4 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(8 \mathrm{~S}_{1 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(8 \mathrm{~S}_{1 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(8 \mathrm{D}_{3 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(8 \mathrm{D}_{3 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(12 \mathrm{D}_{3 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(12 \mathrm{D}_{3 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(4 \mathrm{D}_{5 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(4 \mathrm{D}_{5 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(8 \mathrm{D}_{5 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(8 \mathrm{D}_{5 / 2}\right)$ |
| additive correction to $E_{\mathrm{D}}\left(12 \mathrm{D}_{5 / 2}\right) / h$ | $\delta_{\mathrm{D}}\left(12 \mathrm{D}_{5 / 2}\right)$ |

$\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$. All other input data have values of $r_{i}$ considerably less than 2 , except those for $\nu_{\bar{p}}{ }^{3} \mathrm{He}(32,31: 31,30)$ and $\nu_{\bar{p}}{ }^{3} \mathrm{He}(36,33: 34,32)$, items $C 20$ and $C 22$, for which $r_{20}$ $=2.09$ and $r_{22}=2.06$. However, the self-sensitivity coefficients $S_{\mathrm{c}}$ for these input data are considerably less than 0.01 ; hence, because their contribution to the adjustment is small, their marginally large normalized residuals are of little concern. In this regard, we see from Table XLIV that three of the four inconsistent data have values of $S_{\mathrm{c}}$ considerably larger than 0.01 ; the exception is $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ with $S_{\mathrm{c}}=0.0099$, which is rounded to 0.010 in the table.
Adjustment 2. Since the four direct lattice spacing measurements, items B38.1-B40, are credible, as is the measurement of $h / m_{\mathrm{n}} d_{220}$ (W04), item B55, after due consideration the CODATA Task Group on Fundamental Constants decided that all five of these input data
should be considered for retention, but that each of their a priori assigned uncertainties should be weighted by the multiplicative factor 1.5 to reduce $\left|r_{\mathrm{i}}\right|$ of $h / m_{\mathrm{n}} d_{220}$ (W04) and of $d_{220}(\mathrm{NR} 3)$ to a more acceptable level, that is, to about 2, while maintaining their relative weights. This has been done in adjustment 2 . As can be seen from Table XLIII, this increase of uncertainties has an inconsequential impact on the value of $\alpha$, and no impact on the value of $h$. It does reduce $R_{\mathrm{B}}$, as would be expected.

Adjustment 3. Again, since the measurement of $V_{\mathrm{m}}(\mathrm{Si})$, item $B 53$, as well as the three measurements of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, items B36.1-B36.3, and the two measurements of $K_{\mathrm{J}}$, items B35.1 and B35.2, are credible, the Task Group decided that all six should be considered for retention, but that each of their a priori assigned uncertainties should be weighted by the multiplicative factor 1.5 to reduce $\left|r_{i}\right|$ of $V_{\mathrm{m}}(\mathrm{Si})$ to about 2, while maintaining their relative weights. This has been done in adjustment 3 . Note that this also reduces $\left|r_{i}\right|$ of $h / m_{\mathrm{n}} d_{220}(\mathrm{~W} 04)$ from 2.03 in adjustment 2 to 1.89 in adjustment 3 . We see from Table XLIII that this increase in uncertainty has negligible consequences for the value of $\alpha$, but it does increase the uncertainty of $h$ by about the same factor, as would be expected. Also as would be expected, $R_{\mathrm{B}}$ is further reduced.

It may be recalled that faced with a similar situation in the 2002 adjustment, the Task Group decided to use a multiplicative weighting factor of 2.325 in order to reduce $\left|r_{i}\right|$ of $V_{\mathrm{m}}(\mathrm{Si})$ to 1.50 . The reduced weighting factor of 1.5 in the 2006 adjustment recognizes the new value of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ now available and the excellent agreement with the two earlier values.

Adjustment 4. In adjustment 3, a number of input data, as measured by their self-sensitivity coefficients $S_{c}$, do not contribute in a significant way to the determination of the adjusted constants. We therefore omit in adjustment 4 those input data with $S_{\mathrm{c}}<0.01$ in adjustment 3 unless they are a subset of the data of an experiment that provides other input data with $S_{\mathrm{c}}>0.01$. The 14 input data deleted in adjustment 4 for this reason are $B 31.1-B 35.2, B 37$, and $B 56$, which are the five low- and high-field proton and helion gyromagnetic ratio results; the five calculable capacitor values of $R_{\mathrm{K}}$; both values of $K_{\mathrm{J}}$ as obtained using a Hg electrometer and a voltage balance; the Ag coulometer result for the Faraday constant; and the atom interferometry result for the quotient of the Planck constant and mass of the cesium-133 atom. The respective values of $S_{\mathrm{c}}$ for these data in adjustment 3 are in the range $0.0000-0.0099$. Deleting such marginal data is in keeping with the practice followed in the 1998 and 2002 adjustments; see Sec. I.D of CODATA-98.

Because $h / m\left({ }^{133} \mathrm{Cs}\right)$, item $B 56$, has been deleted as an input datum due to its low weight, $A_{\mathrm{r}}\left({ }^{133} \mathrm{Cs}\right)$, item $B 8$, which is not coupled to any other input datum, has also been omitted as an input datum and as an adjusted constant from adjustment 4. This brings the total number of omitted items to 15 . Table XLIII shows that deleting

TABLE XXXVIII. Observational equations that express the input data related to $R_{\infty}$ in Table XXVIII as functions of the adjusted constants in Table XXXVII. The numbers in the first column correspond to the numbers in the first column of Table XXVIII. Expressions for the energy levels of hydrogenic atoms are discussed in Sec. IV.A.1. As pointed out in Sec. IV.A.1.1, $E_{X}\left(n \mathrm{~L}_{j}\right) / h$ is in fact proportional to $c R_{\infty}$ and independent of $h$, hence $h$ is not an adjusted constant in these equations. The notation for hydrogenic energy levels $E_{X}\left(n \mathrm{~L}_{j}\right)$ and for additive corrections $\delta_{X}\left(n \mathrm{~L}_{j}\right)$ have the same meaning as the notations $E_{n \mathrm{~L} j}^{X}$ and $\delta_{n \mathrm{~L} j}^{X}$ in Sec. IV.A.1.1. See Sec. XII.B for an explanation of the symbol $\doteq$.

Type of input
datum
Observational equation

| A1-A16 | $\delta_{\mathrm{H}}\left(n \mathrm{~L}_{j}\right) \doteq \delta_{\mathrm{H}}\left(n \mathrm{~L}_{j}\right)$ |
| :---: | :---: |
| A17-A25 | $\delta_{\mathrm{D}}\left(n \mathrm{~L}_{j}\right) \doteq \delta_{\mathrm{D}}\left(n \mathrm{~L}_{j}\right)$ |
| $\begin{aligned} & A 26-A 31, \\ & \quad A 38, A 39 \end{aligned}$ | $\begin{aligned} \nu_{\mathrm{H}}\left(n_{1} \mathrm{~L}_{1 j_{1}}-n_{2} \mathrm{~L}_{2 j_{2}}\right) \doteq & {\left[E_{\mathrm{H}}\left(n_{2} \mathrm{~L}_{2 j_{2}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), R_{\mathrm{p}}, \delta_{\mathrm{H}}\left(n_{2} \mathrm{~L}_{2 j_{2}}\right)\right)\right.} \\ & \left.-E_{\mathrm{H}}\left(n_{1} \mathrm{~L}_{1 j_{1}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), R_{\mathrm{p}}, \delta_{\mathrm{H}}\left(n_{1} \mathrm{~L}_{1 j_{1}}\right)\right)\right] / h \end{aligned}$ |
| A32-A37 | $\begin{aligned} \nu_{\mathrm{H}}\left(n_{1} \mathrm{~L}_{1 j_{1}}-n_{2} \mathrm{~L}_{2 j_{2}}\right)-\frac{1}{4} \nu_{\mathrm{H}}\left(n_{3} \mathrm{~L}_{3 j_{3}}-n_{4} \mathrm{~L}_{4 j_{4}}\right) \doteq & \left\{E_{\mathrm{H}}\left(n_{2} \mathrm{~L}_{2 j_{2}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), R_{\mathrm{p}}, \delta_{\mathrm{H}}\left(n_{2} \mathrm{~L}_{2 j_{2}}\right)\right)\right. \\ & -E_{\mathrm{H}}\left(n_{1} \mathrm{~L}_{1 j_{1}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), R_{\mathrm{p}}, \delta_{\mathrm{H}}\left(n_{1} \mathrm{~L}_{1 j_{1}}\right)\right) \\ & -\frac{1}{4}\left[E_{\mathrm{H}}\left(n_{4} \mathrm{~L}_{4 j_{4}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), R_{\mathrm{p}}, \delta_{\mathrm{H}}\left(n_{4} \mathrm{~L}_{4 j_{4}}\right)\right)\right. \\ & \left.\left.-E_{\mathrm{H}}\left(n_{3} \mathrm{~L}_{3 j_{3}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), R_{\mathrm{p}}, \delta_{\mathrm{H}}\left(n_{3} \mathrm{~L}_{3 j_{3}}\right)\right)\right]\right\} / h \end{aligned}$ |
| A40-A44 | $\begin{aligned} \nu_{\mathrm{D}}\left(n_{1} \mathrm{~L}_{1 j_{1}}-n_{2} \mathrm{~L}_{2 j_{2}}\right) \doteq & {\left[E_{\mathrm{D}}\left(n_{2} \mathrm{~L}_{2 j_{j}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{~d}), R_{\mathrm{d}}, \delta_{\mathrm{D}}\left(n_{2} \mathrm{~L}_{2 j_{2}}\right)\right)\right.} \\ & \left.-E_{\mathrm{D}}\left(n_{1} \mathrm{~L}_{1 j_{1}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{~d}), R_{\mathrm{d}}, \delta_{\mathrm{D}}\left(n_{1} \mathrm{~L}_{1 j_{1}}\right)\right)\right] / h \end{aligned}$ |
| A45, $A 46$ | $\begin{aligned} \nu_{\mathrm{D}}\left(n_{1} \mathrm{~L}_{1 j_{1}}-n_{2} \mathrm{~L}_{2 j_{2}}\right)-\frac{1}{4} \nu_{\mathrm{D}}\left(n_{3} \mathrm{~L}_{3 j_{3}}-n_{4} \mathrm{~L}_{4 j_{4}}\right) \doteq & \left\{E_{\mathrm{D}}\left(n_{2} \mathrm{~L}_{2 j_{2}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{~d}), R_{\mathrm{d}}, \delta_{\mathrm{D}}\left(n_{2} \mathrm{~L}_{2 j_{2}}\right)\right)\right. \\ & -E_{\mathrm{D}}\left(n_{1} \mathrm{~L}_{1 j_{j}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{~d}), R_{\mathrm{d}}, \delta_{\mathrm{D}}\left(n_{1} \mathrm{~L}_{1 j_{1}}\right)\right) \\ & -\frac{1}{4}\left[E_{\mathrm{D}}\left(n_{4} \mathrm{~L}_{4 j_{4}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{~d}), R_{\mathrm{d}}, \delta_{\mathrm{D}}\left(n_{4} \mathrm{~L}_{4 j_{4}}\right)\right)\right. \\ & \left.\left.-E_{\mathrm{D}}\left(n_{3} \mathrm{~L}_{3 j_{3}} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{~d}), R_{\mathrm{d}}, \delta_{\mathrm{D}}\left(n_{3} \mathrm{~L}_{3 j_{3}}\right)\right)\right]\right\} / h \end{aligned}$ |
| A47 | $\begin{aligned} \nu_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right)-\nu_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}\right) \doteq & \left\{E_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{~d}), R_{\mathrm{d}}, \delta_{\mathrm{D}}\left(2 \mathrm{~S}_{1 / 2}\right)\right)\right. \\ & -E_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{~d}), R_{\mathrm{d}}, \delta_{\mathrm{D}}\left(1 \mathrm{~S}_{1 / 2}\right)\right) \\ & -\left[E_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), R_{\mathrm{p}}, \delta_{\mathrm{H}}\left(2 \mathrm{~S}_{1 / 2}\right)\right)\right. \\ & \left.\left.-E_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2} ; R_{\infty}, \alpha, A_{\mathrm{r}}(\mathrm{e}), A_{\mathrm{r}}(\mathrm{p}), R_{\mathrm{p}}, \delta_{\mathrm{H}}\left(1 \mathrm{~S}_{1 / 2}\right)\right)\right]\right\} / h \end{aligned}$ |
| A48 | $R_{\mathrm{p}} \doteq R_{\mathrm{p}}$ |
| A49 | $R_{\text {d }} \doteq R_{\text {d }}$ |

these 15 data has virtually no impact on the values of $\alpha$ and $h$.

Adjustment 4 is the adjustment on which the 2006 CODATA recommended values are based, and as such it is referred to as the "final adjustment."

Adjustments 5 and 6. These adjustments are intended to check the robustness of adjustment 4 , the final adjustment, while adjustments 7-11, which are summarized in Table XLV, probe various aspects of the $R_{\infty}$ data in Table XXVIII.

Adjustment 5 only differs from adjustment 3 in that it does not include the input data that lead to the four most accurate values of $\alpha$ : the two measurements of $a_{\mathrm{e}}$,
items $B 11.1$ and $B 11.2$, the measurement of $h / m\left({ }^{133} \mathrm{Cs}\right)$, item $B 56$, and the measurement of $h / m\left({ }^{87} \mathrm{Rb}\right)$, item $B 57$. The $u_{\mathrm{r}}$ of the inferred values of $\alpha$ from these data are $7.0 \times 10^{-10}, 3.7 \times 10^{-9}, 7.7 \times 10^{-9}$, and $6.7 \times 10^{-9}$. We see from Table XLIII that the value of $\alpha$ from adjustment 5 is consistent with the 2006 recommended value from adjustment 4 (the difference is $0.8 u_{\text {diff }}$ ), but its uncertainty is about 20 times larger. Moreover, the resulting value of $h$ is the same as the recommended value.

Adjustment 6 only differs from adjustment 3 in that it does not include the input data that yield the three most accurate values of $h$, namely, the watt-balance measure-

TABLE XXXIX. The 39 adjusted constants (variables) used in the least-squares multivariate analysis of the input data in Table XXX. These adjusted constants appear as arguments of the functions on the right-hand side of the observational equations of Table XL.

| Adjusted constant | Symbol |
| :---: | :---: |
| electron relative atomic mass | $A_{\mathrm{r}}(\mathrm{e})$ |
| proton relative atomic mass | $A_{\mathrm{r}}(\mathrm{p})$ |
| neutron relative atomic mass | $A_{\mathrm{r}}(\mathrm{n})$ |
| deuteron relative atomic mass | $A_{\mathrm{r}}(\mathrm{d})$ |
| triton relative atomic mass | $A_{\mathrm{r}}(\mathrm{t})$ |
| helion relative atomic mass | $A_{\mathrm{r}}(\mathrm{h})$ |
| $\alpha$ particle relative atomic mass | $A_{\mathrm{r}}(\alpha)$ |
| ${ }^{16} \mathrm{O}^{7+}$ relative atomic mass | $A_{\mathrm{r}}\left({ }^{16} \mathrm{O}^{7+}\right)$ |
| ${ }^{87} \mathrm{Rb}$ relative atomic mass | $A_{\mathrm{r}}\left({ }^{87} \mathrm{Rb}\right)$ |
| ${ }^{133} \mathrm{Cs}$ relative atomic mass | $A_{\mathrm{r}}\left({ }^{133} \mathrm{Cs}\right)$ |
| fine-structure constant | $\alpha$ |
| additive correction to $a_{\mathrm{e}}$ (th) | $\delta_{\text {e }}$ |
| additive correction to $a_{\mu}($ th $)$ | $\delta_{\mu}$ |
| additive correction to $g_{C}($ th $)$ | $\delta_{\text {C }}$ |
| additive correction to $g_{O}($ th $)$ | $\delta_{\mathrm{O}}$ |
| electron-proton magnetic moment ratio | $\mu_{\mathrm{e}^{-} /} / \mu_{\mathrm{p}}$ |
| deuteron-electron magnetic moment ratio | $\mu_{\mathrm{d}} / \mu_{\mathrm{e}^{-}}$ |
| triton-proton magnetic moment ratio | $\mu_{\mathrm{t}} / \mu_{\mathrm{p}}$ |
| shielding difference of $d$ and $p$ in HD | $\sigma_{\text {dp }}$ |
| shielding difference of $t$ and $p$ in HT | $\sigma_{\text {tp }}$ |
| electron to shielded proton magnetic moment ratio | $\mu_{\mathrm{e}-} / \mu_{\mathrm{p}}^{\prime}$ |
| shielded helion to shielded proton magnetic moment ratio | $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime}$ |
| neutron to shielded proton magnetic moment ratio | $\mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime}$ |
| electron-muon mass ratio | $m_{\mathrm{e}} / m_{\mu}$ |
| additive correction to $\Delta \nu_{\mathrm{Mu}}$ (th) | $\delta_{\text {Mu }}$ |
| Planck constant | $h$ |
| molar gas constant | $R$ |
| copper $\mathrm{K} \alpha_{1} \mathrm{x}$ unit | $\mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)$ |
| molybdenum $\mathrm{K} \alpha_{1} \mathrm{x}$ unit | $\mathrm{xu}\left(\operatorname{MoK} \alpha_{1}\right)$ |
| ångström star | A* |
| $d_{220}$ of Si crystal ILL | $d_{220}$ (ILL) |
| $d_{220}$ of Si crystal N | $d_{220}(\mathrm{~N})$ |
| $d_{220}$ of Si crystal WASO 17 | $d_{220}$ (W17) |
| $d_{220}$ of Si crystal WASO 04 | $d_{220}$ (W04) |
| $d_{220}$ of Si crystal WASO 4.2a | $d_{220}$ (W4.2a) |
| $d_{220}$ of Si crystal MO* | $d_{220}\left(\mathrm{MO}^{*}\right)$ |
| $d_{220}$ of Si crystal NR3 | $d_{220}$ (NR3) |
| $d_{220}$ of Si crystal NR4 | $d_{220}$ (NR4) |
| $d_{220}$ of an ideal Si crystal | $d_{220}$ |

TABLE XL. Observational equations that express the input data in Table XXX as functions of the adjusted constants in Table XXXIX. The numbers in the first column correspond to the numbers in the first column of Table XXX. For simplicity, the lengthier functions are not explicitly given. See Sec. XII.B for an explanation of the symbol $\doteq$.

| Type of input datum | Observational equation | Sec. |
| :---: | :---: | :---: |
| B1 | $A_{\mathrm{r}}\left({ }^{1} \mathrm{H}\right) \doteq A_{\mathrm{r}}(\mathrm{p})+A_{\mathrm{r}}(\mathrm{e})-E_{\mathrm{b}}\left({ }^{1} \mathrm{H}\right) / m_{\mathrm{u}} c^{2}$ | III.B |
| B2 | $A_{\mathrm{r}}\left({ }^{2} \mathrm{H}\right) \doteq A_{\mathrm{r}}(\mathrm{d})+A_{\mathrm{r}}(\mathrm{e})-E_{\mathrm{b}}\left({ }^{2} \mathrm{H}\right) / m_{\mathrm{u}} c^{2}$ | III.B |
| B3 | $A_{\mathrm{r}}\left({ }^{3} \mathrm{H}\right) \doteq A_{\mathrm{r}}(\mathrm{t})+A_{\mathrm{r}}(\mathrm{e})-E_{\mathrm{b}}\left({ }^{3} \mathrm{H}\right) / m_{\mathrm{u}} c^{2}$ | III.B |
| B4 | $A_{\mathrm{r}}\left({ }^{3} \mathrm{He}\right) \doteq A_{\mathrm{r}}(\mathrm{h})+2 A_{\mathrm{r}}(\mathrm{e})-E_{\mathrm{b}}\left({ }^{3} \mathrm{He}\right) / m_{\mathrm{u}} c^{2}$ | III.B |
| B5 | $A_{\mathrm{r}}\left({ }^{4} \mathrm{He}\right) \doteq A_{\mathrm{r}}(\alpha)+2 A_{\mathrm{r}}(\mathrm{e})-E_{\mathrm{b}}\left({ }^{4} \mathrm{He}\right) / m_{\mathrm{u}} c^{2}$ | III.B |
| B6 | $\begin{aligned} A_{\mathrm{r}}\left({ }^{16} \mathrm{O}\right) \doteq & A_{\mathrm{r}}\left({ }^{(16} \mathrm{O}^{7+}\right)+7 A_{\mathrm{r}}(\mathrm{e})-\left[E_{\mathrm{b}}\left({ }^{16} \mathrm{O}\right)\right. \\ & \left.-E_{\mathrm{b}}\left({ }^{16} \mathrm{O}^{7+}\right)\right] / m_{\mathrm{u}} c^{2} \end{aligned}$ | V.C.2.b |
| B7 | $A_{\mathrm{r}}\left({ }^{87} \mathrm{Rb}\right) \doteq A_{\mathrm{r}}\left({ }^{87} \mathrm{Rb}\right)$ |  |
| B8 | $A_{\mathrm{r}}\left({ }^{133} \mathrm{Cs}\right) \doteq A_{\mathrm{r}}\left({ }^{133} \mathrm{Cs}\right)$ |  |
| B9 | $A_{\mathrm{r}}(\mathrm{e}) \doteq A_{\mathrm{r}}(\mathrm{e})$ |  |
| B10 | $\delta_{\mathrm{e}} \doteq \delta_{\mathrm{e}}$ |  |
| B11 | $a_{\mathrm{e}} \doteq a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)$ | V.A. 1 |
| B12 | $\delta_{\mu} \doteq \delta_{\mu}$ |  |
| B13 | $\bar{R} \doteq-\frac{a_{\mu}\left(\alpha, \delta_{\mu}\right)}{1+a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)} \frac{m_{\mathrm{e}}}{m_{\mu}} \frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}}$ | V.B. 2 |
| B14 | $\delta_{\mathrm{C}} \doteq \delta_{\mathrm{C}}$ |  |
| B15 | $\delta_{\mathrm{O}} \doteq \delta_{\mathrm{O}}$ |  |
| B16 | $\begin{aligned} \frac{f_{\mathrm{s}}\left({ }^{12} \mathrm{C}^{5+}\right)}{f_{\mathrm{c}}\left({ }^{12} \mathrm{C}^{5+}\right)} \doteq & -\frac{g_{\mathrm{C}}\left(\alpha, \delta_{\mathrm{C}}\right)}{10 A_{\mathrm{r}}(\mathrm{e})}\left[12-5 A_{\mathrm{r}}(\mathrm{e})\right. \\ & \left.+\frac{E_{\mathrm{b}}\left({ }^{12} \mathrm{C}\right)-E_{\mathrm{b}}\left({ }^{12} \mathrm{C}^{5+}\right)}{m_{\mathrm{u}} c^{2}}\right] \end{aligned}$ | V.C.2.a |
| B17 | $\frac{f_{\mathrm{s}}\left({ }^{16} \mathrm{O}^{7+}\right)}{f_{\mathrm{c}}\left({ }^{16} \mathrm{O}^{7+}\right)} \doteq-\frac{g_{\mathrm{O}}\left(\alpha, \delta_{\mathrm{O}}\right)}{14 A_{\mathrm{r}}(\mathrm{e})} A_{\mathrm{r}}\left({ }^{16} \mathrm{O}^{7+}\right)$ | V.C.2.b |
| B18 | $\frac{\mu_{\mathrm{e}^{-}}(\mathrm{H})}{\mu_{\mathrm{p}}(\mathrm{H})} \doteq \frac{g_{\mathrm{e}^{-}}(\mathrm{H})}{g_{\mathrm{e}^{-}}}\left(\frac{g_{\mathrm{p}}(\mathrm{H})}{g_{\mathrm{p}}}\right)^{-1} \frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}}$ | VI.A.2.a |

TABLE XL. (Continued.)

| Type of input datum | Observational equation | Sec. |
| :---: | :---: | :---: |
| $B 19$ |  | V.A.2.b |
| $B 20$ | $\frac{\mu_{\mathrm{p}}(\mathrm{HD})}{\mu_{\mathrm{d}}(\mathrm{HD})} \doteq\left[\begin{array}{c} \\ \text { dem }\end{array} \sigma_{\mathrm{dp}}\left(\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}}\right)^{-1}\left(\frac{\mu_{\mathrm{d}}}{\mu_{\mathrm{e}^{-}}}\right)^{-1}\right.$ | VI.A.2.c |
| $B 21$ | $\sigma_{\text {dp }} \doteq \sigma_{\text {dp }}$ |  |
| $B 22$ | $\frac{\mu_{\mathrm{t}}(\mathrm{HT})}{\mu_{\mathrm{p}}(\mathrm{HT})} \doteq\left[1-\sigma_{\mathrm{tp}}\right] \frac{\mu_{\mathrm{t}}}{\mu_{\mathrm{p}}}$ | VI.A.2.c |
| B23 | $\sigma_{\text {tp }} \doteq \sigma_{\text {tp }}$ |  |
| B24 | $\frac{\mu_{\mathrm{e}^{-}}(\mathrm{H})}{\mu_{\mathrm{p}}^{\prime}} \doteq \frac{g_{\mathrm{e}^{-}}(\mathrm{H})}{g_{\mathrm{e}^{-}}} \frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}^{\prime}}$ | VI.A.2.d |
| B25 | $\frac{\mu_{\mathrm{h}}^{\prime}}{\mu_{\mathrm{p}}^{\prime}} \doteq \frac{\mu_{\mathrm{h}}^{\prime}}{\mu_{\mathrm{p}}^{\prime}}$ |  |
| B26 | $\frac{\mu_{\mathrm{n}}}{\mu_{\mathrm{p}}^{\prime}}=\frac{\mu_{\mathrm{n}}}{\mu_{\mathrm{p}}^{\prime}}$ |  |
| B27 | $\delta_{\mathrm{Mu}} \doteq \delta_{\mathrm{Mu}}$ |  |
| B28 | $\Delta \nu_{\mathrm{Mu}} \doteq \Delta \nu_{\mathrm{Mu}}\left(R_{\infty}, \alpha, \frac{m_{\mathrm{e}}}{m_{\mu}}, \delta_{\mu}, \delta_{\mathrm{Mu}}\right)$ | VI.B. 1 |
| B29, B30 | $\nu\left(f_{\mathrm{p}}\right) \doteq \nu\left(f_{\mathrm{p}} ; R_{\infty}, \alpha, \frac{m_{\mathrm{e}}}{m_{\mu}}, \frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}}, \delta_{\mathrm{e}}, \delta_{\mu}, \delta_{\mathrm{Mu}}\right)$ | VI.B |
| B31 | $\begin{aligned} \Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo}) \doteq & -\frac{K_{\mathrm{J}-90} R_{\mathrm{K}-99}\left[1+a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)\right] \alpha^{3}}{2 \mu_{0} R_{\infty}} \\ & \times\left(\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}^{\prime}}\right)^{-1} \end{aligned}$ | VII.A. 1 |
| B32 | $\begin{aligned} \Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo}) \doteq & \frac{K_{\mathrm{J}-90} R_{\mathrm{K}-90}\left[1+a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)\right] \alpha^{3}}{2 \mu_{0} R_{\infty}} \\ & \times\left(\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}^{\prime}}\right)^{-1} \frac{\mu_{\mathrm{h}}^{\prime}}{\mu_{\mathrm{p}}^{\prime}} \end{aligned}$ | VII.A. 1 |
| B33 | $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi}) \doteq-\frac{c\left[1+a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)\right] \alpha^{2}}{K_{\mathrm{J}-90} R_{\mathrm{K}-90} R_{\propto} h}\left(\frac{\mu_{\mathrm{e}-}}{\mu_{\mathrm{p}}^{\prime}}\right)^{-1}$ | VII.A. 2 |
| B34 | $R_{\mathrm{K}} \doteq \frac{\mu_{0} c}{2 \alpha}$ | VII.B |

TABLE XL. (Continued.)

| Type of input datum | Observational equation | Sec. |
| :---: | :---: | :---: |
| $B 35$ | $K_{\mathrm{J}} \doteq\left(\frac{8 \alpha}{\mu_{0} c h}\right)^{1 / 2}$ | VII.C |
| B36 | $K_{\mathrm{J}}^{2} R_{\mathrm{K}} \doteq \frac{4}{h}$ | VII.D |
| B37 | $\mathcal{F}_{90} \doteq \frac{c M_{\mathrm{u}} A_{\mathrm{r}}(\mathrm{e}) \alpha^{2}}{K_{\mathrm{J}-90} R_{\mathrm{K}-90} R_{\infty} h}$ | VII.E |
| B38-B40 | $d_{220}(X) \doteq d_{220}(X)$ |  |
| B41-B52 | $\frac{d_{220}(X)}{d_{220}(Y)}-1 \doteq \frac{d_{220}(X)}{d_{220}(Y)}-1$ |  |
| B53 | $V_{\mathrm{m}}(\mathrm{Si}) \doteq \frac{\sqrt{2} c M_{\mathrm{u}} A_{\mathrm{r}}(\mathrm{e}) \alpha^{2} d_{220}^{3}}{R_{\infty} h}$ | VIII.B |
| B54 | $\begin{aligned} & \frac{\lambda_{\text {meas }}}{d_{220}(\mathrm{ILL})} \\ & \doteq \frac{\alpha^{2} A_{\mathrm{r}}(\mathrm{e})}{R_{\infty} d_{220}(\mathrm{ILL})} \frac{A_{\mathrm{r}}(\mathrm{n})+A_{\mathrm{r}}(\mathrm{p})}{\left[A_{\mathrm{r}}(\mathrm{n})+A_{\mathrm{r}}(\mathrm{p})\right]^{2}-A_{\mathrm{r}}^{2}(\mathrm{~d})} \end{aligned}$ | VIII.C |
| B55 | $\frac{h}{m_{\mathrm{n}} d_{220}(\mathrm{~W} 04)} \doteq \frac{A_{\mathrm{r}}(\mathrm{e})}{A_{\mathrm{r}}(\mathrm{n})} \frac{c \alpha^{2}}{2 R_{\infty} d_{220}(\mathrm{~W} 04)}$ | VIII.D. 1 |
| B56, B57 | $\frac{h}{m(X)} \doteq \frac{A_{\mathrm{r}}(\mathrm{e})}{A_{\mathrm{r}}(X)} \frac{c \alpha^{2}}{2 R_{\infty}}$ | VIII.D |
| B58 | $R \doteq$ ¢ |  |
| B59,B62 | $\frac{\lambda\left(\mathrm{CuK} \alpha_{1}\right)}{d_{220}(X)} \doteq \frac{1537.400 \mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)}{d_{220}(X)}$ | XI.A |
| B60 | $\frac{\lambda\left(\mathrm{WK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})} \doteq \frac{0.2090100 \AA^{*}}{d_{220}(\mathrm{~N})}$ | XI.A |
| B61 | $\frac{\lambda\left(\mathrm{MoK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})} \doteq \frac{707.831 \mathrm{xu}\left(\mathrm{MoK} \alpha_{1}\right)}{d_{220}(\mathrm{~N})}$ | XI.A |

ments of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, items B36.1-B36.3. The $u_{\mathrm{r}}$ of the inferred values of $h$ from these data, as they are used in adjustment 3 (that is, after their uncertainties are multiplied by the weighting factor 1.5 ), are $5.4 \times 10^{-8}, 1.3$ $\times 10^{-7}$, and $3.0 \times 10^{-7}$. From Table XLIII, we see that the value of $h$ from adjustment 6 is consistent with the 2006 recommended value from adjustment 4 (the difference is $1.4 u_{\text {diff }}$, but its uncertainty is well over six times larger. Furthermore, the resulting value of $\alpha$ is the same as the recommended value. Therefore, adjustments 5 and 6 suggest that the less accurate input data are consistent with the more accurate data, thereby providing a consistency check on the 2006 recommended values of the constants.

TABLE XLI. The 12 adjusted constants (variables) relevant to the antiprotonic helium data given in Table XXXII. These adjusted constants appear as arguments of the theoretical expressions on the right-hand side of the observational equations of Table XLII.

| Transition | Adjusted constant |
| :---: | :---: |
| $\overline{\mathrm{p}}{ }^{4} \mathrm{He}^{+}:(32,31) \rightarrow(31,30)$ | $\delta_{\overline{\mathrm{p}}}{ }^{\text {He}}{ }^{+}(32,31: 31,30)$ |
| $\overline{\mathrm{p}}{ }^{4} \mathrm{He}^{+}:(35,33) \rightarrow(34,32)$ | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}{ }^{+}(35,33: 34,32)$ |
| $\overline{\mathrm{p}}{ }^{4} \mathrm{He}^{+}:(36,34) \rightarrow(35,33)$ | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}{ }^{+}(36,34: 35,33)$ |
| $\overline{\mathrm{p}}{ }^{4} \mathrm{He}^{+}:(37,34) \rightarrow(36,33)$ | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}{ }^{+}(37,34: 36,33)$ |
| $\overline{\mathrm{p}}{ }^{4} \mathrm{He}^{+}:(39,35) \rightarrow(38,34)$ | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}{ }^{+}(39,35: 38,34)$ |
| $\overline{\mathrm{p}}^{4} \mathrm{He}^{+}:(40,35) \rightarrow(39,34)$ | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}^{+}(40,35: 39,34)$ |
| $\overline{\mathrm{p}}{ }^{4} \mathrm{He}^{+}:(37,35) \rightarrow(38,34)$ | $\delta_{\overline{\mathrm{p}}}{ }^{4} \mathrm{He}{ }^{+}(37,35: 38,34)$ |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(32,31) \rightarrow(31,30)$ | $\delta_{\overline{\mathrm{p}}}{ }^{3} \mathrm{He}^{+}(32,31: 31,30)$ |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(34,32) \rightarrow(33,31)$ | $\delta_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(34,32: 33,31)}$ |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(36,33) \rightarrow(35,32)$ | $\delta_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(36,33: 35,32)}$ |
| $\bar{p}^{3} \mathrm{He}^{+}:(38,34) \rightarrow(37,33)$ | $\delta_{\overline{\mathrm{p}}}{ }^{3} \mathrm{He}^{+}(38,34: 37,33)$ |
| $\overline{\mathrm{p}}^{3} \mathrm{He}^{+}:(36,34) \rightarrow(37,33)$ | $\delta_{\overline{\mathrm{p}}{ }^{3} \mathrm{He}^{+}(36,34: 37,33)}$ |

Adjustments 7-11. These adjustments differ from adjustment 4, the final adjustment, in the following ways. In adjustment 7, the scattering-data input values for both $R_{\mathrm{p}}$ and $R_{\mathrm{d}}$, items $A 48$ and $A 49$, are omitted; in adjustment 8 , only $R_{\mathrm{p}}$ is omitted, and in adjustment 9 , only $R_{\mathrm{d}}$ is omitted; adjustment 10 includes only the hydrogen data, and adjustment 11 includes only the deuterium data, but for both, the H-D isotope shift, item $A 47$, is omitted. Although a somewhat improved value of the ${ }^{1} \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}$ hydrogen transition frequency and improvements in the theory of H and D energy levels have be-


FIG. 7. Values of the Planck constant $h$ implied by the input data in Table XXX, taken as a weighted mean when more than one measurement of a given type is considered [see Eqs. (367)-(371)], in order of decreasing uncertainty from top to bottom.
come available since the completion of the 2002 adjustment, the value of $R_{\infty}$, which is determined almost entirely by these data, has changed very little. The values of $R_{\mathrm{p}}$ and $R_{\mathrm{d}}$, which are also determined mainly by these data, have changed by less than one-third of their uncertainties. The experimental and theoretical H and D data remain highly consistent.

## 2. Test of the Josephson and quantum Hall effect relations

Investigation of the exactness of the relations $K_{\mathrm{J}}$ $=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ is carried out, as in CODATA-02, by writing

$$
\begin{align*}
& K_{\mathrm{J}}=\frac{2 e}{h}\left(1+\varepsilon_{\mathrm{J}}\right)=\left(\frac{8 \alpha}{\mu_{0} c h}\right)^{1 / 2}\left(1+\varepsilon_{\mathrm{J}}\right),  \tag{372}\\
& R_{\mathrm{K}}=\frac{h}{e^{2}}\left(1+\varepsilon_{\mathrm{K}}\right)=\frac{\mu_{0} c}{2 \alpha}\left(1+\varepsilon_{\mathrm{K}}\right), \tag{373}
\end{align*}
$$

where $\varepsilon_{\mathrm{J}}$ and $\varepsilon_{\mathrm{K}}$ are unknown correction factors taken to be additional adjusted constants determined by leastsquares calculations. Replacing $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ with the generalizations in Eqs. (372) and (373) in the analysis leading to the observational equations in Table XL leads to the modified observational equations given in Table XLVI.

The results of seven different adjustments are presented in Table XLVII. In addition to the adjusted values of $\alpha, h, \varepsilon_{\mathrm{J}}$, and $\varepsilon_{\mathrm{K}}$, we also give the normalized residuals $r_{i}$ of the four input data with the largest values of $\left|r_{i}\right|: V_{\mathrm{m}}(\mathrm{Si})$, item $B 53, h / m_{\mathrm{n}} d_{220}(\mathrm{~W} 04)$, item $B 55$, $d_{220}(\mathrm{NR} 3)$, item $B 39$, and the NIST-89 value for $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$, item B31.1. The residuals are included as additional indicators of whether relaxing the assumption $K_{\mathrm{J}}$ $=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ reduces the disagreements among the data.

The adjusted value of $R_{\infty}$ is not included in Table XLVII, because it remains essentially unchanged from one adjustment to the next and equal to the 2006 recommended value. An entry of 0 in the $\varepsilon_{\mathrm{K}}$ column means that it is assumed that $R_{\mathrm{K}}=h / e^{2}$ in the corresponding adjustment; similarly, an entry of 0 in the $\varepsilon_{\mathrm{J}}$ column means that it is assumed that $K_{\mathrm{J}}=2 e / h$ in the corresponding adjustment. The following comments apply to the adjustments of Table XLVII.

Adjustment (i) is identical to adjustment 1 of Tables XLIII and XLIV in the previous section and is included here simply for reference; all input data are included and multiplicative weighting factors have not been applied to any uncertainties. For this adjustment, $N=150$, $M=79, \nu=71$, and $\chi^{2}=92.1$.

The next three adjustments differ from adjustment (i) in that in adjustment (ii) the relation $K_{\mathrm{J}}=2 e / h$ is relaxed, in adjustment (iii) the relation $R_{\mathrm{K}}=h / e^{2}$ is relaxed, and in adjustment (iv) both relations are relaxed. For these three adjustments, $N=150, M=80, \nu=70$, and $\chi^{2}=91.5$; $N=150, M=80, \nu=70$, and $\chi^{2}=91.3$; and $N=150, M=81$, $\nu=69$, and $\chi^{2}=90.4$, respectively.

TABLE XLII. Observational equations that express the input data related to antiprotonic helium in Table XXXII as functions of adjusted constants in Tables XXXIX and XLI. The numbers in the first column correspond to the numbers in the first column of Table XXXII. Definitions of the symbols and values of the parameters in these equations are given in Sec. IV.B. See Sec. XII.B for an explanation of the symbol $\doteq$.


It is clear from Table XLVII that there is no evidence for the inexactness of either $K_{\mathrm{J}}=2 e / h$ or $R_{\mathrm{K}}=h / e^{2}$. This conclusion is also true if instead of taking adjustment 1 of Table XLIII as our starting point, we had taken adjustment 2 in which the uncertainties of the five x-ray related data are multiplied by the factor 1.5. That is, none of the numbers in Table XLVII would change significantly, except $R_{\mathrm{B}}$ would be reduced from 1.14 to about 1.08. The reason adjustments (iii)-(vii) summarized in Table XLVII give values of $\epsilon_{\mathrm{K}}$ consistent with zero within about 2 parts in $10^{8}$ is mainly because the value of $\alpha$ inferred from the mean of the five measured values of $R_{\mathrm{K}}$ under the assumption $R_{\mathrm{K}}=h / e^{2}$, which has $u_{\mathrm{r}}=1.8 \times 10^{-8}$, agrees with the value of $\alpha$ with $u_{\mathrm{r}}=7.0$ $\times 10^{-10}$ inferred from the Harvard University measured value of $a_{\mathrm{e}}$.

Table XLVI and the uncertainties of the 2006 input data indicate that the values of $\epsilon_{\mathrm{J}}$ from adjustments (ii) and (iv) are determined mainly by the input data for $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ and $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ with observational equations that depend on $\epsilon_{\mathrm{J}}$ but not on $h$; and by the input data for $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi}), K_{\mathrm{J}}, K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, and $\mathcal{F}_{90}$, with observational equations that depend on both $\epsilon_{\mathrm{J}}$ and $h$. Because the value of $h$ in these least-squares calculations arises primarily from the measured value of the molar volume of silicon $V_{\mathrm{m}}(\mathrm{Si})$, the values of $\epsilon_{\mathrm{J}}$ in adjustments (ii) and (iv) arise mainly from a combination of individual values of $\epsilon_{\mathrm{J}}$ that either depend on $V_{\mathrm{m}}(\mathrm{Si})$ or on $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ and $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$. It is therefore of interest to repeat adjustment (iv), first with $V_{\mathrm{m}}(\mathrm{Si})$ deleted but with the $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ and $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ data included, and then with the latter deleted but with

TABLE XLIII. Summary of some of the least-squares adjustments used to analyze all input data given in Tables XXVIII-XXXI. The values of $\alpha$ and $h$ are those obtained in the adjustment, $N$ is the number of input data, $M$ is the number of adjusted constants, $\nu=N-M$ is the degrees of freedom, and $R_{\mathrm{B}}=\sqrt{\chi^{2} / \nu}$ is the Birge ratio. See the text for an explanation and discussion of each adjustment, but, in brief, 1 is all data; 2 is 1 with the uncertainties of the key x-ray and silicon data multiplied by $1.5 ; 3$ is 2 with the uncertainties of the key electrical data also multiplied by 1.5 ; 4 is the final adjustment from which the 2006 recommended values are obtained and is 3 with the input data with low weights deleted; 5 is 3 with the four data that provide the most accurate values of $\alpha$ deleted; and 6 is 3 with the three data that provide the most accurate values of $h$ deleted.

| Adj. | $N$ | $M$ | $\nu$ | $\chi^{2}$ | $R_{\mathrm{B}}$ | $\alpha^{-1}$ | $u_{\mathrm{r}}\left(\alpha^{-1}\right)$ | $h /(\mathrm{J} \mathrm{s})$ | $u_{\mathrm{r}}(h)$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 150 | 79 | 71 | 92.1 | 1.14 | $137.035999687(93)$ | $6.8 \times 10^{-10}$ | $6.62606896(22) \times 10^{-34}$ | $3.4 \times 10^{-8}$ |
| 2 | 150 | 79 | 71 | 82.0 | 1.07 | $137.035999682(93)$ | $6.8 \times 10^{-10}$ | $6.62606896(22) \times 10^{-34}$ | $3.4 \times 10^{-8}$ |
| 3 | 150 | 79 | 71 | 77.5 | 1.04 | $137.035999681(93)$ | $6.8 \times 10^{-10}$ | $6.62606896(33) \times 10^{-34}$ | $5.0 \times 10^{-8}$ |
| 4 | 135 | 78 | 57 | 65.0 | 1.07 | $137.035999679(94)$ | $6.8 \times 10^{-10}$ | $6.62606896(33) \times 10^{-34}$ | $5.0 \times 10^{-8}$ |
| 5 | 144 | 77 | 67 | 72.9 | 1.04 | $137.0360012(19)$ | $1.4 \times 10^{-8}$ | $6.62606896(33) \times 10^{-34}$ | $5.0 \times 10^{-8}$ |
| 6 | 147 | 79 | 68 | 75.4 | 1.05 | $137.035999680(93)$ | $6.8 \times 10^{-10}$ | $6.6260719(21) \times 10^{-34}$ | $3.2 \times 10^{-7}$ |

TABLE XLIV. Normalized residuals $r_{i}$ and self-sensitivity coefficients $S_{\mathrm{c}}$ that result from the six least-squares adjustments summarized in Table XLIII for the four input data whose absolute values of $r_{i}$ in Adj. 1 exceed $1.50 . S_{\mathrm{c}}$ is a measure of how the least-squares estimated value of a given type of input datum depends on a particular measured or calculated value of that type of datum; see Appendix E of CODATA-98. See the text for an explanation and discussion of each adjustment; brief explanations are given at the end of the caption to the previous table.

| Item number | Input quantity | Identification | Adj. 1 |  | Adj. 2 |  | Adj. 3 |  | Adj. 4 |  | Adj. 5 |  | Adj. 6 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | $r_{i}$ | $S_{\text {c }}$ | $r_{i}$ | $S_{\text {c }}$ | $r_{i}$ | $S_{\text {c }}$ | $r_{i}$ | $S_{\text {c }}$ | $r_{i}$ | $S_{\text {c }}$ | $r_{i}$ | $S_{\text {c }}$ |
| B53 | $V_{\mathrm{m}}(\mathrm{Si})$ | N/P/I-05 | -2.82 | 0.065 | -2.68 | 0.085 | -1.86 | 0.046 | -1.86 | 0.047 | -1.79 | 0.053 | -0.86 | 0.556 |
| B55 | $h / m_{\mathrm{n}} \mathrm{d}_{220}(\mathrm{~W} 04)$ | PTB-99 | -2.71 | 0.155 | -2.03 | 0.118 | -1.89 | 0.121 | -1.89 | 0.121 | -1.57 | 0.288 | -1.82 | 0.123 |
| B39 | $d_{220}(\mathrm{NR} 3)$ | NMIJ-04 | 2.37 | 0.199 | 1.86 | 0.145 | 1.74 | 0.148 | 1.74 | 0.148 | 1.78 | 0.151 | -1.00 | 0.353 |
| B31.1 | $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ | NIST-89 | 2.31 | 0.010 | 2.30 | 0.010 | 2.30 | 0.010 | Del | ted | 2.60 | 0.143 | 2.30 | 0.010 |

$V_{\mathrm{m}}(\mathrm{Si})$ included. These are, in fact, adjustments (v) and (vi) of Table XLVII.

In each of these adjustments, the absolute values of $\epsilon_{\mathrm{J}}$ are comparable and significantly larger than the uncertainties, which are also comparable, but the values have different signs. Consequently, when $V_{\mathrm{m}}(\mathrm{Si})$ and the $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ and $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ data are included at the same time as in adjustment (iv), the result for $\epsilon_{\mathrm{J}}$ is consistent with zero.
The values of $\epsilon_{\mathrm{J}}$ from adjustments (v) and (vi) reflect some of the inconsistencies among the data: the disagreement of the values of $h$ implied by $V_{\mathrm{m}}(\mathrm{Si})$ and $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ when it is assumed that $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ are exact; and the disagreement of the values of $\alpha$ implied by the electron magnetic moment anomaly $a_{\mathrm{e}}$ and by $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ and $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ under the same assumption.

In adjustment (vii), the problematic input data for $V_{\mathrm{m}}(\mathrm{Si}), \Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$, and $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ are simultaneously deleted from the calculation. Then the value of $\epsilon_{\mathrm{J}}$ arises mainly from the input data for $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi}), K_{\mathrm{J}}, K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, and $\mathcal{F}_{90}$. Like adjustment (iv), adjustment (vii) shows that $\epsilon_{\mathrm{J}}$ is consistent with zero, although not within 8 parts in $10^{8}$ but within 7 parts in $10^{7}$. However, adjustment (vii) has the advantage of being based on consistent data.

The comparatively narrow range of values of $\alpha$ in Table XLVII is due to the fact that the input data that
mainly determine $\alpha$ do not depend on the Josephson or quantum Hall effects. This is not the case for the input data that primarily determine $h$, hence the values of $h$ vary over a wide range.

## XIII. THE 2006 CODATA RECOMMENDED VALUES

## A. Calculational details

As indicated in Sec. XII.B, the 2006 recommended values of the constants are based on adjustment 4 of Tables XLIII-XLV. This adjustment is obtained by (i) deleting 15 items from the originally considered 150 items of input data of Tables XXVIII, XXX, and XXXII, namely, items $B 8, B 31.1-B 35.2, B 37$, and $B 56$, because of their low weight (self-sensitivity coefficient $S_{\mathrm{c}}<0.01$ ); and (ii) weighting the uncertainties of the nine input data $B 36.1-B 36.3, B 38.1-B 40, B 53$, and $B 55$ by the multiplicative factor 1.5 in order to reduce the absolute values of their normalized residuals $\left|r_{i}\right|$ to less than 2. The correlation coefficients of the data, as given in Tables XXIX, XXXI, and XXXIII, are also taken into account. The 135 final input data are expressed in terms of the 78 adjusted constants of Tables XXXVII, XXXIX, and XLI, corresponding to $N-M=\nu=57$ degrees of freedom. Because $h / m\left({ }^{133} \mathrm{Cs}\right)$, item $B 56$, has been deleted as

TABLE XLV. Summary of the results of some of the least-squares adjustments used to analyze the input data related to $R_{\infty}$. The values of $R_{\infty}, R_{\mathrm{p}}$, and $R_{\mathrm{d}}$ are those obtained in the indicated adjustment, $N$ is the number of input data, $M$ is the number of adjusted constants, $\nu=N-M$ is the degrees of freedom, and $R_{\mathrm{B}}=\sqrt{\chi^{2} / \nu}$ is the Birge ratio. See the text for an explanation and discussion of each adjustment, but, in brief, 4 is the final adjustment; 7 is 4 with the input data for $R_{\mathrm{p}}$ and $R_{\mathrm{d}}$ deleted; 8 is 4 with just the $R_{\mathrm{p}}$ datum deleted; 9 is 4 with just the $R_{\mathrm{d}}$ datum deleted; 10 is 4 but with only the hydrogen data included; and 11 is 4 but with only the deuterium data included.

| Adj. | $N$ | $M$ | $\nu$ | $\chi^{2}$ | $R_{\mathrm{B}}$ | $R_{\infty} / \mathrm{m}^{-1}$ | $u_{\mathrm{r}}\left(R_{\infty}\right)$ | $R_{\mathrm{p}} / \mathrm{fm}$ | $R_{\mathrm{d}} / \mathrm{fm}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 4 | 135 | 78 | 57 | 65.0 | 1.07 | $10973731.568527(73)$ | $6.6 \times 10^{-12}$ | $0.8768(69)$ | $2.1402(28)$ |
| 7 | 133 | 78 | 55 | 63.0 | 1.07 | $10973731.568518(82)$ | $7.5 \times 10^{-12}$ | $0.8760(78)$ | $2.1398(32)$ |
| 8 | 134 | 78 | 56 | 63.8 | 1.07 | $10973731.568495(78)$ | $7.1 \times 10^{-12}$ | $0.8737(75)$ | $2.1389(30)$ |
| 9 | 134 | 78 | 56 | 63.9 | 1.07 | $10973731.568549(76)$ | $6.9 \times 10^{-12}$ | $0.8790(71)$ | $2.1411(29)$ |
| 10 | 117 | 68 | 49 | 60.8 | 1.11 | $10973731.568562(85)$ | $7.8 \times 10^{-12}$ | $0.8802(80)$ |  |
| 11 | 102 | 61 | 41 | 54.7 | 1.16 | $10973731.56839(13)$ | $1.1 \times 10^{-11}$ |  | $2.1286(93)$ |

TABLE XLVI. Generalized observational equations that express input data B31-B37 in Table XXX as functions of the adjusted constants in Tables XXXIX and XXXVII with the additional adjusted constants $\varepsilon_{\mathrm{J}}$ and $\varepsilon_{\mathrm{K}}$ as given in Eqs. (372) and (373). The numbers in the first column correspond to the numbers in the first column of Table XXX. For simplicity, the lengthier functions are not explicitly given. See Sec. XII.B for an explanation of the symbol $\doteq$.

| Type of input datum | Generalized observational equation |
| :---: | :---: |
| B31* | $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo}) \doteq-\frac{K_{\mathrm{J}-90} R_{\mathrm{K}-90}\left[1+a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)\right] \alpha^{3}}{2 \mu_{0} R_{\infty}\left(1+\varepsilon_{\mathrm{J}}\right)\left(1+\varepsilon_{\mathrm{K}}\right)}\left(\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}^{\prime}}\right)^{-1}$ |
| B32* | $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo}) \doteq \frac{K_{\mathrm{J}-90} R_{\mathrm{K}-90}\left[1+a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)\right] \alpha^{3}}{2 \mu_{0} R_{\infty}\left(1+\varepsilon_{\mathrm{J}}\right)\left(1+\varepsilon_{\mathrm{K}}\right)}$ |
|  | $\times\left(\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}^{\prime}}\right)^{-1} \frac{\mu_{\mathrm{h}}^{\prime}}{\mu_{\mathrm{p}}^{\prime}}$ |
| B33* | $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{hi}) \doteq-\frac{c\left[1+a_{\mathrm{e}}\left(\alpha, \delta_{\mathrm{e}}\right)\right] \alpha^{2}}{K_{\mathrm{J}-90} R_{\mathrm{K}-90} R_{\infty} h}\left(1+\varepsilon_{\mathrm{J}}\right)$ |
|  | $\times\left(1+\varepsilon_{\mathrm{K}}\right)\left(\frac{\mu_{\mathrm{e}^{-}}}{\mu_{\mathrm{p}}^{\prime}}\right)^{-1}$ |
| B34* | $R_{\mathrm{K}} \doteq \frac{\mu_{0} c}{2 \alpha}\left(1+\varepsilon_{\mathrm{K}}\right)$ |
| B35* | $K_{\mathrm{J}} \doteq\left(\frac{8 \alpha}{\mu_{0} c h}\right)^{1 / 2}\left(1+\varepsilon_{\mathrm{J}}\right)$ |
| B36* | $K_{\mathrm{J}}^{2} R_{\mathrm{K}} \doteq \frac{4}{h}\left(1+\varepsilon_{\mathrm{J}}\right)^{2}\left(1+\varepsilon_{\mathrm{K}}\right)$ |
| B37* | $\mathcal{F}_{90} \doteq \frac{c M_{\mathrm{u}} A_{\mathrm{r}}(\mathrm{e}) \alpha^{2}}{K_{\mathrm{J}-90} R_{\mathrm{K}-90} R_{\infty 0} h}\left(1+\varepsilon_{\mathrm{J}}\right)\left(1+\varepsilon_{\mathrm{K}}\right)$ |
| B62* | $\varepsilon_{J} \doteq \varepsilon_{J}$ |
| B63* | $\varepsilon_{\mathrm{K}} \doteq \varepsilon_{\mathrm{K}}$ |

an input datum due to its low weight, $A_{\mathrm{r}}\left({ }^{133} \mathrm{Cs}\right)$, item $B 8$, has also been deleted as an input datum and as an adjusted constant.

For the final adjustment, $\chi^{2}=65.0, \sqrt{\chi^{2} / \nu}=R_{\mathrm{B}}=1.04$, and $Q(65.0 \mid 57)=0.22$, where $Q\left(\chi^{2} \mid \nu\right)$ is the probability that the observed value of $\chi^{2}$ for degrees of freedom $\nu$ would have exceeded that observed value (see Appendix E of CODATA-98). Each input datum in the final adjustment has $S_{\mathrm{c}}>0.01$, or is a subset of the data of an experiment that provides an input datum or input data with $S_{\mathrm{c}}>0.01$. Not counting such input data with $S_{\mathrm{c}}$ $<0.01$, the six input data with the largest $\left|r_{i}\right|$ are $B 55$, $B 53, B 39, C 18, B 11.1$, and $B 9$; their values of $r_{i}$ are $-1.89,-1.86,1.74,-1.73,1.69$, and 1.45 , respectively. The next largest $r_{i}$ are 1.22 and 1.11.

The output of the final adjustment is the set of best estimated values, in the least-squares sense, of the 78 adjusted constants and their variances and covariances. Together with (i) those constants that have exact values such as $\mu_{0}$ and $c$; (ii) the value of $G$ obtained in Sec. X; and (iii) the values of $m_{\tau}, G_{\mathrm{F}}$, and $\sin ^{2} \theta_{\mathrm{W}}$ given in Sec. XI.B, all of the 2006 recommended values, including their uncertainties, are obtained from the 78 adjusted constants. How this is done can be found in Sec. V.B of CODATA-98.

## B. Tables of values

The 2006 CODATA recommended values of the basic constants and conversion factors of physics and chemistry and related quantities are given in Tables XLVIII-LV. These tables are very similar in form to their 2002 counterparts; the principal difference is that a number of new recommended values have been included in the 2006 list, in particular, in Table XLIX. These are $m_{\mathrm{P}} c^{2}$ in GeV , where $m_{\mathrm{P}}$ is the Planck mass; the $g$-factor of the deuteron $g_{\mathrm{d}} ; b^{\prime}=\nu_{\max } / T$, the Wien displacement law constant for frequency; and, for the first time, 14 recommended values of a number of constants that characterize the triton, including its mass $m_{\mathrm{t}}$, magnetic moment $\mu_{\mathrm{t}}, g$-factor $g_{\mathrm{t}}$, and the magnetic moment ratios $\mu_{\mathrm{t}} / \mu_{\mathrm{e}}$ and $\mu_{\mathrm{t}} / \mu_{\mathrm{p}}$. The addition of the tritonrelated constants is a direct consequence of the improved measurement of $A_{\mathrm{r}}\left({ }^{3} \mathrm{H}\right)$ (item $B 3$ in Table XXX) and the new NMR measurements on, and reexamined shielding correction differences for, the HT molecule (items $B 22$ and $B 23$ in Table XXX).

Table XLVIII is a highly abbreviated list containing the values of the constants and conversion factors most commonly used. Table XLIX is a much more extensive list of values categorized as follows: universal; electromagnetic; atomic and nuclear; and physicochemical. The atomic and nuclear category is subdivided into 11 subcategories: general; electroweak; electron, $\mathrm{e}^{-}$; muon, $\mu^{-}$; tau, $\tau^{-}$; proton, p ; neutron, n ; deuteron, d ; triton, t ; helion, h ; and alpha particle, $\alpha$. Table L gives the variances, covariances, and correlation coefficients of a selected group of constants. (Application of the covariance matrix is discussed in Appendix E of CODATA-98.) Table LI gives the internationally adopted values of various quantities; Table LII lists the values of a number of x-ray related quantities; Table LIII lists the values of various non-SI units; and Tables LIV and LV give the values of various energy equivalents.

All values given in Tables XLVIII-LV are available on the web pages of the Fundamental Constants Data Center of the NIST Physics Laboratory at physics.nist.gov/ constants. This electronic version of the 2006 CODATA recommended values of the constants also includes a much more extensive correlation coefficient matrix. Indeed, the correlation coefficient of any two constants listed in the tables is accessible on the web site, as well as the automatic conversion of the value of an energyrelated quantity expressed in one unit to the corre-

TABLE XLVII. Summary of the results of several least-squares adjustments carried out to investigate the effect of assuming the relations for $K_{\mathrm{J}}$ and $R_{\mathrm{K}}$ given in Eqs. (372) and (373). The values of $\alpha, h, \varepsilon_{\mathrm{K}}$, and $\varepsilon_{\mathrm{J}}$ are those obtained in the indicated adjustments. The quantity $R_{\mathrm{B}}=\sqrt{\chi^{2} / \nu}$ is the Birge ratio and $r_{i}$ is the normalized residual of the indicated input datum (see Table XXX). These four data have the largest $\left|r_{i}\right|$ of all the input data and are the only data in Adj. (i) with $\left|r_{i}\right|>1.50$. See the text for an explanation and discussion of each adjustment, but in brief, (i) assumes $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ and uses all the data; (ii) is (i) with the relation $K_{\mathrm{J}}=2 e / h$ relaxed; (iii) is (i) with the relation $R_{\mathrm{K}}=h / e^{2}$ relaxed; (iv) is (i) with both relations relaxed; (v) is (iv) with the $V_{\mathrm{m}}(\mathrm{Si})$ datum deleted; (vi) is (iv) with the $\Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$ and $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ data deleted; and (vii) is (iv) with the $V_{\mathrm{m}}(\mathrm{Si}), \Gamma_{\mathrm{p}-90}^{\prime}(\mathrm{lo})$, and $\Gamma_{\mathrm{h}-90}^{\prime}(\mathrm{lo})$ data deleted.

| Adj. | $R_{\mathrm{B}}$ | $\alpha^{-1}$ | $h /(\mathrm{J} \mathrm{s})$ | $\varepsilon_{\mathrm{K}}$ | $\varepsilon_{\mathrm{J}}$ | $r_{B 53}$ | $r_{B 55}$ | $r_{B 39}$ | $r_{B 31.1}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| (i) | 1.14 | $137.035999687(93)$ | $6.62606896(22) \times 10^{-34}$ | 0 | 0 | -2.82 | -2.71 | 2.37 | 2.31 |
| (ii) | 1.14 | $137.035999688(93)$ | $6.6260682(10) \times 10^{-34}$ | 0 | $-61(79) \times 10^{-9}$ | -3.22 | -2.75 | 2.39 | 1.77 |
| (iii) | 1.14 | $137.035999683(93)$ | $6.62606906(25) \times 10^{-34}$ | $16(18) \times 10^{-9}$ | 0 | -2.77 | -2.71 | 2.36 | 2.45 |
| (iv) | 1.14 | $137.035999685(93)$ | $6.6260681(11) \times 10^{-34}$ | $20(18) \times 10^{-9}$ | $-77(80) \times 10^{-9}$ | -3.27 | -2.75 | 2.39 | 1.79 |
| (v) | 1.05 | $137.035999686(93)$ | $6.6260653(13) \times 10^{-34}$ | $23(18) \times 10^{-9}$ | $-281(95) \times 10^{-9}$ | Deleted | -2.45 | 2.19 | 0.01 |
| (vi) | 1.05 | $137.035999686(93)$ | $6.6260744(19) \times 10^{-34}$ | $24(18) \times 10^{-9}$ | $407(143) \times 10^{-9}$ | -0.05 | -2.45 | 2.19 | Deleted |
| (vii) | 1.06 | $137.035999686(93)$ | $6.6260722(95) \times 10^{-34}$ | $24(18) \times 10^{-9}$ | $238(720) \times 10^{-9}$ | Deleted | -2.45 | 2.19 | Deleted |

TABLE XLVIII. An abbreviated list of the CODATA recommended values of the fundamental constants of physics and chemistry based on the 2006 adjustment.

| Quantity | Symbol | Numerical value | Unit | Relative std. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: |
| speed of light in vacuum | $c, c_{0}$ | 299792458 | $\mathrm{m} \mathrm{s}^{-1}$ | (Exact) |
| magnetic constant | $\mu_{0}$ | $4 \pi \times 10^{-7}$ | $\mathrm{NA}^{-2}$ |  |
|  |  | $=12.566370614 \ldots \times 10^{-7}$ | $\mathrm{NA}^{-2}$ | (Exact) |
| electric constant $1 / \mu_{0} c^{2}$ | $\epsilon_{0}$ | $8.854187817 \ldots \times 10^{-12}$ | F m ${ }^{-1}$ | (Exact) |
| Newtonian constant of gravitation | $G$ | $6.67428(67) \times 10^{-11}$ | $\mathrm{m}^{3} \mathrm{~kg}^{-1} \mathrm{~s}^{-2}$ | $1.0 \times 10^{-4}$ |
| Planck constant$h / 2 \pi$ | $h$ | $6.62606896(33) \times 10^{-34}$ | J s | $5.0 \times 10^{-8}$ |
|  | $\hbar$ | $1.054571628(53) \times 10^{-34}$ | J s | $5.0 \times 10^{-8}$ |
| elementary charge | $e$ | $1.602176487(40) \times 10^{-19}$ | C | $2.5 \times 10^{-8}$ |
| magnetic flux quantum $h / 2 e$ | $\Phi_{0}$ | $2.067833667(52) \times 10^{-15}$ | Wb | $2.5 \times 10^{-8}$ |
| conductance quantum $2 e^{2} / h$ | $G_{0}$ | $7.7480917004(53) \times 10^{-5}$ | S | $6.8 \times 10^{-10}$ |
| electron mass | $m_{\text {e }}$ | $9.10938215(45) \times 10^{-31}$ | kg | $5.0 \times 10^{-8}$ |
| proton mass | $m_{\text {p }}$ | $1.672621637(83) \times 10^{-27}$ | kg | $5.0 \times 10^{-8}$ |
| proton-electron mass ratio | $m_{\mathrm{p}} / m_{\mathrm{e}}$ | $1836.15267247(80)$ |  | $4.3 \times 10^{-10}$ |
| fine-structure constant $e^{2} / 4 \pi \epsilon_{0} \hbar c$ inverse fine-structure constant | $\alpha$ | $7.2973525376(50) \times 10^{-3}$ |  | $6.8 \times 10^{-10}$ |
|  | $\alpha^{-1}$ | 137.035999 679(94) |  | $6.8 \times 10^{-10}$ |
| Rydberg constant $\alpha^{2} m_{\mathrm{e}} c / 2 h$ | $R_{\infty}$ | $10973731.568527(73)$ | $\mathrm{m}^{-1}$ | $6.6 \times 10^{-12}$ |
| Avogadro constant | $N_{\text {A }}, L$ | $6.02214179(30) \times 10^{23}$ | $\mathrm{mol}^{-1}$ | $5.0 \times 10^{-8}$ |
| Faraday constant $N_{\mathrm{A}} e$ | $F$ | $96485.3399(24)$ | $\mathrm{C} \mathrm{mol}^{-1}$ | $2.5 \times 10^{-8}$ |
| molar gas constant | $R$ | 8.314 472(15) | $\mathrm{J} \mathrm{mol}^{-1} \mathrm{~K}^{-1}$ | $1.7 \times 10^{-6}$ |
| Boltzmann constant $R / N_{\text {A }}$ | $k$ | $1.3806504(24) \times 10^{-23}$ | J K ${ }^{-1}$ | $1.7 \times 10^{-6}$ |
| Stefan-Boltzmann constant$\left(\pi^{2} / 60\right) k^{4} / \hbar^{3} c^{2}$ | $\sigma$ | $5.670400(40) \times 10^{-8}$ | W m ${ }^{-2} \mathrm{~K}^{-4}$ | $7.0 \times 10^{-6}$ |
|  | Non-SI units accepted for use with the SI |  |  |  |
| electron volt: (e/C) J | eV | $1.602176487(40) \times 10^{-19}$ | J | $2.5 \times 10^{-8}$ |
| (unified) atomic mass unit $\begin{aligned} & 1 \mathrm{u}=m_{\mathrm{u}}=\frac{1}{12} m\left({ }^{12} \mathrm{C}\right) \\ &=10^{-3} \mathrm{~kg} \mathrm{~mol} \\ &-1 \end{aligned} N_{\mathrm{A}} .$ | u | $1.660538782(83) \times 10^{-27}$ | kg | $5.0 \times 10^{-8}$ |

TABLE XLIX. The CODATA recommended values of the fundamental constants of physics and chemistry based on the 2006 adjustment.

| Quantity | Symbol | Numerical value | Unit | Relative std. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: |
| speed of light in vacuum magnetic constant | $c, c_{0}$ | UNIVERSAL |  | (Exact) |
|  |  | 299792458 | $\mathrm{m} \mathrm{s}^{-1}$ |  |
|  | $\mu_{0}$ | $4 \pi \times 10^{-7}$ | $\mathrm{NA}^{-2}$ |  |
|  |  | $=12.566370614 \ldots \times 10^{-7}$ | $\mathrm{NA}^{-2}$ | (Exact) |
| electric constant $1 / \mu_{0} c^{2}$ | $\epsilon_{0}$ | $8.854187817 \ldots \times 10^{-12}$ | F m ${ }^{-1}$ | (Exact) |
| characteristic impedance of vacuum $\sqrt{\mu_{0} / \epsilon_{0}}=\mu_{0} c$ | $Z_{0}$ | $376.730313461 \ldots$ | $\Omega$ | (Exact) |
| Newtonian constant of gravitation | G | $6.67428(67) \times 10^{-11}$ | $\mathrm{m}^{3} \mathrm{~kg}^{-1} \mathrm{~s}^{-2}$ | $1.0 \times 10^{-4}$ |
|  | $G / \hbar c$ | $6.70881(67) \times 10^{-39}$ | $\left(\mathrm{GeV} / \mathrm{c}^{2}\right)^{-2}$ | $1.0 \times 10^{-4}$ |
| Planck constant in eV s | $h$ | $6.62606896(33) \times 10^{-34}$ | J s | $5.0 \times 10^{-8}$ |
|  |  | $4.13566733(10) \times 10^{-15}$ | eV s | $2.5 \times 10^{-8}$ |
| $\begin{aligned} & h / 2 \pi \\ & \quad \text { in } \mathrm{eV} \mathrm{~s} \end{aligned}$ | $\hbar$ | $1.054571628(53) \times 10^{-34}$ | J s | $5.0 \times 10^{-8}$ |
|  |  | $6.58211899(16) \times 10^{-16}$ | eV s | $2.5 \times 10^{-8}$ |
| $\hbar c$ in MeV fm |  | 197.326 9631(49) | MeV fm | $2.5 \times 10^{-8}$ |
| Planck mass $(\hbar c / G)^{1 / 2}$ energy equivalent in GeV | $\begin{aligned} & m_{\mathrm{P}} \\ & m_{\mathrm{P}} c^{2} \end{aligned}$ | $2.17644(11) \times 10^{-8}$ | kg | $5.0 \times 10^{-5}$ |
|  |  | $1.220892(61) \times 10^{19}$ | GeV | $5.0 \times 10^{-5}$ |
| Planck temperature ( $\left.\hbar c^{5} / G\right)^{1 / 2} / k$ | $T_{\mathrm{P}}$ | $1.416785(71) \times 10^{32}$ | K | $5.0 \times 10^{-5}$ |
| Planck length $\hbar / m_{\mathrm{P}} c=\left(\hbar G / c^{3}\right)^{1 / 2}$ | $l_{\mathrm{P}}$ | $1.616252(81) \times 10^{-35}$ | m | $5.0 \times 10^{-5}$ |
| Planck time $l_{\mathrm{P}} / c=\left(\hbar G / c^{5}\right)^{1 / 2}$ | $t_{\mathrm{P}}$ | $5.39124(27) \times 10^{-44}$ | s | $5.0 \times 10^{-5}$ |
| elementary charge | ELECTROMAGNETIC |  |  |  |
|  | $e$ | $1.602176487(40) \times 10^{-19}$ | C | $2.5 \times 10^{-8}$ |
|  | $e / h$ | $2.417989454(60) \times 10^{14}$ | A J ${ }^{-1}$ | $2.5 \times 10^{-8}$ |
| magnetic flux quantum $h / 2 e$ | $\Phi_{0}$ | $2.067833667(52) \times 10^{-15}$ | Wb | $2.5 \times 10^{-8}$ |
| conductance quantum $2 e^{2} / h$ inverse of conductance quantum | $G_{0}$ | $7.7480917004(53) \times 10^{-5}$ | S | $6.8 \times 10^{-10}$ |
|  | $G_{0}^{-1}$ | $12906.4037787(88)$ | $\Omega$ | $6.8 \times 10^{-10}$ |
| Joesphson constant ${ }^{\text {a }} 2 e / h$ | $K_{\text {J }}$ | $483597.891(12) \times 10^{9}$ | $\mathrm{Hz} \mathrm{V}^{-1}$ | $2.5 \times 10^{-8}$ |
| von Klitzing constant ${ }^{\text {b }}$ $h / e^{2}=\mu_{0} c / 2 \alpha$ | $R_{\text {K }}$ | 25812.807 557(18) | $\Omega$ | $6.8 \times 10^{-10}$ |
| Bohr magneton $e \hbar / 2 m_{\mathrm{e}}$ in $\mathrm{eV} \mathrm{T}^{-1}$ | $\mu_{\text {B }}$ | $927.400915(23) \times 10^{-26}$ | $\mathrm{J} \mathrm{T}^{-1}$ | $2.5 \times 10^{-8}$ |
|  |  | $5.7883817555(79) \times 10^{-5}$ | $\mathrm{eV} \mathrm{T}^{-1}$ | $1.4 \times 10^{-9}$ |
|  | $\mu_{\mathrm{B}} / \mathrm{h}$ | $13.99624604(35) \times 10^{9}$ | $\mathrm{Hz} \mathrm{T}^{-1}$ | $2.5 \times 10^{-8}$ |
|  | $\mu_{\mathrm{B}} / h c$ | $46.6864515(12)$ | $\mathrm{m}^{-1} \mathrm{~T}^{-1}$ | $2.5 \times 10^{-8}$ |
|  | $\mu_{\mathrm{B}} / k$ | 0.671 7131(12) | $\mathrm{K} \mathrm{T}^{-1}$ | $1.7 \times 10^{-6}$ |
| $\begin{aligned} & \text { nuclear magneton } e \hbar / 2 m_{\mathrm{p}} \\ & \text { in } \mathrm{eV} \mathrm{~T}^{-1} \end{aligned}$ | $\mu_{\mathrm{N}}$ | $5.05078324(13) \times 10^{-27}$ | $\mathrm{J} ~ T ~^{-1}$ | $2.5 \times 10^{-8}$ |
|  |  | $3.1524512326(45) \times 10^{-8}$ | $\mathrm{eV} \mathrm{T}{ }^{-1}$ | $1.4 \times 10^{-9}$ |
|  | $\mu_{\mathrm{N}} / \mathrm{h}$ | 7.622593 84(19) | MHz T ${ }^{-1}$ | $2.5 \times 10^{-8}$ |
|  | $\mu_{\mathrm{N}} / h c$ | $2.542623616(64) \times 10^{-2}$ | $\mathrm{m}^{-1} \mathrm{~T}^{-1}$ | $2.5 \times 10^{-8}$ |
|  | $\mu_{\mathrm{N}} / k$ | $3.6582637(64) \times 10^{-4}$ | $\mathrm{K} \mathrm{T}^{-1}$ | $1.7 \times 10^{-6}$ |
|  | ATOMIC AND NUCLEAR |  |  |  |
|  |  | General |  |  |
| fine-structure constant $e^{2} / 4 \pi \epsilon_{0} \hbar c$ inverse fine-structure constant | $\alpha$ | $7.2973525376(50) \times 10^{-3}$ |  | $6.8 \times 10^{-10}$ |
|  | $\alpha^{-1}$ | $137.035999679(94)$ |  | $6.8 \times 10^{-10}$ |
| Rydberg constant $\alpha^{2} m_{\mathrm{e}} c / 2 h$ | $R_{\infty}$ | $10973731.568527(73)$ | $\mathrm{m}^{-1}$ | $6.6 \times 10^{-12}$ |
|  | $R_{\infty} C$ | $3.289841960361(22) \times 10^{15}$ | Hz | $6.6 \times 10^{-12}$ |

TABLE XLIX. (Continued.)

| Quantity | Symbol | Numerical value | Unit | Relative std. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: |
|  | $R_{\infty} h c$ | $2.17987197(11) \times 10^{-18}$ | J | $5.0 \times 10^{-8}$ |
| $R_{\infty} h c$ in eV |  | $13.60569193(34)$ | eV | $2.5 \times 10^{-8}$ |
| Bohr radius $\alpha / 4 \pi R_{\infty}=4 \pi \epsilon_{0} \hbar^{2} / m_{\mathrm{e}} e^{2}$ | $a_{0}$ | $0.52917720859(36) \times 10^{-10}$ | m | $6.8 \times 10^{-10}$ |
| Hartree energy $e^{2} / 4 \pi \epsilon_{0} a_{0}=2 R_{\propto} h c$ $=\alpha^{2} m_{\mathrm{e}} c^{2}$ | $E_{\mathrm{h}}$ | $4.35974394(22) \times 10^{-18}$ | J | $5.0 \times 10^{-8}$ |
| in eV |  | 27.211383 86(68) | eV | $2.5 \times 10^{-8}$ |
| quantum of circulation | $h / 2 m_{\text {e }}$ | $3.6369475199(50) \times 10^{-4}$ | $\mathrm{m}^{2} \mathrm{~s}^{-1}$ | $1.4 \times 10^{-9}$ |
|  | $h / m_{\mathrm{e}}$ | $7.273895040(10) \times 10^{-4}$ | $\mathrm{m}^{2} \mathrm{~s}^{-1}$ | $1.4 \times 10^{-9}$ |
|  |  | Electroweak |  |  |
| Fermi coupling constant ${ }^{\text {c }}$ | $G_{\mathrm{F}} /(\hbar c)^{3}$ | $1.16637(1) \times 10^{-5}$ | $\mathrm{GeV}^{-2}$ | $8.6 \times 10^{-6}$ |
| weak mixing angle ${ }^{\mathrm{d}} \theta_{\mathrm{W}}$ (on-shell scheme) $\sin ^{2} \theta_{\mathrm{W}}=s_{\mathrm{W}}^{2} \equiv 1-\left(m_{\mathrm{W}} / m_{\mathrm{Z}}\right)^{2}$ | $\sin ^{2} \theta_{\mathrm{W}}$ | $0.22255(56)$ |  | $2.5 \times 10^{-3}$ |
|  |  | Electron, $\mathrm{e}^{-}$ |  |  |
| electron mass | $m_{\text {e }}$ | $9.10938215(45) \times 10^{-31}$ | kg | $5.0 \times 10^{-8}$ |
| in $\mathrm{u}, m_{\mathrm{e}}=A_{\mathrm{r}}(\mathrm{e}) \mathrm{u}$ (electron relative atomic mass times $u$ ) |  | $5.4857990943(23) \times 10^{-4}$ | u | $4.2 \times 10^{-10}$ |
| energy equivalent | $m_{\mathrm{e}} c^{2}$ | $8.18710438(41) \times 10^{-14}$ | J | $5.0 \times 10^{-8}$ |
| in MeV |  | 0.510998910 (13) | MeV | $2.5 \times 10^{-8}$ |
| electron-muon mass ratio | $m_{\mathrm{e}} / m_{\mu}$ | $4.83633171(12) \times 10^{-3}$ |  | $2.5 \times 10^{-8}$ |
| electron-tau mass ratio | $m_{\mathrm{e}} / m_{\tau}$ | $2.87564(47) \times 10^{-4}$ |  | $1.6 \times 10^{-4}$ |
| electron-proton mass ratio | $m_{\mathrm{e}} / m_{\mathrm{p}}$ | $5.4461702177(24) \times 10^{-4}$ |  | $4.3 \times 10^{-10}$ |
| electron-neutron mass ratio | $m_{\mathrm{e}} / m_{\mathrm{n}}$ | $5.4386734459(33) \times 10^{-4}$ |  | $6.0 \times 10^{-10}$ |
| electron-deuteron mass ratio | $m_{\mathrm{e}} / m_{\text {d }}$ | $2.7244371093(12) \times 10^{-4}$ |  | $4.3 \times 10^{-10}$ |
| electron to $\alpha$ particle mass ratio | $m_{\mathrm{e}} / m_{\alpha}$ | $1.37093355570(58) \times 10^{-4}$ |  | $4.2 \times 10^{-10}$ |
| electron charge to mass quotient | $-e / m_{\text {e }}$ | $-1.758820150(44) \times 10^{11}$ | C kg ${ }^{-1}$ | $2.5 \times 10^{-8}$ |
| electron molar mass $N_{\text {A }} m_{\mathrm{e}}$ | $M(\mathrm{e}), M_{\text {e }}$ | $5.4857990943(23) \times 10^{-7}$ | $\mathrm{kg} \mathrm{mol}^{-1}$ | $4.2 \times 10^{-10}$ |
| Compton wavelength $h / m_{\mathrm{e}} c$ | $\lambda_{\text {C }}$ | $2.4263102175(33) \times 10^{-12}$ | m | $1.4 \times 10^{-9}$ |
| $\lambda_{\mathrm{C}} / 2 \pi=\alpha a_{0}=\alpha^{2} / 4 \pi R_{\infty}$ | $\chi_{\text {C }}$ | $386.15926459(53) \times 10^{-15}$ | m | $1.4 \times 10^{-9}$ |
| classical electron radius $\alpha^{2} a_{0}$ | $r_{\text {e }}$ | $2.8179402894(58) \times 10^{-15}$ | m | $2.1 \times 10^{-9}$ |
| Thomson cross section $(8 \pi / 3) r_{\mathrm{e}}^{2}$ | $\sigma_{\text {e }}$ | $0.6652458558(27) \times 10^{-28}$ | $\mathrm{m}^{2}$ | $4.1 \times 10^{-9}$ |
| electron magnetic moment | $\mu_{\text {e }}$ | $-928.476377(23) \times 10^{-26}$ | $\mathrm{J} \mathrm{T}^{-1}$ | $2.5 \times 10^{-8}$ |
| to Bohr magneton ratio | $\mu_{\mathrm{e}} / \mu_{\mathrm{B}}$ | -1.001 $15965218111(74)$ |  | $7.4 \times 10^{-13}$ |
| to nuclear magneton ratio | $\mu_{\mathrm{e}} / \mu_{\mathrm{N}}$ | -1838.281970 92(80) |  | $4.3 \times 10^{-10}$ |
| electron magnetic moment |  |  |  |  |
| anomaly $\left\|\mu_{\mathrm{e}}\right\| / \mu_{\mathrm{B}}-1$ | $a_{\text {e }}$ | $1.15965218111(74) \times 10^{-3}$ |  | $6.4 \times 10^{-10}$ |
| electron $g$-factor $-2\left(1+a_{\mathrm{e}}\right)$ | $g_{\text {e }}$ | $-2.0023193043622(15)$ |  | $7.4 \times 10^{-13}$ |
| electron-muon magnetic moment ratio | $\mu_{\mathrm{e}} / \mu_{\mu}$ | 206.766 9877(52) |  | $2.5 \times 10^{-8}$ |
| electron-proton magnetic moment ratio | $\mu_{\mathrm{e}} / \mu_{\mathrm{p}}$ | -658.210 6848(54) |  | $8.1 \times 10^{-9}$ |
| electron to shielded proton magnetic moment ratio $\left(\mathrm{H}_{2} \mathrm{O}\right.$, sphere, $25^{\circ} \mathrm{C}$ ) | $\mu_{\mathrm{e}} / \mu_{\mathrm{p}}^{\prime}$ | -658.227 5971(72) |  | $1.1 \times 10^{-8}$ |
| electron-neutron magnetic moment ratio | $\mu_{\mathrm{e}} / \mu_{\mathrm{n}}$ | 960.920 50(23) |  | $2.4 \times 10^{-7}$ |
| electron-deuteron magnetic moment ratio | $\mu_{\mathrm{e}} / \mu_{\mathrm{d}}$ | -2143.923 498(18) |  | $8.4 \times 10^{-9}$ |

TABLE XLIX. (Continued.)

|  |  |  |  |
| :--- | :--- | :--- | :--- |
| Quantity | Symbol | Numerical value | Relative std. |
| uncert. $u_{\mathrm{r}}$ |  |  |  |

TABLE XLIX. (Continued.)

| Quantity | Symbol | Numerical value | Unit | Relative std. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: |
| proton-neutron mass ratio proton charge to mass quotient proton molar mass $N_{\mathrm{A}} m_{\mathrm{p}}$ proton Compton wavelength $h / m_{p} c$ | $m_{\mathrm{p}} / m_{\mathrm{n}}$ | 0.998623478 24(46) |  | $4.6 \times 10^{-10}$ |
|  | $e / m_{\mathrm{p}}$ | $9.57883392(24) \times 10^{7}$ | C kg ${ }^{-1}$ | $2.5 \times 10^{-8}$ |
|  | $M(\mathrm{p}), M_{\mathrm{p}}$ | $1.00727646677(10) \times 10^{-3}$ | $\mathrm{kg} \mathrm{mol}^{-1}$ | $1.0 \times 10^{-10}$ |
|  | $\lambda_{\text {C, }}$ | $1.3214098446(19) \times 10^{-15}$ | m | $1.4 \times 10^{-9}$ |
| $\lambda_{\mathrm{C}, \mathrm{p}} / 2 \pi$ | $\lambda_{\text {C,p }}$ | $0.21030890861(30) \times 10^{-15}$ | m | $1.4 \times 10^{-9}$ |
| proton rms charge radius | $R_{\text {p }}$ | $0.8768(69) \times 10^{-15}$ | m | $7.8 \times 10^{-3}$ |
| proton magnetic moment | $\mu_{\mathrm{p}}$ | $1.410606662(37) \times 10^{-26}$ | $\mathrm{J} \mathrm{T}^{-1}$ | $2.6 \times 10^{-8}$ |
| to Bohr magneton ratio | $\mu_{\mathrm{p}} / \mu_{\mathrm{B}}$ | $1.521032209(12) \times 10^{-3}$ |  | $8.1 \times 10^{-9}$ |
| to nuclear magneton ratio | $\mu_{\mathrm{p}} / \mu_{\mathrm{N}}$ | 2.792847 356(23) |  | $8.2 \times 10^{-9}$ |
| proton $g$-factor $2 \mu_{\mathrm{p}} / \mu_{\mathrm{N}}$ | $g_{\text {p }}$ | $5.585694713(46)$ |  | $8.2 \times 10^{-9}$ |
| proton-neutron magnetic moment ratio | $\mu_{\mathrm{p}} / \mu_{\mathrm{n}}$ | -1.459 $89806(34)$ |  | $2.4 \times 10^{-7}$ |
| shielded proton magnetic moment $\left(\mathrm{H}_{2} \mathrm{O}\right.$, sphere, $25^{\circ} \mathrm{C}$ ) | $\mu_{\mathrm{p}}^{\prime}$ | $1.410570419(38) \times 10^{-26}$ | $\mathrm{J} \mathrm{T}^{-1}$ | $2.7 \times 10^{-8}$ |
| to Bohr magneton ratio | $\mu_{\mathrm{p}}^{\prime} / \mu_{\mathrm{B}}$ | $1.520993128(17) \times 10^{-3}$ |  | $1.1 \times 10^{-8}$ |
| to nuclear magneton ratio | $\mu_{\mathrm{p}}^{\prime} / \mu_{\mathrm{N}}$ | 2.792775 598(30) |  | $1.1 \times 10^{-8}$ |
| proton magnetic shielding correction $1-\mu_{\mathrm{p}}^{\prime} / \mu_{\mathrm{p}}$ $\left(\mathrm{H}_{2} \mathrm{O}\right.$, sphere, $\left.25^{\circ} \mathrm{C}\right)$ | $\sigma_{\mathrm{p}}^{\prime}$ | $25.694(14) \times 10^{-6}$ |  | $5.3 \times 10^{-4}$ |
| proton gyromagnetic ratio $2 \mu_{\mathrm{p}} / \hbar$ | $\gamma_{\mathrm{p}}$ | $2.675222099(70) \times 10^{8}$ | $\mathrm{s}^{-1} \mathrm{~T}^{-1}$ | $2.6 \times 10^{-8}$ |
|  | $\gamma_{p} / 2 \pi$ | 42.577 4821(11) | MHz T ${ }^{-1}$ | $2.6 \times 10^{-8}$ |
| $\begin{aligned} & \text { shielded proton gyromagnetic } \\ & \text { ratio } 2 \mu_{\mathrm{p}}^{\prime} / \hbar \\ & \left(\mathrm{H}_{2} \mathrm{O}, \text { sphere, } 25^{\circ} \mathrm{C}\right) \end{aligned}$ | $\gamma_{\mathrm{p}}^{\prime}$ | $2.675153362(73) \times 10^{8}$ | $\mathrm{s}^{-1} \mathrm{~T}^{-1}$ | $2.7 \times 10^{-8}$ |
|  | $\gamma_{p}^{\prime} / 2 \pi$ | $\begin{gathered} 42.5763881(12) \\ \text { Neutron, } \mathrm{n} \end{gathered}$ | $\mathrm{MHz} \mathrm{T}{ }^{-1}$ | $2.7 \times 10^{-8}$ |
| neutron mass | $m_{\mathrm{n}}$ | $1.674927211(84) \times 10^{-27}$ | kg | $5.0 \times 10^{-8}$ |
| in $\mathrm{u}, m_{\mathrm{n}}=A_{\mathrm{r}}(\mathrm{n}) \mathrm{u}$ (neutron relative atomic mass times $u$ ) |  | $1.00866491597(43)$ | u | $4.3 \times 10^{-10}$ |
| energy equivalent | $m_{\mathrm{n}} c^{2}$ | $1.505349505(75) \times 10^{-10}$ | J | $5.0 \times 10^{-8}$ |
| in MeV |  | $939.565346(23)$ | MeV | $2.5 \times 10^{-8}$ |
| neutron-electron mass ratio | $m_{\mathrm{n}} / m_{\mathrm{e}}$ | 1838.683 6605(11) |  | $6.0 \times 10^{-10}$ |
| neutron-muon mass ratio | $m_{\mathrm{n}} / m_{\mu}$ | $8.89248409(23)$ |  | $2.5 \times 10^{-8}$ |
| neutron-tau mass ratio | $m_{\mathrm{n}} / m_{\tau}$ | 0.528740 (86) |  | $1.6 \times 10^{-4}$ |
| neutron-proton mass ratio | $m_{\mathrm{n}} / m_{\mathrm{p}}$ | 1.001378419 18(46) |  | $4.6 \times 10^{-10}$ |
| neutron molar mass $N_{\mathrm{A}} m_{\mathrm{n}}$ | $M(\mathrm{n}), M_{\mathrm{n}}$ | $1.00866491597(43) \times 10^{-3}$ | $\mathrm{kg} \mathrm{mol}^{-1}$ | $4.3 \times 10^{-10}$ |
| neutron Compton wavelength $h / m_{\mathrm{n}} c$ | $\lambda_{\text {C, } n}$ | $1.3195908951(20) \times 10^{-15}$ | m | $1.5 \times 10^{-9}$ |
| $\lambda_{\mathrm{C}, \mathrm{n}} / 2 \pi$ | $\lambda_{\text {C,n }}$ | $0.21001941382(31) \times 10^{-15}$ | m | $1.5 \times 10^{-9}$ |
| neutron magnetic moment | $\mu_{\mathrm{n}}$ | $-0.96623641(23) \times 10^{-26}$ | $\mathrm{J} \mathrm{T}^{-1}$ | $2.4 \times 10^{-7}$ |
| to Bohr magneton ratio | $\mu_{\mathrm{n}} / \mu_{\mathrm{B}}$ | $-1.04187563(25) \times 10^{-3}$ |  | $2.4 \times 10^{-7}$ |
| to nuclear magneton ratio | $\mu_{\mathrm{n}} / \mu_{\mathrm{N}}$ | $-1.91304273(45)$ |  | $2.4 \times 10^{-7}$ |
| neutron $g$-factor $2 \mu_{\mathrm{n}} / \mu_{\mathrm{N}}$ | $g_{\text {n }}$ | -3.826 085 45(90) |  | $2.4 \times 10^{-7}$ |
| neutron-electron magnetic moment ratio | $\mu_{\mathrm{n}} / \mu_{\mathrm{e}}$ | $1.04066882(25) \times 10^{-3}$ |  | $2.4 \times 10^{-7}$ |
| neutron-proton magnetic moment ratio | $\mu_{\mathrm{n}} / \mu_{\mathrm{p}}$ | -0.684 979 34(16) |  | $2.4 \times 10^{-7}$ |
| neutron to shielded proton magnetic moment ratio $\left(\mathrm{H}_{2} \mathrm{O}\right.$, sphere, $25^{\circ} \mathrm{C}$ ) | $\mu_{\mathrm{n}} / \mu_{\mathrm{p}}^{\prime}$ | -0.684 996 94(16) |  | $2.4 \times 10^{-7}$ |

TABLE XLIX. (Continued.)

| Quantity | Symbol | Numerical value | Unit | Relative std. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: |
| neutron gyromagnetic ratio $2\left\|\mu_{\mathrm{n}}\right\| / \hbar$ | $\gamma_{\mathrm{n}}$ | $1.83247185(43) \times 10^{8}$ | $\mathrm{s}^{-1} \mathrm{~T}^{-1}$ | $2.4 \times 10^{-7}$ |
|  | $\gamma_{\mathrm{n}} / 2 \pi$ | 29.164 6954(69) | $\mathrm{MHz} \mathrm{T}{ }^{-1}$ | $2.4 \times 10^{-7}$ |
|  |  | Deuteron, d |  |  |
| deuteron mass | $m_{\text {d }}$ | $3.34358320(17) \times 10^{-27}$ | kg | $5.0 \times 10^{-8}$ |
| in $\mathrm{u}, m_{\mathrm{d}}=A_{\mathrm{r}}(\mathrm{d}) \mathrm{u}$ (deuteron relative atomic mass times $u$ ) |  | 2.013553212 724(78) | u | $3.9 \times 10^{-11}$ |
| energy equivalent in MeV | $m_{\mathrm{d}} c^{2}$ | $3.00506272(15) \times 10^{-10}$ | J | $5.0 \times 10^{-8}$ |
|  |  | 1875.612 793(47) | MeV | $2.5 \times 10^{-8}$ |
| deuteron-electron mass ratio | $m_{\text {d }} / m_{\text {e }}$ | 3670.482 9654(16) |  | $4.3 \times 10^{-10}$ |
| deuteron-proton mass ratio | $m_{\mathrm{d}} / m_{\mathrm{p}}$ | $1.99900750108(22)$ |  | $1.1 \times 10^{-10}$ |
| deuteron molar mass $N_{\mathrm{A}} m_{\mathrm{d}}$ | $M(\mathrm{~d}), M_{\mathrm{d}}$ | $2.013553212724(78) \times 10^{-3}$ | $\mathrm{kg} \mathrm{mol}^{-1}$ | $3.9 \times 10^{-11}$ |
| deuteron rms charge radius | $R_{\text {d }}$ | $2.1402(28) \times 10^{-15}$ | m | $1.3 \times 10^{-3}$ |
| deuteron magnetic moment | $\mu_{\text {d }}$ | $0.433073465(11) \times 10^{-26}$ | $\mathrm{J} \mathrm{T}^{-1}$ | $2.6 \times 10^{-8}$ |
| to Bohr magneton ratio | $\mu_{\mathrm{d}} / \mu_{\mathrm{B}}$ | $0.4669754556(39) \times 10^{-3}$ |  | $8.4 \times 10^{-9}$ |
| to nuclear magneton ratio | $\mu_{\mathrm{d}} / \mu_{\mathrm{N}}$ | 0.8574382308 (72) |  | $8.4 \times 10^{-9}$ |
| deuteron $g$-factor $\mu_{\mathrm{d}} / \mu_{\mathrm{N}}$ | $g_{\text {d }}$ | $0.8574382308(72)$ |  | $8.4 \times 10^{-9}$ |
| deuteron-electon magnetic moment ratio | $\mu_{\mathrm{d}} / \mu_{\mathrm{e}}$ | -4.664 $345537(39) \times 10^{-4}$ |  | $8.4 \times 10^{-9}$ |
| deuteron-proton magnetic moment ratio | $\mu_{\mathrm{d}} / \mu_{\mathrm{p}}$ | 0.307012 2070(24) |  | $7.7 \times 10^{-9}$ |
| deuteron-neutron magnetic moment ratio | $\mu_{\mathrm{d}} / \mu_{\mathrm{n}}$ | -0.448 $20652(11)$ |  | $2.4 \times 10^{-7}$ |
|  |  | Triton, t |  |  |
| triton mass <br> in $\mathrm{u}, m_{\mathrm{t}}=A_{\mathrm{r}}(\mathrm{t}) \mathrm{u}$ (triton relative atomic mass times $\mathbf{u}$ ) | $m_{\text {t }}$ | $5.00735588(25) \times 10^{-27}$ | kg | $5.0 \times 10^{-8}$ |
|  |  | $3.0155007134(25)$ | u | $8.3 \times 10^{-10}$ |
| energy equivalent | $m_{\mathrm{t}} c^{2}$ | $4.50038703(22) \times 10^{-10}$ | J | $5.0 \times 10^{-8}$ |
| in MeV |  | $2808.920906(70)$ | MeV | $2.5 \times 10^{-8}$ |
| triton-electron mass ratio | $m_{\mathrm{t}} / m_{\mathrm{e}}$ | $5496.9215269(51)$ |  | $9.3 \times 10^{-10}$ |
| triton-proton mass ratio | $m_{\mathrm{t}} / m_{\mathrm{p}}$ | $2.9937170309(25)$ |  | $8.4 \times 10^{-10}$ |
| triton molar mass $N_{\text {A }} m_{\mathrm{t}}$ | $M(\mathrm{t}), M_{\mathrm{t}}$ | $3.0155007134(25) \times 10^{-3}$ | $\mathrm{kg} \mathrm{mol}^{-1}$ | $8.3 \times 10^{-10}$ |
| triton magnetic moment | $\mu_{\text {t }}$ | $1.504609361(42) \times 10^{-26}$ | $\mathrm{J} \mathrm{T}^{-1}$ | $2.8 \times 10^{-8}$ |
| to Bohr magneton ratio | $\mu_{\mathrm{t}} / \mu_{\mathrm{B}}$ | $1.622393657(21) \times 10^{-3}$ |  | $1.3 \times 10^{-8}$ |
| to nuclear magneton ratio | $\mu_{\mathrm{t}} / \mu_{\mathrm{N}}$ | 2.978962 448(38) |  | $1.3 \times 10^{-8}$ |
| triton $g$-factor $2 \mu_{\mathrm{t}} / \mu_{\mathrm{N}}$ | $g_{\text {t }}$ | 5.957924 896(76) |  | $1.3 \times 10^{-8}$ |
| triton-electron magnetic moment ratio | $\mu_{t} / \mu_{\mathrm{e}}$ | $-1.620514423(21) \times 10^{-3}$ |  | $1.3 \times 10^{-8}$ |
| triton-proton magnetic moment ratio | $\mu_{\mathrm{t}} / \mu_{\mathrm{p}}$ | 1.066639 908(10) |  | $9.8 \times 10^{-9}$ |
| triton-neutron magnetic moment ratio | $\mu_{\mathrm{t}} / \mu_{\mathrm{n}}$ | -1.557185 53(37) |  | $2.4 \times 10^{-7}$ |
|  |  | Helion, h |  |  |
| helion ( ${ }^{3} \mathrm{He}$ nucleus) mass in $\mathrm{u}, m_{\mathrm{h}}=A_{\mathrm{r}}(\mathrm{h}) \mathrm{u}$ (helion relative atomic mass times $u$ ) | $m_{\text {h }}$ | $5.00641192(25) \times 10^{-27}$ | kg | $5.0 \times 10^{-8}$ |
|  |  | 3.014932 2473(26) | u | $8.6 \times 10^{-10}$ |
| energy equivalent in MeV | $m_{\mathrm{h}} c^{2}$ | $4.49953864(22) \times 10^{-10}$ | J | $5.0 \times 10^{-8}$ |
|  |  | $2808.391383(70)$ | MeV | $2.5 \times 10^{-8}$ |
| helion-electron mass ratio | $m_{\mathrm{h}} / m_{\mathrm{e}}$ | 5495.885 2765(52) |  | $9.5 \times 10^{-10}$ |

TABLE XLIX. (Continued.)

| Quantity | Symbol | Numerical value | Unit | Relative std. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: |
| helion-proton mass ratio | $m_{\mathrm{h}} / m_{\mathrm{p}}$ | 2.9931526713 (26) |  | $8.7 \times 10^{-10}$ |
| helion molar mass $N_{\mathrm{A}} m_{\mathrm{h}}$ | $M(\mathrm{~h}), M_{\mathrm{h}}$ | $3.0149322473(26) \times 10^{-3}$ | $\mathrm{kg} \mathrm{mol}^{-1}$ | $8.6 \times 10^{-10}$ |
| shielded helion magnetic moment (gas, sphere, $25^{\circ} \mathrm{C}$ ) | $\mu_{\text {h }}^{\prime}$ | $-1.074552982(30) \times 10^{-26}$ | $\mathrm{J} \mathrm{T}^{-1}$ | $2.8 \times 10^{-8}$ |
| to Bohr magneton ratio | $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{B}}$ | $-1.158671471(14) \times 10^{-3}$ |  | $1.2 \times 10^{-8}$ |
| to nuclear magneton ratio | $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{N}}$ | -2.127 $497718(25)$ |  | $1.2 \times 10^{-8}$ |
| shielded helion to proton magnetic moment ratio (gas, sphere, $25^{\circ} \mathrm{C}$ ) | $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}$ | -0.761 $766558(11)$ |  | $1.4 \times 10^{-8}$ |
| shielded helion to shielded proton magnetic moment ratio (gas $/ \mathrm{H}_{2} \mathrm{O}$, spheres, $25^{\circ} \mathrm{C}$ ) | $\mu_{\mathrm{h}}^{\prime} / \mu_{\mathrm{p}}^{\prime}$ | -0.7617861313(33) |  | $4.3 \times 10^{-9}$ |
| shielded helion gyromagnetic ratio $2\left\|\mu_{\mathrm{h}}^{\prime}\right\| / \hbar$ (gas, sphere, $25^{\circ} \mathrm{C}$ ) | $\gamma_{\mathrm{h}}^{\prime}$ | $2.037894730(56) \times 10^{8}$ | $\mathrm{s}^{-1} \mathrm{~T}^{-1}$ | $2.8 \times 10^{-8}$ |
|  | $\gamma_{\mathrm{h}}^{\prime} / 2 \pi$ | $32.43410198(90)$ | $\mathrm{MHz} \mathrm{T}{ }^{-1}$ | $2.8 \times 10^{-8}$ |
|  |  | Alpha particle, $\alpha$ |  |  |
| alpha particle mass | $m_{\alpha}$ | $6.64465620(33) \times 10^{-27}$ | kg | $5.0 \times 10^{-8}$ |
| in $\mathrm{u}, m_{\alpha}=A_{\mathrm{r}}(\alpha) \mathrm{u}$ (alpha particle relative atomic mass times $u$ ) |  | 4.001506179 127(62) | u | $1.5 \times 10^{-11}$ |
| energy equivalent | $m_{\alpha} c^{2}$ | $5.97191917(30) \times 10^{-10}$ | J | $5.0 \times 10^{-8}$ |
| in MeV |  | 3727.379 109(93) | MeV | $2.5 \times 10^{-8}$ |
| $\alpha$ particle to electron mass ratio | $m_{\alpha} / m_{\text {e }}$ | $7294.2995365(31)$ |  | $4.2 \times 10^{-10}$ |
| $\alpha$ particle to proton mass ratio | $m_{\alpha} / m_{\mathrm{p}}$ | 3.972599689 51(41) |  | $1.0 \times 10^{-10}$ |
| $\alpha$ particle molar mass $N_{\text {A }} m_{\alpha}$ | $M(\alpha), M_{\alpha}$ | $4.001506179127(62) \times 10^{-3}$ | $\mathrm{kg} \mathrm{mol}^{-1}$ | $1.5 \times 10^{-11}$ |
| PHYSICOCHEMICAL |  |  |  |  |
| Avogadro constant | $N_{\text {A }}, L$ | $6.02214179(30) \times 10^{23}$ | $\mathrm{mol}^{-1}$ | $5.0 \times 10^{-8}$ |
| atomic mass constant $\begin{aligned} m_{\mathrm{u}} & =\frac{1}{12} m\left({ }^{12} \mathrm{C}\right)=1 \mathrm{u} \\ & =10^{-3} \mathrm{~kg} \mathrm{~mol}^{-1} / N_{\mathrm{A}} \end{aligned}$ | $m_{\mathrm{u}}$ | $1.660538782(83) \times 10^{-27}$ | kg | $5.0 \times 10^{-8}$ |
| energy equivalent | $m_{\mathrm{u}} c^{2}$ | $1.492417830(74) \times 10^{-10}$ | J | $5.0 \times 10^{-8}$ |
| in MeV |  | $931.494028(23)$ | MeV | $2.5 \times 10^{-8}$ |
| Faraday constant ${ }^{\text {f }} N_{\mathrm{A}} e$ | F | 96485.3399 (24) | $\mathrm{C} \mathrm{mol}^{-1}$ | $2.5 \times 10^{-8}$ |
| molar Planck constant | $N_{\text {A }} h$ | $3.9903126821(57) \times 10^{-10}$ | $\mathrm{J} \mathrm{s} \mathrm{mol}{ }^{-1}$ | $1.4 \times 10^{-9}$ |
|  | $N_{\text {A }} h c$ | $0.11962656472(17)$ | $\mathrm{J} \mathrm{m} \mathrm{mol}^{-1}$ | $1.4 \times 10^{-9}$ |
| molar gas constant | $R$ | 8.314 472(15) | $\mathrm{J} \mathrm{mol}^{-1} \mathrm{~K}^{-1}$ | $1.7 \times 10^{-6}$ |
| Boltzmann constant $R / N_{\mathrm{A}}$ in $\mathrm{eV} \mathrm{K}^{-1}$ | $k$ | $1.3806504(24) \times 10^{-23}$ | $\mathrm{JK}^{-1}$ | $1.7 \times 10^{-6}$ |
|  |  | $8.617343(15) \times 10^{-5}$ | $\mathrm{eV} \mathrm{K}{ }^{-1}$ | $1.7 \times 10^{-6}$ |
|  | $k / h$ | $2.0836644(36) \times 10^{10}$ | Hz K ${ }^{-1}$ | $1.7 \times 10^{-6}$ |
|  | $k / h c$ | 69.503 56(12) | $\mathrm{m}^{-1} \mathrm{~K}^{-1}$ | $1.7 \times 10^{-6}$ |
| molar volume of ideal gas $R T / p$ $T=273.15 \mathrm{~K}, p=101.325 \mathrm{kPa}$ | $V_{\mathrm{m}}$ | $22.413996(39) \times 10^{-3}$ | $\mathrm{m}^{3} \mathrm{~mol}^{-1}$ | $1.7 \times 10^{-6}$ |
| Loschmidt constant $N_{\mathrm{A}} / V_{\mathrm{m}}$ | $n_{0}$ | $2.6867774(47) \times 10^{25}$ | $\mathrm{m}^{-3}$ | $1.7 \times 10^{-6}$ |
| $T=273.15 \mathrm{~K}, p=100 \mathrm{kPa}$ | $V_{\mathrm{m}}$ | $22.710981(40) \times 10^{-3}$ | $\mathrm{m}^{3} \mathrm{~mol}^{-1}$ | $1.7 \times 10^{-6}$ |
| Sackur-Tetrode constant (absolute entropy constant) ${ }^{g}$ $\frac{5}{2}+\ln \left[\left(2 \pi m_{\mathrm{u}} k T_{1} / h^{2}\right)^{3 / 2} k T_{1} / p_{0}\right]$ |  |  |  |  |
| $T_{1}=1 \mathrm{~K}, p_{0}=100 \mathrm{kPa}$ | $S_{0} / R$ | -1.151 7047(44) |  | $3.8 \times 10^{-6}$ |
| $T_{1}=1 \mathrm{~K}, p_{0}=101.325 \mathrm{kPa}$ |  | -1.164 8677(44) |  | $3.8 \times 10^{-6}$ |

TABLE XLIX. (Continued.)

| Quantity | Symbol | Numerical value | Unit | Relative std. <br> uncert. $u_{\mathrm{r}}$ |
| :--- | :--- | :--- | :--- | :--- |
| Stefan-Boltzmann constant <br> $\left(\pi^{2} / 60\right) k^{4} / \hbar^{3} c^{2}$ | $\sigma$ | $5.670400(40) \times 10^{-8}$ | $\mathrm{~W} \mathrm{~m}^{-2} \mathrm{~K}^{-4}$ | $7.0 \times 10^{-6}$ |
| first radiation constant $2 \pi h c^{2}$ | $c_{1}$ | $3.74177118(19) \times 10^{-16}$ | $\mathrm{~W} \mathrm{~m}^{2}$ | $5.0 \times 10^{-8}$ |
| first radiation constant for spectral <br> radiance $2 h c^{2}$ | $c_{1 \mathrm{~L}}$ | $1.191042759(59) \times 10^{-16}$ | $\mathrm{~W} \mathrm{~m}^{2} \mathrm{sr}^{-1}$ | $5.0 \times 10^{-8}$ |
| second radiation constant $h c / k$ | $c_{2}$ | $1.4387752(25) \times 10^{-2}$ | m K | $1.7 \times 10^{-6}$ |
| Wien displacement law constants <br> $b=\lambda_{\text {max }} T=c_{2} / 4.965114231 \ldots$ | $b$ | $2.8977685(51) \times 10^{-3}$ | m K | $1.7 \times 10^{-6}$ |
| $b^{\prime}=\nu_{\text {max }} / T=2.821439372 \ldots c / c_{2}$ | $b^{\prime}$ | $5.878933(10) \times 10^{10}$ | $\mathrm{~Hz} \mathrm{~K}{ }^{-1}$ | $1.7 \times 10^{-6}$ |

${ }^{\text {a }}$ See Table LI for the conventional value adopted internationally for realizing representations of the volt using the Josephson effect.
${ }^{\mathrm{b}}$ See Table LI for the conventional value adopted internationally for realizing representations of the ohm using the quantum Hall effect.
${ }^{\text {c }}$ Value recommended by the Particle Data Group (Yao et al., 2006).
${ }^{\mathrm{d}}$ Based on the ratio of the masses of the W and Z bosons $m_{\mathrm{W}} / m_{\mathrm{Z}}$ recommended by the Particle Data Group (Yao et al., 2006). The value for $\sin ^{2} \theta_{\mathrm{W}}$ they recommend, which is based on a particular variant of the modified minimal subtraction scheme $(\overline{\mathrm{MS}})$, is $\sin ^{2} \hat{\theta}_{\mathrm{W}}\left(M_{\mathrm{Z}}\right)=0.23122(15)$.
${ }^{\text {e }}$ This and all other values involving $m_{\tau}$ are based on the value of $m_{\tau} c^{2}$ in MeV recommended by the Particle Data Group (Yao et al., 2006), but with a standard uncertainty of 0.29 MeV rather than the quoted uncertainty of $-0.26 \mathrm{MeV},+0.29 \mathrm{MeV}$.
${ }^{\mathrm{f}}$ The numerical value of $F$ to be used in coulometric chemical measurements is $96485.3401(48)$ [5.0×10 $0^{-8}$ ] when the relevant current is measured in terms of representations of the volt and ohm based on the Josephson and quantum Hall effects and the internationally adopted conventional values of the Josephson and von Klitzing constants $K_{\mathrm{J}-90}$ and $R_{\mathrm{K}-90}$ given in Table LI.
${ }^{\mathrm{g}}$ The entropy of an ideal monatomic gas of relative atomic mass $A_{\mathrm{r}}$ is given by $S=S_{0}+\frac{3}{2} R \ln A_{\mathrm{r}}-R \ln \left(p / p_{0}\right)+\frac{5}{2} R \ln (T / K)$.
sponding value expressed in another unit (in essence, an automated version of Tables LIV and LV).

As discussed in Sec. V, well after the 31 December 2006 closing date of the 2006 adjustment and the 29 March 2007 distribution date of the 2006 recommended values on the web, Aoyama et al. (2007) reported their discovery of an error in the coefficient $A_{1}^{(8)}$ in the theoretical expression for the electron magnetic moment anomaly $a_{\mathrm{e}}$. Use of the new coefficient would lead to an increase in the 2006 recommended value of $\alpha$ by 6.8
times its uncertainty, and an increase of its uncertainty by a factor of 1.02 . The recommended values and uncertainties of constants that depend solely on $\alpha$, or on $\alpha$ in combination with other constants with $u_{\mathrm{r}}$ no larger than a few parts in $10^{10}$, would change in the same way. However, the changes in the recommended values of the vast majority of the constants listed in the tables would lie in the range 0 to 0.5 times their 2006 uncertainties, and their uncertainties would remain essentially unchanged.

TABLE L. The variances, covariances, and correlation coefficients of the values of a selected group of constants based on the 2006 CODATA adjustment. The numbers in bold above the main diagonal are $10^{16}$ times the numerical values of the relative covariances, the numbers in bold on the main diagonal are $10^{16}$ times the numerical values of the relative variances, and the numbers in italics below the main diagonal are the correlation coefficients. ${ }^{\text {a }}$

|  | $\alpha$ | $h$ | $e$ | $m_{\mathrm{e}}$ | $N_{\mathrm{A}}$ | $m_{\mathrm{e}} / m_{\mu}$ | $F$ |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: | ---: |
| $\alpha$ | $\mathbf{0 . 0 0 4 7}$ | $\mathbf{0 . 0 0 0 2}$ | $\mathbf{0 . 0 0 2 4}$ | $-\mathbf{0 . 0 0 9 2}$ | $\mathbf{0 . 0 0 9 2}$ | $-\mathbf{0 . 0 0 9 2}$ | $\mathbf{0 . 0 1 1 6}$ |
| $h$ | 0.0005 | $\mathbf{2 4 . 8 6 1 4}$ | $\mathbf{1 2 . 4 3 0 8}$ | $\mathbf{2 4 . 8 6 1 1}$ | $-\mathbf{2 4 . 8 6 1 0}$ | $-\mathbf{0 . 0 0 0 3}$ | $\mathbf{- 1 2 . 4 3 0 2}$ |
| $e$ | 0.0142 | 0.9999 | $\mathbf{6 . 2 1 6 6}$ | $\mathbf{1 2 . 4 2 5 9}$ | $-\mathbf{1 2 . 4 2 5 9}$ | $-\mathbf{0 . 0 0 4 8}$ | $-\mathbf{- 6 . 2 0 9 3}$ |
| $m_{\mathrm{e}}$ | -0.0269 | 0.9996 | 0.9992 | $\mathbf{2 4 . 8 7 9 5}$ | $-\mathbf{2 4 . 8 7 9 4}$ | $\mathbf{0 . 0 1 8 0}$ | $-\mathbf{1 2 . 4 5 3 5}$ |
| $N_{\mathrm{A}}$ | 0.0269 | -0.9996 | -0.9991 | -1.0000 | $\mathbf{2 4 . 8 8 1 1}$ | $-\mathbf{0 . 0 1 8 0}$ | $\mathbf{1 2 . 4 5 5 2}$ |
| $m_{\mathrm{e}} / m_{\mu}$ | -0.0528 | 0.0000 | -0.0008 | 0.0014 | -0.0014 | $\mathbf{6 . 4 2 9 6}$ | $-\mathbf{0 . 0 2 2 7}$ |
| $F$ | 0.0679 | -0.9975 | -0.9965 | -0.9990 | 0.9991 | -0.0036 | $\mathbf{6 . 2 4 5 9}$ |

[^3]TABLE LI. Internationally adopted values of various quantities.

| Quantity | Symbol | Numerical value | Unit | Relative std. <br> uncert. $u_{\mathrm{r}}$ |
| :--- | :--- | :--- | :--- | :--- |
| relative atomic mass ${ }^{\mathrm{a}}$ of ${ }^{12} \mathrm{C}$ | $\left.A_{\mathrm{r}}{ }^{(12} \mathrm{C}\right)$ | 12 |  | (Exact) |
| molar mass constant | $M_{\mathrm{u}}$ | $1 \times 10^{-3}$ | $\mathrm{~kg} \mathrm{~mol}^{-1}$ | (Exact) |
| molar mass of ${ }^{12} \mathrm{C}$ | $M\left({ }^{12} \mathrm{C}\right)$ | $12 \times 10^{-3}$ | $\mathrm{~kg} \mathrm{~mol}^{-1}$ | (Exact) |
| conventional value of Josephson constant ${ }^{\mathrm{b}}$ | $K_{\mathrm{J}-90}$ | 483597.9 | $\mathrm{GHz} \mathrm{V}^{-1}$ | (Exact) |
| conventional value of von Klitzing constant ${ }^{\mathrm{c}}$ | $R_{\mathrm{K}-90}$ | 25812.807 | $\Omega$ | (Exact) |
| standard atmosphere |  | 101325 | Pa | (Exact) |

${ }^{\text {a }}$ The relative atomic mass $A_{\mathrm{r}}(X)$ of particle $X$ with mass $m(X)$ is $A_{\mathrm{r}}(X)=m(X) / m_{\mathrm{u}}$, where $m_{\mathrm{u}}=m\left({ }^{12} \mathrm{C}\right) / 12=M_{\mathrm{u}} / N_{\mathrm{A}}=1 \mathrm{u}$ is the atomic mass constant, $M_{\mathrm{u}}$ is the molar mass constant, $N_{\mathrm{A}}$ is the Avogadro constant, and u is the unified atomic mass unit. Thus the mass of particle $X$ is $m(X)=A_{\mathrm{r}}(X) \mathrm{u}$ and the molar mass of $X$ is $M(X)=A_{\mathrm{r}}(X) M_{\mathrm{u}}$.
${ }^{\mathrm{b}}$ This is the value adopted internationally for realizing representations of the volt using the Josephson effect.
${ }^{\text {c }}$ This is the value adopted internationally for realizing representations of the ohm using the quantum Hall effect.

## XIV. SUMMARY AND CONCLUSION

We conclude by (i) comparing the 2006 and 2002 CODATA recommended values of the constants and identifying those new results that have contributed most to the changes from the 2002 values, (ii) presenting some conclusions that can be drawn from the 2006 recommended values and analysis of the 2006 input data, and (iii) looking to the future and identifying experimental and theoretical work that can advance our knowledge of the values of the constants.

## A. Comparison of 2006 and 2002 CODATA recommended values

The 2006 and 2002 recommended values of a representative group of constants are compared in Table LVI. Regularities in the numbers in columns 2-4 arise because many constants are obtained from expressions
proportional to $\alpha, h$, or $R$ raised to various powers. Thus, the first six quantities are calculated from expressions proportional to $\alpha^{a}$, where $|a|=1,2,3$, or 6 . The next 15 quantities, $h$ through $\mu_{\mathrm{p}}$, are calculated from expressions containing the factor $h^{a}$, where $|a|=1$ or $\frac{1}{2}$. And the five quantities $R$ through $\sigma$ are proportional to $R^{a}$, where $|a|=1$ or 4 .

Further comments on the entries in Table LVI are as follows.
(i) The uncertainty of the 2002 recommended value of $\alpha$ has been reduced by nearly a factor of 5 by the measurement of $a_{\mathrm{e}}$ at Harvard University and the improved theoretical expression for $a_{\mathrm{e}}(\mathrm{th})$. The difference between the Harvard result and the earlier University of Washington result, which played a major role in the determination of $\alpha$ in the 2002 adjustment, accounts for most of the change in the recommended value of $\alpha$ from 2002 to 2006.

TABLE LII. Values of some x-ray-related quantities based on the 2006 CODATA adjustment of the values of the constants.

| Quantity | Symbol | Numerical value | Unit | Relative std. uncert. $u_{\mathrm{r}}$ |
| :---: | :---: | :---: | :---: | :---: |
| Cux unit: $\lambda\left(\mathrm{CuK} \alpha_{1}\right) / 1537.400$ | $\mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)$ | $1.00207699(28) \times 10^{-13}$ | m | $2.8 \times 10^{-7}$ |
| Mo x unit: $\lambda\left(\operatorname{MoK} \alpha_{1}\right) / 707.831$ | $\mathrm{xu}\left(\mathrm{MoK} \alpha_{1}\right)$ | $1.00209955(53) \times 10^{-13}$ | m | $5.3 \times 10^{-7}$ |
| ångström star: $\lambda\left(\mathrm{WK} \alpha_{1}\right) / 0.2090100$ | A* | $1.00001498(90) \times 10^{-10}$ | m | $9.0 \times 10^{-7}$ |
| lattice parameter ${ }^{\mathrm{a}}$ of Si (in vacuum, $22.5^{\circ} \mathrm{C}$ ) | $a$ | $543.102064(14) \times 10^{-12}$ | m | $2.6 \times 10^{-8}$ |
| $\{220\}$ lattice spacing of Si $a / \sqrt{8}$ (in vacuum, $22.5^{\circ} \mathrm{C}$ ) | $d_{220}$ | $192.0155762(50) \times 10^{-12}$ | m | $2.6 \times 10^{-8}$ |
| molar volume of Si $M(\mathrm{Si}) / \rho(\mathrm{Si})=N_{\mathrm{A}} a^{3} / 8$ <br> (in vacuum, $22.5^{\circ} \mathrm{C}$ ) | $V_{\mathrm{m}}(\mathrm{Si})$ | $12.0588349(11) \times 10^{-6}$ | $\mathrm{m}^{3} \mathrm{~mol}^{-1}$ | $9.1 \times 10^{-8}$ |

[^4]TABLE LIII. The values in SI units of some non-SI units based on the 2006 CODATA adjustment of the values of the constants.

| Quantity |  |  |  |  |
| :--- | :--- | :--- | :--- | :--- |
|  | Symbol | Numerical value | Relative std. |  |
| uncert. $u_{\mathrm{r}}$ |  |  |  |  |

TABLE LIII. (Continued.)

| Quantity | Symbol | Numerical value | Unit | Relative std. <br> uncert. $u_{\mathrm{r}}$ |
| :--- | :--- | :--- | :--- | :--- |
| a.u. of magnetizability | $e^{2} a_{0}^{2} / m_{\mathrm{e}}$ | $7.891036433(27) \times 10^{-29}$ | $\mathrm{~J} \mathrm{~T}^{-2}$ | $3.4 \times 10^{-9}$ |
| a.u. of permittivity $\left(10^{7} / c^{2}\right)$ | $e^{2} / a_{0} E_{\mathrm{h}}$ | $1.112650056 \ldots \times 10^{-10}$ | $\mathrm{~F} \mathrm{~m}^{-1}$ | (Exact) |

(ii) The uncertainty of the 2002 recommended value of $h$ has been reduced by over a factor of 3 due to the new NIST watt-balance result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ and because the factor used to increase the uncertainties of the data related to $h$ (applied to reduce the inconsistencies among the data) was reduced from 2.325 in the 2002 adjustment to 1.5 in the 2006 adjustment. That the change in value from 2002 to 2006 is small is due to the excellent agreement between the new value of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ and the earlier NIST and NPL values, which played a major role in the determination of $h$ in the 2002 adjustment.
(iii) The updating of two measurements that contributed to the determination of the 2002 recommended value of $G$ reduced the spread in the values and reinforced the most accurate result, that from the University of Washington. On this basis, the Task Group reduced
the assigned uncertainty from $u_{\mathrm{r}}=1.5 \times 10^{-4}$ in 2002 to $u_{\mathrm{r}}=1.0 \times 10^{-4}$ in 2006. This uncertainty reflects the historical difficulty of measuring $G$. Although the recommended value is the weighted mean of the eight available values, the assigned uncertainty is still over four times the uncertainty of the mean multiplied by the corresponding Birge ratio $R_{\mathrm{B}}$.
(iv) The large shift in the recommended value of $d_{220}$ from 2002 to 2006 is due to the fact that in the 2002 adjustment only the NMIJ result for $d_{220}$ (NR3) was included, while in the 2006 adjustment this result (but updated by more recent NMIJ measurements) was included together with the PTB result for $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ and the new INRIM results for $d_{220}(\mathrm{~W} 4.2 \mathrm{a})$ and $d_{220}\left(\mathrm{MO}^{*}\right)$. Moreover, the NMIJ value of $d_{220}$ inferred from

TABLE LIV. The values of some energy equivalents derived from $E=m c^{2}=h c / \lambda=h \nu=k T$, and based on the 2006 CODATA adjustment of the values of the constants; $1 \mathrm{eV}=(e / \mathrm{C}) \mathrm{J}, 1 \mathrm{u}=m_{\mathrm{u}}=\frac{1}{12} m\left({ }^{12} \mathrm{C}\right)=10^{-3} \mathrm{~kg} \mathrm{~mol}^{-1} / N_{\mathrm{A}}$, and $E_{\mathrm{h}}=2 R_{\propto} h c=\alpha^{2} m_{\mathrm{e}} c^{2}$ is the Hartree energy (hartree).

| Relevant unit |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
|  | J | kg | $\mathrm{m}^{-1}$ | Hz |
| 1 J | $(1 \mathrm{~J})=$ | (1 J) $/ c^{2}=$ | (1 J) $/ h c=$ | (1 J) $/ h=$ |
|  | 1 J | $1.112650056 \ldots \times 10^{-17} \mathrm{~kg}$ | $5.03411747(25) \times 10^{24} \mathrm{~m}^{-1}$ | $1.509190450(75) \times 10^{33} \mathrm{~Hz}$ |
| 1 kg | $(1 \mathrm{~kg}) c^{2}=$ | $(1 \mathrm{~kg})=$ | $(1 \mathrm{~kg}) c / h=$ | $(1 \mathrm{~kg}) c^{2} / h=$ |
|  | $8.987551787 \ldots \times 10^{16} \mathrm{~J}$ | 1 kg | $4.52443915(23) \times 10^{41} \mathrm{~m}^{-1}$ | $1.356392733(68) \times 10^{50} \mathrm{~Hz}$ |
| $1 \mathrm{~m}^{-1}$ | $\left(1 \mathrm{~m}^{-1}\right) h c=$ | $\left(1 \mathrm{~m}^{-1}\right) h / c=$ | $\left(1 \mathrm{~m}^{-1}\right)=$ | $\left(1 \mathrm{~m}^{-1}\right) c=$ |
|  | $1.986445501(99) \times 10^{-25} \mathrm{~J}$ | $2.21021870(11) \times 10^{-42} \mathrm{~kg}$ | $1 \mathrm{~m}^{-1}$ | 299792458 Hz |
| 1 Hz | $(1 \mathrm{~Hz}) h=$ | $(1 \mathrm{~Hz}) h / c^{2}=$ | $(1 \mathrm{~Hz}) / \mathrm{c}=$ | $(1 \mathrm{~Hz})=$ |
|  | $6.62606896(33) \times 10^{-34} \mathrm{~J}$ | $7.37249600(37) \times 10^{-51} \mathrm{~kg}$ | $3.335640951 \ldots \times 10^{-9} \mathrm{~m}^{-1}$ | 1 Hz |
| 1 K | $(1 \mathrm{~K}) k=$ | $(1 \mathrm{~K}) k / c^{2}=$ | $(1 \mathrm{~K}) \mathrm{k} / \mathrm{hc}=$ | $(1 \mathrm{~K}) \mathrm{k} / \mathrm{h}=$ |
|  | $1.3806504(24) \times 10^{-23} \mathrm{~J}$ | $1.5361807(27) \times 10^{-40} \mathrm{~kg}$ | $69.50356(12) \mathrm{m}^{-1}$ | $2.0836644(36) \times 10^{10} \mathrm{~Hz}$ |
| 1 eV | $(1 \mathrm{eV})=$ | $(1 \mathrm{eV}) / \mathrm{c}^{2}=$ | $(1 \mathrm{eV}) / h c=$ | $(1 \mathrm{eV}) / h=$ |
|  | $1.602176487(40) \times 10^{-19} \mathrm{~J}$ | $1.782661758(44) \times 10^{-36} \mathrm{~kg}$ | $8.06554465(20) \times 10^{5} \mathrm{~m}^{-1}$ | $2.417989454(60) \times 10^{14} \mathrm{~Hz}$ |
| 1 u | $(1 \mathrm{u}) \mathrm{c}^{2}=$ | $(1 \mathrm{u})=$ | $(1 \mathrm{u}) c / \mathrm{h}=$ | $(1 \mathrm{u}) c^{2} / h=$ |
|  | $1.492417830(74) \times 10^{-10} \mathrm{~J}$ | $1.660538782(83) \times 10^{-27} \mathrm{~kg}$ | $7.513006671(11) \times 10^{14} \mathrm{~m}^{-1}$ | $2.2523427369(32) \times 10^{23} \mathrm{~Hz}$ |
| $1 E_{\mathrm{h}}$ | $\left(1 E_{\mathrm{h}}\right)=$ | $\left(1 E_{\mathrm{h}}\right) / c^{2}=$ | $\left(1 E_{\mathrm{h}}\right) / h c=$ | $\left(1 E_{\mathrm{h}}\right) / \mathrm{h}=$ |
|  | $4.35974394(22) \times 10^{-18} \mathrm{~J}$ | $4.85086934(24) \times 10^{-35} \mathrm{~kg}$ | $2.194746313705(15) \times 10^{7} \mathrm{~m}^{-1}$ | $6.579683920722(44) \times 10^{15} \mathrm{~Hz}$ |

TABLE LV. The values of some energy equivalents derived from $E=m c^{2}=h c / \lambda=h \nu=k T$, and based on the 2006 CODATA adjustment of the values of the constants; $1 \mathrm{eV}=(e / \mathrm{C}) \mathrm{J}, 1 \mathrm{u}=m_{\mathrm{u}}=\frac{1}{12} m\left({ }^{12} \mathrm{C}\right)=10^{-3} \mathrm{~kg} \mathrm{~mol}^{-1} / N_{\mathrm{A}}$, and $E_{\mathrm{h}}=2 R_{\infty} h c=\alpha^{2} m_{\mathrm{e}} c^{2}$ is the Hartree energy (hartree).

| Relevant unit |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
|  | K | eV | u | $E_{\text {h }}$ |
| 1 J | $\begin{aligned} & (1 \mathrm{~J}) / k= \\ & 7.242963(13) \times 10^{22} \mathrm{~K} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~J})= \\ & 6.24150965(16) \times 10^{18} \mathrm{eV} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~J}) / c^{2}= \\ & 6.70053641(33) \times 10^{9} \mathrm{u} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~J})= \\ & 2.29371269(11) \times 10^{17} E_{\mathrm{h}} \end{aligned}$ |
| 1 kg | $\begin{aligned} & (1 \mathrm{~kg}) c^{2} / k= \\ & 6.509651(11) \times 10^{39} \mathrm{~K} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~kg}) c^{2}= \\ & 5.60958912(14) \times 10^{35} \mathrm{eV} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~kg})= \\ & 6.02214179(30) \times 10^{26} \mathrm{u} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~kg}) c^{2}= \\ & 2.06148616(10) \times 10^{34} E_{\mathrm{h}} \end{aligned}$ |
| $1 \mathrm{~m}^{-1}$ | $\begin{aligned} & \left(1 \mathrm{~m}^{-1}\right) h c / k= \\ & 1.4387752(25) \times 10^{-2} \mathrm{~K} \end{aligned}$ | $\begin{aligned} & \left(1 \mathrm{~m}^{-1}\right) h c= \\ & 1.239841875(31) \times 10^{-6} \mathrm{eV} \end{aligned}$ | $\begin{aligned} & \left(1 \mathrm{~m}^{-1}\right) h / c= \\ & 1.3310250394(19) \times 10^{-15} \mathrm{u} \end{aligned}$ | $\begin{aligned} & \left(1 \mathrm{~m}^{-1}\right) h c= \\ & 4.556335252760(30) \times 10^{-8} E_{\mathrm{h}} \end{aligned}$ |
| 1 Hz | $\begin{aligned} & (1 \mathrm{~Hz}) h / k= \\ & 4.7992374(84) \times 10^{-11} \mathrm{~K} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~Hz}) h= \\ & 4.13566733(10) \times 10^{-15} \mathrm{eV} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~Hz}) h / c^{2}= \\ & 4.4398216294(64) \times 10^{-24} \mathrm{u} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~Hz}) h= \\ & 1.519829846006(10) \times 10^{-16} E_{\mathrm{h}} \end{aligned}$ |
| 1 K | $\begin{aligned} & (1 \mathrm{~K})= \\ & 1 \mathrm{~K} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~K}) k= \\ & 8.617343(15) \times 10^{-5} \mathrm{eV} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~K}) k / c^{2}= \\ & 9.251098(16) \times 10^{-14} \mathrm{u} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{~K}) k= \\ & 3.1668153(55) \times 10^{-6} E_{\mathrm{h}} \end{aligned}$ |
| 1 eV | $\begin{aligned} & (1 \mathrm{eV}) / k= \\ & 1.1604505(20) \times 10^{4} \mathrm{~K} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{eV})= \\ & 1 \mathrm{eV} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{eV}) / c^{2}= \\ & 1.073544188(27) \times 10^{-9} \mathrm{u} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{eV})= \\ & 3.674932540(92) \times 10^{-2} E_{\mathrm{h}} \end{aligned}$ |
| 1 u | $\begin{aligned} & (1 \mathrm{u}) c^{2} / k= \\ & 1.0809527(19) \times 10^{13} \mathrm{~K} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{u}) c^{2}= \\ & 931.494028(23) \times 10^{6} \mathrm{eV} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{u})= \\ & 1 \mathrm{u} \end{aligned}$ | $\begin{aligned} & (1 \mathrm{u}) c^{2}= \\ & 3.4231777149(49) \times 10^{7} E_{\mathrm{h}} \end{aligned}$ |
| $1 E_{\mathrm{h}}$ | $\begin{aligned} & \left(1 E_{\mathrm{h}}\right) / k= \\ & 3.1577465(55) \times 10^{5} \mathrm{~K} \end{aligned}$ | $\begin{aligned} & \left(1 E_{\mathrm{h}}\right)= \\ & 27.21138386(68) \mathrm{eV} \end{aligned}$ | $\begin{aligned} & \left(1 E_{\mathrm{h}}\right) / c^{2}= \\ & 2.9212622986(42) \times 10^{-8} \mathrm{u} \end{aligned}$ | $\begin{aligned} & \left(1 E_{\mathrm{h}}\right)= \\ & 1 E_{\mathrm{h}} \end{aligned}$ |

$d_{220}(\mathrm{NR} 3)$ strongly disagrees with the values of $d_{220}$ inferred from the other three results.
(v) The marginally significant shift in the recommended value of $g_{\mu}$ from 2002 to 2006 is mainly due to the following: In the 2002 adjustment, the principal hadronic contribution to the theoretical expression for $a_{\mu}$ was based on both a calculation that included only $\mathrm{e}^{+} \mathrm{e}^{-}$ annihilation data and a calculation that used data from hadronic decays of the $\tau$ in place of some of the $\mathrm{e}^{+} \mathrm{e}^{-}$ annihilation data. In the 2006 adjustment, the principal hadronic contribution was based on a calculation that used only annihilation data because of various concerns that subsequently arose about the reliability of incorporating the $\tau$ data in the calculation; the calculation based on both $\mathrm{e}^{+} \mathrm{e}^{-}$annihilation data and $\tau$ decay data was only used to estimate the uncertainty of the hadronic contribution. Because the results from the two calculations are in significant disagreement, the uncertainty of $a_{\mu}($ th) is comparatively large: $u_{\mathrm{r}}=1.8 \times 10^{-6}$.
(vi) The reduction of the uncertainties of the magnetic moment ratios $\mu_{\mathrm{p}} / \mu_{\mathrm{B}}, \mu_{\mathrm{p}} / \mu_{\mathrm{N}}, \mu_{\mathrm{d}} / \mu_{\mathrm{N}}, \mu_{\mathrm{e}} / \mu_{\mathrm{p}}$, and $\mu_{\mathrm{d}} / \mu_{\mathrm{p}}$ are due to the new NMR measurement of $\mu_{\mathrm{p}}(\mathrm{HD}) / \mu_{\mathrm{d}}(\mathrm{HD})$ and careful reexamination of the calculation of the D-H shielding correction difference $\sigma_{\mathrm{dp}}$. Because the value of the product $\left(\mu_{\mathrm{p}} / \mu_{\mathrm{e}}\right)\left(\mu_{\mathrm{e}} / \mu_{\mathrm{d}}\right)$ im-
plied by the new measurement is highly consistent with the same product implied by the individual measurements of $\mu_{\mathrm{e}}(\mathrm{H}) / \mu_{\mathrm{p}}(\mathrm{H})$ and $\mu_{\mathrm{d}}(\mathrm{D}) / \mu_{\mathrm{e}}(\mathrm{D})$, the changes in the values of the ratios are small.

In summary, the most important differences between the 2006 and 2002 adjustments are that the 2006 adjustment had available new experimental and theoretical results for $a_{\mathrm{e}}$, which provided a dramatically improved value of $\alpha$, and a new result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, which provided a significantly improved value of $h$. These two advances from 2002 to 2006 have resulted in major reductions in the uncertainties of many of the 2006 recommended values compared with their 2002 counterparts.

## B. Some implications of the 2006 CODATA recommended values and adjustment for physics and metrology

A number of conclusions that can be drawn from the 2006 adjustment concerning metrology and the basic theories and experimental methods of physics are presented here, where the focus is on those conclusions that are new or are different from those drawn from the 2002 and 1998 adjustments.

Conventional electric units. One can interpret the adoption of the conventional values $K_{\mathrm{J}-90}$

TABLE LVI. Comparison of the 2006 and 2002 CODATA adjustments of the values of the constants by comparison of the corresponding recommended values of a representative group of constants. Here $D_{\mathrm{r}}$ is the 2006 value minus the 2002 value divided by the standard uncertainty $u$ of the 2002 value (i.e., $D_{\mathrm{r}}$ is the change in the value of the constant from 2002 to 2006 relative to its 2002 standard uncertainty).

| Quantity | 2006 rel. std. uncert. $u_{\mathrm{r}}$ | $\begin{aligned} & \text { Ratio } 2002 u_{\mathrm{r}} \\ & \text { to } 2006 u_{\mathrm{r}} \end{aligned}$ | $D_{\text {r }}$ |
| :---: | :---: | :---: | :---: |
| $\alpha$ | $6.8 \times 10^{-10}$ | 4.9 | -1.3 |
| $R_{\text {K }}$ | $6.8 \times 10^{-10}$ | 4.9 | 1.3 |
| $a_{0}$ | $6.8 \times 10^{-10}$ | 4.9 | -1.3 |
| $\lambda_{\mathrm{C}}$ | $1.4 \times 10^{-9}$ | 4.9 | -1.3 |
| $r_{\text {e }}$ | $2.1 \times 10^{-9}$ | 4.9 | -1.3 |
| $\sigma_{\text {e }}$ | $4.1 \times 10^{-9}$ | 4.9 | -1.3 |
| $h$ | $5.0 \times 10^{-8}$ | 3.4 | -0.3 |
| $m_{\text {e }}$ | $5.0 \times 10^{-8}$ | 3.4 | -0.3 |
| $m_{\text {h }}$ | $5.0 \times 10^{-8}$ | 3.4 | -0.3 |
| $m_{\alpha}$ | $5.0 \times 10^{-8}$ | 3.4 | -0.3 |
| $N_{\text {A }}$ | $5.0 \times 10^{-8}$ | 3.4 | 0.3 |
| $E_{\mathrm{h}}$ | $5.0 \times 10^{-8}$ | 3.4 | -0.3 |
| $c_{1}$ | $5.0 \times 10^{-8}$ | 3.4 | -0.3 |
| $e$ | $2.5 \times 10^{-8}$ | 3.4 | -0.3 |
| $K_{\text {J }}$ | $2.5 \times 10^{-8}$ | 3.4 | 0.3 |
| $F$ | $2.5 \times 10^{-8}$ | 3.4 | 0.2 |
| $\gamma_{\mathrm{p}}^{\prime}$ | $2.7 \times 10^{-8}$ | 3.2 | 0.2 |
| $\mu_{\text {B }}$ | $2.5 \times 10^{-8}$ | 3.4 | -0.4 |
| $\mu_{\text {N }}$ | $2.5 \times 10^{-8}$ | 3.4 | -0.4 |
| $\mu_{\text {e }}$ | $2.5 \times 10^{-8}$ | 3.4 | 0.4 |
| $\mu_{\mathrm{p}}$ | $2.6 \times 10^{-8}$ | 3.3 | -0.4 |
| $R$ | $1.7 \times 10^{-6}$ | 1.0 | 0.0 |
| $k$ | $1.7 \times 10^{-6}$ | 1.0 | 0.0 |
| $V_{\text {m }}$ | $1.7 \times 10^{-6}$ | 1.0 | 0.0 |
| $c_{2}$ | $1.7 \times 10^{-6}$ | 1.0 | 0.0 |
| $\sigma$ | $7.0 \times 10^{-6}$ | 1.0 | 0.0 |
| G | $1.0 \times 10^{-4}$ | 1.5 | 0.1 |
| $R_{\infty}$ | $6.6 \times 10^{-12}$ | 1.0 | 0.0 |
| $m_{\mathrm{e}} / m_{\mathrm{p}}$ | $4.3 \times 10^{-10}$ | 1.1 | 0.2 |
| $m_{\mathrm{e}} / m_{\mu}$ | $2.5 \times 10^{-8}$ | 1.0 | 0.3 |
| $A_{\mathrm{r}}(\mathrm{e})$ | $4.2 \times 10^{-10}$ | 1.0 | -0.1 |
| $A_{\mathrm{r}}(\mathrm{p})$ | $1.0 \times 10^{-10}$ | 1.3 | -0.9 |
| $A_{\mathrm{r}}(\mathrm{n})$ | $4.3 \times 10^{-10}$ | 1.3 | 0.7 |
| $A_{\mathrm{r}}$ (d) | $3.9 \times 10^{-11}$ | 4.5 | 0.1 |
| $A_{\mathrm{r}}(\mathrm{h})$ | $8.6 \times 10^{-10}$ | 2.3 | 0.7 |
| $A_{\mathrm{r}}(\alpha)$ | $1.5 \times 10^{-11}$ | 0.9 | -0.4 |
| $d_{220}$ | $2.6 \times 10^{-8}$ | 1.4 | -2.9 |
| $g_{\text {e }}$ | $7.4 \times 10^{-13}$ | 5.0 | 1.3 |
| $g_{\mu}$ | $6.0 \times 10^{-10}$ | 1.0 | -1.4 |
| $\mu_{\mathrm{p}} / \mu_{\mathrm{B}}$ | $8.1 \times 10^{-9}$ | 1.2 | 0.2 |
| $\mu_{\mathrm{p}} / \mu_{\mathrm{N}}$ | $8.2 \times 10^{-9}$ | 1.2 | 0.2 |
| $\mu_{\mathrm{n}} / \mu_{\mathrm{N}}$ | $2.4 \times 10^{-7}$ | 1.0 | 0.0 |
| $\mu_{\mathrm{d}} / \mu_{\mathrm{N}}$ | $8.4 \times 10^{-9}$ | 1.3 | -0.2 |
| $\mu_{\mathrm{e}} / \mu_{\mathrm{p}}$ | $8.1 \times 10^{-9}$ | 1.2 | 0.2 |

TABLE LVI. (Continued.)

| Quantity | 2006 rel. std. <br> uncert. $u_{\mathrm{r}}$ | Ratio $2002 u_{\mathrm{r}}$ <br> to $2006 u_{\mathrm{r}}$ | $D_{\mathrm{r}}$ |
| :--- | :---: | :---: | :---: |
| $\mu_{\mathrm{n}} / \mu_{\mathrm{p}}$ | $2.4 \times 10^{-7}$ | 1.0 | 0.0 |
| $\mu_{\mathrm{d}} / \mu_{\mathrm{p}}$ | $7.7 \times 10^{-9}$ | 1.9 | -0.3 |

$=483597.9 \mathrm{GHz} / \mathrm{V}$ and $R_{\mathrm{K}-90}=25812.807 \Omega$ for the Josephson and von Klitzing constants as establishing conventional, practical units of voltage and resistance, $V_{90}$ and $\Omega_{90}$, given by $V_{90}=\left(K_{\mathrm{J}-90} / K_{\mathrm{J}}\right) \mathrm{V}$ and $\Omega_{90}$ $=\left(R_{\mathrm{K}} / R_{\mathrm{K}-90}\right) \Omega$. Other conventional electric units follow from $V_{90}$ and $\Omega_{90}$, for example, $A_{90}=V_{90} / \Omega_{90}, C_{90}$ $=A_{90} \mathrm{~s}, \quad W_{90}=A_{90} V_{90}, \quad F_{90}=C_{90} / V_{90}, \quad$ and $\quad H_{90}=\Omega_{90} \mathrm{~s}$, which are the conventional, practical units of current, charge, power, capacitance, and inductance, respectively (Taylor and Mohr, 2001). For the relations between $K_{\mathrm{J}}$ and $K_{\mathrm{J}-90}$, and $R_{\mathrm{K}}$ and $R_{\mathrm{K}-90}$, the 2006 adjustment gives

$$
\begin{align*}
& K_{\mathrm{J}}=K_{\mathrm{J}-90}\left[1-1.9(2.5) \times 10^{-8}\right],  \tag{374}\\
& R_{\mathrm{K}}=R_{\mathrm{K}-90}\left[1+2.159(68) \times 10^{-8}\right], \tag{375}
\end{align*}
$$

which lead to

$$
\begin{align*}
& V_{90}=\left[1+1.9(2.5) \times 10^{-8}\right] \mathrm{V},  \tag{376}\\
& \Omega_{90}=\left[1+2.159(68) \times 10^{-8}\right] \Omega,  \tag{377}\\
& A_{90}=\left[1-0.3(2.5) \times 10^{-8}\right] \mathrm{A},  \tag{378}\\
& C_{90}=\left[1-0.3(2.5) \times 10^{-8}\right] \mathrm{C},  \tag{379}\\
& W_{90}=\left[1+1.6(5.0) \times 10^{-8}\right] \mathrm{W},  \tag{380}\\
& F_{90}=\left[1-2.159(68) \times 10^{-8}\right] \mathrm{F},  \tag{381}\\
& H_{90}=\left[1+2.159(68) \times 10^{-8}\right] \mathrm{H} . \tag{382}
\end{align*}
$$

Equations (376) and (377) show that $V_{90}$ exceeds V and $\Omega_{90}$ exceeds $\Omega$ by $1.9(2.5) \times 10^{-8}$ and $2.159(68) \times 10^{-8}$, respectively. This means that measured voltages and resistances traceable to the Josephson effect and $K_{\mathrm{J}-90}$ and the quantum Hall effect and $R_{\mathrm{K}-90}$, respectively, are too small relative to the SI by these same fractional amounts. However, these differences are well within the $40 \times 10^{-8}$ uncertainty assigned to $V_{90} / \mathrm{V}$ and the 10 $\times 10^{-8}$ uncertainty assigned to $\Omega_{90} / \Omega$ by the Consultative Committee for Electricity and Magnetism (CCEM) of the CIPM (Quinn, 1989, 2001).

Josephson and quantum Hall effects. The study in Sec. XII.B. 2 provides no statistically significant evidence that the fundamental Josephson and quantum Hall effect relations $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ are not exact. The two theories of the most important phenomena of condensed-matter physics are thereby further supported.

Antiprotonic helium. The good agreement between the value of $A_{\mathrm{r}}(\mathrm{e})$ obtained from the measured values
and theoretical expressions for a number of transition frequencies in antiprotonic ${ }^{4} \mathrm{He}$ and ${ }^{3} \mathrm{He}$ with three other values obtained by entirely different methods indicates that these rather complex atoms are reasonably well understood both experimentally and theoretically.

Newtonian constant of gravitation. Although the inconsistencies among the values of $G$ have been reduced somewhat as a result of modifications to two of the eight results available in 2002, the situation remains problematic; there is no evidence that the historic difficulty of measuring $G$ has been overcome.

Tests of $Q E D$. The good agreement of highly accurate values of $\alpha$ inferred from $h / m\left({ }^{133} \mathrm{Cs}\right)$ and $h / m\left({ }^{87} \mathrm{Rb}\right)$, which are only weakly dependent on QED theory, with the values of $\alpha$ inferred from $a_{\mathrm{e}}$, muonium transition frequencies, and H and D transition frequencies, provide support for the QED theory of $a_{\mathrm{e}}$ as well as the bound-state QED theory of muonium and H and D . In particular, the weighted mean of the two $\alpha$ values inferred from $h / m\left({ }^{133} \mathrm{Cs}\right)$ and $h / m\left({ }^{87} \mathrm{Rb}\right)$, $\alpha^{-1}$ $=137.03599934(69)\left[5.0 \times 10^{-9}\right]$, and the weighted mean of the two $\alpha$ values inferred from the two experimental values of $a_{\mathrm{e}}, \alpha^{-1}=137.035999680(94)\left[6.9 \times 10^{-10}\right]$, differ by only $0.5 u_{\text {diff }}$, with $u_{\text {diff }}=5.1 \times 10^{-9}$. This is a truly impressive confirmation of QED theory.

Physics beyond the Standard Model. If the principal hadronic contribution to $a_{\mu}\left(\right.$ th) obtained from the $\mathrm{e}^{+} \mathrm{e}^{-}$ annihilation-data plus $\tau$ hadronic-decay-data calculation (see the previous section) is completely ignored, and the value based on the annihilation-data-only calculation with its uncertainty of $45 \times 10^{-11}$ is used in $a_{\mu}($ th $)$, then the value of $\alpha$ inferred from the BNL experimentally determined value of $a_{\mu}(\exp ), \quad \alpha^{-1}=137.035670(91)$ $\left[6.6 \times 10^{-7}\right]$, differs from the $h / m\left({ }^{133} \mathrm{Cs}\right)-h / m\left({ }^{87} \mathrm{Rb}\right)$ mean value of $\alpha$ by $3.6 u_{\text {diff. Although such a large dis- }}$ crepancy may suggest "new physics," the consensus is that such a view is premature (Davier, 2006).

Electrical and silicon crystal-related measurements. The previously discussed inconsistencies involving the watt-balance determinations of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, the mercury electrometer and voltage balance measurements of $K_{\mathrm{J}}$, the XROI determinations of the $\{220\}$ lattice spacing of various silicon crystals, the measurement of $h / m_{\mathrm{n}} d_{220}$ (W04), and the measurement of $V_{\mathrm{m}}(\mathrm{Si})$ hint at possible problems with one or more of these these rather complex experiments. This suggests that some of the many different measurement techniques required for their execution may not be as well understood as is currently believed.

Redefinition of the kilogram. There has been considerable recent discussion about the possibility of the 24th General Conference on Weights and Measures (CGPM), which convenes in 2011, redefining the kilogram, ampere, kelvin, and mole by linking these SI base units to fixed values of $h, e, k$, and $N_{\mathrm{A}}$, respectively (Mills et al., 2006; Stock and Witt, 2006), in much the same way that the current definition of the meter is linked to a fixed value of $c$ (BIPM, 2006). Before such a definition of the kilogram can be accepted, $h$ should be known with a $u_{\mathrm{r}}$
of a few parts in $10^{8}$. It is therefore noteworthy that the 2006 CODATA recommended value of $h$ has $u_{\mathrm{r}}=5.0$ $\times 10^{-8}$ and the most accurate measured value of $h$ (the 2007 NIST watt-balance result) has $u_{\mathrm{r}}=3.6 \times 10^{-8}$.

## C. Outlook and suggestions for future work

Because there is little redundancy among some of the key input data, the 2006 CODATA set of recommended values, like its 2002 and 1998 predecessors, does not rest on as solid a foundation as one might wish. The constants $\alpha, h$, and $R$ play a critical role in determining many other constants, yet the recommended value of each is determined by a severely limited number of input data. Moreover, some input data for the same quantity have uncertainties of considerably different magnitudes and hence these data contribute to the final adjustment with considerably different weights.

The input datum that primarily determines $\alpha$ is the 2006 experimental result for $a_{\mathrm{e}}$ from Harvard University with $u_{\mathrm{r}}=6.5 \times 10^{-10}$; the uncertainty $u_{\mathrm{r}}=37 \times 10^{-10}$ of the next most accurate experimental result for $a_{\mathrm{e}}$, that reported by the University of Washington in 1987, is 5.7 times larger. Furthermore, there is only a single value of the eighth-order coefficient $A_{1}^{(8)}$, that due to Kinoshita and Nio; it plays a critical role in the theoretical expression for $a_{\mathrm{e}}$ from which $\alpha$ is obtained and requires lengthy QED calculations.

The 2007 NIST watt-balance result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ with $u_{\mathrm{r}}$ $=3.6 \times 10^{-8}$ is the primary input datum that determines $h$, since the uncertainty of the next most accurate value of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$, the NIST 1998 result, is 2.4 times larger. Further, the 2005 consensus value of $V_{\mathrm{m}}(\mathrm{Si})$ disagrees with all three high accuracy measurements of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ currently available.

For $R$, the key input datum is the 1998 NIST value based on speed-of-sound measurements in argon using a spherical acoustic resonator with $u_{\mathrm{r}}=1.7 \times 10^{-6}$. The uncertainty of the next most accurate value, the 1979 NPL result, also obtained from speed of sound measurements in argon but using an acoustic interferometer, is 4.7 times larger.

Lack of redundancy is, of course, not the only difficulty with the 2006 adjustment. An equally important but not fully independent issue is the several inconsistencies involving some of the electrical and silicon crystal-related input data as already discussed, including the recently reported preliminary result for $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ from the NPL watt balance given in Sec VII.D.1. There is also the issue of the recently corrected (but still tentative) value for the coefficient $A_{1}^{(8)}$ in the theoretical expression for $a_{\mathrm{e}}$ given in Sec. V, which would directly effect the recommended value of $\alpha$.

With these problems in mind, some of which impact the possible redefinition of the kilogram, ampere, kelvin, and mole in terms of exact values of $h, e, k$, and $N_{\mathrm{A}}$ in 2011, we offer the following "wish list" for new work. If these needs, some of which appeared in our similar 2002 list, are successfully met, the key issues facing the preci-
sion measurement-fundamental constants and fundamental metrology fields should be resolved. As a consequence, our knowledge of the values of the constants, together with the International System of Units (SI), would be significantly advanced.
(i) A watt-balance determination of $K_{\mathrm{J}}^{2} R_{\mathrm{K}}$ from a laboratory other than NIST or NPL with a $u_{\mathrm{r}}$ fully competitive with $u_{\mathrm{r}}=3.6 \times 10^{-8}$, the uncertainty of the most accurate value currently available from NIST.
(ii) A timely completion of the current international effort to determine $N_{\mathrm{A}}$ with a $u_{\mathrm{r}}$ of a few parts in $10^{8}$ using highly enriched silicon crystals with $x\left({ }^{28} \mathrm{Si}\right)$ $>0.99985$ (Becker et al., 2006). This will require major advances in determining the $\{220\}$ lattice spacing, density, and molar mass of silicon.
(iii) A determination of $R$ (or Boltzmann constant $k$ $=R / N_{\mathrm{A}}$ ) with a $u_{\mathrm{r}}$ fully competitive with $u_{\mathrm{r}}=1.7 \times 10^{-6}$, the uncertainty of the most accurate value of $R$ currently available, preferably using a method other than measuring the velocity of sound in argon.
(iv) An independent calculation of the eighth-order coefficient $A_{1}^{(8)}$ in the QED theoretical expression for $a_{\mathrm{e}}$.
(v) A determination of $\alpha$ that is only weakly dependent on QED theory with a value of $u_{\mathrm{r}}$ fully competitive with $u_{\mathrm{r}}=7.0 \times 10^{-10}$, the uncertainty of the most accurate value currently available as obtained from $a_{\mathrm{e}}(\exp )$ and $a_{\mathrm{e}}(\mathrm{th})$.
(vi) A determination of the Newtonian constant of gravitation $G$ with a $u_{\mathrm{r}}$ fully competitive with $u_{\mathrm{r}}=1.4$ $\times 10^{-5}$, the uncertainty of the most accurate value of $G$ currently available.
(vii) A measurement of a transition frequency in hydrogen or deuterium, other than the already well-known hydrogen $1 \mathrm{~S}_{1 / 2}-2 \mathrm{~S}_{1 / 2}$ frequency, with an uncertainty within an order of magnitude of the current uncertainty of that frequency, $u_{\mathrm{r}}=1.4 \times 10^{-14}$, thereby providing an improved value of the Rydberg constant $R_{\infty}$.
(viii) Improved theory of the principal hadronic contribution to the theoretical expression for the muon magnetic moment anomaly $a_{\mu}$ (th) and improvements in the experimental data underlying the calculation of this contribution so that the origin of the current disagreement between $a_{\mu}(\mathrm{th})$ and $a_{\mu}(\exp )$ can be better understood.
(ix) Although there is no experimental or theoretical evidence that $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ are not exact, improved calculable-capacitor measurements of $R_{\mathrm{K}}$ and low-field measurements of the gyromagnetic ratios of the shielded proton and shielded helion, which could provide further tests of the exactness of these relations, would not be unwelcome, nor would high accuracy results ( $u_{\mathrm{r}} \approx 10^{-8}$ ) from experiments to close the "quantum electrical triangle" (Gallop, 2005; Piquemal et al., 2007).

It will be most interesting to see what portion, if any, of this ambitious program of work is completed by the 31 December 2010 closing date of the next CODATA adjustment of the values of the constants. Indeed, the progress made, especially in meeting needs (i)-(iii), may likely determine whether the 24th CGPM, which con-
venes in October 2011, will approve new definitions of the kilogram, ampere, kelvin, and mole as discussed. If such new definitions are adopted, $h, e, k$, and $N_{\mathrm{A}}$ as well as a number of other fundamental constants, for example, $K_{\mathrm{J}}, R_{\mathrm{K}}$ (assuming $K_{\mathrm{J}}=2 e / h$ and $R_{\mathrm{K}}=h / e^{2}$ ), $R$, and $\sigma$, would be exactly known, and many others would have significantly reduced uncertainties. The result would be a significant advance in our knowledge of the values of the constants.

## ACKNOWLEDGMENTS

We gratefully acknowledge the help of our many colleagues who provided us with results prior to formal publication and for promptly and patiently answering our many questions about their work.

## NOMENCLATURE

| AMDC | Atomic Mass Data Center, Centre de Spectrométrie Nucléaire et de Spectrométrie de Masse (CSNSM), Orsay, France |
| :---: | :---: |
| AME2003 | 2003 atomic mass evaluation of the AMDC |
| $A_{\mathrm{r}}(X)$ | Relative atomic mass of $X$ : |
|  | $A_{\mathrm{r}}(X)=m(X) / m_{\mathrm{u}}$ |
| $A_{90}$ | Conventional unit of electric current: |
|  | $A_{90}=V_{90} / \Omega_{90}$ |
| A* | Ångström-star: $\lambda\left(\mathrm{WK} \alpha_{1}\right)=0.2090100 \AA^{*}$ |
| $a_{\text {e }}$ | Electron magnetic moment anomaly: $a_{\mathrm{e}}=\left(\left\|g_{\mathrm{e}}\right\|-2\right) / 2$ |
| $a_{\mu}$ | Muon magnetic moment anomaly: $a_{\mu}=\left(\left\|g_{\mu}\right\|-2\right) / 2$ |
| BIPM | International Bureau of Weights and Measures, Sèvres, France |
| BNL | Brookhaven National Laboratory, Upton, New York, USA |
| CERN | European Organization for Nuclear Research, Geneva, Switzerland |
| CIPM | International Committee for Weights and Measures |
| CODATA | Committee on Data for Science and Technology of the International Council for Science |
| CPT | Combined charge conjugation, parity inversion, and time reversal |
| $c$ | Speed of light in vacuum |
| cw | Continuous wave |
| d | Deuteron (nucleus of deuterium D or ${ }^{2} \mathrm{H}$ ) |
| $d_{220}$ | \{220\} lattice spacing of an ideal crystal of naturally occurring silicon |
| $d_{220}(X)$ | $\{220\}$ lattice spacing of crystal $X$ of naturally occurring silicon |
| $E_{\mathrm{b}}$ | Binding energy |
| e | Symbol for either member of the electronpositron pair; when necessary, $\mathrm{e}^{-}$or $\mathrm{e}^{+}$is used to indicate the electron or positron |
| $e$ | Elementary charge: absolute value of the charge of the electron |


| $F$ | Faraday constant: $F=N_{\mathrm{A}} e$ | $M_{\text {u }}$ | Molar mass constant: $M_{\mathrm{u}}=10^{-3} \mathrm{~kg} \mathrm{~mol}^{-1}$ |
| :---: | :---: | :---: | :---: |
| FCDC | Fundamental Constants Data Center, NIST | $m_{\mathrm{u}}$ | Unified atomic mass constant: |
| FSU | Friedrich-Schiller University, Jena, Germany |  | $m_{\mathrm{u}}=m\left({ }^{12} \mathrm{C}\right) / 12$ |
| $\mathcal{F}_{90}$ | $\mathcal{F}_{90}=\left(F / A_{90}\right) \mathrm{A}$ |  | other elementary particles, the first symbol |
| G | Newtonian constant of gravitation |  | is used, i.e., $m_{\mathrm{e}}, m_{\mathrm{p}}$, etc.) |
| $g$ | Local acceleration of free fall | $N_{\text {A }}$ | Avogadro constant |
| $g_{\text {d }}$ | Deuteron $g$-factor: $g_{\mathrm{d}}=\mu_{\mathrm{d}} / \mu_{\mathrm{N}}$ | N/P/I | NMIJ-PTB-IRMM combined result |
| $g_{\text {e }}$ | Electron $g$-factor: $g_{\mathrm{e}}=2 \mu_{\mathrm{e}} / \mu_{\mathrm{B}}$ | NIM | National Institute of Metrology, Beijing, |
| $g_{\text {p }}$ | Proton $g$-factor: $g_{p}=2 \mu_{p} / \mu_{N}$ |  | China (People's Republic of) |
| $g_{\mathrm{p}}^{\prime}$ | Shielded proton $g$-factor: $g_{\mathrm{p}}^{\prime}=2 \mu_{\mathrm{p}}^{\prime} / \mu_{\mathrm{N}}$ | NIST | National Institute of Standards and Technology Gaithersburg Maryland USA and |
| $g_{\text {t }}$ | Triton $g$-factor: $g_{\mathrm{t}}=2 \mu_{\mathrm{t}} / \mu_{\mathrm{N}}$ |  | nology, Gaithersburg, Maryland, USA and Boulder, Colorado, USA |
| $g_{X}(Y)$ | $g$-factor of particle $X$ in the ground (1S) state of hydrogenic atom $Y$ | NMI | National Metrology Institute, Lindfield, Australia |
| $g_{\mu}$ | Muon $g$-factor: $g_{\mu}=2 \mu_{\mu} /\left(e \hbar / 2 m_{\mu}\right)$ | NMIJ | National Metrology Institute of Japan, |
| GSI | Gesellschaft für Schwerionenforschung, Darmstadt, Germany |  | Tsukuba, Japan |
| HD | HD molecule (bound state of hydrogen and deuterium atoms) | $\begin{aligned} & \text { NMR } \\ & \text { NPL } \end{aligned}$ | Nuclear magnetic resonance <br> National Physical Laboratory, Teddington, UK |
| HT | HT molecule (bound state of hydrogen and tritium atoms) | NRLM | National Research Laboratory of Metrology, Tsukuba, Japan |
| h | Helion (nucleus of ${ }^{3} \mathrm{He}$ ) | n | Neutron |
| $h$ | Planck constant; $\hbar=h / 2 \pi$ | PRC | People's Republic of China |
| Harvard; <br> HarvU | Harvard University, Cambridge, Massachusetts, USA | PTB | Physikalisch-Technische Bundesanstalt, Braunschweig and Berlin, Germany |
| ILL | Institut Max von Laue-Paul Langevin, Grenoble, France |  | Proton |
| IMGC | Istituto di Metrologia "T. Colonetti," Torino, Italy | $\bar{p}^{A} \mathrm{He}^{+}$ | Antiprotonic helium $\left({ }^{A} \mathrm{He}^{+}+\overline{\mathrm{p}}\right.$ atom, $A=3$ or 4) |
| INRIM | Istituto Nazionale di Ricerca Metrologica, Torino, Italy | QED $Q\left(\chi^{2} \mid \nu\right)$ | Quantum electrodynamics <br> Probability that an observed value of chi- |
| IRMM | Institute for Reference Materials and Measurements, Geel, Belgium |  | square for $\nu$ degrees of freedom would exceed $\chi^{2}$ |
| JINR | Joint Institute for Nuclear Research, Dubna, Russian Federation | $\frac{R}{R}$ | Molar gas constant <br> Ratio of muon anomaly difference fre- |
| KRISS | Dubna, Russian Federation Korea Research Institute of Standards and |  | quency to free proton NMR frequency |
|  | Science, Taedok Science Town, Republic of | $R_{\text {B }}$ | Birge ratio: $R_{\mathrm{B}}=\left(\chi^{2} / \nu\right)^{1 / 2}$ |
|  | Korea | $R_{\text {d }} ; \mathrm{Rd}$ | Bound-state rms charge radius of the deu- |
| KR/VN | KRISS-VNIIM Collaboration |  | teron |
| $K_{\text {J }}$ | Josephson constant: $K_{\mathrm{J}}=2 e / h$ | $R_{\text {K }}$ | von Klitzing constant: $R_{\mathrm{K}}=h / e^{\prime}$ |
| $K_{\text {J-90 }}$ | Conventional value of the Josephson constant $K_{\mathrm{J}}: K_{\mathrm{J}-90}=483597.9 \mathrm{GHz} \mathrm{V}^{-1}$ | $R_{\text {K-90 }}$ | Conventional value of the von Klitzing constant $R_{\mathrm{K}}: R_{\mathrm{K}-90}=25812.807 \Omega$ |
| $k$ | Boltzmann constant: $k=R / N_{\text {A }}$ | $R_{\mathrm{p}} ; \mathrm{Rp}$ | Bound-state rms charge radius of the pro- |
| LAMPF | Clinton P. Anderson Meson Physics Facility at Los Alamos National Laboratory, Los Alamos, New Mexico, USA | $\begin{aligned} & R_{\infty} \\ & r\left(x_{i}, x_{j}\right) \end{aligned}$ | ton <br> Rydberg constant: $R_{\infty}=m_{\mathrm{e}} c \alpha^{2} / 2 h$ <br> Correlation coefficient of estimated values |
| LKB | Laboratoire Kastler-Brossel, Paris, France |  | $x_{i}$ and $x_{j}: r\left(x_{i}, x_{j}\right)=u\left(x_{i}, x_{j}\right) /\left[u\left(x_{i}\right) u\left(x_{j}\right)\right]$ |
| LK/SY | LKB and SYRTE Collaboration | $r_{i}$ | Normalized residual of $x_{i}: r_{i}=\left(x_{i}-\hat{x}_{i}\right) / u\left(x_{i}\right)$, |
| LNE | Laboratoire national de métrologie et d'essais, Trappes, France | rms | $\hat{x}_{i}$ is the adjusted value of $x_{i}$ Root mean square |
| MIT | Massachusetts Institute of Technology, | $S_{\text {c }}$ | Self-sensitivity coefficient |
|  | Cambridge, Massachusetts, USA | SI | Système international d'unités (Interna- |
| MPQ | Max-Planck-Institut für Quantenoptik, |  | tional System of Units) |
|  | Garching, Germany | Stanford; | Stanford University, Stanford, California, |
| MSL | Measurement Standards Laboratory, Lower | StanfU | USA |
|  | Hutt, New Zealand | StPtrsb | St. Petersburg, Russian Federation |
| $M(X)$ | Molar mass of $X: M(X)=A_{\mathrm{r}}(X) M_{\mathrm{u}}$ | SYRTE | Systèmes de référence Temps Espace, Paris, |
| Mu | Muonium ( $\mu^{+} \mathrm{e}^{-}$atom) |  | France |


| $T$ | Thermodynamic temperature | $\delta_{\text {e }}$ | Additive correction to the theoretical ex- |
| :---: | :---: | :---: | :---: |
| t | Triton (nucleus of tritium T or ${ }^{3} \mathrm{H}$ ) |  | pression for the electron magnetic moment |
| th | Theory |  | anomaly $a_{\text {e }}$ |
| Type A | Uncertainty evaluation by the statistical analysis of a series of observations | $\delta_{\text {Mu }}$ | Additive correction to the theoretical expression for the ground-state hyperfine |
| Type B | Uncertainty evaluation by means other than the statistical analysis of a series of observations | $\delta_{\overline{\mathrm{p}} \mathrm{He}}$ | splitting of muonium $\Delta \nu_{\mathrm{Mu}}$ <br> Additive correction to the theoretical expression for a particular transition fre- |
| $t_{90}$ | Celsius temperature on the International |  | quency of antiprotonic helium |
|  | Temperature Scale of 1990 (ITS-90) | $\delta_{X}\left(n \mathrm{~L}_{j}\right)$ | Additive correction to the theoretical ex- |
| U.Sussex; USus | University of Sussex, Sussex, UK |  | pression for an energy level of either hydrogen H or deuterium D with quantum num- |
| UK | United Kingdom |  | bers $n$, L, and $j$ |
| USA | United States of America | $\delta_{\mu}$ | Additive correction to the theoretical ex- |
| UWash | University of Washington, Seattle, Washington, USA |  | pression for the muon magnetic moment anomaly $a_{\mu}$ |
| u | Unified atomic mass unit (also called the dalton, Da$): 1 \mathrm{u}=m_{\mathrm{u}}=m\left({ }^{12} \mathrm{C}\right) / 12$ |  | Electric constant: $\epsilon_{0}=1 / \mu_{0} c^{2}$ <br> Wavelength of $\mathrm{K} \alpha_{1}$ x-ray line of element $X$ |
| $u\left(x_{i}\right)$ | Standard uncertainty (i.e., estimated standard deviation) of an estimated value $x_{i}$ of a quantity $X_{i}$ (also simply $u$ ) | $\lambda_{\text {meas }}$ | Measured wavelength of the 2.2 MeV capture $\gamma$ ray emitted in the reaction $\mathrm{n}+\mathrm{p} \rightarrow \mathrm{d}$ $+\gamma$ |
| $u\left(x_{i}, x_{j}\right)$ | Covariance of estimated values $x_{i}$ and $x_{j}$ | $\mu$ | Symbol for either member of the muon- |
| $u_{\text {diff }}$ | Standard uncertainty of the difference $x_{i}$ $-x_{j}: u_{\text {diff }}^{2}=u^{2}\left(x_{i}\right)+u^{2}\left(x_{j}\right)-2 u\left(x_{i}, x_{j}\right)$ |  | antimuon pair; when necessary, $\mu^{-}$or $\mu^{+}$is used to indicate the negative muon or posi- |
| $u_{\mathrm{r}}\left(x_{i}\right)$ | Relative standard uncertainty of an estimated value $x_{i}$ of a quantity $X_{i}$ : | $\mu_{\text {B }}$ | tive muon Bohr magneton: $\mu_{\mathrm{B}}=e \hbar / 2 m_{\mathrm{e}}$ |
|  | $u_{\mathrm{r}}\left(x_{i}\right)=u\left(x_{i}\right) /\left\|x_{i}\right\|, x_{i} \neq 0$ (also simply $u_{\mathrm{r}}$ ) | $\mu_{\mathrm{N}}$ | Nuclear magneton: $\mu_{\mathrm{N}}=e \hbar / 2 m_{\mathrm{p}}$ |
| $u_{\mathrm{r}}\left(x_{i}, x_{j}\right)$ | Relative covariance of estimated values $x_{i}$ and $x_{j}: u_{\mathrm{r}}\left(x_{i}, x_{j}\right)=u\left(x_{i}, x_{j}\right) /\left(x_{i} x_{j}\right)$ | $\mu_{X}(Y)$ | Magnetic moment of particle $X$ in atom or molecule $Y$ |
| $V_{\mathrm{m}}(\mathrm{Si})$ | Molar volume of naturally occurring silicon | $\mu_{0}$ | Magnetic constant: $\mu_{0}=4 \pi \times 10^{-7} \mathrm{~N} / \mathrm{A}^{2}$ |
| VNIIM | D. I. Mendeleyev All-Russian Research Institute for Metrology, St. Petersburg, Russian Federation | $\mu_{X}, \mu_{X}^{\prime}$ | Magnetic moment, or shielded magnetic moment, of particle $X$ |
| $V_{90}$ | Conventional unit of voltage based on the | $\nu$ | Degrees of freedom of a particular adjustment |
|  | Josephson effect and $K_{\mathrm{J}-90}$ : $V_{90}=\left(K_{\mathrm{J}-90} / K_{\mathrm{J}}\right) \mathrm{V}$ | $\nu\left(f_{\mathrm{p}}\right)$ | Difference between muonium hyperfine splitting Zeeman transition frequencies $\nu_{34}$ |
| WGAC | Working Group on the Avogadro Constant of the CIPM Consultative Committee for Mass and Related Quantities (CCM) |  | and $\nu_{12}$ at a magnetic flux density $B$ corresponding to the free proton NMR frequency $f_{\mathrm{p}}$ |
| $W_{90}$ | Conventional unit of power: $W_{90}=V_{90}^{2} / \Omega_{90}$ | $\sigma$ | Stefan-Boltzmann constant: |
| XROI | Combined x-ray and optical interferometer |  | $\sigma=2 \pi^{5} k^{4} / 15 h^{3} c^{2}$ |
| $\mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)$ | $\mathrm{Cu} x$ unit: $\lambda\left(\mathrm{CuK} \alpha_{1}\right)=1537.400 \mathrm{xu}\left(\mathrm{CuK} \alpha_{1}\right)$ | $\tau$ | Symbol for either member of the tau- |
| $\mathrm{xu}\left(\mathrm{MoK} \alpha_{1}\right)$ | Mo x unit: $\lambda\left(\operatorname{MoK} \alpha_{1}\right)=707.831 \mathrm{xu}\left(\operatorname{MoK} \alpha_{1}\right)$ |  | antitau pair; when necessary, $\tau^{-}$or $\tau^{+}$is used |
| $x(X)$ | Amount-of-substance fraction of $X$ |  | to indicate the negative tau or positive tau |
| YAG | Yttrium aluminium garnet; $\mathrm{Y}_{3} \mathrm{Al}_{5} \mathrm{O}_{12}$ | $\chi^{2}$ | The statistic "chi square" |
| Yale; YaleU | Yale University, New Haven, Connecticut, USA | $\Omega_{90}$ | Conventional unit of resistance based on the quanum Hall effect and $R_{\mathrm{K}-90}$ : |
| $\alpha$ | Fine-structure constant: $\alpha=e^{2} / 4 \pi \epsilon_{0} \hbar c$ $\approx 1 / 137$ | $\doteq$ | $\Omega_{90}=\left(R_{\mathrm{K}} / R_{\mathrm{K}-90}\right) \Omega$ |
| $\alpha$ | Alpha particle (nucleus of ${ }^{4} \mathrm{He}$ ) |  | observational equation |
| $\Gamma_{X-90}^{\prime}(\mathrm{lo})$ | $\Gamma_{X-90}^{\prime}(\mathrm{lo})=\left(\gamma_{X}^{\prime} A_{90}\right) \mathrm{A}^{-1}, X=\mathrm{p}$ or h |  |  |

## REFERENCES

Abramowitz, M., and I. A. Stegun, 1965, Handbook of Mathematical Functions (Dover, New York).
Adkins, G. S., R. N. Fell, and P. M. Mitrikov, 2002, Phys. Rev. A 65, 042103.
Aitken, A. C., 1934, Proc. R. Soc. Edinburgh 55, 42.

AMDC, 2003, http://amdc.in2p3.fr/masstables/Ame2003/ a0p4sqza.cov03
AMDC, 2006, http://amdc.in2p3.fr
Angeli, I., 2004, At. Data Nucl. Data Tables 87, 185.
Aoyama, T., M. Hayakawa, T. Kinoshita, and M. Nio, 2007, Phys. Rev. Lett. 99, 110406.
Armstrong, T. R., and M. P. Fitzgerald, 2003, Phys. Rev. Lett. 91, 201101.
Audi, G., A. H. Wapstra, and C. Thibault, 2003, Nucl. Phys. A 729, 337.
Bagley, C. H., and G. G. Luther, 1997, Phys. Rev. Lett. 78, 3047.

Baikov, P. A., and D. J. Broadhurst, 1995, in New Computing Techniques in Physics Research IV, International Workshop on Software Engineering and Artificial Intelligence for High Energy and Nuclear Physics, edited by B. Denby and D. Perret-Gallix (World Scientific, Singapore), pp. 167-172.
Bailey, J. et al., 1979, Nucl. Phys. B150, 1.
Barker, W. A., and F. N. Glover, 1955, Phys. Rev. 99, 317.
Basile, G., A. Bergamin, G. Cavagnero, G. Mana, E. Vittone, and G. Zosi, 1994, Phys. Rev. Lett. 72, 3133.
Becker, P., 2001, Rep. Prog. Phys. 64, 1945.
Becker, P., 2003, Metrologia 40, 366.
Becker, P., H. Bettin, H.-U. Danzebrink, M. Gläser, U. Kuetgens, A. Nicolaus, D. Schiel, P. De Bièvre, S. Valkiers, and P. Taylor, 2003, Metrologia 40, 271.
Becker, P., G. Cavagnero, U. Kuetgens, G. Mana, and E. Massa, 2007, IEEE Trans. Instrum. Meas. 56, 230.
Becker, P., K. Dorenwendt, G. Ebeling, R. Lauer, W. Lucas, R. Probst, H.-J. Rademacher, G. Reim, P. Seyfried, and H. Siegert, 1981, Phys. Rev. Lett. 46, 1540.
Becker, P., et al., 2006, Meas. Sci. Technol. 17, 1854.
Beer, W., A. L. Eichenberger, B. Jeanneret, B. Jeckelmann, A. R. Pourzand, P. Richard, and J. P. Schwarz, 2003, IEEE Trans. Instrum. Meas. 52, 626.
Beier, T., 2000, Phys. Rep. 339, 79.
Beier, T., H. Häffner, N. Hermanspahn, S. G. Karshenboim, H.-J. Kluge, W. Quint, S. Stahl, J. Verdú, and G. Werth, 2002, Phys. Rev. Lett. 88, 011603.
Beier, T., I. Lindgren, H. Persson, S. Salomonson, P. Sunnergren, H. Häffner, and N. Hermanspahn, 2000, Phys. Rev. A 62, 032510.
Ben Dahan, M., E. Peik, J. Reichel, Y. Castin, and C. Salomon, 1996, Phys. Rev. Lett. 76, 4508.
Bennett, G. W., et al., 2006, Phys. Rev. D 73, 072003.
Benz, S. P., C. A. Hamilton, C. J. Burroughs, T. E. Harvey, and L. A. Christian, 1997, Appl. Phys. Lett. 71, 1866.

Berkeland, D. J., E. A. Hinds, and M. G. Boshier, 1995, Phys. Rev. Lett. 75, 2470.
Bertoldi, A., G. Lamporesi, L. Cacciapuoti, M. de Angelis, M. Fattori, T. Petelski, A. Peters, M. Prevedelli, J. Stuhler, and G. M. Tino, 2006, Eur. Phys. J. D 40, 271.

Bijnens, J., and J. Prades, 2007, Mod. Phys. Lett. A 22, 767.
BIPM, 2006, International System of Units (SI), 8th ed. (Bureau International des Poids et Mesures, Sèvres, France).
Blundell, S. A., K. T. Cheng, and J. Sapirstein, 1997, Phys. Rev. Lett. 78, 4914.
Borisoglebsky, L. A., and E. E. Trofimenko, 1979, Phys. Lett. 81B, 175.
Bourzeix, S., B. de Beauvoir, F. Nez, M. D. Plimmer, F. de Tomasi, L. Julien, F. Biraben, and D. N. Stacey, 1996, Phys. Rev. Lett. 76, 384.
Bower, V. E., and R. S. Davis, 1980, J. Res. Natl. Bur. Stand.

85, 175.
Breit, G., 1928, Nature (London) 122, 649.
Brown, L. S., and G. Gabrielse, 1986, Rev. Mod. Phys. 58, 233. Cage, M. E., R. F. Dziuba, R. E. Elmquist, B. F. Field, G. R. Jones, Jr., P. T. Olsen, W. D. Phillips, J. Q. Shields, R. L. Steiner, B. N. Taylor, and E. R. Williams, 1989, IEEE Trans. Instrum. Meas. 38, 284.
Campbell, G. K., A. E. Leanhardt, J. Mun, M. Boyd, E. W. Streed, W. Ketterle, and D. E. Pritchard, 2005, Phys. Rev. Lett. 94, 170403.
Carlson, C. E., 2007, Can. J. Phys. 85, 429.
Cavagnero, G., H. Fujimoto, G. Mana, E. Massa, K. Nakayama, and G. Zosi, 2004a, Metrologia 41, 445(E).
Cavagnero, G., H. Fujimoto, G. Mana, E. Massa, K. Nakayama, and G. Zosi, 2004b, Metrologia 41, 56.
Cladé, P., E. de Mirandes, M. Cadoret, S. Guellati-Khélifa, C. Schwob, F. Nez, L. Julien, and F. Biraben, 2006a, Phys. Rev. Lett. 96, 033001.
Cladé, P., E. de Mirandes, M. Cadoret, S. Guellati-Khélifa, C. Schwob, F. Nez, L. Julien, and F. Biraben, 2006b, Phys. Rev. A 74, 052109.
Cladé, P., S. Guellati-Khélifa, C. Schwob, F. Nez, L. Julien, and F. Biraben, 2005, Europhys. Lett. 71, 730.

Close, F. E., and H. Osborn, 1971, Phys. Lett. 34B, 400.
Clothier, W. K., G. J. Sloggett, H. Bairnsfather, M. F. Currey, and D. J. Benjamin, 1989, Metrologia 26, 9.
Colclough, A. R., 1973, Metrologia 9, 75.
Colclough, A. R., 1984, in Precision Measurement and Fundamental Constants II, edited by B. N. Taylor and W. D. Phillips NBS Spec. Pub. 617 (U.S. GPO, Washington, D.C.), pp. 263275.

Colclough, A. R., T. J. Quinn, and T. R. D. Chandler, 1979, Proc. R. Soc. London, Ser. A 368, 125.
Cook, A. H., 1961, Philos. Trans. R. Soc. London, Ser. A 254, 125.

Cook, A. H., and N. W. B. Stone, 1957, Philos. Trans. R. Soc. London, Ser. A 250, 279.
Czarnecki, A., U. D. Jentschura, and K. Pachucki, 2005, Phys. Rev. Lett. 95, 180404.
Czarnecki, A., B. Krause, and W. J. Marciano, 1996, Phys. Rev. Lett. 76, 3267.
Czarnecki, A., W. J. Marciano, and A. Vainshtein, 2003, Phys. Rev. D 67, 073006.
Czarnecki, A., W. J. Marciano, and A. Vainshtein, 2006, Phys. Rev. D 73, 119901(E).
Czarnecki, A., and K. Melnikov, 2001, Phys. Rev. Lett. 87, 013001.

Czarnecki, A., K. Melnikov, and A. Yelkhovsky, 2001, Phys. Rev. A 63, 012509.
Daussy, C., M. Guinet, A. Amy-Klein, K. Djerroud, Y. Hermier, S. Briaudeau, C. J. Bordé, and C. Chardonnet, 2007, Phys. Rev. Lett. 98, 250801.
Davier, M., 2006, private communication.
Davier, M., 2007, Nucl. Phys. B, Proc. Suppl. 169, 288.
Davier, M., S. Eidelman, A. Höcker, and Z. Zhang, 2003, Eur. Phys. J. C 31, 503.
Davier, M., and A. Höcker, 1998, Phys. Lett. B 435, 427.
Davier, M., A. Höcker, and Z. Zhang, 2006, Rev. Mod. Phys. 78, 1043.
Davier, M., and W. J. Marciano, 2004, Annu. Rev. Nucl. Part. Sci. 54, 115.
Davis, R. S., 1995, Meas. Sci. Technol. 6, 227.
de Beauvoir, B., F. Nez, L. Julien, B. Cagnac, F. Biraben, D.

Touahri, L. Hilico, O. Acef, A. Clairon, and J. J. Zondy, 1997, Phys. Rev. Lett. 78, 440.
de Trocóniz, J. F., and F. J. Ynduráin, 2005, Phys. Rev. D 71, 073008.

Dewey, M. S., E. G. Kessler, Jr., R. D. Deslattes, H. G. Börner, M. Jentschel, C. Doll, and P. Mutti, 2006, Phys. Rev. C 73, 044303.

Ding, T., D. Wan, R. Bai, Z. Zhang, Y. Shen, and R. Meng, 2005, Geochim. Cosmochim. Acta 69, 5487.
Drake, G. W. F., 2002, Can. J. Phys. 80, 1195.
Drake, G. W. F., and R. A. Swainson, 1990, Phys. Rev. A 41, 1243.

Eidelman, S., and M. Passera, 2007, Mod. Phys. Lett. A 22, 159.

Eidelman, S. I., S. G. Karshenboim, and V. A. Shelyuto, 2002, Can. J. Phys. 80, 1297.
Eides, M., 2002, private communication.
Eides, M. I., 1996, Phys. Rev. A 53, 2953.
Eides, M. I., and H. Grotch, 1995a, Phys. Rev. A 52, 3360.
Eides, M. I., and H. Grotch, 1995b, Phys. Rev. A 52, 1757.
Eides, M. I., and H. Grotch, 1997a, Ann. Phys. (N.Y.) 260, 191.
Eides, M. I., and H. Grotch, 1997b, Phys. Rev. A 56, R2507.
Eides, M. I., and H. Grotch, 1997c, Phys. Rev. A 55, 3351.
Eides, M. I., H. Grotch, and V. A. Shelyuto, 1997, Phys. Rev. A 55, 2447.
Eides, M. I., H. Grotch, and V. A. Shelyuto, 2001a, Phys. Rev. A 63, 052509.
Eides, M. I., H. Grotch, and V. A. Shelyuto, 2001b, Phys. Rep. 342, 63.
Eides, M. I., H. Grotch, and V. A. Shelyuto, 2002, Phys. Rev. D 65, 013003.
Eides, M. I., H. Grotch, and V. A. Shelyuto, 2003, Phys. Rev. D 67, 113003.
Eides, M. I., H. Grotch, and V. A. Shelyuto, 2004, Phys. Rev. D 70, 073005.
Eides, M. I., H. Grotch, and V. A. Shelyuto, 2005, Can. J. Phys. 83, 363.
Eides, M. I., and V. A. Shelyuto, 1995, Phys. Rev. A 52, 954.
Eides, M. I., and V. A. Shelyuto, 2003, Phys. Rev. A 68, 042106.
Eides, M. I., and V. A. Shelyuto, 2004, Phys. Rev. A 70, 022506.
Eides, M. I., and V. A. Shelyuto, 2007, Can. J. Phys. 85, 509.
Elkhovskiĭ, A. S., 1996, Zh. Eksp. Teor. Fiz. 110, 431 [JETP 83, 230 (1996)].
Erickson, G. W., 1977, J. Phys. Chem. Ref. Data 6, 831.
Erickson, G. W., and D. R. Yennie, 1965, Ann. Phys. (N.Y.) 35, 271.

Erler, J., and G. T. Sánchez, 2006, Phys. Rev. Lett. 97, 161801.
Farnham, D. L., R. S. Van Dyck, Jr., and P. B. Schwinberg, 1995, Phys. Rev. Lett. 75, 3598.
Faustov, R., 1970, Phys. Lett. 33B, 422.
Fellmuth, B., C. Gaiser, and J. Fischer, 2006, Meas. Sci. Technol. 17, R145.
Fischer, M., et al., 2004a, Phys. Rev. Lett. 92, 230802.
Fischer, M., et al., 2004b, in Astrophysics, Clocks, and Fundamental Constants, edited by S. Karshenboim and E. Peik, Lecture Notes in Physics Vol. 648 (Springer, Berlin), pp. 209227.

Fixler, J. B., G. T. Foster, J. M. McGuirk, and M. A. Kasevich, 2007, Science 315, 74.
Flowers, J. L., B. W. Petley, and M. G. Richards, 1993, Metrologia 30, 75.
Fortier, T. M., et al., 2007, Phys. Rev. Lett. 98, 070801.
Friar, J. L., 1979a, Z. Phys. A 292, 1.

Friar, J. L., 1979b, Ann. Phys. (N.Y.) 122, 151.
Friar, J. L., 1981, Z. Phys. A 303, 84(E).
Friar, J. L., J. Martorell, and D. W. L. Sprung, 1999, Phys. Rev. A 59, 4061.
Friar, J. L., and G. L. Payne, 1997a, Phys. Rev. C 56, 619.
Friar, J. L., and G. L. Payne, 1997b, Phys. Rev. A 56, 5173.
Fujii, K., et al., 2005, IEEE Trans. Instrum. Meas. 54, 854.
Fujimoto, H., G. Mana, and K. Nakayama, 2006, private communication.
Fujimoto, H., G. Mana, and K. Nakayama, 2007, IEEE Trans. Instrum. Meas. 56, 351.
Funck, T., and V. Sienknecht, 1991, IEEE Trans. Instrum. Meas. 40, 158.
Gabrielse, G., 2006, Int. J. Mass. Spectrom. 251, 273.
Gabrielse, G., D. Hanneke, T. Kinoshita, M. Nio, and B. Odom, 2006, Phys. Rev. Lett. 97, 030802.
Gabrielse, G., D. Hanneke, T. Kinoshita, M. Nio, and B. Odom, 2007, Phys. Rev. Lett. 99, 039902(E).
Gallop, J., 2005, Philos. Trans. R. Soc. London, Ser. A 363, 2221.

Genevès, G., P. Gournay, A. Gosset, M. Lecollinet, F. Villar, P. Pinot, P. Juncar, A. Clairon, A. Landragin, D. Holleville, F. Pereira Dos Santos, J. David, et al., 2005, IEEE Trans. Instrum. Meas. 54, 850.
George, M. C., L. D. Lombardi, and E. A. Hessels, 2001, Phys. Rev. Lett. 87, 173002.
Gerginov, V., K. Calkins, C. E. Tanner, J. J. Mcferran, S. Diddams, A. Bartels, and L. Hollberg, 2006, Phys. Rev. A 73, 032504.

Giusfredi, G., P. C. Pastor, P. De Natale, D. Mazzotti, C. De Mauro, L. Fallani, G. Hagel, V. Krachmalnicoff, and M. Inguscio, 2005, Can. J. Phys. 83, 301.
Glazov, D. A., and V. M. Shabaev, 2002, Phys. Lett. A 297, 408.
Golosov, É. A., A. S. Elkhovskiŭ, A. I. Mil'shteı̆n, and I. B. Khriplovich, 1995, Zh. Eksp. Teor. Fiz. 107, 393 [ JETP 80, 208 (1995)].
Gorshkov, M. V., Y. I. Neronov, E. N. Nikolaev, Y. V. Tarbeev, and V. L. Tal'roze, 1989, Dokl. Akad. Nauk SSSR 305, 1362 [Sov. Phys. Dokl. 34, 362 (1989)].
Greene, G. L., N. F. Ramsey, W. Mampe, J. M. Pendlebury, K. Smith, W. B. Dress, P. D. Miller, and P. Perrin, 1977, Phys. Lett. 71B, 297.
Greene, G. L., N. F. Ramsey, W. Mampe, J. M. Pendlebury, K. Smith, W. B. Dress, P. D. Miller, and P. Perrin, 1979, Phys. Rev. D 20, 2139.
Grotch, H., 1970, Phys. Rev. Lett. 24, 39.
Gundlach, J. H., and S. M. Merkowitz, 2000, Phys. Rev. Lett. 85, 2869.
Gundlach, J. H., and S. M. Merkowitz, 2002, private communication.
Häffner, H., T. Beier, S. Djekić, N. Hermanspahn, H.-J. Kluge, W. Quint, S. Stahl, J. Verdú, T. Valenzuela, and G. Werth, 2003, Eur. Phys. J. D 22, 163.
Hagiwara, K., A. D. Martin, D. Nomura, and T. Teubner, 2004, Phys. Rev. D 69, 093003.
Hagiwara, K., A. D. Martin, D. Nomura, and T. Teubner, 2007, Phys. Lett. B 649, 173.
Hagley, E. W., and F. M. Pipkin, 1994, Phys. Rev. Lett. 72, 1172.

Hanke, M., and E. G. Kessler, 2005, J. Phys. D 38, A117.
Hänsch, T., J. Alnis, P. Fendel, M. Fischer, C. Gohle, M. Herrmann, R. Holzwarth, N. Kolachevsky, T. Udem, and M. Zimmermann, 2005, Philos. Trans. R. Soc. London, Ser. A

363, 2155.
Hartland, A., R. G. Jones, and D. J. Legg, 1988, Document CCE/88-9 submitted to the 18th meeting of the Comité Consultatif d'Électricité of the CIPM.
Härtwig, J., S. Grosswig, P. Becker, and D. Windisch, 1991, Phys. Status Solidi A 125, 79.
Hayano, R. S., 2007, Can. J. Phys. 85, 453.
He, M.-G., and Z.-G. Liu, 2002, Acta Phys. Sin. 51, 1004.
Hertzog, D. W., 2007, Nucl. Phys. B, Proc. Suppl. 169, 255.
Hori, M., 2006, private communication.
Hori, M., A. Dax, J. Eades, K. Gomikawa, R. S. Hayano, N. Ono, W. Pirkl, E. Widmann, H. A. Torii, B. Juhász, D. Barna, and D. Horváth, 2006, Phys. Rev. Lett. 96, 243401.
Hori, M., J. Eades, R. S. Hayano, T. Ishikawa, W. Pirkl, E. Widmann, H. Yamaguchi, H. A. Torii, B. Juhász, D. Horváth, and T. Yamazaki, 2003, Phys. Rev. Lett. 91, 123401.
Hori, M., J. Eades, R. S. Hayano, T. Ishikawa, J. Sakaguchi, E. Widmann, H. Yamaguchi, H. A. Torii, B. Juhász, D. Horváth, and T. Yamazaki, 2001, Phys. Rev. Lett. 87, 093401.
Hu, Z.-K., J.-Q. Guo, and J. Luo, 2005, Phys. Rev. D 71, 127505.

Huber, A., T. Udem, B. Gross, J. Reichert, M. Kourogi, K. Pachucki, M. Weitz, and T. W. Hänsch, 1998, Phys. Rev. Lett. 80, 468.
Hylton, D. J., 1985, Phys. Rev. A 32, 1303.
Jeffery, A., R. E. Elmquist, J. Q. Shields, L. H. Lee, M. E. Cage, S. H. Shields, and R. F. Dziuba, 1998, Metrologia 35, 83.

Jeffery, A.-M., R. E. Elmquist, L. H. Lee, J. Q. Shields, and R. F. Dziuba, 1997, IEEE Trans. Instrum. Meas. 46, 264.

Jeffery, A.-M., L. H. Lee, and J. Q. Shields, 1999, IEEE Trans. Instrum. Meas. 48, 356.
Jegerlehner, F., 2007, Acta Phys. Pol. B 38, 3021.
Jentschura, U., and K. Pachucki, 1996, Phys. Rev. A 54, 1853. Jentschura, U. D., 2003, J. Phys. A 36, L229.
Jentschura, U. D., 2004, Phys. Rev. A 70, 052108.
Jentschura, U. D., A. Czarnecki, and K. Pachucki, 2005, Phys. Rev. A 72, 062102.
Jentschura, U. D., A. Czarnecki, K. Pachucki, and V. A. Yerokhin, 2006, Int. J. Mass. Spectrom. 251, 102.
Jentschura, U. D., and P. J. Mohr, 2004, Phys. Rev. A 69, 064103.

Jentschura, U. D., and P. J. Mohr, 2005, Phys. Rev. A 72, 014103.

Jentschura, U. D., P. J. Mohr, and G. Soff, 1999, Phys. Rev. Lett. 82, 53.
Jentschura, U. D., P. J. Mohr, and G. Soff, 2001, Phys. Rev. A 63, 042512.
Jentschura, U. D., and I. Nándori, 2002, Phys. Rev. A 66, 022114.

Jentschura, U. J., S. Kotochigova, E.-O. Le Bigot, and P. J. Mohr, 2005, NIST Technical Note 1469, http:// physics.nist.gov/hdel
Karagioz, O. V., and V. P. Izmailov, 1996, Izmer. Tekh. 39(10), 3 [ Meas. Tech. 39, 979 (1996)].
Karagioz, O. V., V. P. Izmaylov, and G. T. Gillies, 1998, Gravitation Cosmol. 4, 239.
Karshenboim, S. G., 1993, Zh. Eksp. Teor. Fiz. 103, 1105
[JETP 76, 541 (1993)].
Karshenboim, S. G., 1995, J. Phys. B 28, L77.
Karshenboim, S. G., 1996, J. Phys. B 29, L29.
Karshenboim, S. G., 1997, Z. Phys. D: At., Mol. Clusters 39, 109.

Karshenboim, S. G., 2000, Phys. Lett. A 266, 380.
Karshenboim, S. G., and V. G. Ivanov, 2002, Can. J. Phys. 80, 1305.

Karshenboim, S. G., V. G. Ivanov, Y. I. Neronov, B. P. Nikolaev, and Y. N. Tolparov, 2005, Can. J. Phys. 83, 405.
Karshenboim, S. G., V. G. Ivanov, and V. M. Shabaev, 1999, Phys. Scr., T T80, 491.
Karshenboim, S. G., V. G. Ivanov, and V. M. Shabaev, 2000, Zh. Eksp. Teor. Fiz. 117, 67 [ JETP 90, 59 (2000)].
Karshenboim, S. G., V. G. Ivanov, and V. M. Shabaev, 2001a, Can. J. Phys. 79, 81.
Karshenboim, S. G., V. G. Ivanov, and V. M. Shabaev, 2001b, Zh. Eksp. Teor. Fiz. 120, 546 [ JETP 93, 477 (2001)].
Karshenboim, S. G., and A. I. Milstein, 2002, Phys. Lett. B 549, 321.

Kataev, A. L., 2006, Phys. Rev. D 74, 073011.
Kessler, E., 2006, private communication.
Kessler, E. G., Jr., M. S. Dewey, R. D. Deslattes, A. Henins, H. G. Börner, M. Jentschel, C. Doll, and H. Lehmann, 1999, Phys. Lett. A 255, 221.
Khriplovich, I. B., and R. A. Sen'kov, 1998, Phys. Lett. A 249, 474.

Khriplovich, I. B., and R. A. Sen'kov, 2000, Phys. Lett. B 481, 447.

Kibble, B. P., 1975, in Atomic Masses and Fundamental Constants 5, edited by J. H. Sanders and A. H. Wapstra (Plenum, New York), pp. 545-551.
Kibble, B. P., and G. J. Hunt, 1979, Metrologia 15, 5.
Kibble, B. P., and I. Robinson, 1977, Feasibility Study for a Moving Coil Apparatus to Relate the Electrical and Mechanical SI Units (National Physical Laboratory, Teddington, Middlesex, UK) (NPL Report DES 40).
Kibble, B. P., and I. A. Robinson, 2003, Meas. Sci. Technol. 14, 1243.

Kibble, B. P., I. A. Robinson, and J. H. Belliss, 1990, Metrologia 27, 173.
Kim, C. G., B. C. Woo, P. G. Park, K. S. Ryu, and C. S. Kim, 1995, IEEE Trans. Instrum. Meas. 44, 484.
Kino, Y., H. Kudo, and M. Kamimura, 2003, Mod. Phys. Lett. A 18, 388.
Kinoshita, T., T. Aoyama, M. Hayakawa, and M. Nio, 2006, Nucl. Phys. B, Proc. Suppl. 160, 235.
Kinoshita, T., and M. Nio, 2004, Phys. Rev. D 70, 113001.
Kinoshita, T., and M. Nio, 2006a, Phys. Rev. D 73, 013003.
Kinoshita, T., and M. Nio, 2006b, Phys. Rev. D 73, 053007.
Kinoshita, T., B. Nizic, and Y. Okamoto, 1990, Phys. Rev. D 41, 593.

Kleinevoß, U., 2002, "Bestimmung der Newtonschen Gravitationskonstanten G," Ph.D. thesis (University of Wuppertal).
Kleinvoß, U., H. Meyer, H. Piel, and S. Hartmann, 2002, private communication.
Kleppner, D., 1997, private communication.
Kolachevsky, N., M. Fischer, S. G. Karshenboim, and T. W. Hänsch, 2004, Phys. Rev. Lett. 92, 033003.
Korobov, V. I., 2003, Phys. Rev. A 67, 062501.
Korobov, V. I., 2005, in Proceedings of the EXA05 International Conference on Exotic Atoms and Related Topics, edited by A. Hirtl, J. Marton, E. Widmann, and J. Zmeskal (Austrian Academy of Sciences Press, Vienna), pp. 391-400.
Korobov, V. I., 2006, private communication.
Kotochigova, S., 2006, private communication.
Kotochigova, S., and P. J. Mohr, 2006, (unpublished).
Kotochigova, S., P. J. Mohr, and B. N. Taylor, 2002, Can. J.

Phys. 80, 1373.
Krause, B., 1997, Phys. Lett. B 390, 392.
Krüger, E., W. Nistler, and W. Weirauch, 1999, Metrologia 36, 147.

Łach, G., B. Jeziorski, and K. Szalewicz, 2004, Phys. Rev. Lett. 92, 233001.
Lampard, D. G., 1957, Proc. IEE (London) 104C, 271, Monograph 216 M .
Laporta, S., 2001, Phys. Lett. B 523, 95.
Laporta, S., P. Mastrolia, and E. Remiddi, 2004, Nucl. Phys. B 688, 165.
Laporta, S., and E. Remiddi, 1993, Phys. Lett. B 301, 440.
Laporta, S., and E. Remiddi, 1996, Phys. Lett. B 379, 283.
Laporta, S., and E. Remiddi, 2006, private communication.
Lea, S. N., 2007, Rep. Prog. Phys. 70, 1473.
Lee, R. N., A. I. Milstein, I. S. Terekhov, and S. G. Karshenboim, 2005, Phys. Rev. A 71, 052501.
Li, G., M. A. Samuel, and M. I. Eides, 1993, Phys. Rev. A 47, 876.

Liu, R., H. Liu, T. Jin, Z. Lu, X. Du, S. Xue, J. Kong, B. Yu, X. Zhou, T. Liu, and W. Zhang, 1995, Acta Metrologia Sinica 16, 161.

Liu, W., et al., 1999, Phys. Rev. Lett. 82, 711.
Lundeen, S. R., and F. M. Pipkin, 1986, Metrologia 22, 9.
Luo, J., Z.-K. Hu, X.-H. Fu, S.-H. Fan, and M.-X. Tang, 1999, Phys. Rev. D 59, 042001.
Luther, H., K. Grohmann, and B. Fellmuth, 1996, Metrologia 33, 341.
Mana, G., 2001, in Proceedings of the International School of Physics "Enrico Fermi," Course CXLVI, Recent Advances in Metrology and Fundamental Constants, edited by T. J. Quinn, S. Leschiutta, and P. Tavella (IOS Press, Amsterdam), pp. 177-193.
Mana, G., and E. Massa, 2006, private communication.
Manohar, A. V., and I. W. Stewart, 2000, Phys. Rev. Lett. 85, 2248.

Mariam, F. G., 1981, "High Precision Measurement of the Muonium Ground State Hyperfine Interval and the Muon Magnetic Moment," Ph.D. thesis (Yale University).
Mariam, F. G., W. Beer, P. R. Bolton, P. O. Egan, C. J. Gardner, V. W. Hughes, D. C. Lu, P. A. Souder, H. Orth, J. Vetter, U. Moser, and G. zu Putlitz, 1982, Phys. Rev. Lett. 49, 993.

Martin, J., U. Kuetgens, J. Stümpel, and P. Becker, 1998, Metrologia 35, 811.
Martynenko, A. P., and R. N. Faustov, 2001, Zh. Eksp. Teor. Fiz. 120, 539 [ JETP 93, 471 (2001)].
Martynenko, A. P., and R. N. Faustov, 2002, Yad. Fiz. 65, 297 [ Phys. At. Nucl. 65, 265 (2002)].
Mastrolia, P., and E. Remiddi, 2001, in The Hydrogen Atom: Precision Physics of Simple Atomic Systems, edited by S. G. Karshenboim, F. S. Pavone, G. F. Bassani, M. Inguscio, and T. W. Hänsch (Springer, Berlin), pp. 776-783.

McGuirk, J. M., G. T. Foster, J. B. Fixler, M. J. Snadden, and M. A. Kasevich, 2002, Phys. Rev. A 65, 033608.

Melnikov, K., and A. Vainshtein, 2004, Phys. Rev. D 70, 113006.

Melnikov, K., and A. Vainshtein, 2006, Theory of the Muon Anomalous Magnetic Moment, Springer Tracts in Modern Physics Vol. 216 (Springer, Berlin).
Melnikov, K., and T. van Ritbergen, 2000, Phys. Rev. Lett. 84, 1673.

Melnikov, K., and A. Yelkhovsky, 1999, Phys. Lett. B 458, 143.
Miller, J. P., E. de Rafael, and B. Lee Roberts, 2007, Rep. Prog.

Phys. 70, 795.
Mills, I. M., P. J. Mohr, T. J. Quinn, B. N. Taylor, and E. R. Williams, 2006, Metrologia 43, 227.
Milstein, A. I., O. P. Sushkov, and I. S. Terekhov, 2002, Phys. Rev. Lett. 89, 283003.
Milstein, A. I., O. P. Sushkov, and I. S. Terekhov, 2003a, Phys. Rev. A 67, 062103.
Milstein, A. I., O. P. Sushkov, and I. S. Terekhov, 2003b, Phys. Rev. A 67, 062111.
Mohr, P. J., 1975, in Beam-Foil Spectroscopy, edited by I. A.
Sellin and D. J. Pegg (Plenum, New York), Vol. 1, pp. 89-96. Mohr, P. J., 1982, Phys. Rev. A 26, 2338.
Mohr, P. J., 1983, At. Data Nucl. Data Tables 29, 453.
Mohr, P. J., and B. N. Taylor, 2000, Rev. Mod. Phys. 72, 351.
Mohr, P. J., and B. N. Taylor, 2005, Rev. Mod. Phys. 77, 1.
Mohr, P. J., B. N. Taylor, and D. B. Newell, 2007, http:// physics.nist.gov/constants
Moldover, M. R., J. P. M. Trusler, T. J. Edwards, J. B. Mehl, and R. S. Davis, 1988, J. Res. Natl. Bur. Stand. 93, 85.
Müller, H., S.-W. Chiow, Q. Long, C. Vo, and S. Chu, 2006, Appl. Phys. B: Lasers Opt. 84, 633.
Nagy, S., T. Fritioff, M. Björkhage, I. Bergström, and R. Schuch, 2006, Europhys. Lett. 74, 404.
Nakayama, K., and H. Fujimoto, 1997, IEEE Trans. Instrum. Meas. 46, 580.
Nebel, T., et al. 2007, Can. J. Phys. 85, 469.
Neronov, Y. I., and A. E. Barzakh, 1977, Zh. Eksp. Teor. Fiz. 72, 1659 [Sov. Phys. JETP 45, 871 (1977)].
Neronov, Y. I., A. E. Barzakh, and K. Mukhamadiev, 1975, Zh. Eksp. Teor. Fiz. 69, 1872 [Sov. Phys. JETP 42, 950 (1975)].
Neronov, Y. I., and S. G. Karshenboim, 2003, Phys. Lett. A 318, 126.
Newton, G., D. A. Andrews, and P. J. Unsworth, 1979, Philos. Trans. R. Soc. London, Ser. A 290, 373.
Niering, M., et al., 2000, Phys. Rev. Lett. 84, 5496.
Nio, M., 2001, in Quantum Electrodynamics and Physics of the Vacuum, edited by G. Cantatore, AIP Conf. Proc. No. 564 (American Institute of Physics, Melville, NY), pp. 178-185.
Nio, M., 2002, private communication.
Odom, B., D. Hanneke, B. D'Urso, and G. Gabrielse, 2006, Phys. Rev. Lett. 97, 030801.
Pachucki, K., 1993a, Phys. Rev. A 48, 2609.
Pachucki, K., 1993b, Phys. Rev. A 48, 120.
Pachucki, K., 1994, Phys. Rev. Lett. 72, 3154.
Pachucki, K., 1995, Phys. Rev. A 52, 1079.
Pachucki, K., 2001, Phys. Rev. A 63, 042503.
Pachucki, K., 2006, Phys. Rev. Lett. 97, 013002.
Pachucki, K., 2007, Phys. Rev. A 76, 022508.
Pachucki, K., A. Czarnecki, U. D. Jentschura, and V. A. Yerokhin, 2005, Phys. Rev. A 72, 022108.
Pachucki, K., and H. Grotch, 1995, Phys. Rev. A 51(3), 1854.
Pachucki, K., and U. D. Jentschura, 2003, Phys. Rev. Lett. 91, 113005.

Pachucki, K., U. D. Jentschura, and V. A. Yerokhin, 2004, Phys. Rev. Lett. 93, 150401.
Pachucki, K., U. D. Jentschura, and V. A. Yerokhin, 2005, Phys. Rev. Lett. 94, 229902(E).
Pachucki, K., and S. G. Karshenboim, 1999, Phys. Rev. A 60, 2792.

Palomar, J. E., 2003, Nucl. Phys. B, Proc. Suppl. 121, 183.
Park, P. G., C. H. Choi, C. S. Kim, V. Y. Shifrin, and V. N. Khorev, 1999, J. Korean Phys. Soc. 34, 327.
Passera, M., 2005, J. Phys. G 31, R75.

Passera, M., 2007, Phys. Rev. D 75, 013002.
Peik, E., M. Ben Dahan, I. Bouchoule, Y. Castin, and C. Salomon, 1997, Phys. Rev. A 55, 2989.
Penin, A. A., 2004, Int. J. Mod. Phys. A 19, 3897.
Peters, A., K. Y. Chung, and S. Chu, 2001, Metrologia 38, 25.
Petley, B. W., and R. W. Donaldson, 1984, Metrologia 20, 81.
Phillips, W. D., W. E. Cooke, and D. Kleppner, 1977, Metrologia 13, 179.
Phillips, W. D., D. Kleppner, and F. G. Walther, 1984, private communication.
Picard, A., M. Stock, H. Fang, T. J. Witt, and D. Reymann, 2007, IEEE Trans. Instrum. Meas. 56, 538.
Piquemal, F., L. Devoille, N. Feltin, and B. Steck, 2007, in Metrology and Fundamental Constants, Vol. 166 of International School of Physics "Enrico Fermi," Course CLXVI, edited by T. W. Hänsch, S. Leschiutta, A. J. Wallard, and M. L. Rastello (IOS Press, Amsterdam), pp. 181-210.
Preston-Thomas, H., 1990a, Metrologia 27, 107(E).
Preston-Thomas, H., 1990b, Metrologia 27, 3.
Quinn, T. J., 1989, Metrologia 26, 69.
Quinn, T. J., 2001, Metrologia 38, 89.
Quinn, T. J., A. R. Colclough, and T. R. D. Chandler, 1976, Philos. Trans. R. Soc. London, Ser. A 283, 367.
Quinn, T. J., C. C. Speake, S. J. Richman, R. S. Davis, and A. Picard, 2001, Phys. Rev. Lett. 87, 111101.
Rainville, S., J. K. Thompson, E. G. Myers, J. M. Brown, M. S. Dewey, E. G. Kessler Jr, R. D. Deslattes, H. G. Börner, M. Jentschel, P. Mutti, and D. E. Pritchard, 2005, Nature (London) 438, 1096.
Ramsey, N. F., 1990, in Quantum Electrodynamics, edited by T. Kinoshita (World Scientific, Singapore), Chap. 13, pp. 673695.

Reynolds, B. C., R. B. Georg, F. Oberli, U. Wiechert, and A. N. Halliday, 2006, J. Anal. At. Spectrom. 21, 266.

Ritter, M. W., P. O. Egan, V. W. Hughes, and K. A. Woodle, 1984, Phys. Rev. A 30, 1331.
Robinson, I. A., and B. P. Kibble, 1997, IEEE Trans. Instrum. Meas. 46, 596.
Robinson, I. A., and B. P. Kibble, 2002, in 2002 Conference on Precision Electromagnetic Measurements Digest, edited by T. McComb (IEEE, Piscataway, NJ), pp. 574-575.
Robinson, I. A., and B. P. Kibble, 2007, Metrologia 44, 427.
Sapirstein, J. R., and D. R. Yennie, 1990, in Quantum Electrodynamics, edited by T. Kinoshita (World Scientific, Singapore), Chap. 12, pp. 560-672.
Sarajlic, E., S. Chu, and A. Wicht, 2006, private communication.
Schlamminger, S., 2007, private communication.
Schlamminger, S., E. Holzschuh, and W. Kündig, 2002, Phys. Rev. Lett. 89, 161102.
Schlamminger, S., E. Holzschuh, W. Kündig, F. Nolting, R. E. Pixley, J. Schurr, and U. Staumann, 2006, Phys. Rev. D 74, 082001.

Schmidt, J. W., R. M. Gavioso, E. F. May, and M. R. Moldover, 2007, Phys. Rev. Lett. 98, 254504.
Schwarz, J. P., R. Liu, D. B. Newell, R. L. Steiner, E. R. Williams, D. Smith, A. Erdemir, and J. Woodford, 2001, J. Res. Natl. Inst. Stand. Technol. 106, 627.
Schwob, C., L. Jozefowski, B. de Beauvoir, L. Hilico, F. Nez, L. Julien, F. Biraben, O. Acef, and A. Clairon, 1999, Phys. Rev. Lett. 82, 4960.
Schwob, C., L. Jozefowski, B. de Beauvoir, L. Hilico, F. Nez, L. Julien, F. Biraben, O. Acef, and A. Clairon, 2001, Phys. Rev.

Lett. 86, 4193(E).
Shabaev, V. M., A. N. Artemyev, T. Beier, and G. Soff, 1998, J. Phys. B 31, L337.
Sheppard, W. F., 1912, in Proceedings of the Fifth International Congress of Mathematicians, edited by E. W. Hobson and A. E. H. Love (Cambridge University Press, Cambridge), Vol. ii, pp. 348-384.
Shifrin, V. Y., V. N. Khorev, P. G. Park, C. H. Choi, and C. S. Kim, 1998a, Izmer. Tekh. 1998, 68.
Shifrin, V. Y., P. G. Park, V. N. Khorev, C. H. Choi, and C. S.
Kim, 1998b, IEEE Trans. Instrum. Meas. 47, 638.
Shifrin, V. Y., P. G. Park, V. N. Khorev, C. H. Choi, and S. Lee, 1999, IEEE Trans. Instrum. Meas. 48, 196.
Sick, I., 2003, Phys. Lett. B 576, 62.
Sick, I., 2007, Can. J. Phys. 85, 409.
Sick, I., and D. Trautmann, 1998, Nucl. Phys. A 637, 559.
Sienknecht, V., and T. Funck, 1985, IEEE Trans. Instrum. Meas. IM-34, 195.
Sienknecht, V., and T. Funck, 1986, Metrologia 22, 209.
Small, G. W., B. W. Ricketts, P. C. Coogan, B. J. Pritchard, and M. M. R. Sovierzoski, 1997, Metrologia 34, 241.

Steiner, R., D. Newell, and E. Williams, 2005, J. Res. Natl. Inst. Stand. Technol. 110, 1.
Steiner, R. L., E. R. Williams, R. Liu, and D. B. Newell, 2007, IEEE Trans. Instrum. Meas. 56, 592.
Steiner, R. L., E. R. Williams, D. B. Newell, and R. Liu, 2005, Metrologia 42, 431.
Stock, M., and T. J. Witt, 2006, Metrologia 43, 583.
Stöckinger, D., 2007, J. Phys. G 34, R45.
Tarbeev, Y. V., V. Y. Shifrin, V. N. Khorev, and N. V. Studentsov, 1989, Izmer. Tekh. 32(4), 3 [ Meas. Tech. 32, 279 (1989)].

Taylor, B. N., and P. J. Mohr, 2001, IEEE Trans. Instrum. Meas. 50, 563.
Thompson, A. M., 1959, Proc. IRE 106B, 307, Paper No. 2887 M.

Thompson, A. M., and D. G. Lampard, 1956, Nature (London) 177, 888.
Trapon, G., O. Thévenot, J. C. Lacueille, and W. Poirier, 2003, Metrologia 40, 159.
Trapon, G., O. Thévenot, J.-C. Lacueille, W. Poirier, H. Fhima, and G. Genevès, 2001, IEEE Trans. Instrum. Meas. 50, 572.
Udem, T., 2006, private communication.
Udem, T., J. Reichert, R. Holzwarth, and T. W. Hänsch, 1999, Phys. Rev. Lett. 82, 3568.
Valkiers, S., K. Ruße, P. Taylor, T. Ding, and M. Inkret, 2005, Int. J. Mass. Spectrom. 242, 321.
Van Dyck, R. S., Jr., 2006, private communication.
Van Dyck, R. S., Jr., D. B. Pinegar, S. V. Liew, and S. L. Zafonte, 2006, Int. J. Mass. Spectrom. 251, 231.
Van Dyck, R. S., Jr., P. B. Schwinberg, and H. G. Dehmelt, 1987, Phys. Rev. Lett. 59, 26.
Verdú, J., 2006, private communication.
Verdú, J., T. Beier, S. Djekić, H. Häffner, H.-J. Kluge, W. Quint, T. Valenzuela, M. Vogel, and G. Werth, 2003, J. Phys. B 36, 655.
Verdú, J., S. Djekić, S. Stahl, T. Valenzuela, M. Vogel, and G. Werth, 2004, Phys. Rev. Lett. 92, 093002.
Verdú, J. L., T. Beier, S. Djekic, H. Häffner, H.-J. Kluge, W. Quint, T. Valenzuela, and G. Werth, 2002, Can. J. Phys. 80, 1233.

Vitushkin, L. et al., 2007, in Proceedings of the 1st International Symposium of the International Gravity Field Service "Grav-
ity Field of the Earth," General Command of Mapping, special issue 18, pp. 382-387.
Wapstra, A. H., G. Audi, and C. Thibault, 2003, Nucl. Phys. A 729, 129.
Weitz, M., A. Huber, F. Schmidt-Kaler, D. Leibfried, W. Vassen, C. Zimmermann, K. Pachucki, T. W. Hänsch, L. Julien, and F. Biraben, 1995, Phys. Rev. A 52, 2664.
Werth, G., 2003, private communication.
Werth, G., J. Alonso, T. Beier, K. Blaum, S. Djekic, H. Häffner, N. Hermanspahn, W. Quint, S. Stahl, J. Verdú, T. Valenzuela, and M. Vogel, 2006, Int. J. Mass. Spectrom. 251, 152.

Wichmann, E. H., and N. M. Kroll, 1956, Phys. Rev. 101, 843.
Wicht, A., J. M. Hensley, E. Sarajlic, and S. Chu, 2002, Phys. Scr. T102, 82.
Wicht, A., E. Sarajlic, J. M. Hensley, and S. Chu, 2005, Phys. Rev. A 72, 023602.
Wieser, M. E., 2006, Pure Appl. Chem. 78, 2051.
Williams, E. R., G. R. Jones, Jr., S. Ye, R. Liu, H. Sasaki, P. T. Olsen, W. D. Phillips, and H. P. Layer, 1989, IEEE Trans. Instrum. Meas. 38, 233.
Williams, E. R., R. L. Steiner, D. B. Newell, and P. T. Olsen, 1998, Phys. Rev. Lett. 81, 2404.

Wimett, T. F., 1953, Phys. Rev. 91, 499.
Winkler, P. F., D. Kleppner, T. Myint, and F. G. Walther, 1972, Phys. Rev. A 5, 83.
Yamazaki, T., N. Morita, R. S. Hayano, E. Widmann, and J. Eades, 2002, Phys. Rep. 366, 183.
Yao, W.-M., et al., 2006, J. Phys. G 33, 1.
Yerokhin, V. A., 2000, Phys. Rev. A 62, 012508.
Yerokhin, V. A., P. Indelicato, and V. M. Shabaev, 2002, Phys. Rev. Lett. 89, 143001.
Yerokhin, V. A., P. Indelicato, and V. M. Shabaev, 2003, Phys. Rev. Lett. 91, 073001.
Yerokhin, V. A., P. Indelicato, and V. M. Shabaev, 2005a, Phys. Rev. A 71, 040101.
Yerokhin, V. A., P. Indelicato, and V. M. Shabaev, 2005b, Zh. Eksp. Teor. Fiz. 128, 322 [ JETP 101, 280 (2005)].
Yerokhin, V. A., P. Indelicato, and V. M. Shabaev, 2007, Can. J. Phys. 85, 521.
Yerokhin, V. A., and V. M. Shabaev, 2001, Phys. Rev. A 64, 012506.

Zelevinsky, T., D. Farkas, and G. Gabrielse, 2005, Phys. Rev. Lett. 95, 203001.
Zhang, Z., X. Wang, D. Wang, X. Li, Q. He, and Y. Ruan, 1995, Acta Metrologia Sinica 16, 1.


[^0]:    *This report was prepared by the authors under the auspices of the CODATA Task Group on Fundamental Constants. The members of the task group are:
    F. Cabiati, Istituto Nazionale di Ricerca Metrologica, Italy
    K. Fujii, National Metrology Institute of Japan, Japan
    S. G. Karshenboim, D. I. Mendeleyev All-Russian Research Institute for Metrology, Russian Federation
    I. Lindgren, Chalmers University of Technology and Göteborg University, Sweden
    B. A. Mamyrin (deceased), A. F. Ioffe Physical-Technical Institute, Russian Federation
    W. Martienssen, Johann Wolfgang Goethe-Universität, Germany
    P. J. Mohr, National Institute of Standards and Technology, United States of America
    D. B. Newell, National Institute of Standards and Technology, United States of America
    F. Nez, Laboratoire Kastler-Brossel, France
    B. W. Petley, National Physical Laboratory, United Kingdom
    T. J. Quinn, Bureau international des poids et mesures
    B. N. Taylor, National Institute of Standards and Technology, United States of America
    W. Wöger, Physikalisch-Technische Bundesanstalt, Germany
    B. M. Wood, National Research Council, Canada
    Z. Zhang, National Institute of Metrology, China (People's Republic of)
    This review is being published simultaneously by Reviews of Modern Physics; see Rev. Mod. Phys. 80, 633 (2008).
    ${ }^{\dagger}$ mohr@nist.gov
    *barry.taylor@nist.gov
    ${ }^{\S}$ dnewell@nist.gov

[^1]:    ${ }^{\text {a }}$ TR\&D: Tribotech Research and Development Company, Moscow, Russian Federation; LANL: Los Alamos National Laboratory, Los Alamos, New Mexico, USA; UWash: University of Washington, Seattle, Washington, USA; BIPM: International Bureau of Weights and Measures, Sèvres, France; UWup: University of Wuppertal, Wuppertal, Germany; MSL: Measurement Standards Laboratory, Lower Hutt, New Zealand; HUST: Huazhong University of Science and Technology, Wuhan, PRC; UZur: University of Zurich, Zurich, Switzerland.

[^2]:    ${ }^{\text {a }}$ The values in brackets are relative to the corresponding transition frequency.

[^3]:    ${ }^{\text {a }}$ The relative covariance is $u_{\mathrm{r}}\left(x_{i}, x_{j}\right)=u\left(x_{i}, x_{j}\right) /\left(x_{i} x_{j}\right)$, where $u\left(x_{i}, x_{j}\right)$ is the covariance of $x_{i}$ and $x_{j}$; the relative variance is $u_{\mathrm{r}}^{2}\left(x_{i}\right)=u_{\mathrm{r}}\left(x_{i}, x_{i}\right)$; and the correlation coefficient is $r\left(x_{i}, x_{j}\right)=u\left(x_{i}, x_{j}\right) /\left[u\left(x_{i}\right) u\left(x_{j}\right)\right]$.

[^4]:    ${ }^{\text {a }}$ This is the lattice parameter (unit cell edge length) of an ideal single crystal of naturally occurring Si free of impurities and imperfections, and is deduced from measurements on extremely pure and nearly perfect single crystals of Si by correcting for impurity effects.

